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3 1 **Cope’s rule and the adaptive landscape of dinosaur body size**
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5 2 **evolution**
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43 19 **Abstract:** The largest known dinosaurs weighed at least 20 million times as
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45 20 much as the smallest, [indicating](#) exceptional phenotypic divergence. Previous
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47 21 studies have focused on extreme giant sizes, tests of Cope’s rule, and
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49 22 miniaturization on the line leading to birds. We use non-uniform
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51 23 macroevolutionary models [based on Ornstein-Uhlenbeck and trend processes](#)
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53 24 to unify these observations, asking: what patterns of evolutionary rates,
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55 25 directionality and constraint explain the diversification of dinosaur body mass?
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3 26 We find that dinosaur evolution is constrained by attraction to discrete body
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5 27 size optima that undergo rare, but abrupt, evolutionary shifts. This model
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7 28 explains both the rarity of multi-lineage directional trends, and the occurrence
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10 29 of abrupt directional excursions during the origins of groups such as tiny
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12 30 pygostylian birds and giant sauropods. Most expansion of trait space results
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14 31 from rare, constraint-breaking innovations in just a small number of lineages.
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16 32 These lineages shifted rapidly into novel regions of trait space, occasionally to
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18 33 small sizes, but most often to large or giant sizes. As with Cenozoic
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20 34 mammals, intermediate body sizes were typically attained only transiently by
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22 35 lineages on a trajectory from small to large size. This demonstrates that
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24 36 bimodality in the macroevolutionary adaptive landscape for land vertebrates
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27 37 has existed for more than 200 million years.
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32 39 **Key words:** dinosaur, body size, Cope's rule, adaptive landscape, Ornstein-
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34 40 Uhlenbeck models, trend models, phylogenetic Bayesian information criterion.
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38 42 CRETACEOUS dinosaurs spanned more than six orders of magnitude in
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40 43 body size, from an estimated 15 g in some birds to at least 40 tonnes (Bakker
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42 44 1971; Anderson *et al.* 1985; Bates *et al.* 2015) and perhaps as much as 90
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44 45 tonnes (Colbert 1962; Mazzetta *et al.* 2004; Benson *et al.* 2014a; Lacovara *et*
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46 46 *al.* 2014; Carballido *et al.* 2017) in sauropods. This represents extraordinary
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48 47 phenotypic divergence from a Triassic ancestor living 140 million years earlier
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50 48 and weighing 10–30 kg (Sereno 1997; Benson *et al.* 2014a). Body size
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52 49 influences many aspects of animal biology including physiology, ecology, and
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54 50 life history energetics (e.g. Brown 1995), so exceptional variation in body size
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signifies considerable variation in biological processes. *Mellisuga helenae* (bee hummingbird), the smallest living dinosaur, weighs 2 g (del Hoyo *et al.* 1999) and extends the range of body sizes achieved by dinosaurs to 20 million-fold, underscoring the evolutionary versatility of dinosaurs, and of the vertebrate bauplan in general.

Significant research has focused on estimating the extremely large body masses attained by some dinosaurs (Colbert 1962; Bakker 1971; Anderson *et al.* 1985; Burness *et al.* 2001; Mazetta *et al.* 2004; Carpenter 2006; Campione & Evans 2012; Benson *et al.* 2014a; Woodruff & Foster 2014; Bates *et al.* 2015) and on framing hypotheses of the physiological, environmental, ecological, and life history factors that made large sizes possible (Alexander 1989; Janis & Carrano 1992; Burness *et al.* 2001; Sander & Clauss 2008; O'Connor 2009; Sander *et al.* 2010; Werner & Griebeler 2011; Sookias *et al.* 2012; Erickson 2014). Because some taxa attained giant sizes, and because much smaller sizes appeared on the evolutionary line leading to birds, quantitative research into dinosaur body size evolution has generally been divided between studies of avian miniaturization (Turner *et al.* 2007; Novas *et al.* 2012; Lee *et al.* 2014; Puttick *et al.* 2014) and studies examining multi-lineage directional trends, especially trends of body size increase ('Cope's rule'; Hone *et al.* 2005; Carrano 2006; Zanno & Makovicky 2013; De Souza & Santucci 2014).

However, this division is artificial. In fact, these topics represent facets of a single, broader goal to characterize patterns of body size evolution in dinosaurs and their underlying macroevolutionary adaptive landscape. Characterization of these patterns for dinosaurs has lagged significantly

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3 76 behind that for land mammals (e.g. Alroy 1999; Smith *et al.* 2012; Saarinen *et*
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5 77 *al.* 2014; Baker *et al.* 2015), which also evolved to a large range of body
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7 78 sizes. Comparison of these patterns provides a test of the hypothesis that a
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9 79 distinct dinosaurian life history (Janis & Carrano 1992; Varicchio 2011)
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11 80 resulted in a unique adaptive landscape that drove idiosyncratic
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13 81 macroevolutionary patterns during the Mesozoic (Codron *et al.* 2012;
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15 82 O'Gorman & Hone 2013).

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18 83 Characterizing patterns of dinosaur body size evolution is also relevant
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20 84 to longstanding questions about how evolution has generated the
21
22 85 phenomenal disparity of organismal phenotypes observed both today and in
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24 86 the geological past (Foote 1997a). For example, palaeontological time series
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26 87 indicate that early rapid increases in disparity are common among major
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28 88 animal groups (e.g. Hughes *et al.* 2013), but it is not clear whether these
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30 89 patterns primarily result from high early rates of evolution (an 'early burst'
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32 90 model; Harmon *et al.* 2010) or from the existence of constraints on the range
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34 91 of phenotypes attainable by a clade, such that trait space becomes rapidly
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36 92 saturated (e.g. Slater 2013; Oyston *et al.* 2015). Dinosaurs have well-resolved
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38 93 phylogenies compared to many fossil groups, and so provide a model system
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40 94 for addressing this question using phylogenetic comparative methods.

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43 95 Here, we use non-uniform, model-based approaches to quantify
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45 96 patterns of dinosaur body size evolution, asking what macroevolutionary
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47 97 processes drove the diversification of dinosaur body sizes during the
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49 98 Mesozoic? These models are non-uniform because they allow multiple
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51 99 macroevolutionary regimes to exist on a phylogeny, each with its own set of
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53 100 model parameters. We specifically compare models based on Ornstein-

Uhlenbeck (OU) dynamics (Hansen 1997; Butler & King 2004; Beaulieu *et al.* 2012) to those based on directional trend-like dynamics (Pagel 2002; Hunt 2008; Hunt & Carrano 2010). These models imply different interpretations of long-term shifts in the body size distribution of species. Trend models encompass a style of macroevolution that is unbounded and based on directional evolution, whereas OU models describe constrained phenotypic divergence within adaptive zones (Hansen 2013). Trend models ascribe long-term directionality to a pervasive tendency for trait values to increase (or decrease) over time, and are supported when set of independently evolving lineages show changes in trait values that go preferentially in one direction over another. This is consistent with many explanations of Cope’s rule, which focus on the broad advantages of ever larger body size (e.g., Brown & Maurer 1986; Van Valkenburgh *et al* 2004; Kingsolver & Pfennig 2004). In contrast, multi-peak OU models construe divergence as resulting from relatively few discrete shifts to new adaptive zones, with constrained evolutionary change occurring within these zones. This is consistent with more nuanced attempts to explain the frequent evolution of large size in land vertebrates (Stanley 1973; Hansen 1997; Alroy 1999). Within this framework, directionality occurs when a shift to a new adaptive zone occurs, and is limited to just a few instances or branches in the clade, rather than reflecting a general evolutionary tendency across multiple lineages.

Abbreviations.

Measurements. FAP, minimum anteroposterior diameter of femoral shaft; FC, minimum circumference around femoral shaft; FL, proximodistal length of

126 femur; FML, minimum mediolateral diameter of femoral shaft; TAP, minimum
 127 anteroposterior diameter of tibia shaft; HAP, minimum anteroposterior
 128 diameter of humeral shaft; HC, minimum circumference around humeral shaft;
 129 HL, proximodistal length of humerus; HML, minimum mediolateral diameter of
 130 humeral shaft; RAP, minimum anteroposterior diameter of radial shaft; RC,
 131 minimum circumference around radial shaft; RL, proximodistal length of
 132 radius; RML, minimum mediolateral diameter of radial shaft; TC, minimum
 133 circumference around tibia shaft; TL, proximodistal length of tibia; TML,
 134 minimum mediolateral diameter of tibia shaft;

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136 *Models of trait evolution.* BM, Brownian motion model; OU, Ornstein-
 137 Uhlenbeck model; OU1, single-peak Ornstein-Uhlenbeck model (single, fixed
 138 θ); OUM, multi-peak Ornstein-Uhlenbeck model (multiple regimes within
 139 individual θ values) with fixed values of α and σ ; OUMV, multi-peak Ornstein-
 140 Uhlenbeck model with fixed value of α and σ as a free parameter; OUMA,
 141 multi-peak Ornstein-Uhlenbeck model with fixed values of σ and α as a free
 142 parameter; OUMVA, multi-peak Ornstein-Uhlenbeck model with both σ and α
 143 as free parameters.

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145 *Model parameters.* α , constraint or attraction parameter of OU model; θ , trait
 146 mean, or 'optimum' of OU model; λ , Pagel's λ , phylogenetic signal parameter;
 147 σ , Brownian variance or rate parameter of a BM or OU model.

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3 149 *Optimality criteria*. AIC, Akaike's information criterion; AICc, Akaike's
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5 150 information criterion for finite sample sizes; BIC, Bayesian information
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7 151 criterion; pBIC, phylogenetic Bayesian information criterion.
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11 153 **METHODS**

12 154

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15 155 *Phylogeny*.

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18 156 All our analyses were conducted using an updated version of the composite
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20 157 phylogeny of Benson *et al.* (2014a; see Appendix S1). Preliminary analyses
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22 158 indicated highly complex patterns of body size evolution in Late Cretaceous
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24 159 theropods. This is consistent with the high frequency of large changes in trait
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26 160 value on single-branches documented among Late Cretaceous theropods in
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28 161 our previous study (Benson *et al.* 2014a). Because of this complexity, when
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30 162 post-Aptian theropods (i.e. those occurring from the late Early Cretaceous
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32 163 onwards) were included in our analyses, meaningful regime shifts could not
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34 164 be recognized. In principle, frequent, large changes in trait values could result
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36 165 from either fast background rates and low constraint (under Brownian motion)
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38 166 or from frequent shifts between short-lived, constrained regimes (under multi-
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40 167 peak Ornstein-Uhlenbeck models; explained below. It should be difficult to
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42 168 distinguish between the two models when regime shifts are frequent, and the
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44 169 equilibrium, constrained phase become difficult to recognise. Therefore we
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46 170 limited analyses of Theropoda to taxa occurring only up to the end of the
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48 171 Aptian. This amounts to analyzing a large time slice, that extends from the
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50 172 Triassic up to the Aptian. Time-slicing of fossil phylogenies is appropriate if
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52 173 patterns of evolution change through time and younger patterns have the
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3 174 potential to over-write the signatures of older patterns. This has been
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5 175 proposed for cladistic biogeographic methods (Hunn & Upchurch 2001;
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7 176 Upchurch & Hunn 2002) and approaches to estimating diversification rate
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9 177 shifts using tree symmetry (Tarver & Donoghue 2011), and is also appropriate
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11 178 here because it is impossible for the evolutionary patterns of lineages
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13 179 occurring in a later interval to actually change those occurring in an earlier
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15 180 interval. Nevertheless, complexity is evident in patterns of post-Aptian
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17 181 theropod evolution (Benson *et al.* 2014a), and we take account of this in our
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19 182 interpretations (see *Discussion*).

