

# 1 Long-Range Lateral Dyke Propagation is Independent of the 2 Level of Neutral Buoyancy

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## 6 **ABSTRACT**

7 Lateral dyke propagation is a fundamental component of magmatic plumbing systems and a key  
8 process in crustal construction at mid-ocean ridges and large igneous provinces. Classical models  
9 attribute the tendency for lateral magma transport to the Level of Neutral Buoyancy (LNB),  
10 where magma density equals host rock density and vertical ascent stalls, forcing lateral  
11 propagation. While this concept explains many observations in shallow magmatic systems, its  
12 applicability to long-range dyke swarms remains uncertain. The Mull Dyke Swarm (MDS)  
13 located in northwest Britain provides a critical test case. Extending over 600 km from its source  
14 on Mull, the swarm exhibits a consistently shallow upper tipline, sill emplacement near the  
15 surface, and no evidence of eruption along its trajectory. Pressure reconstructions based on  
16 associated sill complexes indicate that magma pressures were sufficient to permit eruption over  
17 the first 150 km, yet lateral propagation continued. Moreover, calculated LNB depths diverge  
18 markedly from observed intrusion geometries. These mismatches suggest that the LNB concept  
19 is neither a necessary nor sufficient control on long-range lateral dyke propagation. Instead,  
20 propagation is better described as a fracture mechanics problem governed by the competition  
21 between magma overpressure, host rock strength, and confining stress. The MDS thus  
22 exemplifies a broader class of magmatic systems in which buoyancy plays a minor role in  
23 fracture propagation. We argue that lateral dyke emplacement on Earth and other planets should  
24 be reframed within the physics of large aspect-ratio hydraulic fractures, rather than buoyancy-  
25 driven ascent.

## 26 27 **INTRODUCTION**

28 Lateral propagation of dykes is a fundamental aspect of magmatic plumbing systems and is  
29 particularly significant for the formation of oceanic crust (Ryan, 1987). This phenomenon has  
30 been studied for dyke intrusion events associated with magma reservoirs at shallow depths  
31 beneath shield volcanoes in intra-plate settings (Rubin and Pollard, 1987) or at mid ocean ridges  
32 (Woods et al., 2019) where the lateral propagation is of the order of 10s of km and aspect ratios  
33 of the blade-like dykes is c. 10:1. Long-range lateral propagation exceeding distances of 100s to

34 1000s of km is, however, widely observed in hundreds of giant radial dyke swarms on Earth,  
35 Venus and Mars with aspect ratios of  $> 50:1$ , but is less well understood (Ernst et al., 1995).  
36 Long-range lateral propagation of dykes requires that they do not break the surface before  
37 reaching their ultimate length (Ernst et al., 1995).  
38 Theoretical models accounting for lateral propagation of dykes have typically attributed the lack  
39 of surface eruption to buoyancy with a central axis of propagation at the level of neutral  
40 buoyancy (LNB) (Lister, 1990) (Fig. 1). The LNB is typically defined at the cross-over in crustal  
41 density and magma density where the magma encounters a region of local gravitational  
42 equilibrium (local magma density equals local wall rock density) (Walker, 1974; Wilson and  
43 Head, 1981; Ryan, 1987). This density contrast is considered to arrest upward propagation and  
44 promote lateral propagation (Ryan, 1987; Lister and Kerr, 1991; Fialko and Rubin, 1999).  
45 Lateral propagation is thus envisaged to be centered on the LNB precisely because this is the  
46 depth where excess magma pressure  $\Delta P$  (difference between magma pressure and the lithostatic  
47 stress) is at its maximum (Fig. 1A) (Ryan, 1987).  
48 Vertically propagating dykes can propagate above the LNB depending on the integrated density  
49 profile over the depth range of the magma column (Fig. 1A-B) (Lister and Kerr, 1991; Taisne  
50 and Jaupart, 2009). Dykes can overshoot the LNB if they have a deep enough source depth from  
51 the dyke initiation position (Wilson and Head, 1981). It is this source depth and magma column  
52 height that allows the excess pressure of the magma to remain positive up to the surface and  
53 therefore leads to surface eruption (Fig. 1A). Too shallow a source, and there is insufficient  
54 buoyant drive gained from denser host rocks to allow the magma to intrude across near surface  
55 rocks of lower than magma density (Fig. 1B) (Walker, 1989).

