

Measurement of Muon Antineutrino Oscillations with an Accelerator-Produced Off-Axis Beam

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T2K reports its first measurements of the parameters governing the disappearance of $\bar{\nu}_\mu$ in an off-axis beam due to flavor change induced by neutrino oscillations. The quasimonochromatic $\bar{\nu}_\mu$ beam, produced with a peak energy of 0.6 GeV at J-PARC, is observed at the far detector Super-Kamiokande, 295 km away, where the $\bar{\nu}_\mu$ survival probability is expected to be minimal. Using a data set corresponding to 4.01×10^{20} protons on target, 34 fully contained μ -like events were observed. The best-fit oscillation parameters are $\sin^2(\theta_{23}) = 0.45$ and $|\Delta\bar{m}_{32}^2| = 2.51 \times 10^{-3} \text{ eV}^2$ with 68% confidence intervals of 0.38–0.64 and $2.26\text{--}2.80 \times 10^{-3} \text{ eV}^2$, respectively. These results are in agreement with existing antineutrino parameter measurements and also with the ν_μ disappearance parameters measured by T2K.

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Introduction.—In the three-flavor framework, neutrino oscillation can be described by the unitary Pontecorvo-Maki-Nakagawa-Sakata matrix, which is parameterized by three angles θ_{12} , θ_{23} , θ_{13} and a CP -violating phase δ_{CP} [1–3]. Given a neutrino propagation distance, L (km), and energy, E_ν (GeV), such that $L/E_\nu \sim O(1000)$, the survival probability for a muon neutrino propagating through vacuum can be approximated by

$$P(\nu_\mu \rightarrow \nu_\mu) \simeq 1 - 4\cos^2(\theta_{13})\sin^2(\theta_{23}) \times [1 - \cos^2(\theta_{13})\sin^2(\theta_{23})]\sin^2\left(\frac{1.267\Delta m_{32}^2 L}{E_\nu}\right), \quad (1)$$

where $\Delta m_{32}^2(\text{eV}^2)$ is the neutrino mass squared splitting, defined as $m_3^2 - m_2^2$. Equation (1) shows that measuring the disappearance probability as a function of L/E_ν leads to a measurement of the oscillation parameters. In this model of neutrino oscillation, the disappearance probability in vacuum is identical for neutrinos and antineutrinos. The disappearance probabilities in matter can differ by as much as 0.1% for the T2K baseline and neutrino flux, but our data set is not sensitive to this small effect. Observing a significant difference between the disappearance probabilities of neutrinos and antineutrinos would, therefore, be evidence for new physics [3]. Results from the MINOS [4] and Super-Kamiokande (SK) Collaborations [5] indicate no significant difference between muon antineutrino oscillations and muon neutrino oscillations.

In this Letter, we present the first measurement of $P(\bar{\nu}_\mu \rightarrow \bar{\nu}_\mu)$ by the T2K Collaboration. This analysis allows

the dominant antineutrino oscillation parameters for $\bar{\nu}_\mu$ disappearance to vary independently from those describing neutrino oscillations, i.e., $\theta_{23} \neq \bar{\theta}_{23}$ and $\Delta m_{32}^2 \neq \Delta\bar{m}_{32}^2$, where the barred parameters refer to antineutrino oscillations. $\bar{\theta}_{13}$, $\bar{\theta}_{12}$, and $\Delta\bar{m}_{21}^2$ are assumed to be identical to their matter counterparts since our data set cannot constrain them. This ensures that the expected background at the far detector is consistent with the current knowledge of neutrino oscillations, while allowing us to use the T2K antineutrino-mode data to measure $\bar{\theta}_{23}$ and $\Delta\bar{m}_{32}^2$.

T2K experiment.—The T2K experiment [6] is composed of a neutrino beam line, a suite of near detectors, and the far detector, Super-Kamiokande. Both the far detector and one of the near detectors are placed 2.5° off the neutrino beam axis and so observe a narrow-band beam [7]. This “off-axis” method reduces backgrounds from higher-energy neutrinos, producing a neutrino flux that peaks around 0.6 GeV, the energy at which the first minimum in the $\bar{\nu}_\mu$ survival probability is expected to occur at the T2K baseline.