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23 183 To reduce the size of the dataset and make our analyses
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25 184 computationally tractable, we subsetting our data as follows: (1) Triassic–
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27 185 Jurassic Dinosauria (intended only to establish the existence of a
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29 186 plesiomorphic dinosaur body size regime); (2) Triassic–Cretaceous
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31 187 Sauropodomorpha; (3) Triassic–Cretaceous Ornithischia; (4) Triassic–Aptian
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33 188 Theropoda. Subsetting and other phylogenetic functions were performed
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35 189 using functions from the package ape version 4.1 (Paradis *et al.* 2004) and
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37 190 phytools version 0.6-00 (Revell 2012) in R version 3.3.3 (R Core Team 2017).

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41 191 Our trees contain polytomies that represent areas of continuing
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43 192 uncertainty in dinosaur phylogeny. To accommodate this uncertainty,
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45 193 analyses were conducted multiple times across a set of 20 phylogenies in
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47 194 which these polytomies were resolved at random. Furthermore, two
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49 195 alternative topologies were used for early sauropodomorphs, those of Yates
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51 196 (2007; Apaldetti *et al.* 2013) and Upchurch *et al.* (2007), which vary in the
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53 197 number of taxa included in a monophyletic ‘prosauropod’ clade (more taxa
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55 198 were found as pectinate outgroups to Sauropoda by Yates [2007]). A second
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199 source of uncertainty concerns the ages of terminal taxa in our tree, which are
200 frequently known only within bounds of several million years. Accordingly, we
201 drew ages for each phylogeny from uniform distributions between the
202 maximum and minimum possible ages for each taxon using a custom script.
203
204 *Time-scaling the phylogeny.*
205 In contrast to the ages of terminal taxa in our phylogeny, which are
206 constrained by the stratigraphy of their occurrences, we can only reconstruct
207 the ages of the nodes in our phylogeny based on indirect information. Various
208 methods have been proposed to assign node ages (= divergence times) to
209 trees of fossil taxa (Bapst 2012, 2013, 2014a; Lloyd *et al.* 2016). Bapst
210 (2014b) used simulations to determine how well these methods performed
211 compared to the true tree when estimating rates and modes of univariate trait
212 evolution, and recommended using the *ca/3* probabilistic method instead of
213 the minimum branch length (mbL) method. Considering only the simulation
214 scenario most similar to our data (Bapst 2014b, fig. 6F: fossils occur as
215 terminal taxa with random times of observation between their apparent first
216 and last appearance dates), the median AICc weight for Brownian motion
217 (compared to OU; models described below) when Brownian motion (BM) was
218 in fact the true generating model was approximately 0.3 when node ages
219 were estimated using mbl, compared to approximately 0.5 when *ca/3* was
220 used. Therefore, *ca/3* is marginally less biased towards supporting OU than is
221 mbl. Note, however, that these simulations represent essentially a best-case
222 scenario for *ca/3*, because phylogenies were simulated under a time-
223 homogeneous birth-death-sampling model (as assumed by *ca/3*), whereas it

224 is likely that real sampling rates, speciation, and extinction rates are highly
225 heterogeneous (e.g. Bapst & Hopkins 2017).

226 Following the above consideration, we are uncertain as to the best
227 approach to estimating divergence times in dinosaur phylogeny. Here, and
228 previously, we calibrated our tree to stratigraphy using the "mb1" (minimum
229 branch length) method (e.g. Laurin 2004; Bapst 2012), setting a minimum
230 branch duration of 1 Ma ("mb1"), which results in post-Palaeozoic divergence
231 times for Dinosauria (Benson *et al.* 2014a). We also perform our initial
232 analyses (i.e. those using SURFACE and comparing to trend-based models)
233 using two other methods (*ca/3* and the extended Hedman method of Lloyd *et*
234 *al.* 2016). All three methods were applied using the R packages paleotree
235 version 2.9 (Bapst 2012; *ca/3*, mbl) and a custom script provided by G. T
236 Lloyd (http://www.graemetlloyd.com/pubdata/functions_7.r).

237 The *ca/3* method uses a birth-death-sampling model (similar to the
238 fossilized birth-death process of Heath *et al.* 2014) to estimate node ages.
239 However, this is only possible when sampling rates, speciation rates, and
240 extinction rates can be estimated *a priori* (Bapst 2013, 2014b). We believe
241 this to be difficult for dinosaurs, in which vanishingly few genera or species
242 have occurrences in multiple time intervals (i.e. most dinosaurs are singleton
243 occurrences without meaningful range data).

244 In spite of this difficulty, Lloyd *et al.* (2016) recently used *ca/3* to
245 estimate node ages for a large phylogeny of dinosaurs. Lloyd *et al.* (2016)
246 obtained their sampling, speciation and extinction rates (following Foote
247 1997b) from the *apparent* range-frequency distribution of dinosaur taxa as
248 represented in the Paleobiology Database (paleobiodb.org/). It is likely these

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249 'range' data used by Lloyd *et al.* (2016) at least indirectly reflect variation in
250 the intervals of stratigraphic uncertainty in the placement of specimens, or the
251 occurrence of wastebasket taxa or species misidentification. We do not
252 advocate using this as a substitute for quality-controlled range data on fossil
253 taxon ranges. Nevertheless, the resulting parameter estimates are
254 qualitatively reasonable and we used them to test the sensitivity of our results
255 to choice of time-scaling method, by calibrating a set of dinosaur trees to
256 stratigraphy using *cal3* (extinction and speciation rate = 0.935; sampling rate
257 = 0.018; D. Bapst, pers. comm. 17 April 2017).

258 Lloyd *et al.* (2016) also presented a modified probabilistic method
259 based on Hedman (2010), which uses the ages of a sequence of outgroups to
260 a node to estimate the age of that node. Lloyd *et al.* (2016) compared the
261 performance of this method to that of *cal3*, demonstrating that it could yield
262 similar estimates of divergence times that are also similar to those obtained
263 by our previous work and here using *mb1* (Benson *et al.* 2014a; e.g. an Early
264 Triassic age for root of Dinosauria).

265 The simulations of Bapst (2014b) suggest that all methods for node
266 age estimation in fossil trees result in a bias towards finding support for the
267 Ornstein-Uhlenbeck (OU) model of evolution relative to Brownian motion
268 (Bapst 2014b), and in weak overestimation of rates of evolution. Furthermore,
269 both *mb1* and *cal3* performed poorly compared to the true tree (median AICc
270 weight for OU of approximately 0.8 when BM is the true model). This occurs
271 because mis-estimation of the phylogeny (including its branch lengths) leads
272 to inflated support for OU (Bapst 2014b; Bapst & Hopkins 2017), and no
273 method for estimating the node ages of fossil trees proposed so far performs

perfectly. Nevertheless, depending on the specific analyses that are carried out, this bias might be considered to be small: an AICc weight of 0.3 for BM (described above) amounts to an AICc score for OU that is only 1.7 points better than that for BM. This is a small difference (Burnham & Anderson 2004). Furthermore, BM is a nested case of OU (the constraint parameter $\alpha = 0$ in BM, but is free to vary in OU), and OU with a near-zero α parameter can be essentially identical to BM (described below), irrespective of its level of support from AICc weights. Therefore, the distinction between the values of α estimated from phylogenies with node ages calibrated using different methods might be a more important than their AICc weights. Because we find extremely strong support for the OU model in most clades, and because we find evidence of generally high values of α , we do not consider that choice of time-scaling method has been influential specifically on our finding of support for OU over BM.

Body mass estimates.

We used the non-phylogenetic versions of scaling equations provided by Campione & Evans (2012) and Campione *et al.* (2014) to estimate dinosaur body masses. These equations estimate tetrapod body mass using the minimum shaft circumferences of the humerus and femur (in quadrupeds) or that of the femur only (in bipeds):

$$[1] \text{ mass}_{\text{quadruped}} = (10^{(2.749 \cdot \log_{10}(\text{FC} + \text{HC}) - 1.104)}) / 1000$$

$$[2] \text{ mass}_{\text{biped}} = (10^{(2.749 \cdot \log_{10}(\text{FC} \cdot 2^{0.5}) - 1.104)}) / 1000.$$

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300 The decision to use non-phylogenetic equations resulted from

301 comparison of non-phylogenetic (i.e. ordinary least squares; OLS) regression

302 models to phylogenetic generalised least squares regression models (Garland

303 & Ives 2001; implemented using the R packages ape version 4.1 and nlme

304 version 3.1-131; Paradis *et al.* 2004; Pinheiro *et al.* 2013; using the tree of

305 Campione & Evans [2012]). Ordinary least squares regression provides a

306 substantially better explanation of the quadrupedal extant tetrapod data than

307 does phylogenetic regression ($AIC_{COLS} = -268$; $AIC_{C_{phylogenetic}} = -232$; AIC_c is

308 Akaike's information criterion for finite sample sizes; Sugiura 1978; Burnham

309 & Anderson 2004). This is consistent with the lack of support for differing

310 relationships of body mass with stylopodial shaft circumferences among

311 different clades and among taxa with different stances (Campione & Evans

312 2012). It indicates either that (1) stylopod shaft circumferences and tetrapod

313 body mass are related to each other via a strong functional linkage that is

314 constrained by the physical laws of the universe, with coefficients that do not

315 vary substantially across the phylogeny (Motani & Schmitz 2011; Campione &

316 Evans 2012), or (2) that the relationship between these variables evolves in a

317 non-Brownian fashion. Examination of the residuals of these relationships

318 supports the former hypothesis (physical constraint: the residuals are

319 homoskedastic and phylogenetically normally-distributed, therefore providing

320 no evidence of non-Brownian dynamics).