56 A number of studies have challenged the role of the LNB in magma plumbing in the lithosphere  
57 in diverse tectonic settings and magmatic systems. These include the emplacement of granitic  
58 magmas (Vigneresse and Clemens, 2000), the intrusion and extrusion of intermediate magmas in  
59 arc settings (Loucks, 2014) and the role of volatiles on magma ascent in a range of arc settings  
60 (Rasmussen et al., 2022). Most recently, the commonly stated view that the lateral propagation of  
61 giant dyke swarms is linked to the LNB has been challenged by a study that modelled the magma  
62 pressure distribution in the Mull Dyke Swarm (MDS) in NW Britain (Foschi and Cartwright,  
63 2025).

64 The aim here is to analyze the MDS and its deviation from theoretically expected behavior to  
65 discuss more widely and fundamentally the question of whether the LNB is at all relevant to the  
66 physics of long-range lateral dyke propagation. From this, we argue that the physics of fluid  
67 pressure driven fractures, such as dykes, is solely a contest between magma pressure and rock  
68 strength and that the concept of buoyancy force acting to control propagation only strictly applies  
69 to vertically dominated dyke propagation.

70

## 71 **THE MULL DYKE SWARM**

72 The MDS was intruded during the later stages of the main period of activity in the British and  
73 Irish Igneous Province and is dated at around 58 Ma (Kerr et al., 1999; Chambers and Pringle,  
74 2001). The NW-SE to WNW-ESE trending MDS emanated from a sub-volcanic magma source  
75 on the island of Mull (Fig. 2). Geochemical modelling shows that this source consisted of two  
76 magma reservoirs, a shallow chamber at c. 7-8 km and a deeper one at or close to the Moho  
77 (Morrison et al., 1985; Macdonald et al., 2010, 2015; Ishizuka et al., 2017). Three major sub-  
78 swarms of dykes cross major crustal domain boundaries in their > 500 km ESE trajectory to their

79 terminus in the North Sea, making them amongst the longest continuously mapped dykes on  
80 Earth (Fig. 2A) (Carver et al., 2023). Intrusive volumes for these three sub-swarms range from c.  
81 100 to c. 200 km<sup>3</sup>, implying an almost complete evacuation of the sub-volcanic reservoir,  
82 expressed in the remains of the volcanic superstructure as ring complexes (Carver et al., 2023).  
83 A reconstruction of the simplified geology along the strike of the MDS shows that the dykes  
84 transected major crustal domains with contrasting crustal density profiles (Fig. 2B), from  
85 crystalline dominated rocks to thick sedimentary basin fills with lower densities than the typical  
86 range of basaltic magma densities. Critically, the upper tips of the dykes approached to within a  
87 kilometer or less of the contemporaneous surface along the entire strike length with no evidence  
88 of surface eruption of basaltic magma. Had there been eruption *en-route*, the dykes would almost  
89 certainly have terminated close to the eruption site (Ernst et al., 1995).

90 At least five separate dyke-fed sills or sill complexes were emplaced at shallow crustal levels  
91 along the trajectory of the sub-swarms during dyke propagation (Foschi and Cartwright, 2025).  
92 They used these sills as piezometers to reconstruct the minimum magma pressure required for  
93 their formation and extrapolated upwards to find the excess magma pressure at surface ( $P_0$ ).  
94 They found that  $P_0$  was positive for the first 150 km of the dyke trajectory implying that surface  
95 eruption should have occurred on purely buoyancy grounds (Fig. 2C). Using known crustal  
96 density values, they computed the depth range for the LNB along the path followed by the main  
97 sub-swarms. They found that the LNB was at the surface close to Mull and plunged to ~8 km at  
98 the distal limits of the swarm (Fig. 2B).

99 Here, we were motivated to explore the probable range of upper arrest depths (tipline) if  
100 propagation was driven by buoyancy using the concept of the LNB as a basis for the analysis.  
101 Using a realistic range of magma and crustal densities and a maximum reservoir depth of 10 km

102 (Ishizuka et al., 2017), we computed the range of depths at which dykes would arrest upwards if  
103 the vertical component of dyke propagation was driven by buoyancy (see Supplementary  
104 Information). Importantly, this Monte Carlo simulation showed that these arrest depths are  
105 consistently several kilometers deeper than the inferred depths of the upper tips derived from  
106 seismic interpretation and aeromagnetic modelling, validated by direct well calibration (Carver et  
107 al., 2023; Pryce et al., 2025; Cartwright et al. 2025, their Fig. 9) (Fig. 2D).