The J-PARC main ring provides a 30-GeV proton beam which impinges upon a graphite target, producing pions and kaons. The target is held inside the first of three magnetic horns which focus charged particles into a 96-m-long decay volume, where they decay and produce neutrinos. The polarity of the horn current determines whether positive or negative mesons are focused, which in turn determines whether the neutrino beam is largely composed of muon neutrinos or muon antineutrinos. The decay volume ends in a beam dump followed by the muon monitor, which measures the neutrino beam direction on a bunch-by-bunch basis using muons from the meson decays.

The near-detector complex [6] consists of the on-axis Interactive Neutrino GRID detector (INGRID) [8] and the off-axis detector (ND280), both 280 m downstream of the proton-beam target. INGRID is a $7 + 7$ array of iron-scintillator detectors, arranged in a “cross” configuration at the beam center. INGRID provides high-statistics monitoring of the neutrino beam intensity, direction, profile, and

stability and has shown that the neutrino beam direction is controlled to 0.4 mrad. ND280 consists of a number of subdetectors installed inside the refurbished UA1/NOMAD magnet, which provides a 0.2 T field. The near-detector analysis described here uses the tracker region of ND280, which consists of three time projection chambers (TPC1, 2, 3) [9] interleaved with two fine-grained detectors (FGD1, 2) [10]. The FGDs are the neutrino target and track charged particles coming from the interaction vertex, while the TPCs perform 3D tracking and determine the charge, momentum, and energy loss of each charged particle traversing them. The observed energy loss is used for particle identification which, when combined with particle charge information, allows a precise separation and measurement of the $\bar{\nu}_\mu$ (right-sign) and ν_μ (wrong-sign) interactions in the antineutrino-mode beam.

The far detector is a 50-kt (22.5-kt fiducial mass) water Cherenkov detector [11,12], where the volume is divided into an outer detector (OD) with 1885 outward-facing 20-cm-diameter photomultiplier tubes and an inner detector (ID) with 11 129 inward-facing 50-cm-diameter photomultiplier tubes. The events arriving at SK from the J-PARC beam spill are synchronized with a global positioning system with <150 ns precision.

The results presented here are based on data taken in three periods: two where the beam operated in antineutrino mode, (1) June 2014 and (2) November 2014–June 2015, and one in neutrino mode, (3) November 2010–May 2013. The oscillation analysis uses periods (1) and (2), while the near-detector analysis uses data from periods (1) and (3). This corresponds to an exposure of 4.01×10^{20} protons on target (POT) in antineutrino mode for the oscillation analysis, and an exposure of 0.43×10^{20} POT in antineutrino mode plus 5.82×10^{20} POT in neutrino mode for the near-detector analysis.

Analysis strategy.—This analysis resembles that of Ref. [13], fitting samples of charged-current (CC) interactions at ND280 to produce a tuned prediction of the unoscillated antineutrino spectrum at the far detector, including its associated uncertainty. This analysis differs from Ref. [13] in that both ν_μ and $\bar{\nu}_\mu$ samples at ND280 are fitted. This ensures that the neutrino interaction model is consistent between both neutrino- and antineutrino-beam-mode data sets and provides a constraint on both the right-sign signal and the wrong-sign background in the antineutrino-mode beam.

Flux simulation.—The nominal neutrino flux at ND280 and SK (without oscillation) is predicted by simulating the secondary beam line [14] using FLUKA2011 [15,16] and GEANT3 with G4CALOR [17,18]. The simulated hadronic interactions are tuned to external hadron-production data. The unoscillated neutrino flux prediction at SK is shown in Fig. 1 for each neutrino type and for both neutrino- and antineutrino-mode beams. At the peak energy of the T2K beam, the ν_μ flux in the neutrino-mode beam is 20% higher