321 The ability of stylopodial circumferences to predict live body mass in

322 tetrapods was initially documented by Anderson *et al.* (1985; primarily in

323 mammals) and Campbell & Marcus (1992; in birds). Using a large dataset of

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3 324 extant reptiles and mammals, Campione & Evans (2012) and Campione *et al.*
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5 325 (2014) showed that the combined humeral and femoral circumference is a
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7 326 robust proxy for estimating body mass that is largely independent of
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10 327 phylogenetic history, gait, and limb posture in non-avian tetrapods. Extant
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12 328 birds have a different scaling relationship between femoral shaft
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14 329 circumference and body mass than do other bipedal tetrapods, possibly due
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16 330 to their subhorizontal femoral orientation (Campione *et al.* 2014). However,
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18 331 the body proportions of non-avian dinosaurs and most other Mesozoic stem-
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20 332 group birds indicate that they did not possess the apomorphic femoral
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22 333 orientation of extant birds (e.g. Carrano 1998, 2001; Campione *et al.* 2014),
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24 334 suggesting that the non-avian bipedal tetrapod scaling relationship of
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26 335 Campione *et al.* (2014) is appropriate for estimating the body masses of
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28 336 bipedal non-avian dinosaurs (Campione *et al.* 2014).
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32 Previously, we were able to estimate the masses of either 441 or 426
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34 338 dinosaur specimens (depending on whether questionably facultative
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36 339 quadrupeds were treated as bipeds [requiring less data for their mass
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38 340 estimation] or quadrupeds [requiring more data for mass estimation]). Of
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40 341 these, 310 were included in our phylogeny (Benson *et al.* 2014a). In the
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43 342 present study, we extended our dataset of mass estimates to 584 dinosaur
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45 343 specimens by estimating unknown femoral and humeral minimum shaft
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47 344 circumferences using other limb bone measurements. This was done using
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49 345 AICc-based comparisons to find the best generalized least squares
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51 346 regression model for each combination of variables from among the following
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53 347 options (Appendix S1): (1) varying the strength of phylogenetic signal using
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55 348 Pagel's lambda (Pagel 1999); (2) estimating a non-zero intercept or setting
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349 the intercept to zero; and where relevant (3) including stance (quadrupedal or
350 bipedal) or clade assignment (e.g. titanosaur | non-titanosaur; hadrosauroid |
351 non-hadrosauroid; stegosaur | ankylosaur) as a covariate or interaction term.

352 Sets of model comparisons were conducted across all bipedal
353 dinosaurs, across quadrupedal dinosaurs, and within groups of quadrupedal
354 dinosaurs. Where multiple models of approximately equal goodness were
355 available to predict unknown stylopod minimum shaft circumferences in a
356 single dinosaur specimen, we used the model with the smallest estimated
357 prediction error. The full set of estimates and their prediction errors are
358 provided in Dataset S1.

359 It is important to account for errors in tip values when evaluating OU
360 models, because failure to account for the error in estimated trait values can
361 lead to spurious favoring of OU over BM-like models (Silvestro *et al.* 2015;
362 Cooper *et al.* 2016). However, the calculation of prediction errors of some of
363 our mass estimates was complicated by the fact that the limb shaft
364 circumference measurements upon which they were based were also
365 estimates, with associated prediction errors of their own.

366 We calculated the total error of each mass estimate using simulations
367 that accounted for error propagation through multiple rounds of regression.
368 The standard error of masses estimated directly from femoral and humeral
369 circumferences has two contributing sources: (i) error in estimating that
370 species mean femoral and humeral dimensions, and (ii) the error of the
371 regressions used to estimate individual body masses. For nearly all
372 dinosaurs, this species mean is estimated from a single individual animal. We
373 assume within species that the limb dimensions are normally distributed with

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3 374 a coefficient of variation equal to 5, which translates to a standard deviation
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5 375 on a \log_{10} scale of 0.0217. This is a reasonable magnitude of variation for
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7 376 size-related traits in vertebrates (Yablokov 1974; see also Hunt & Carrano
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9 377 2010, p. 256). Simulations combine this error in limb dimensions with the error
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11 378 in estimating the regression of body mass on limb shaft circumferences.
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13 379 When femoral/humeral circumference measurements were estimated from
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15 380 other variables via regression, the error was propagated through this
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17 381 regression, and then through the regression of the imputed limb data to
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19 382 produce mass estimates. In this way, the mass estimates are analyzed with
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21 383 uncertainties that reflect how they were calculated. These standard errors are
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23 384 incorporated into the likelihood functions in the standard way by increasing
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25 385 the expected variances of the tips by an amount equal to their squared
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27 386 standard error (O'Meara *et al.* 2006; Ives *et al.* 2007). R functions used to
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29 387 compute these standard errors are provided on Dryad (Benson *et al.* 2017).
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34 388 Of our 584 mass estimates, 526 were from adult individuals,
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36 389 representing a total of 393 taxa included in our phylogeny. Only mass
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38 390 estimates of adult individuals were used in our analyses. Skeletal maturity
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40 391 was assessed from published histological studies (e.g., Lee & Werning 2008;
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42 392 Benton *et al.* 2010; Erickson *et al.* 2006, 2009a,b, 2010; Osi *et al.* 2012;
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44 393 Werning 2012) and qualitative indicators such as the fusion of neurocentral
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46 394 and neurocranial sutures.
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52 396 *Ornstein-Uhlenbeck (OU) models.*
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54 397 Our macroevolutionary analyses make use of Ornstein-Uhlenbeck (OU), or
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56 398 'Hansen' models (Hansen 1997; Butler & King 2004; Beaulieu *et al.* 2012),
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399 which include the parameters: Z_0 , the estimated trait value at the root of the
400 tree; β or σ^2 , the Brownian variance, which describes the rate at which trait
401 variance is expected to accumulate along phylogenetic lineages in Brownian
402 motion models (Felsenstein 1985), and is a measure of stochastic spread of
403 trait values over time (Hansen 1997; Beaulieu *et al.* 2012; Hunt 2012); θ , a
404 macroevolutionary trait ‘optimum’; and α , the strength of attraction to θ . Under
405 OU, the expected change in a trait X within an infinitesimal time interval
406 between t and $t + dt$ is $dX(t)$, and:

407
408 [3] $dX(t) = \alpha[\theta - X(t)]dt + \sigma dB(t)$.

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410 Where $X(t)$ is the value of X at time t , and the term $dB(t)$ is a random variable
411 with a mean of zero and variance of $\sigma^2 dt$ (Butler & King 2004; Beaulieu *et al.*
412 2012). This formulation includes a term describing trait attraction towards θ ,
413 which is the product of α and the difference between $X(t)$ and θ :

414
415 [4] $\alpha[\theta - X(t)]dt$.

416
417 It also includes an independent term describing stochastic evolution in the
418 form of Brownian motion (Felsenstein 1985):

419
420 [5] $\sigma dB(t)$.

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422 When $\alpha = 0$, term [4] becomes zero, resulting in Brownian motion (term [5]),
423 as a special case of OU (Fig. 1A–B; e.g. Butler & King 2004; Slater 2013).

Interpretation of OU models can be complicated (Cooper *et al.* 2016) because OU models can also simulate the behavior of several other commonly used macroevolutionary models (Fig. 1):

(1) A 'trend' model in which Brownian motion is modified by the addition of a variable μ , describing the expected amount of directional change in trait values through time (Pagel 2002; Hunt & Carrano 2010). Trend-like behaviour is described by OU models when α becomes small and θ is outside the range of observed trait values, in which case $\alpha\theta$ approximates μ (Fig. 1E; Hansen 1997). Notably, trend models (and other models in which θ is different to Z_0) cannot be identified without the inclusion of fossil data (e.g. Slater *et al.* 2012).

(2) A stasis or 'white noise' model, in which trait values are drawn from a normal distribution with mean θ and a stable variance, independent of the phylogeny (Sheets & Mitchell 2001; Hunt 2006). OU models converge to stasis-like behavior through time when α is high, in which case instantaneous trait values $[X(t)]$ are approximately equal to θ with a variance equal to $\sigma^2/2\alpha$. OU models describe stasis-like behavior from $t = 0$ when α is high and $\theta \approx Z_0$ (Fig. 1C; Hansen 1997).

The parameters of OU models therefore provide information on the mode of evolution. The phylogenetic half-life, $t_{0.5} = \ln(2) / \alpha$, is particularly useful value in this context as it describes the time taken for θ to become more influential than Z_0 in determining trait values within a regime (Hansen 1997; see also Slater 2015).

A key difference between BM and OU is that, under OU, non-zero values of α act to constrain trait values around θ , thereby limiting the

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449 accumulation of trait variance through time (e.g. Hansen 1997; Butler & King
450 2004; Slater 2013). The expected variance in trait values among descendants
451 (living at time = t) of a single common ancestor (that lived at time = 0) is
452 $(\sigma^2/2\alpha)*(1 - \exp(-2\alpha t))$ (Hansen 1997), which asymptotes at $\sigma^2/2\alpha$ when t
453 exceeds several phylogenetic half lives. This contrasts with the linear increase
454 in variance with time under Brownian motion according to $\sigma^2 t$ (e.g. Felsenstein
455 1985; Hunt 2012; and compare Fig. 1A with Fig. 1C).

456

457 *Characterising the macroevolutionary landscape of dinosaur body size—*
458 *SURFACE algorithm.*

459 We used a two-step approach to characterize dinosaur body size evolution.
460 First, we used the R package SURFACE version 0.4-1 (Ingram & Mahler
461 2013). SURFACE implements an approach that locates a set of
462 macroevolutionary regimes characterized by OU models with distinct trait
463 optima (θ) on a phylogeny. To reduce computational demands, SURFACE
464 assumes conserved, single values of α and σ^2 across the entire phylogeny.
465 The locations of regime shifts are estimated using stepwise AICc (AICc =
466 Akaike's information criterion for finite sample sizes; Akaike 1974; Sugiura
467 1978; Burnham & Anderson 2004), without prior specification of how the
468 regimes should be distributed on the phylogeny (Ingram & Mahler 2013).
469 SURFACE initially undertakes a forward phase, first fitting a two-regime
470 model by identifying the best node at which to specify a regime shift using
471 AICc. It then holds the position of that regime shift and iteratively searches for
472 further shifts until no improvement in AICc can be attained. The algorithm then
473 undertakes a backwards phase in which phylogenetic regimes are merged

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3 474 together if there is a resulting improvement in AICc score, allowing the
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5 475 detection of evolutionary convergence (Ingram & Mahler 2013).
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7 476 Ho & Ané (2014) demonstrated the existence of a 'large p small n'
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9 477 problem for fitting multi-regime OU models to comparative data. Because the
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11 478 number of possible shift configurations increases dramatically as shifts are
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13 479 added, and because AIC, AICc, and BIC (Bayesian information criterion) do
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15 480 not address the issue of false positives due to multiple comparisons, the
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17 481 SURFACE algorithm is liberal, and tends to support overly complex models
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19 482 (Ho & Ané 2014; Khabbazian *et al.* 2016; Davis & Betancur-R 2017). This has
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21 483 been forcefully demonstrated for ultrametric trees comprising only extant taxa
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23 484 (Ho & Ané 2014). Nevertheless, adding fossil taxa (i.e. analysing non-
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25 485 ultrametric trees, as done here), improves identifiability of the parameters of
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27 486 OU models (Slater 2013; Ho & Ané 2014), and it is possible that it also
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29 487 facilitates accurate regime shifts determinations. To address the problems
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31 488 with existing optimality criteria such as AICc, Khabbazian *et al.* (2016)
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33 489 proposed a new information criterion, "pBIC" (phylogenetic Bayesian
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35 490 information criterion). pBIC makes use of the effective sample size of taxa
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37 491 providing information about the trait optimum at each node on a phylogeny
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39 492 (Ané 2008), which is often considerably smaller than the number of taxa
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41 493 descended from that node (e.g. Ho & Ané 2014). The pBIC is conservative,
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43 494 with low rates of false positive identification of OU model regime shifts.
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45 495 However, until now it has not been implemented for non-ultrametric trees. We
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47 496 implemented a set of functions that calculate pBIC for SURFACE model fits,
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49 497 and conduct stepwise, SURFACE-like searches using pBIC instead of AICc.
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51 498 These were used to test whether our SURFACE fits were over-parameterized
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499 by (1) Conducting fully-conservative stepwise-pBIC searches; (2) Conducting
500 liberal forward-phase stepwise AICc searches and conservative backward-
501 phase pBIC searches; and (3) By calculating pBIC for the shift configurations
502 returned by (fully-liberal) forward- and backward-phase stepwise AICc
503 searches. Ideally, pBIC would be used for all model comparisons throughout
504 this paper (and others). However, because it is not implemented for most
505 model implementations, we predominantly use AICc for model comparisons,
506 using pBIC only to ensure that AICc does not unduly favour complex model
507 fits for our data. Our pBIC functions are available on Dryad (Benson *et al.*
508 2017).

