

Probing the cool ISM in galaxies via 21 cm H I absorption

J. R. Allison^{1,*}, E. M. Sadler^{1,2}, S. J. Curran^{1,2} and S. N. Reeves^{1,2,3}

¹Sydney Institute for Astronomy, School of Physics A28,
 University of Sydney, NSW 2006, Australia

²ARC Centre of Excellence for All-sky Astrophysics (CAASTRO)

³CSIRO Astronomy & Space Science, P.O. Box 76, Epping NSW 1710, Australia

*email: jra@physics.usyd.edu.au

Abstract. Recent targeted studies of associated H I absorption in radio galaxies are starting to map out the location, and potential cosmological evolution, of the cold gas in the host galaxies of Active Galactic Nuclei (AGN). The observed 21 cm absorption profiles often show two distinct spectral-line components: narrow, deep lines arising from cold gas in the extended disc of the galaxy, and broad, shallow lines from cold gas close to the AGN (e.g. Morganti et al. 2011). Here, we present results from a targeted search for associated H I absorption in the youngest and most recently-triggered radio AGN in the local universe (Allison et al. 2012b). So far, by using the recently commissioned Australia Telescope Compact Array Broadband Backend (CABB; Wilson et al. 2011), we have detected two new absorbers and one previously-known system. While two of these show both a broad, shallow component and a narrow, deep component (see Fig. 1), one of the new detections has only a single broad, shallow component. Interestingly, the host galaxies of the first two detections are classified as gas-rich spirals, while the latter is an early-type galaxy. These detections were obtained using a spectral-line finding method, based on Bayesian inference, developed for future large-scale absorption surveys (Allison et al. 2012a).

Keywords. methods: data analysis, galaxies: active, radio lines: galaxies

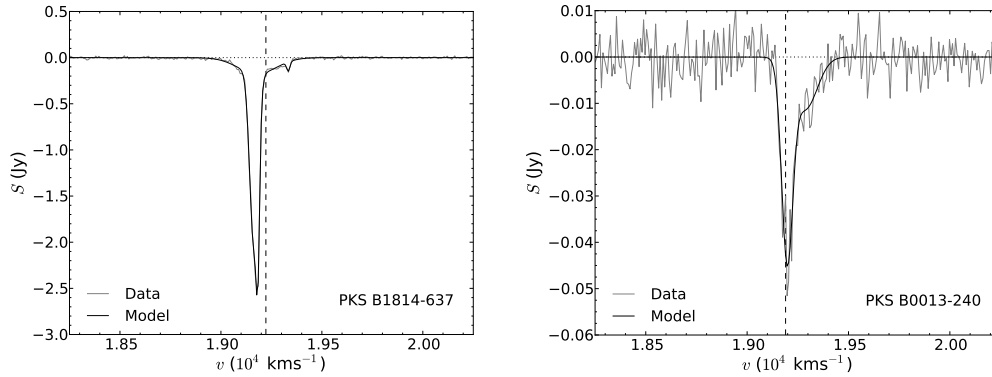


Figure 1. 21 cm absorption in the late-type host galaxies of two compact radio sources.

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