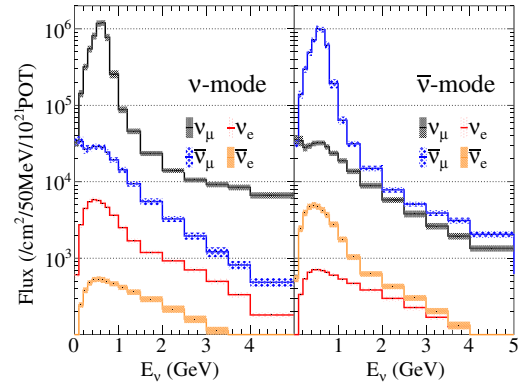


FIG. 1. The nominal unoscillated neutrino flux prediction at SK for each neutrino type in the neutrino-mode beam (left) and antineutrino-mode beam (right). The shaded boxes indicate the total systematic uncertainty on each energy bin.

than the $\bar{\nu}_\mu$ flux in the antineutrino-mode beam, due to the larger production cross section for π^+ compared to π^- in proton-carbon interactions. The ratio of the wrong-sign component (ν_μ in the $\bar{\nu}_\mu$ beam), mainly coming from forward-going high-energy pions, to the right-sign component ($\bar{\nu}_\mu$) at the peak energy is 3%. The largest sources of neutrino flux uncertainty are from beam-line and hadron-production modeling uncertainties, which are common to ND280 and SK. The new NA61/SHINE 2009 thin-target data [19] are included in the hadron-production tuning for this analysis, reducing the total flux uncertainty from between 12%–15% to 10% around 0.6 GeV.

Neutrino interaction simulation.—Neutrino interactions are modeled with the NEUT Monte Carlo event generator [20–24]. The generator uses the same model with common parameters to describe both ν and $\bar{\nu}$ interactions. In the case of CC quasielastic (CCQE) reactions ($\nu_\mu + n \rightarrow \mu^- + p$ or $\bar{\nu}_\mu + p \rightarrow \mu^+ + n$) neutrino and antineutrino cross sections differ by the sign of the vector-axial interference term [25,26]. At a neutrino energy of 0.6 GeV, this makes the neutrino-oxygen CCQE cross section a factor of ~ 4 larger than that of antineutrinos.

To set the initial values and uncertainties of some parameters, such as the CCQE axial mass and the normalization of the multinucleon contribution, results from the MiniBooNE and MINERvA experiments [27–30] on CH_2 and CH targets are used. These parameters are then tuned by the near-detector fit.

Near-detector fit.—The seven samples used in the near-detector fit are summarized in Table I. Muon-neutrino-induced CC interactions in the neutrino beam mode are found by requiring that the highest-momentum, negative-curvature track in an event starts within the upstream FGD (FGD1) fiducial volume (FV) and has an energy deposit in TPC2 consistent with a muon. Events with a TPC track that starts upstream of the start point of the muon candidate are rejected, and the remaining ν_μ CC candidates are divided into three subsamples according to

TABLE I. Data and MC-predicted event rates for the different ND280 samples before and after the fit. Errors indicate systematic uncertainties only.

Sample	Data	Prefit	Postfit
ν beam mode			
ν_μ CC 0π	17 362	$15\,625 \pm 1663$	$17\,248 \pm 133$
ν_μ CC $1\pi^+$	3988	4748 ± 686	4190 ± 60
ν_μ CC other	4219	3772 ± 431	4079 ± 62
$\bar{\nu}$ beam mode			
$\bar{\nu}_\mu$ CC 1 track	435	387 ± 41	438 ± 13
$\bar{\nu}_\mu$ CC N tracks	136	128 ± 17	129 ± 5
ν_μ CC 1 track	131	141 ± 15	147 ± 6
ν_μ CC N tracks	145	147 ± 17	144 ± 6

the number of associated pions: ν_μ CC 0π , ν_μ CC $1\pi^+$, and ν_μ CC other, which are dominated by CCQE, CC resonant pion production, and deep inelastic scattering interactions, respectively [13]. For the antineutrino-beam-mode samples, the selection of $\bar{\nu}_\mu$ (ν_μ) CC interactions is similar to that used in the neutrino beam mode, except the positive (negative) track must be the highest-momentum track in the event. The selected $\bar{\nu}_\mu$ (ν_μ) CC candidate events are divided into two subsamples rather than three, due to the small amount of antineutrino-mode data used in this analysis. These are defined by the number of reconstructed tracks crossing TPC2: $\bar{\nu}_\mu$ (ν_μ) CC 1 track, dominated by CCQE interactions; and $\bar{\nu}_\mu$ (ν_μ) CC N tracks ($N > 1$), a mixture of resonant production and deep inelastic scattering.