509

510 *Characterising the macroevolutionary landscape of dinosaur body size—*
511 *OUwie algorithm.*

512 Having identified candidate macroevolutionary regimes using SURFACE, we
513 estimated the full set of parameters (Z_0 , θ , α , σ^2) of those regimes using
514 maximum-likelihood in OUwie version 1.50 (Beaulieu *et al.* 2012). OUwie
515 employs a model-fitting algorithm that potentially allows all key parameters to
516 vary freely, including α and σ^2 (Beaulieu *et al.* 2012; models described in
517 Table 1), and further differs from SURFACE in fixing the locations of regime
518 shifts on the tree a priori. Our SURFACE results suggest that, in general,
519 body size evolution occurs in a step-wise fashion, characterized by substantial
520 values of α , and attraction to a set of distinct optima (θ) in trait space.
521 However, if α actually varied among regimes, then we might incorrectly fit
522 regime shifts to Brownian-like or trend-like portions of our phylogeny when
523 holding α constant across regimes, as done by SURFACE. Furthermore,

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3 524 SURFACE does not currently allow the inclusion of estimated measurement
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5 525 errors, whereas OUwie does allow these to be taken into account. For these
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7 526 two reasons, we used OUwie to estimate the full set of parameters for each
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9 527 regime, and to compare the fits of models in which different sets of
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11 528 parameters were allowed to vary freely, using AICc weights. This allowed us,
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13 529 for example, to determine whether allowing a distinct root node trait value (Z_0)
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15 530 improved upon a model in which this was set equal to the trait optimum (θ) for
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17 531 the regime present at the root of the tree, and whether allowing θ and α both
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19 532 to vary among regimes resulted in a better model than one in which only θ
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21 533 was allowed to vary.
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24 534 Estimating the full set of parameters independently among regimes is
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26 535 computationally intensive (Beaulieu *et al.* 2012). Indeed, our analyses
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28 536 frequently recovered nonsensical parameter estimates for the most complex
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30 537 models (especially for OUMA and OUMVA; Table 1). For this reason the fits
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32 538 of complex models allowing most or all parameters to vary among regimes
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34 539 sometimes had to be discarded, and this was done using the following criteria:
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36 540 (1) Model fits that returned any highly precise (s.e. = 0) or imprecise (s.e. > 2)
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38 541 parameter estimates, except when implying trend-like dynamics with
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40 542 unrealized values of θ and low values of α (as described above; this occurs,
41
42 543 e.g. on the single ornithischian lineage leading to the small-bodied ankylosaur
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44 544 *Struthiosaurus*). (2) Model fits that returned highly erroneous estimates of the
45
46 545 ancestral body mass ($Z_0 < 0$ [= 1 kg] or $Z_0 > 2$ [= 100 kg]), as indicated by
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48 546 comparisons with other analyses (for example, analyses of Triassic–Jurassic
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50 547 Dinosauria were considered to provide robust estimates of the ancestral body
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52 548 mass of Ornithischia that should be replicated by analyses of Ornithischia). (3)
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Entire sets of model fits for phylogenies for which all of the complex OU models (OUMV, OUMA, OUMVA) returned nonsensical parameter estimates by the preceding two criteria. Therefore, we simplified the models prior to analysis by analysing subtrees that contained relatively fewer shifts, broadly comprising Sauropodomorpha, Thyreophora, Marginocephalia, Iguanodontia and pre-Albian Coelurosauria. We also deleted taxa that were characterized by single, terminal-branch regimes not present at any internal branches of the phylogeny. All subtrees analysed included several early dinosaur taxa that provide information on the ancestral body size for Dinosauria: the early saurischians *Pampadromeus*, *Saturnalia*, *Chromogisaurus*, *Staurikosaurus*, *Eoraptor*, *Tawa*, and the early ornithischian *Pisanosaurus*.

Elaborations of the trend model.

The model of BM with a trend, as described above, posits that evolutionary dynamics hold uniformly over time and across branches of the tree (e.g. Hunt & Carrano 2010). However, it is possible that allowing for shifts in the trend parameter, μ , may provide a better account of the macroevolutionary processes operating in a clade, especially given the complexity of body size dynamics previously documented in dinosaurs (Carrano 2006; Benson *et al.* 2014a). We explored two kinds of elaborations of the uniform trend model: allowing for temporal shifts in μ (*time shift* models) and allowing for shifts in μ at specific nodes on the tree (*node shift* models) (Fig. 2). With temporal shifts, trend dynamics are uniform across all branches at any given instant of time, and, when a shift to a new value of μ occurs, it applies to all lineages alive at that time (Fig. 2A). Such a model might offer improvement if, for example,

body size increases are concentrated early or late in the history of a clade. Trends with dynamics that shift at nodes (Fig. 1E–F; Fig. 2B) allow for heterogeneity in body size evolution across subclades, describing a situation in which some clades evolve towards larger or smaller sizes through time whereas others do not.

The trend-based models were fit via maximum likelihood. Under a uniform trend model, tip values from a tree have a joint multivariate normal distribution (Hansen & Martins 1996). The multivariate vector of means, \mathbf{m} , is equal to $Z_0 + \mu \mathbf{t}$, where, \mathbf{t} is a vector of time spans between each terminal taxon and the root of the tree (e.g., t_i is the difference in age between the root of the tree and the i^{th} tip) and Z_0 is the trait value at the root. The covariance matrix among tips is the same as that for Brownian motion, $\sigma^2 \mathbf{V}$, where \mathbf{V} is the matrix that represents shared branch lengths among the tips (Martins & Hansen 1997). Thus, to compute the likelihood for a particular combination of parameters values (Z_0 , μ , σ^2), one calculates \mathbf{m} and \mathbf{V} from the parameters, and then evaluates the density function of the multivariate normal distribution with inputs \mathbf{m} and \mathbf{V} .

The likelihood calculations are only slightly altered when μ varies over time or across branches. Now, the vector of means sums over multiple different μ along the path from the root to a particular tip such that the mean of the i^{th} terminal taxon is

$$m_i = Z_0 + \sum_j \mu_j t_{ij}$$

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598 In which j indexes the (possibly) different trend regimes on the path from the
599 root to the i^{th} tip, with μ_j as the trend parameter and t_j as the duration of time in
600 each regime. The covariance computation is unchanged from uniform BM.

601 We found that trend models with temporal shifts often have multiple,
602 local optima in the likelihood surface corresponding to different combinations
603 of timings for the shifts in the μ parameter. These optima complicate hill-
604 climbing searches, so we used a grid search to explore the space of temporal
605 shift points while using hill-climbing algorithms to find maximum likelihood
606 estimates for the remaining parameters.

607 The positions of branch shifts were identified using a stepwise
608 procedure similar to forward phase of SURFACE. Starting with a uniform
609 trend, a shift in the μ parameter was tested for each internal node of the tree
610 and the node leading to the highest increase in AICc was retained. Next, a
611 second shift point was searched for by testing each remaining node, and so
612 on until no more complex model improves AICc. To simplify the parameter
613 search, it was assumed that the shift occurred at the base of the branch
614 leading to the node being tested.

615 We constrained the Brownian variance (σ^2) to be shared across all
616 regimes in all models. When each regime was allowed to have its own μ and
617 σ^2 parameters, the method would sometimes return shifts at nodes with near
618 zero values for the Brownian variance when two sister taxa happened to have
619 very similar body sizes. Constraining the models to a single σ^2 parameter
620 shared across all trend regimes prevents such unrealistic parameter values.

Functions to fit these trend models were written in custom R code and rely heavily on functions from the ape (Paradis *et al.* 2004) and phytools (Revell 2012) packages. This code is provided on Dryad (Benson *et al.* 2017).

624

625 RESULTS

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627 *Body mass estimates—predicting unknown stylopodial shaft circumferences.*

628 Femoral and humeral minimum shaft circumferences (FC and HC) can be
629 estimated as the perimeters of ovals with diameters equal to the measured
630 minimum anteroposterior and mediolateral bone shaft diameters (equation [3]
631 in Appendix S1). The relationships between these ‘oval circumference
632 estimates’ and measured femoral and humeral shaft circumferences is
633 described best by non-phylogenetic regression models (Table S1 and
634 equations [4]–[5] in Appendix S1), which explain more than 99% of the
635 variance in measured minimum shaft circumferences. These regression
636 models can therefore be safely applied as a ‘correction factor’, allowing us to
637 reliably estimate unknown stylopodial minimum shaft circumferences.

638 However, it is only possible to use this approach when both shaft diameters
639 are known, motivating the search for other relationships to predict FC and HC.
640 These relationships are described below, and were used to generate the full
641 set of mass estimates presented in Dataset S1. More detailed explanations of
642 the model fits are provided in Appendix S1.

643 *Bipedal dinosaurs.* In bipedal dinosaurs, including all theropods, many
644 early sauropodomorphs, and many ornithischians, FC can be predicted using
645 other measurements of the femur and using measurements of the tibia,

demonstrating that bipedal taxa had relatively conservative hindlimb proportions (Table S2; equations [6]–[14] in Appendix S1). These measurements are, in order of predictive strength (as indicated by AICc-weights): femoral shaft minimum mediolateral diameter (FML), tibia minimum shaft circumference (TC), femoral shaft minimum anteroposterior diameter (FAP), and femoral length (FL). AICc weights also indicate that non-phylogenetic models predict FC better than phylogenetic models. Nevertheless, the relationships of FC with FL and with TL include clade assignment to Ornithischia, Sauropoda or Theropoda as a categorical covariate. This suggests that differences in the robustness (length:circumference ratio) of the femoral shaft evolved rapidly during basal divergences among major dinosaur clades, and were subsequently conserved, rather than evolving in a Brownian motion-like way across the phylogeny. The coefficients of the ‘clade’ covariate (equations [9]–[11] in Appendix S1) indicate that ornithischians and sauropodomorphs have proportionally longer tibiae compared to FC than do theropods. Furthermore, relative to shaft circumference, ornithischians have proportionally the shortest femora, and theropods have proportionally the longest femora. Insufficient femoral circumference measurements are available to determine whether, as in extant birds, Mesozoic birds (Avialae) had an abbreviated femur compared to other theropods, but this is unlikely in most cases as only a few Mesozoic taxa within Ornithuromorpha deviated from the typical hindlimb segment proportions of other small-bodied theropods (Benson & Choiniere 2013).