108

## 109 **DISCUSSION**

110 Three main arguments can be advanced from the evidence provided by the MDS that lateral  
111 propagation is not dependent on the LNB. Firstly, the geometry of the upper tip of the main  
112 dykes does not follow the reconstructed depth for the LNB along the dyke transect. If lateral  
113 propagation was centered at the LNB as represented in earlier models (e.g., Ryan, 1987; Fialko  
114 and Rubin, 1999), the dyke center line would plunge by up to 8 km over the strike length of the  
115 major dykes (Fig. 2B). This seems highly unlikely given the contradicting geological evidence  
116 showing an almost horizontal upper tipline (Foschi and Cartwright, 2025) (Fig. 2B).

117 Secondly, if magma ascent was purely by buoyancy in this area, then based on the LNB position  
118 and a full range of magma densities, the dykes would have stalled at depths  $> \sim 3\text{--}5.6$  km below  
119 the observed dyke-fed sill complexes in the distal parts of the MDS (Fig. 2D). These sills provide  
120 very good estimates of the minimum magma pressure and indicate upper tip positions 6–8 km  
121 above the LNB and 2.5–4.5 km above the predicted buoyancy-driven arrest depth.

122 Thirdly, proximal to the Mull magma reservoir, the LNB lies above the ground surface for nearly  
123 50 km. If dykes were driven by buoyancy they should have breached the surface and magma  
124 would have erupted. Given that the major dykes failed to erupt and continued to propagate over

125 500 km laterally to the southeast, it is evident that the LNB did not control the propagation  
126 geometry.

127 These arguments demonstrate that the LNB exerted little to no control on dyke emplacement of  
128 the MDS. Instead, we argue that dyke geometry reflects the interplay between magma pressure,  
129 host-rock strength, and confining stress. This interplay of stresses and pressures is instead  
130 consistent with the mechanical frameworks of Rubin (1995) and Buck et al. (2004). As Rubin  
131 (1995) emphasized, “magma within a dyke senses the density of the host rock only insofar as  
132 that density contributes to the ambient stress normal to the dyke plane.” Therefore, for lateral  
133 propagation, the critical parameters are magma pressure distribution and stress heterogeneity, not  
134 buoyancy relative to crustal density structure.

135 The difference in the potential for the LNB to influence a dominantly vertically propagating dyke  
136 as opposed to one that is dominantly laterally propagating is illustrated with a conceptual figure  
137 (Fig. 3A-B). In the vertical case (Fig. 3A-B, location  $X_0$ ), dyke initiation above a sill-like  
138 chamber requires magma pressure ( $P_M$ ) to exceed the total vertical stress ( $\sigma_V$ ). Following Rubin  
139 and Pollard (1987), a lower bound for the reservoir magma pressure required for failure of the  
140 overlying rock units is equal to the total vertical stress of the rocks capping the reservoir,  $\sigma_V$ .  
141 Here, buoyancy and rock-column density govern whether the dyke can reach the surface and  
142 result into a subaerial eruption of magma and a formation of a volcanic edifice (Fig. 3A-B,  $X_0$ ).

143 This model requires also that the LNB is positioned at a depth that allows an excess pressure at  
144 the surface to be greater than zero. Conversely, far away from the volcanic source the pressure at  
145  $X_1$  is independent of the density structure of the rock column above the source even if the actual  
146 density structure has changed over distance. The only condition necessary to sustain lateral dyke

147 propagation is that magma pressure exceeds the minimum horizontal stress, a fraction of the total  
148 vertical stress.

149 In a dominantly laterally propagating dyke, nucleation occurs at the margins of a sill-like magma  
150 chamber where tensile stress concentrates during sill inflation (Gudmundsson, 2020). The  
151 injection pressure ( $P_{inj}$ ) at the chamber margin is related to the vertical confining stress there,  
152 but once the dyke propagates away from the reservoir, the magma pressure at a given depth is  
153 determined by  $P_{inj}$  minus lateral pressure losses due to viscous flow and fracture propagation  
154 (Macdonald et al., 1988; Lister and Kerr, 1991). Consequently, the pressure distribution in a  
155 lateral dyke becomes largely independent of the local overburden density and of the LNB; the  
156 vertical magma gradient is then controlled by magma density, height above the injection point  
157 (e.g., the depth of the sill-like magma chamber), viscous losses, and work done in propagation  
158 against the minimum horizontal stress,  $\sigma_{Hmin}$ .