The fit uses a binned likelihood, with the samples binned according to the muon momentum and angle (θ) relative to the central axis of the detector, roughly 1.7° away from the incident (anti)neutrino direction. The TPCs calculate the muon momentum from the curvature of the lepton in the ND280 magnetic field, with a resolution of 6% at 1 GeV/ c [9]. Figure 2 shows the 1D projections of these distributions for the ν_μ CC 0π , the $\bar{\nu}_\mu$ CC 1 track, and the ν_μ CC 1 track samples for both data and the post-fit MC model. The p -value of the data fit likelihood ratio was found to be 0.05, and the agreement between the ND280 data and the MC model was judged to be acceptable. The fit gives estimates for 25 antineutrino beam flux parameters at SK, 12 cross-section parameters (including 4 specific to oxygen), and their covariance. There are also additional parameters to control pion final state interactions (FSI) and reinteractions within the detector, which are independent for ND280 and SK.

To decouple the properties of the carbon target at ND280 from those of the oxygen target at SK, separate Fermi momentum, binding energy, multinucleon event normalization, and CC coherent pion-production normalization parameters are introduced for interactions on oxygen. Since oxygen comprises only 3.6% by mass of the FGD1 target, this near-detector analysis is insensitive to

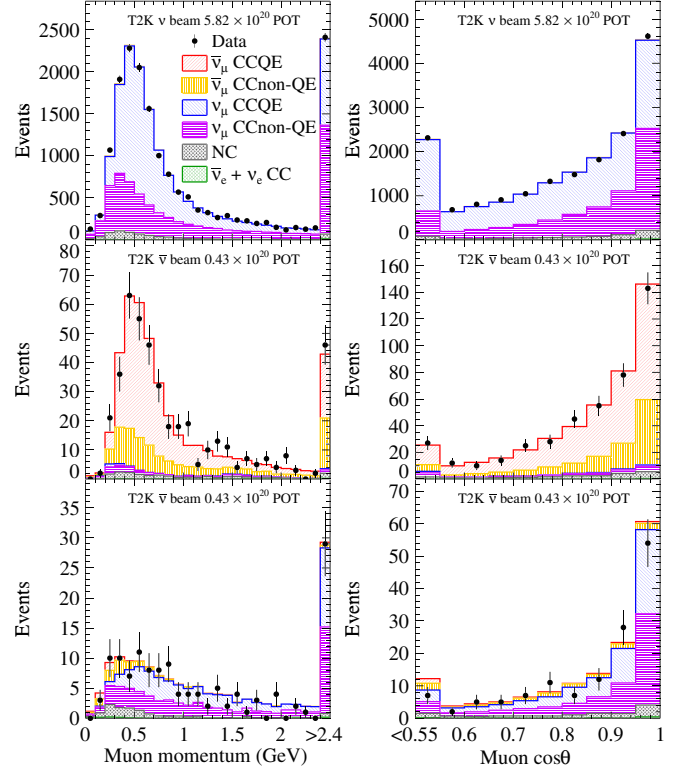


FIG. 2. The momentum (left) and angular (right) distributions of the muon candidates at ND280 from the ν_μ CC 0π (top), the $\bar{\nu}_\mu$ CC 1 track (center) and the ν_μ CC 1 track (bottom) samples. The data are superimposed on the post-fit MC prediction, separated by interaction mode.

these parameters. A conservative (100% uncertainty) ansatz is adopted for the normalization of multinucleon ejection oxygen events, giving a 9.5% uncertainty on the number of events at SK. For the parameters that ND280 can constrain, the fit reduces their effect on the uncertainty on the expected number of events at SK from 9.2% to 3.4%.