Quadrupedal dinosaurs. Dinosaurs evolved quadrupedal gaits from bipedal ancestors independently within Sauropodomorpha (Cooper 1984;

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3 671 Yates *et al.* 2010; Bonaparte & Pumares 1995; Bonnan 2003; Yates &
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5 672 Kitching 2003; Yates 2007) and the ornithischian clades Thyreophora,
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7 673 Ceratopsia (Chinnery 2004; Chinnery & Horner 2007; Zhao *et al.* 2013) and
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9 674 Iguanodontia (Galton 1970, 1974; Norman 1980, 1986; Maidment *et al.* 2012)
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11 675 (Serenó 1999). Among transitional taxa within Sauropodomorpha,
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13 676 Thyreophora and Ceratopsia, and among iguanodontians, there is some
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15 677 ambiguity as to which are facultative rather than obligate quadrupeds.
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18 678 Furthermore, it is not clear what equation should be used to estimate body
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20 679 mass in facultative quadrupeds/bipeds. We previously estimated the masses
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22 680 of such taxa using the equation for quadrupeds (Campione & Evans 2012)
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24 681 and the equation for bipeds (Campione *et al.* 2014) and found that it made
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26 682 little difference in large-scale analytical studies (Benson *et al.* 2014a).
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29 683 Furthermore, Campione *et al.* (2014) also showed that when estimating the
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31 684 masses of facultative bipedal animals (such as macropods and primates), the
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33 685 quadrupedal equation still performed better than the bipedal correction,
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35 686 supporting use of the quadrupedal equation in those instances. Therefore, in
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37 687 the present work we estimate them all using the quadrupedal equation.

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40 688 Substantial variation in interlimb and intra-hindlimb proportions has
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42 689 previously been noted among groups of quadrupedal dinosaurs (Maidment *et*
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44 690 *al.* 2012). It is not surprising, therefore, that our analyses find intermediate
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46 691 levels of phylogenetic signal, or the importance of clade assignment as a
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48 692 covariate term, when analysing all quadrupedal dinosaur together. This was
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50 693 found for the relationships between stylopodial shaft circumferences and other
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52 694 limb measurements, including those comparing forelimb measurements with
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54 695 hindlimb measurements (Tables S3, S4 in Appendix S1). Nevertheless, there
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696 are some exceptions. Femoral circumferences of quadrupedal dinosaurs can
697 be predicted using non-phylogenetic relationships with either femoral
698 minimum mediolateral shaft diameter (FML) or femoral length (FL), and
699 models for this relationship that include phylogenetic signal or clade
700 assignment as a covariate term have negligible AICc weights (indicating that
701 clade membership is not important for this relationship; Table S3; equations
702 [15]-[16] in Appendix S1). The same is true when humeral minimum
703 mediolateral (HML) or anteroposterior (HAP) shaft diameters are used to
704 predict humeral minimum shaft circumferences (HC) (Table S4; equations
705 [17]-[18] in Appendix S1). These relationships indicate that the proportions of
706 both the femur and humerus are relatively conserved among quadrupedal
707 dinosaurs, even though the proportional lengths of limb segments vary within
708 and among limbs.

709 *Quadrupedal sauropodomorphs.* Many limb measurements provide
710 poor predictions of stylopod minimum shaft circumferences in quadrupedal
711 sauropodomorphs (Tables S5, S6). Nevertheless, FML and HC have strong
712 relationships with FC (Table S5; equations [19]-[20] in Appendix S1), and
713 HML and FC have strong relationships with HC (Table S6; equations [21]-[22]
714 in Appendix S1). The best models of these relationships, according to AICc
715 weights, lack phylogenetic signal. This indicates that the relative
716 circumferences of the femoral and humeral shafts (FC and HC), and the
717 mediolateral diameters of the shafts of these bones (FML, HML) do not
718 change substantially across the phylogeny of quadrupedal sauropodomorphs.

719 *Thyreophoran ornithischians.* Many limb measurements provide poor
720 predictions of stylopod minimum shaft circumferences in thyreophoran

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3 721 ornithischians, regardless of whether phylogenetic or non-phylogenetic
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5 722 regression models are used (Tables S7, S8). Nevertheless, FML and HC
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7 723 have strong relationships with FC (Table S7; equations [23]-[24] in Appendix
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9 724 S1), and FC has a strong relationship with HC (Table S8; equation [25] in
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11 725 Appendix S1). The best models for regression of FC on HC and HC on FC are
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13 726 non-phylogenetic models. Models of the relationship between HC and FC
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15 727 including phylogenetic signal, or assignment to Stegosauria or Ankylosauria
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17 728 as a covariate or interaction term have negligible AICc weights (Tables S7,
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19 729 S8). However, The relationship between FML and FC has strong phylogenetic
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21 730 signal ($\lambda = 1.00$ or 0.98 ; Table S7), indicating that the eccentricity of the femur
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23 731 varies not only between Stegosauria and Ankylosauria, but also among more
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25 732 closely related taxa within those clades.

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27 733 *Ceratopsian ornithischians*. Measurements of both the femur and tibia
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29 734 do a poor job of predicting FC in quadrupedal ceratopsians (Table S9), and so
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31 735 were not used. However, HC and HL have strong, non-phylogenetic
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33 736 relationships with FC (Table S9; equations [26]-[27] in Appendix S1). Humeral
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35 737 circumference (HC) is predicted well by non-phylogenetic relationships with
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37 738 HL and FC (Table S10; equations [28]-[29] in Appendix S1). Overall, these
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39 739 relationships indicate a conserved numerical relationship between FC, HC
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41 740 and HL in quadrupedal ceratopsians.

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43 741 *Iguanodontian ornithischians*. Femoral measurements provide good
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45 742 estimates of FC in quadrupedal iguanodontians, and humeral measurements
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47 743 provide good estimates of HC. By contrast measurements of other bones
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49 744 provide poor estimates (Tables S11, S12; equations [30]-[35] in Appendix S1).
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745 The best model of the relationship between FC and FML is non-
746 phylogenetic (AICc-weight = 0.66), although the best models including a
747 ‘clade’ covariate specifying assignment to Hadrosauroidea or to non-
748 hadrosauroid iguanodontians have non-negligible AICc weights (= 0.14; Table
749 S11). The best model of the relationship between FC and FL includes this
750 ‘clade’ term (AICc-weight = 0.26; Table S11), although models lacking this
751 term have similar AICc weights (best AICc-weight = 0.20). This indicates that
752 hadrosauroids have proportionally longer femora than do non-hadrosauroid
753 iguanodontians (equation [30] in Appendix S1).

754 The best models of the relationships between HC and other humeral
755 measurements are non-phylogenetic regression models. However, a model
756 including a ‘clade’ term, specifying assignment to Hadrosauroidea or to non-
757 hadrosauroid iguanodontians has a non-negligible AICc-weight (= 0.10) for
758 the relationship between HC and HL (best, non-phylogenetic model: AICc-
759 weight = 0.54), and models including strong phylogenetic signal ($\lambda = 1.00$)
760 have comparable AICc weights to non-phylogenetic models of the relationship
761 between HC and HAP. It is of note that igunodontians are the only clade of
762 possibly quadrupedal dinosaurs that do not exhibit a conserved relationship
763 between the femoral and humeral shaft circumferences. This occurs because
764 hadrosauroids have proportionally gracile humeri compared to
765 iguanodontians.

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767 *Characterising the macroevolutionary landscape of dinosaur body size*
768 *evolution.*

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3 770 *Locating macroevolutionary regimes using stepwise AIC*. Comparison of AICc
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5 771 scores for multi-regime OU models fit using SURFACE to those of our trend-
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7 772 based models demonstrates overwhelming support for OU on phylogenies
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9 773 calibrated using mbl1 (Fig. 3), *cal3* and the Hedman method (Figs S1, S2). In
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11 774 fact, trend-based models have comparable AICc weights to single-regime
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13 775 Brownian motion models for Theropoda and Dinosauria on many of our trees
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15 776 (Fig. 3; Dataset S2). In contrast, for Ornithischia, and especially
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17 777 Sauropodomorpha, node shift trend-based models have better (i.e. lower)
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19 778 AICc weights than either Brownian motion or temporal shift models for almost
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21 779 all phylogenies, especially when calibrated using mbl1 (Fig. 3; Dataset S2).
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23 780 Nevertheless, even for these clades, trend-based models perform very poorly
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25 781 compared to OU-based models (Figs 3, S1, S2).

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29 782 Stepwise AICc searches from SURFACE (Ingram & Mahler 2013) for
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31 783 multi-regime OU models recovered different distributions of macroevolutionary
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33 784 regimes across alternative phylogenies used in our analyses (Appendix S2).
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35 785 However, much of this variation was trivial. The only substantive element of
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37 786 variation concerns bimodal distributions of AICc values obtained across
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39 787 phylogenies for Ornithischia, Triassic–Aptian Theropoda, and Triassic–
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41 788 Jurassic Dinosauria (shown for mbl trees in Fig. 4, and for *cal3* and Hedman
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43 789 trees in Appendix S2). These formed two clusters: one with low AICc values
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45 790 and high α , and another with high AICc values and low α (Fig. 4). The
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47 791 populations of high-AICc/small α model fits are characterised by the
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49 792 occurrence of fewer macroevolutionary regimes shifts than are present in the
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51 793 population of better models (e.g. Fig. 5; full results are presented in Appendix
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53 794 S2). The most extreme of these models are very weakly constrained and so
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795 approximate Brownian motion ($\alpha < 0.005$ [small]; phylogenetic half life > 60
796 Ma; $\theta = Z_0$), with a small number of regime shifts capturing ephemeral, high-
797 magnitude trend-like dynamics within some groups ($\alpha = \text{small}$; θ is outside the
798 range of observed trait values) (Fig. 2B).