159 This conceptual model for lateral dyke pressure distributions can be extended to giant radial dyke  
160 swarms on other terrestrial planets (Ernst et al., 1995). Although reservoir dimensions are poorly  
161 known, their diameters are typically small compared to the lateral lengths of the dykes that they  
162 feed (c.f., Wilson and Head, 1992).

163 To reinforce how challenging it would be to explain long-range lateral propagation with a  
164 dominantly buoyant driver, we directly contrast these two propagation mechanisms. In the case  
165 of a buoyancy-driven dyke, vertical ascent stalls at the LNB and the dyke then spreads laterally.  
166 This process requires large density contrasts and buoyancy forces to drive magma hundreds or  
167 thousands of kilometers laterally (Lister and Kerr, 1991). This propagation geometry with an  
168 inverted tear drop shape (buoyant plume) is physically analogous to  $\text{CO}_2$  ascending in a water

169 saturated reservoir and spreading laterally beneath a regional seal (e.g., Huppert and Woods,  
170 1995).

171 In the case of a dominantly laterally propagating dyke, lateral propagation initiates at the  
172 reservoir margin (similar to that shown in Fig. 3A) and is governed from the outset by the  
173 injection pressure and sustained magma supply. The lateral variations in density only affect  
174 resistance to fracture, not the driving pressure.

175 In summary, there is no doubt that buoyancy influences magma plumbing where magma ascends  
176 dominantly vertically. As elegantly argued by Ryan (1987) and Walker (1989) this accounts for  
177 observations of both inactive and active magmatic plumbing systems. However, the concept of  
178 the LNB does not explain our observations of long-range lateral propagation of the MDS.

179 We argue that magmatic plumbing systems cannot all be viewed in the conceptual framework of  
180 buoyant magma ascent. The sources of magma pressure are undoubtedly dominated by vertical  
181 loading onto confined magma source regions, but long-range lateral dyke propagation should not  
182 be analyzed as a buoyancy-driven process. Similar reservations undoubtedly apply to many other  
183 types of intrusive and extrusive magmatism and warrant a renewed research effort directed at  
184 identifying the contributory factors governing magma pressure distributions.

185

## 186 **ACKNOWLEDGMENTS**

187 We are grateful to SLB for provision of an academic license of Petrel. The paper contains  
188 information provided by the Oil and Gas Authority and/or other third parties.