Far-detector selection.—At the far detector, fully contained fiducial volume (FCFV) events are selected by requiring no hit clusters in the OD, that the reconstructed interaction vertex is more than 2 m away from the ID wall, and that the visible energy in the event is larger than 30 MeV. The last criterion requires that the amount of Cherenkov light is more than that of a 30-MeV electromagnetic shower.

To enhance the $\bar{\nu}_\mu$ CCQE purity of the sample, selected events must have a single, μ -like Cherenkov ring, no more than one decay electron, and a muon momentum greater than 0.2 GeV [13]. The number of data and MC events passing each selection criterion are shown in Table II and the reconstructed energy spectrum of the 34 selected events is plotted in Fig. 3. The reconstructed neutrino energy is calculated using the muon momentum and production angle, under the assumption that a CCQE interaction occurred on a nucleon at rest. The selection efficiency for $\bar{\nu}_\mu$ CCQE is estimated to be 77% while backgrounds

TABLE II. The number of events observed at the far detector in the antineutrino-beam-mode data after applying each selection cut. MC expectation is calculated assuming oscillations with $\sin^2(\theta_{23}) = \sin^2(\bar{\theta}_{23}) = 0.5$, $|\Delta m_{32}^2| = |\Delta \bar{m}_{32}^2| = 2.4 \times 10^{-3} \text{ eV}^2$, and $\sin^2(\theta_{13}) = \sin^2(\bar{\theta}_{13}) = 0.0257$. The “ $\bar{\nu}_e + \nu_e + \text{NC}$ ” column includes the NC interactions of all the (anti)neutrino flavors. Efficiency numbers are calculated with respect to the number of MC events generated in the fiducial volume (FV interaction).

	Data	Total MC	CCQE $\bar{\nu}_\mu$	CCQE ν_μ	CCnonQE $\bar{\nu}_\mu$	CCnonQE ν_μ	$\bar{\nu}_e + \nu_e + \text{NC}$
FV interaction	...	186.7	17.8	11.4	20.0	36.5	101
FCFV	90	99.7	14.4	8.6	15.1	26.6	35.1
Single ring	50	52.2	14.0	7.7	8.1	8.7	13.8
μ -like	40	39.4	13.8	7.6	7.8	8.0	2.2
$P_\mu > 0.2 \text{ GeV}$	40	39.3	13.8	7.6	7.8	8.0	2.2
$N_{\text{decay}-e} < 2$	34	36.1	13.7	7.5	7.3	5.6	2.1
Efficiency (%)			77.1	65.7	36.6	15.3	2.0

from neutral-current (NC), ν_e , and $\bar{\nu}_e$ interactions are reduced by a factor of 50. The systematic uncertainties in the detector response are evaluated using atmospheric neutrinos, cosmic-ray muons, and their decay electrons [13].

Oscillation fit.—The oscillation parameters $\sin^2(\bar{\theta}_{23})$ and $\Delta \bar{m}_{32}^2$ are estimated using a maximum-likelihood fit to the measured reconstructed energy spectrum in the far detector. All other oscillation parameters are fixed as shown in Table III. Oscillation probabilities are calculated using the full three-flavor oscillation framework [31], assuming the normal mass hierarchy ($\Delta m_{32}^2 > 0$). Matter effects are included with an Earth density of $\rho = 2.6 \text{ g/cm}^3$ [32].

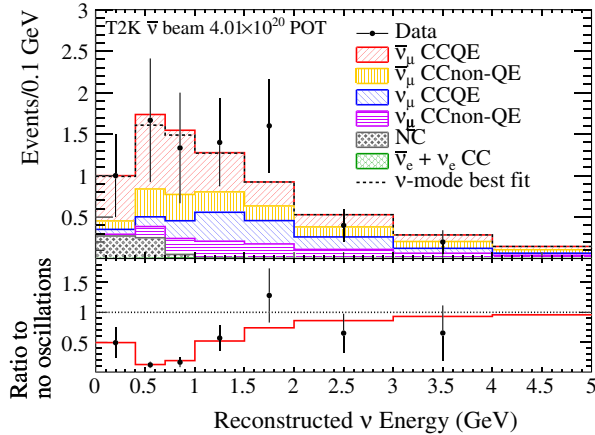


FIG. 3. Top: The reconstructed energy distribution of the 34 far-detector $\bar{\nu}_\mu$ candidates and the best-fit prediction, separated by interaction mode. This is compared to the predicted spectrum assuming the antineutrino oscillation parameters are identical to the neutrino parameters measured by T2K [13]. Bottom: The observed data and $\bar{\nu}_\mu$ -mode best-fit prediction as a ratio to the unoscillated prediction.