799 The recovery of two sets of model fits for some groups illustrates the
800 difficulty of fitting complex phylogenetic models to phenotypic datasets. It is
801 not possible to manually explore the full range of candidate models that could
802 be fit to each phylogeny. Nevertheless, because we calibrated the same
803 topologies using each method of divergence time estimation (*mb1*, *cal3*,
804 Hedman), and because the divergence time methods yielded different regime
805 configurations, we are able to show using OUwie that SURFACE recovered
806 suboptimal models fits for at least some of our phylogenies when it returned
807 results in the high-AICc/low- α class (Appendix S3). Focusing on our *mb1*
808 trees: low- α solutions were initially fit to tree topologies 4, 15 and 18 for
809 Ornithischia. However, a high- α solution was found for all these topologies
810 when timescaled using the Hedman algorithm (henceforth: the Hedman
811 regimes), and both the AICc and pBIC scores for the Hedman regimes
812 mapped to the *mb1* timescaled phylogeny were better than for the *mb1* regimes
813 mapped to the same topology using any of the timescaling methods
814 (Appendix S3). The same result was also found for almost all the tree
815 topologies that had initially recovered low- α solutions from stepwise AICc
816 searches on *mb1* trees for Dinosauria (2, 3, 10, 16, 18, 20) and Theropoda
817 (1, 4, 9, 12, 16, 17, 19; and *cal3* regimes provided the best fit to topology 7).
818 Only in a small number of cases did the low- α solution perform better than any
819 other found by our searches either by a substantial (topology 5 for Dinosauria)

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3 820 or negligible margin (topology 6 for Dinosauria; topology 15 for Theropoda).
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5 821 Difficulty finding the best fits might be expected for complex datasets:
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7 822 stepwise optimisation methods are not guaranteed to find the best solution to
8
9 823 complex model fits, and some statisticians have cautioned against their use
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11 824 (e.g. Burnham *et al.* 2011). Our result in general suggest that the high-
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13 825 AICc/low- α class of SURFACE model fits are likely to be suboptimal and are
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15 826 not discussed further here.
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18 827 The low-AICc/high- α class of SURFACE model fits includes relatively
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20 828 large numbers of regime shifts for most clades (5-7 for Sauropodomorpha; 6-
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22 829 10 for Theropoda; 15–18 for Ornithischia; 11–15 for Dinosauria). However,
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24 830 stepwise AICc is known to provide support for overly complex models
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26 831 compared to the conservative pBIC criterion (Ho & Ané 2014; Khabbazian *et*
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28 832 *al.* 2016; Davis & Betancur-R 2017). We therefore also implemented modified
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30 833 versions of the SURFACE algorithm using pBIC in both the forward and
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32 834 backward phases, and using AICc in the forward phase followed by pBIC in
33
34 835 the backward phase (mixed algorithm). Comparable results were found for
35
36 836 Sauropodomorpha by all search methods (Appendix S4; small differences
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38 837 noted below). For the other clades (Dinosauria, Ornithischia, Theropoda), the
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40 838 mixed algorithm finds identical models to our stepwise AICc searches for
41
42 839 almost 100% of cases (Appendix S4), whereas full stepwise pBIC searches
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44 840 find considerably simpler models (Appendix S4) than the low-AICc/high- α
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46 841 class of stepwise AICc model fits. Nevertheless, in all of these cases, the
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48 842 pBIC values of the simpler models are worse than those of the complex
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50 843 models, indicating a failure of the algorithm to sufficiently explore parameter
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52 844 space. These are therefore not discussed further. Visualisation of the search
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845 paths (Appendix S4) shows that the stepwise pBIC searches are unable to
846 cross a central region of the model space characterised by intermediate
847 complexity and suboptimal pBIC, but highly optimal AICc. Stepwise AICc
848 searches can cross this central valley and discover more complex models
849 with highly optimal pBIC scores. In summary, both the AICc and pBIC
850 optimality criteria provide support the selection of complex multi-regime
851 models, although hill-climbing searches using pBIC do not generally discover
852 these models (Appendix S4).

853 Our SURFACE results across multiple phylogenies for each dinosaur
854 group show congruent distributions of macroevolutionary regimes with only
855 minor variations (Appendix S2). The descriptions given here are based on our
856 mbl1 trees. However, model fits to our other trees are shown in Appendix S2,
857 and show few differences. All model fits for Triassic–Jurassic Dinosauria
858 specify a small-bodied ancestral regime, with estimates of θ ($= Z_0$ for
859 Dinosauria) ranging from 14–24 kg (Fig. 6). This ‘basal dinosaur’ regime was
860 inherited by the earliest theropods (e.g. *Coelophysis*, *Staurikosaurus*, *Tawa*),
861 sauropodomorphs (*Chromogisaurus*, *Efraasia*, *Pampadromaeus*, *Saturnalia*)
862 and ornithischians (*Pisanosaurus*, *Scutellosaurus*, *Stormbergia*). This is
863 congruent with estimates of the primitive dinosaurian body mass presented in
864 our previous work (16–24 kg for Dinosauria and 27 kg for Theropoda; Benson
865 *et al.* 2014a, table 2 [Z_0 values]), and is approximately one order of magnitude
866 smaller than the estimated mass of the theropod ancestor presented by Lee
867 *et al.* (2014; 175 kg).

868 Sauropodomorph evolution is characterized by a Triassic regime shift
869 to larger body masses in the grade that includes *Plateosaurus* or *Ruehleia*

870 and taxa that are more closely related to sauropods, such as *Massospondylus*
 871 and *Lufengosaurus* (Fig. 7; 'prosauropods' herein; $\theta = 1,100$ kg to 1,900 kg;
 872 Appendix S2; phenograms shown in Appendix S5), and subsequently to even
 873 larger masses in the clade comprising *Isanosaurus*, *Vulcanodon* or
 874 *Tazoudasaurus*, and all more derived taxa, including *Vulcanodon* (Fig. 7;
 875 'sauropods' herein; $\theta = 15,000$ kg to 17,000 kg). All members of the
 876 'prosauropod' regime became extinct during the Early Jurassic in a size-
 877 selective extinction that was only survived by giant sauropodomorphs of the
 878 sauropod regime. There is some variance within the 'prosauropod' regime due
 879 to the occurrence of smaller body sizes in taxa such as *Anchisaurus* and
 880 *Sarahsaurus*, which are assigned a separate macroevolutionary regime in
 881 some iterations of our SURFACE analyses. Furthermore, the 'sauropod'
 882 regime shift could be located at a slightly more or less inclusive node than that
 883 defined by *Isanosaurus*, and includes terminal single-branch regimes that
 884 model the occurrence of body size reduction in the island dwarf sauropod
 885 *Europasaurus* (1,000 kg), and body size increase in the gigantic taxon
 886 *Argentinosaurus* (95,000 kg), and sometimes in *Ruyangosaurus* (54,000 kg)
 887 (Bonaparte & Coria 1993; Sander *et al.* 2006; Lü *et al.* 2009). Unlike for the
 888 other clades, stepwise pBIC searches for Sauropodomorpha generally find
 889 slightly better models (according to pBIC) than stepwise AICc searches.
 890 These models include slightly fewer regimes by generally including
 891 *Argentinosaurus* (and *Ruyangosaurus*) in the same regime as other
 892 sauropods, and *Anchisaurus* and *Sarahsaurus* with other Late Triassic/Early
 893 Jurassic non-sauropodan sauropodomorphs.

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894 Ornithischian body size evolution (Fig. 8) is characterized by a Triassic
895 shift to small body sizes within Heterodontosauridae ($\theta = 0.7$ to 1.6 kg),
896 Middle–Late Jurassic shifts to larger body sizes in thyreophorans
897 (Stegosauria $\theta = 3,100$ to 11,000 kg; Ankylosauria $\theta = 1,000$ to 1,200 kg) and
898 iguanodontians (which convergently share the regime seen in ankylosaurs in
899 many solutions), and to smaller body sizes in early ceratopsians such as
900 *Psittacosaurus* ($\theta = 6.5$ to 8.4 kg) (phenograms shown in Appendix S5). The
901 Cretaceous saw further shifts to larger body sizes in the ceratopsian clade
902 that includes leptoceratopsids and ceratopsids ($\theta = 190$ to 350 kg; and $\theta =$
903 4,500 to 9,900 kg in Ceratopsidae), and in euhadrosaurian iguanodontians
904 (which convergently share the regime seen in Ceratopsidae in many
905 solutions).

906 Against the background patterns of ornithischian body size evolution
907 (Fig. 8), several single-lineage regimes shifts are seen consistently on the
908 single branches leading to exceptionally large- or small-bodied taxa
909 (Appendices S2, S5): the small Triassic taxa *Lesothosaurus* (6.3 kg) and
910 *Abrictosaurus* (1.4 kg), the large-bodied non-hadrosauroid iguanodontians
911 *Iguanodon* (8,300 kg) and *Muttaborrasaurus* (5,200 kg), the rhabdodontid
912 iguanodontian and possible island dwarf *Mocholodon* (41 kg), the gigantic
913 hadrosaurine iguanodontian *Shantungosaurus* (17,400 kg), the large-bodied
914 thescelosaurid, *Thescelosaurus neglectus* (340 kg), the small-bodied non-
915 ceratopsid ceratopsians *Graciliceratops* (6.7 kg) and *Bagaceratops* (5.7 kg),
916 the giant pachycephalosaur *Pachycephalosaurus* (370 kg), the possible island
917 dwarf ankylosaur *Struthiosaurus* (130 kg), and the largest Early Jurassic
918 ornithischian: the thyreophoran *Scelidosaurus* (650 kg) (Appendix S5; see

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3 919 Benton *et al.* 2010; Osi *et al.* 2012; for discussions of island dwarfism in
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5 920 dinosaurs).

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7 921 The overall pattern in Triassic–Aptian theropod body size evolution
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9 922 (Figs 9–10) starts with a shift from the ‘basal dinosaur’ regime to a regime
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11 923 characterized by larger body size in the clade comprising *Liliensternus* plus
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13 924 the speciose, long-lived clades Tetanurae and Ceratosauria ($\theta = 780$ to 960
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15 925 kg or 130 to 150 kg; explained below), which originated in the Late Triassic.
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17 926 This was followed by a series of shifts towards smaller body sizes on the line
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19 927 leading to birds, within Coelurosauria.

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22 928 The earliest-diverging coelurosaurs, including early tyrannosauroids
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24 929 and taxa such as *Zuolong*, have smaller body masses than those seen in
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26 930 allosauroids, megalosauroids, and many ceratosaurs. In many of our
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28 931 phylogenies, this is identified as a shift to a smaller body size regime at the
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30 932 base of Coelurosauria ($\theta = 120$ to 150 kg) from a primitive large body size
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32 933 regime ($\theta = 780$ to 960 kg) (Fig. 9). However, in other phylogenies the basal
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34 934 coelurosaurian regime is inherited unmodified from the Late Triassic regime (θ
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36 935 = 130 to 150 kg), and the larger body sizes seen in some ceratosaurs, and
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38 936 non-coelurosaurian tetanurans (allosauroids and megalosauroids) represent
39
40 937 three separate evolutionary shifts to a shared large body size regime ($\theta =$
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42 938 1100 to 1200 kg) in those clades (Fig. 10).

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45 939 Subsequently, a shift to a smaller body size regime shared with early
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47 940 dinosaurs occurred in the clade comprising ornithomimosaurs and all more
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49 941 derived coelurosaurs ($\theta = 11$ to 14 kg). A further shift to a smaller body size
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51 942 regime occurred in Paraves ($\theta = 1.0$ to 1.2 kg), consistent with the
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53 943 evolutionary miniaturization event proposed by Turner *et al.* (2007) (Figs 9–
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944 10). This ~ 1 kg regime was inherited by long-tailed, early birds (Avialae) such
945 as *Archaeopteryx* (Appendix S2). A second shift to a smaller body size regime
946 occurred in pygostylian birds ($\theta = 0.093$ to 0.110 kg), and a further shift to an
947 even smaller size regime in the enantiornithine clade including *Iberomesornis*
948 and more derived taxa is found on some phylogenies (Appendices S2, S5;
949 pygostylian $\theta = 0.170$ to 0.180 ; enantiornithine $\theta = 0.053$ to 0.055).