189

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279

280 FIGURE CAPTIONS

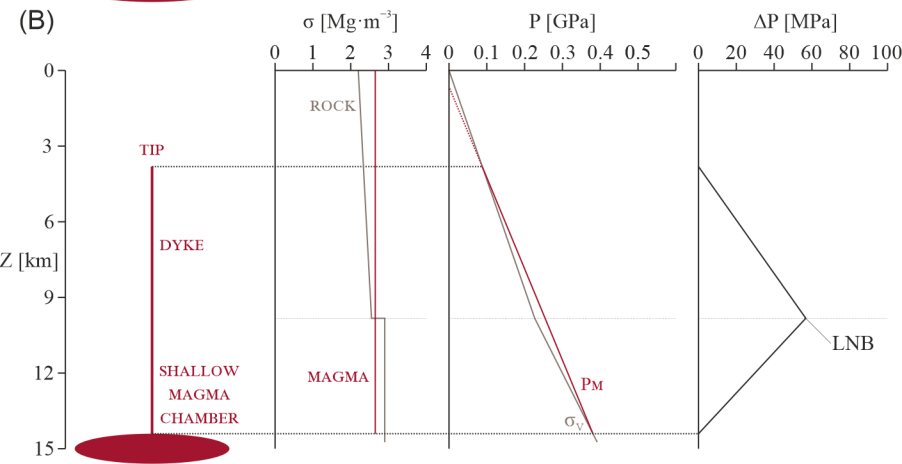
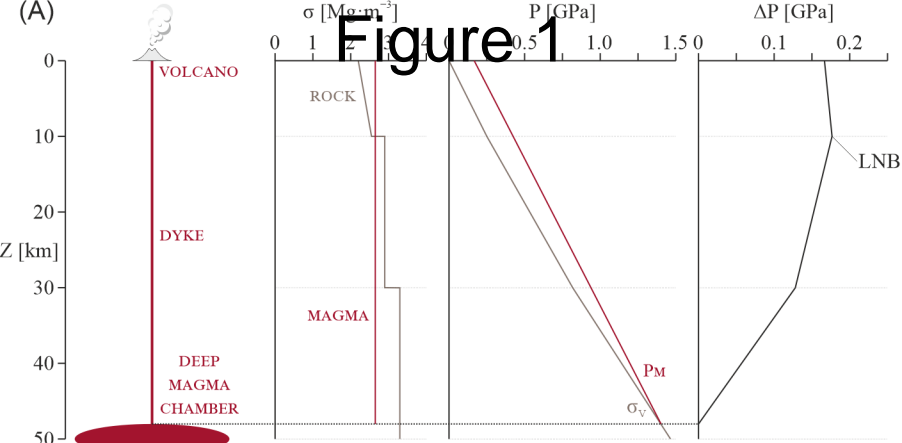
281 **Figure 1.** Definition of the Level of Neutral Buoyancy (LNB) as a function of depth. A: a dyke  
282 sourced within the mantle (here at 50 km below the surface) has enough excess pressure to  
283 overshoot the LNB and erupt magma at surface (modified from Lister and Kerr, 1991). B: a dyke  
284 sourced at a shallow magma reservoir (here at 15 km below the surface) is unable to reach the  
285 surface.

286  
287 **Figure 2.** A: Simplified outline map of the Mull Dyke Swarm from the onshore and offshore  
288 areas over which it extends. Onshore dyke traces are from the British Geological Survey  
289 1:250,000 sheets of N. England and S. Scotland; offshore dyke traces are from Carver et al.,  
290 2023). B: Simplified geology along the approximate mid-position of the MDS in its transect  
291 from Mull to the southern North Sea (modified after Foschi and Cartwright, 2025) with dyke  
292 outline, Level of Neutral Buoyancy (LNB) boxplots at known sill locations, and indicative  
293 continuous depth of the LNB depth over the distance of the dyke. C: Excess magma pressure at  
294 surface  $P_0$  at six locations referenced to distance from Mull. The first 3 locations from Mull are  
295 consistent with subaerial eruption (inconsistent with observations, see text). D: Depth of dyke tip  
296 line relative to distance from Mull based on buoyancy only (see Supplementary Information).

297  
298 **Figure 3.** A: Schematic profile of two dykes emanating from a sill-like magma chamber (source)  
299 from its crest and its lateral margin, with indicative time evolution contours of the dyke body,  
300 depth of the Level of Neutral Buoyancy (LNB), and host rock consistent with increasing density  
301 structure with depth (not to scale). The dyke from the crest results into an eruption and the  
302 formation of a volcanic edifice. The dyke from the lateral margin of the chamber results into a  
303 laterally propagating dyke without eruption. B: Pressure and depth plots and schematic profiles  
304 for location  $X_0$  and  $X_1$ . At  $X_0$  the magma pressure  $P_M$  is larger than the  $\sigma_V$  and results in an  
305 eruption. At  $X_1$  the magma pressure  $P_M$  is larger than the  $\sigma_{Hmin}$  and results in an advancement of  
306 the dyke laterally (into the page). The dyke does not erupt because the pressure is equal to the  
307  $\sigma_{Hmin}$  at its tipline.

308  
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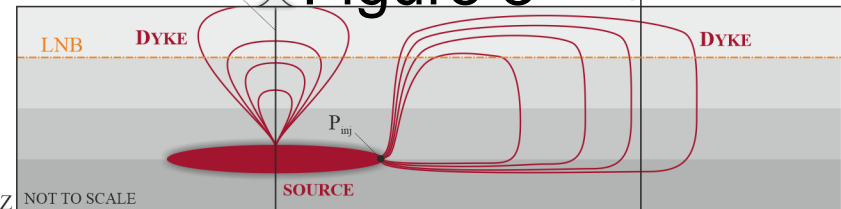






# Figure 3

(A)



(B)