TABLE III. Oscillation parameters used for the fit. The parameters $\sin^2(\bar{\theta}_{23})$ and $\Delta \bar{m}_{32}^2$ were allowed to fit in the ranges given. All other parameters were fixed to the values shown, taken from previous T2K fits [13] and the Particle Data Group review [33].

Parameter	ν	$\bar{\nu}$
$\sin^2(\theta_{23})$	0.527	fit 0–1
$\Delta m_{32}^2 (\times 10^{-3} \text{ eV}^2)$	2.51	fit 0–20
$\sin^2(\theta_{13})$		0.0248
$\sin^2(\theta_{12})$		0.304
$\Delta m_{21}^2 (\times 10^{-5} \text{ eV}^2)$		7.53
$\delta_{CP} \text{ (rad)}$		−1.55

Confidence regions are constructed for the oscillation parameters using the constant $\Delta\chi^2$ method [33]. A marginal likelihood is used for this, integrating over the nuisance parameters $\mathbf{f}$ with prior probability functions $\pi(\mathbf{f})$ to find the likelihood as a function of only the relevant oscillation parameters $\mathbf{o}$,

$$\mathcal{L}(\mathbf{o}) = \int \prod_i^{\text{Ebins}} \mathcal{L}_i(\mathbf{o}, \mathbf{f}) \times \pi(\mathbf{f}) d\mathbf{f}, \quad (2)$$

where Ebins denotes the number of reconstructed neutrino energy bins.

We define $\Delta\chi^2 = -2 \ln[\mathcal{L}(\mathbf{o}) / \max(\mathcal{L})]$ as the ratio of the marginal likelihood at a point $\mathbf{o}$ in the $\sin^2(\bar{\theta}_{23}) - \Delta \bar{m}_{32}^2$ oscillation parameter space and the maximum marginal likelihood. The confidence region is then defined as the area of the oscillation parameter space for which $\Delta\chi^2$ is less than a standard critical value. The Feldman-Cousins critical chi-square value was calculated for a coarse set of points in the oscillation parameter space. The difference in the confidence region calculated from these points and that from the standard chi-square values was found to be negligible.

Table IV summarizes the fractional error on the expected number of SK events from a 1σ variation of the flux, cross-section, and far-detector systematic parameters. Although the fractional error on the expected number of events due to

TABLE IV. Percentage change in the number of one-ring μ -like events before the oscillation fit from 1σ systematic parameter variations, assuming the oscillation parameters listed in Table III and that the antineutrino and neutrino oscillation parameters are identical.

Source of uncertainty (number of parameters)	$\delta n_{\text{SK}}^{\text{exp}} / n_{\text{SK}}^{\text{exp}} (\%)$
ND280-unconstrained cross section (6)	10.0
Flux and ND280-constrained cross section (31)	3.4
Super-Kamiokande detector systematics (6)	3.8
Pion FSI and reinteractions (6)	2.1
Total (49)	11.6