950 An evolutionary shift to a larger body size regime is consistently
951 identified in the herbivorous therizinosaurs ($\theta = 120$ to 150 kg), on the terminal
952 branches leading to the giant Triassic taxon *Herrerasaurus* (270 kg), in the
953 large-bodied dromaeosaurids *Utahraptor* (250 kg) and *Tianyuraptor* (20 kg),
954 and sometimes in the large ornithuromorph bird *Yanornis* (1.4 kg) (Figs 9–10;
955 Appendices S2, S5). In model fits in which tetanurans have a primitively large
956 body size regime, terminal-branch regimes are also required to describe the
957 origins of small-body size in the ceratosaur *Limusaurus* (20 kg) and the basal
958 tetanuran *Chuandongocoelurus* (18 kg) (Appendix S2).

959 *Parameterisations of macroevolutionary regimes.* Fully parameterized
960 multi-regime OU models fit using OUwie (Beaulieu *et al.* 2012), and taking
961 into account the errors associated with our body mass estimates, were
962 generally much better than single-regime Brownian motion models, single-
963 regime OU models, or our trend-based models. This is indicated, for example,
964 by differences in AICc between multi-regime OU models and trend-based
965 models (Fig. 11). Of these comparisons, BM1 was only supported for one tree
966 in Thyreophora (tree 11), and one tree in Theropoda (tree 11), and trend-
967 based models were only supported for tree 9 in Ornithopoda and tree 10 in
968 Theropoda.

969 The solution for theropod tree 10 includes a small, but positive trend of
 970 body mass increase in non-pygostylian coelurosaurs ($\mu = 0.019$ per Myr), with
 971 a transition to weakly decreasing body size in pygostylians ($\mu = -0.022$ per
 972 Myr) (Fig. 12A). Among non-pygostylians, the genus *Microraptor* exhibits a
 973 strong trend towards miniaturization ($\mu = -0.15$ per Myr). Among pygostylians,
 974 the Songlinornithidae, a clade of ornithuromorphs that includes *Yanornis* and
 975 *Yixianornis* in our trees, exhibit a strong trend of body size increase ($\mu = 0.098$
 976 per Myr). The solution for ornithopod tree 9 is typical of the well-supported
 977 trend models for ornithopods (Fig. 12B). The base of the clade experiences
 978 BM-like evolution with a trend parameter very close to zero. There is a shift to
 979 a moderately increasing trend ($\mu = 0.15$ per Myr) within Thescelosauridae,
 980 giving rise to relatively large body sizes in this clade by the end of the
 981 Cretaceous, and a parallel shift to a similar increasing trend ($\mu = 0.10$ per Myr)
 982 near the base of Iguanodontia.

983 The three main macroevolutionary regimes identified by SURFACE for
 984 sauropodomorphs ('basal dinosaur' | 'prosauropod' | 'sauropod') are best
 985 characterized by complex OU-based models such as OUMV ($\alpha > 0$; θ and σ
 986 vary among regimes), OUMA (θ and α vary among regimes) and OUMVA (θ ,
 987 σ and α vary among regimes), with weakly non-negligible fits for variable rate
 988 Brownian motion models (BMS; $\alpha = 0$ and σ varies among regimes) on two
 989 trees (tree 8: AICc-weight = 0.09 for BMS compared to 0.79 for OUMV; tree
 990 15: AICc-weight = 0.11 compared to 0.65 for OUMV).

991 OUMA (θ and α vary among regimes; Table 1) and OUMVA (θ , σ and α
 992 vary among regimes; Table 1) models frequently returned nonsensical
 993 parameter estimates (defined above) and could not be compared to other

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994 models on all trees. In such cases, OUMV (θ and σ vary among regimes;
995 Table 1) was the most complex model that could be tested. Across the results
996 for all trees, a clear pattern is evident in which OUMV, OUMA or OUMVA
997 models of sauropodomorph body size evolution were overwhelmingly
998 supported over simpler models by AICc weights (Fig. 11A; Dataset S3). OUM,
999 OU1, BM1 and BMS models (defined in Table 1) are never present among the
1000 set of non-negligible models (i.e. those with AICc weights at least 10% those
1001 of the best model; Dataset S3). Estimates of α for regimes other than the
1002 primitive dinosaur regime ranged from 0.008 to 0.085, indicating phylogenetic
1003 half-lives ($t_{0.5}$) ranging from 86.6 Ma to 8.25 Ma, which is shorter than the 140
1004 Ma duration of Sauropoda, which exhibit the longest-lived macroevolutionary
1005 regime within Sauropodomorpha. This indicates the predominance of
1006 constrained evolution, in which trait optima (θ) are more influential than
1007 variance (σ) in determining individual body masses. Consistent with this
1008 inference, substantial body size disparity failed to accrue though the entire
1009 evolutionary history of sauropods, and sauropod body size disparity was
1010 remarkably consistent throughout this time (Fig. 13A)

1011 Our analyses of the ornithischian subclades Thyreophora,
1012 Marginocephalia and Ornithopoda provide further support for OU models on
1013 most trees (Dataset S3). For thyreophorans (Ankylosauria + Stegosauria),
1014 OU-based models received overwhelming support from AICc-weights for
1015 almost all trees (BM1 was best supported on tree 11; Fig. 11B), including
1016 OUM, OUMV, and OUMA models with values of α ranging from 0.032 ($t_{0.5}$ =
1017 22 Ma) to 0.089 ($t_{0.5}$ = 7.8 Ma) for regimes other than the primitive dinosaur
1018 regime. For marginocephalians (Ceratopsia + Pachycephalosauria), OUM

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3 1019 models with values of α ranging from 0.012 ($t_{0.5} = 58$ Ma) and 0.172 ($t_{0.5} = 4.0$
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5 1020 Ma) received overwhelmingly the best AICc-weights. Although OU-based
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7 1021 models also generally provided the best explanation of ornithopod body size,
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9 1022 many of these had values of α that were approximately equal to zero, and
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11 1023 BMS (variable-rate Brownian motion) models received the best AICc-weights
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13 1024 for trees 8 and 20, and non-negligible AICc-weights for tree 2 (Dataset S3).
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16 It was difficult to fit models to theropod body size evolution using the
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18 1026 entire tree of Theropoda, and many preliminary searches yielded nonsensical
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20 1027 parameter values. However, models were fit successfully for six trees, using
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22 1028 the portion of the tree that includes ornithomimosaur and all more derived
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24 1029 coelurosaurs, excluding *Utahraptor* and therizinosaurs (when these were
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26 1030 placed in the same regime as *Utahraptor* by SURFACE, as for most trees),
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28 1031 which were identified by SURFACE as having distinct, large-bodied tip
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30 1032 regimes. Model comparisons indicate strong support for complex OU-based
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32 1033 models (OUMA and OUMVA) across at least this portion of the theropod tree.
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34 1034 Values of α range from 0.014 ($t_{0.5} = 50$ Ma) to 1.165 ($t_{0.5} = 0.6$ Ma) (Dataset
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36 1035 S3).
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42 43 1037 **DISCUSSION**

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47 1039 Phylogenetic studies of trait evolution in extinct groups have largely focused
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49 1040 on two questions. (1) Are patterns of trait evolution consistent with niche-filling
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51 1041 models of adaptive radiation (e.g. Slater 2013, 2015; Benson *et al.* 2014a;
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53 1042 Close *et al.* 2015; Hopkins & Smith 2015; Stubbs & Benton 2016;
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55 1043 Cantalapiedra *et al.* 2017)? And (2) do phylogenetic lineages of animals,
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1044 especially of tetrapods, collectively show a directional tendency to increase in
1045 body mass through evolutionary time (i.e. Cope's rule: Stanley 1973; McShea
1046 1994; Van Valkenburgh *et al.* 2004; Hone & Benton 2005; Carrano 2006; Hunt
1047 & Roy 2006; Hunt & Carrano 2010; Benson *et al.* 2012; Sookias *et al.* 2012;
1048 Zanno & Makovicky 2013; Benson *et al.* 2014*b*; Slater *et al.* 2017)? Our
1049 results provide a single framework under which both questions can be
1050 addressed.

1051 In this discussion, we initially focus on previous hypotheses of trend-
1052 like dynamics, including Cope's Rule and evolutionary patterns along the
1053 avian stem lineage. We then consider the broader patterns of dinosaurian
1054 body mass evolution and their implications for adaptive landscapes, and the
1055 roles of niche-filling models and constraint in generating observed patterns of
1056 phenotypic disparity.

1057
1058 *Cope's Rule.*

1059 Given the astounding sizes reached by some Mesozoic dinosaurs, it is not
1060 surprising that there has been extensive interest in testing Cope's Rule (or
1061 Depéret's Rule; sensu Simpson 1953) in the group (Sereno 1997; Hone &
1062 Benton 2005; Hone *et al.* 2005; Carrano 2006; Butler & Goswami 2008; Hone
1063 *et al.* 2008; Hunt & Carrano 2010; Sookias *et al.* 2012*a*; Zanno & Makovicky
1064 2013). Edward D. Cope informally discussed a general accumulation of sizes
1065 over evolutionary time (Cope, 1887), and Cope's Rule has commonly been
1066 defined as an active directional trend towards larger body sizes (e.g. Depéret
1067 1907; McShea 1994; Carrano 2006). 'Active' trends involve similar directional
1068 changes occurring simultaneously among sets of independent lineages.

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3 1069 Under this paradigm, the observation of a large sample of lineages, most of
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5 1070 which show change in the same direction, provides statistical evidence of
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7 1071 selection towards larger size through time (McShea 1994). The outcome of
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9 1072 this process therefore involves increases in both the minimum and maximum
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11 1073 body sizes of descendants (McShea 1994). 'Passive' expansion instead
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13 1074 involves an increase in the total range of trait values seen among the
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15 1075 descendants of a small-bodied ancestor due to non-directional, diffusive
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17 1076 changes in body size. This model, which resembles Brownian motion also
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19 1077 involves increased maximum body sizes. However, it differs from an 'active'
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21 1078 trend model because the minimum body sizes of descendants are decreased
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23 1079 or unaltered (McShea 1994). Alroy (2000) has noted that a wider range of
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25 1080 dynamics are possible, a point with which we agree (and see Carrano 2006).
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29 1081 Previous studies on non-avian dinosaurs have recovered support for
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31 1082 overall increases in mean body size through time (Sereno 1997; Hone *et al.*
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33 1083 2005), and detailed study has indicated that this generally resulted from
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35 1084 passive expansion (Carrano 2006; Sookias *et al.* 2012). Evidence of active
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37 1085 trends, embodied by 'trend-based' models here, has been scarce, although it
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39 1086 has been demonstrated among ornithischians (Hunt & Carrano 2010; and
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41 1087 ornithopods [herein]; and also in pterodactyloid pterosaurs; Benson *et al.*
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43 1088 2014b).