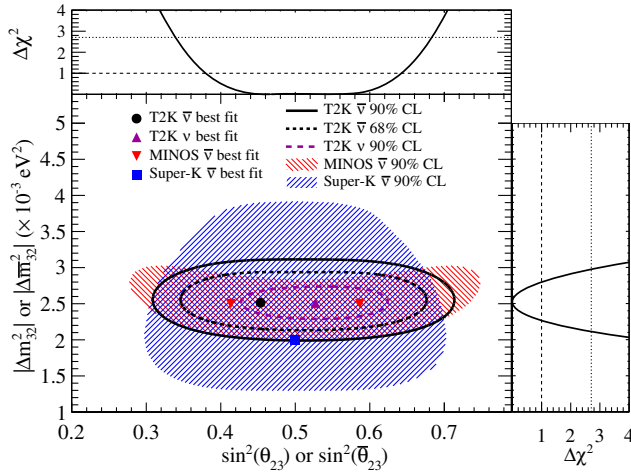


FIG. 4. The 68% and 90% confidence regions for the $\sin^2(\bar{\theta}_{23}) - |\Delta\bar{m}_{32}^2|$ plane assuming normal hierarchy, alongside the T2K ν [13], SK Collaboration $\bar{\nu}$ [5], and MINOS $\bar{\nu}$ [4] 90% confidence regions. One-dimensional $\Delta\chi^2$ profiles for the two parameters are shown at the top and right, overlaid with lines representing the 1D $\Delta\chi^2$ values for the 68% and 90% confidence intervals.

systematic errors is large, the effect of systematic parameters on the confidence regions found in this fit is negligible due to the limited data statistics. The impact of fixing the values of $\sin^2(\theta_{23})$ and Δm_{32}^2 in the fit is also negligible.

The observed $\bar{\nu}_\mu$ reconstructed energy spectrum from the antineutrino-beam-mode data is shown in the upper plot of Fig. 3, overlaid with the best-fit spectrum assuming normal hierarchy, separated by interaction mode. The lower plot in Fig. 3 is the ratio of data to the expected, unoscillated spectrum.

The best-fit values obtained are $\sin^2(\bar{\theta}_{23}) = 0.45$ and $|\Delta\bar{m}_{32}^2| = 2.51 \times 10^{-3} \text{ eV}^2$, with 68% confidence intervals of 0.38–0.64 and $2.26\text{--}2.80 \times 10^{-3} \text{ eV}^2$, respectively. A goodness-of-fit test was performed by comparing this fit to an ensemble of toy experiments, giving a p -value of 0.38.

The fit results are shown in Fig. 4 as 68% and 90% confidence regions in the $\sin^2(\bar{\theta}_{23}) - \Delta\bar{m}_{32}^2$ plane. The 90% confidence regions from the T2K neutrino-beam-mode joint disappearance and appearance fit [13], the SK fit to $\bar{\nu}_\mu$ in atmospheric neutrino data [5], and the MINOS fit to $\bar{\nu}_\mu$ beam and atmospheric data [4] are also shown for comparison. A second, fully Bayesian, analysis was also performed, producing a credible region matching the confidence regions presented above.

Conclusions.—We report the first study of $\bar{\nu}_\mu$ disappearance using an off-axis beam and present measurements of $\sin^2(\bar{\theta}_{23}) = 0.45$ and $\Delta\bar{m}_{32}^2 = 2.51 \times 10^{-3} \text{ eV}^2$. These results are consistent with the values of $\sin^2(\theta_{23})$ and Δm_{32}^2 observed previously by T2K [13], providing no indication of new physics, and are also in good agreement with similar measurements from MINOS [4] and the SK

Collaboration [5]. The results presented here, with the first T2K antineutrino data set, are competitive with those from both the MINOS and SK Collaborations, demonstrating the effectiveness of the off-axis beam technique.

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- [1] Z. Maki, M. Nakagawa, and S. Sakata, *Prog. Theor. Phys.* **28**, 870 (1962).
- [2] B. Pontecorvo, *Zh. Eksp. Teor. Fiz.* **53**, 1717 (1967) [*Sov. Phys. JETP* **26**, 984 (1968)].
- [3] J. Beringer *et al.* (Particle Data Group), *Phys. Rev. D* **86**, 010001 (2012).
- [4] P. Adamson *et al.* (MINOS Collaboration), *Phys. Rev. Lett.* **108**, 191801 (2012).
- [5] K. Abe *et al.* (Super-Kamiokande Collaboration), *Phys. Rev. Lett.* **107**, 241801 (2011).
- [6] K. Abe *et al.* (T2K Collaboration), *Nucl. Instrum. Methods Phys. Res., Sect. A* **659**, 106 (2011).
- [7] D. Beavis, A. Carroll, I. Chiang *et al.* (E889 Collaboration), Brookhaven National Laboratory Report No. 52459, 1995.
- [8] K. Abe *et al.* (T2K Collaboration), *Nucl. Instrum. Methods Phys. Res., Sect. A* **694**, 211 (2012).
- [9] N. Abgrall *et al.* (T2K ND280 TPC Collaboration), *Nucl. Instrum. Methods Phys. Res., Sect. A* **637**, 25 (2011).
- [10] P. Amaudruz *et al.* (T2K ND280 FGD Collaboration), *Nucl. Instrum. Methods Phys. Res., Sect. A* **696**, 1 (2012).
- [11] S. Fukuda *et al.* (Super-Kamiokande Collaboration), *Nucl. Instrum. Methods Phys. Res., Sect. A* **501**, 418 (2003).

- [12] K. Abe *et al.* (Super-Kamiokande Collaboration), *Nucl. Instrum. Methods Phys. Res., Sect. A* **737**, 253 (2014).
- [13] K. Abe *et al.* (T2K Collaboration), *Phys. Rev. D* **91**, 072010 (2015).
- [14] K. Abe *et al.* (T2K Collaboration), *Phys. Rev. D* **87**, 012001 (2013); **87**, 019902 (2013).
- [15] A. Ferrari, P.R. Sala, A. Fasso, and J. Ranft, Reports No. CERN-2005-010, SLAC-R-773, and INFN-TC-05-11, 2005.
- [16] T. T. Böhlen, F. Cerutti, M. P. W. Chin, A. Fassò, A. Ferrari, P. G. Ortega, A. Mairani, P. R. Sala, G. Smirnov, and V. Vlachoudis, *Nucl. Data Sheets* **120**, 211 (2014).
- [17] R. Brun, F. Carminati, and S. Giani, Report No. CERN-W5013, 1994.
- [18] C. Zeitnitz and T. A. Gabriel, in *Proceedings of International Conference on Calorimetry in High Energy Physics, Tallahassee, FL, USA, February 1993* (World Scientific, Singapore, 2014).
- [19] N. Abgrall *et al.* (NA61/SHINE Collaboration), *Eur. Phys. J. C* **76**, 84 (2016).
- [20] Y. Hayato, *Acta Phys. Pol. B* **40**, 2477 (2009).
- [21] Version 5.3.2 of the NEUT library is used, which includes (i) the multinucleon ejection model of Nieves *et al.* [22] and (ii) nuclear long-range correlations for CCQE interactions, treated in the random phase approximation [23].
- [22] J. Nieves, I. R. Simo, and M. J. V. Vacas, *Phys. Rev. C* **83**, 045501 (2011).
- [23] J. Nieves, J. E. Amaro, and M. Valverde, *Phys. Rev. C* **70**, 055503 (2004); **72**, 019902(E) (2005).
- [24] C. Wilkinson, in *Proceedings of 16th International Workshop on Neutrino Factories and Future Neutrino Beam Facilities (NUFACT 2014)*, (Glasgow, Scotland, UK, August 2014).
- [25] C. H. Llewellyn Smith, *Phys. Rep.* **3**, 261 (1972).
- [26] M. Jacob, *Gauge Theories and Neutrino Physics* (North-Holland, Amsterdam, 1978).
- [27] A. A. Aguilar-Arevalo *et al.* (MiniBooNE Collaboration), *Phys. Rev. D* **81**, 092005 (2010).
- [28] A. A. Aguilar-Arevalo *et al.* (MiniBooNE Collaboration), *Phys. Rev. D* **88**, 032001 (2013).
- [29] G. Fiorentini *et al.* (MINERvA Collaboration), *Phys. Rev. Lett.* **111**, 022502 (2013).
- [30] L. Fields *et al.* (MINERvA Collaboration), *Phys. Rev. Lett.* **111**, 022501 (2013).
- [31] V. Barger, K. Whisnant, S. Pakvasa, and R. J. N. Phillips, *Phys. Rev. D* **22**, 2718 (1980).
- [32] K. Hagiwara, N. Okamura, and K. Senda, *J. High Energy Phys.* **11** (2011) 082.
- [33] K. A. Olive *et al.* (Particle Data Group), *Chin. Phys. C* **38**, 090001 (2014).