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47 1089 Notably, few of our analyses find support for unconstrained, 'active'
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49 1090 trend models in dinosaur body mass evolution, which classically have been
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51 1091 associated with 'Cope's rule'. Furthermore, where it occurs, this support is
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53 1092 equivocal (Fig. 11). Nevertheless, we also reject strictly uniform passive
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55 1093 expansion (i.e. Brownian motion dynamics). Instead, our results provide
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1094 strong support for multi-peak Ornstein-Uhlenbeck models, which describe the
1095 exploration of a macroevolutionary landscape by phylogenetically defined
1096 regimes that undergo constrained evolution around distinct trait optima (e.g.
1097 Hansen 1997, 2013; Butler & King 2004; Beaulieu *et al.* 2012; Ingram &
1098 Mahler 2013; Mahler & Ingram 2014; Lapiedra *et al.* 2013).

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1100 *Miniaturisation on the avian stem lineage.*

1101 Modern and Mesozoic birds (Avialae) include small-bodied taxa often
1102 weighing only tens of grams. Small body size is likely associated with the
1103 origin of bird flight, as the size of aerodynamic surfaces scale with the square
1104 of linear dimensions, whereas body mass scales with their cube. This makes
1105 it easier for smaller animals to fly. Therefore, patterns of body size
1106 miniaturisation on the avian stem lineage have received significant research
1107 attention (Turner *et al.* 2007; Novas *et al.* 2012; Benson *et al.* 2014a; Lee *et*
1108 *al.* 2014; Puttick *et al.* 2014). Previous studies have documented the origins of
1109 body masses around 1 kg in early paravians such as *Archaeopteryx* (Turner
1110 *et al.* 2007), and also present in several other dinosaur lineages (e.g. Butler *et*
1111 *al.* 2009; Novas *et al.* 2015). In fact, 1 kg is large for extant bird species,
1112 which have a modal body mass around 100 g (Brown 1995). Such small body
1113 sizes are absent among adult dinosaurs (including *Archaeopteryx*), other than
1114 in Pygostylia (and possibly also Alvarezsauroidea; Table 2). Therefore,
1115 because of a focus on *Archaeopteryx*, evolutionary patterns associated with
1116 the attainment of uniquely bird-like small body masses in Mesozoic dinosaurs
1117 have not been addressed until now.

1118 We find that avian miniaturization resulted from a serial pattern of shifts
 1119 to macroevolutionary regimes with smaller body size optima (Figs 9–10). This
 1120 model describes a stepwise, rather than gradual, pattern of body size
 1121 decreases along the avian stem lineage (also documented by Novas *et al.*
 1122 2012). Regime shifts are concentrated during two intervals: (1) the Early
 1123 Jurassic/Middle Jurassic boundary interval, which saw decreases in body size
 1124 from the tetanuran or coelurosaurian ancestor, culminating in a paravian
 1125 ancestor weighing an estimated 1 kg (e.g. Turner *et al.* 2007), and (2) the
 1126 latest Jurassic or Early Cretaceous, giving rise to considerably smaller body
 1127 sizes in Pygostylia. Most Early Cretaceous pygostylians have body masses in
 1128 the range of 13 g (*Iberomesornis*) to 307 g (*Pengornis*) (Figs 9–10), similar to
 1129 those of many extant birds, but smaller than all other dinosaurs.

1130 Lee *et al.* (2014) reported sustained, gradual evolutionary body size
 1131 reduction along the avian stem lineage, from an ancestral theropod dinosaur
 1132 weighing 175 kg to masses of 1 kg in Avialae. This contrasts strongly with our
 1133 findings, which indicate a stepwise pattern and involve at least some body
 1134 size increases during early theropod evolution, from an ancestor weighing 10–
 1135 30 kg (Figs 9–10).

1136 In fact, the higher estimated body mass obtained by Lee *et al.* (2014)
 1137 for the theropod ancestor is an artefact of incomplete and biased sampling of
 1138 early dinosaur taxa in their analysis. As evidence of this: we can replicate their
 1139 result by imposing their taxon sample on our dataset. Lee *et al.* (2014)
 1140 selected only two Triassic dinosaurs for inclusion in their analysis, both of
 1141 which were theropods and have anomalously high body masses compared to
 1142 other early dinosaurs (*Herrerasaurus*, 274 kg, *Liliensternus*, 84 kg; Table 3;

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1143 and see Appendices S2, S5). Most other Triassic theropods have masses of
1144 30 kg or less (i.e. *Coelophysis*, *Staurikosaurus*, *Tawa*; excepting
1145 *Gojirasaurus*; Table 3), and similar small body masses are seen among the
1146 earliest sauropodomorphs and ornithischians (Table 3: Carnian). These small-
1147 bodied early dinosaurs are informative outgroups to Theropoda, but were
1148 omitted from the analysis of Lee *et al.* (2014).

1149 Maximum-likelihood estimation (Felsenstein 1973; Schluter *et al.* 1997;
1150 Paradis *et al.* 2004) using our dataset and a pruned tree of theropod
1151 dinosaurs that excludes the smaller-bodied Triassic dinosaur taxa (mimicking
1152 the analysis of Lee *et al.* 2014), finds the mass of the ancestral theropod to be
1153 253 kg (95% confidence interval: 45.9–1390 kg). This is similar to the value
1154 obtained by Lee *et al.* (2014), but differs from the 10–30 kg estimate obtained
1155 by our full analysis, and clearly demonstrates the importance of representative
1156 taxon sampling in comparative analyses of trait evolution. Our finding that the
1157 ancestral theropod was much smaller than proposed by Lee *et al.* (2014)
1158 clearly refutes the hypothesis of sustained, directional body size reduction
1159 from this ancestral theropod dinosaur along the lineage leading to birds.

1160 *Paravian miniaturization event.* Miniaturization in early paravians was
1161 first proposed by Turner *et al.* (2007; see also Novas *et al.* 2012; Benson *et*
1162 *al.* 2014a; Puttick *et al.* 2014). This miniaturization resulted from several
1163 downwards regime shifts from ancestral tetanurans or coelurosaurs, which
1164 were characterized by larger body sizes (Figs 9–10). The body masses of
1165 tetanuran and coelurosaurian ancestors are uncertain: our multi-peak OU
1166 models suggest two alternative possible optima for the primitive tetanuran
1167 macroevolutionary regime. The first possibility suggests a relatively larger

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3 1168 optimum ~1,000 kg, with a shift to a smaller optimum within Coelurosauria.
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5 1169 This is congruent with classical notions of theropod body size evolution (Fig.
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7 1170 9; Novas *et al.* 2012). Alternatively, a smaller optimum of 150 kg is possible
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9 1171 among early tetanurans, and was maintained in coelurosaurs. Under this
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11 1172 second possibility, the larger body sizes observed in non-coelurosaurian
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13 1173 tetanurans such as allosauroids and megalosauroids represent independent
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15 1174 instances of clade-specific body size increases (Fig. 10).
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18 1175 Miniaturization to sizes of ~1 kg is not unique to paravians, or even to
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20 1176 theropods. Similar body masses evolved under similar evolutionary dynamics
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22 1177 in heterodontosaurid ornithischians (Butler *et al.* 2010), the ceratopsian
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24 1178 *Graciliceratops* (Maryanska & Osmólska 1975), mononychine theropods
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26 1179 (which have small body sizes but occur after the Aptian and so could not be
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28 1180 analysed in our study), and also the early possible tetanuran *Chilesaurus* (~ 1
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30 1181 kg; immature individual; Novas *et al.* 2015) (Table 2). *Chilesaurus* was not
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32 1182 included in our analyses because adult mass estimates are not available.
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34 1183 However, although another small-bodied basal tetanuran,
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36 1184 *Chuandongocoelurus* (18 kg) is currently accommodated by our models as
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38 1185 the product of a lineage-specific miniaturization event (Appendices S2, S5).
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40 1186 Nevertheless, it is still possible that continued sampling of the Early–Middle
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42 1187 Jurassic theropod fossil record will indicate much smaller body size for the
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44 1188 earliest tetanurans, more congruent with the masses of these apparently
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46 1189 exceptional, small-bodied taxa. This is especially probable given the known
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48 1190 biases against preservation of small-bodied species in at least parts of the
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50 1191 dinosaurian fossil record (Brown *et al.* 2013).
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3 1192 *Pygostylian miniaturization event*. The second set of either one or two
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5 1193 regime shifts on the line leading to birds occurred during the latest Jurassic or
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7 1194 Early Cretaceous. This culminated in body sizes predominantly in the range of
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9 1195 15 to 300 g in Pygostylia (Figs 9–10; Appendixes S2, S5). This event has
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11 1196 significance as the evolutionary origin of the modal body size observed
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13 1197 among extant birds (100 kg; Brown 1995), and is unique among dinosaurs
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15 1198 (i.e. such small body sizes only occur within Pygostylia; Fig. 14F).
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17 1199 Furthermore, small body size in Pygostylia co-occurs alongside several other
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19 1200 features of the skeleton and integument that improve flight ability (O'Connor *et*
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21 1201 *al.* 2011), and with relaxed constraints on hindlimb morphology (Benson &
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23 1202 Choiniere 2013). Pygostylian miniaturization, and not paravian miniaturization
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25 1203 to 1 kg, might therefore be classed among the key features that facilitated
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27 1204 evolutionary advancements in powered flight on the avian stem lineage.
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31 1205 Apparently in contradiction of our results, Puttick *et al.* (2014)
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33 1206 demonstrated that Avialae shared a macroevolutionary regime with Paraves.
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35 1207 If this were correct, then we would expect Mesozoic birds to occupy the same
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37 1208 range of body sizes as their close dinosaurian relatives. This is not the case.
38
39 1209 Instead, pygostylian body masses are distinct from those of other paravians
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41 1210 (430 g [*Microraptor*] – 600 kg [*Austroraptor*]; Tables 2, 4). Puttick *et al.* (2014)
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43 1211 rejected the possibility of a distinct regime within Avialae based on tests of a
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45 1212 set of a priori hypotheses that did not include the possibility of a distinct
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47 1213 pygostylian regime, as detected here by our exploratory SURFACE searches.
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49 1214 We recommend as good practice the examination of datasets using
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51 1215 exploratory methods during a hypothesis-formulation phase in future studies.
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1217 *Explaining dinosaur body size diversification.*

1218 Dinosaur body size evolution is best explained by multi-regime OU models;
1219 models applying a single regime to the entire phylogeny are poorly supported.
1220 Our results therefore suggest that dinosaur body mass evolution can be
1221 described by a set of constrained models within phylogenetically determined
1222 regimes, each characterized by attraction to a distinct macroevolutionary body
1223 size optimum.

1224 A recent paper cautioned users against uncritical acceptance of results
1225 that favour OU models (Cooper *et al.* 2016). When they examined the
1226 performance of statistical tests of OU versus BM, these authors reported
1227 elevated rejection rates for BM in favor of OU, especially when error in tip
1228 values was not accounted for. Given that we accounted for error in body mass
1229 estimates in a realistic manner, we do not expect that this bias applies to our
1230 analyses. Nevertheless, Cooper *et al.* (2016) also noted that, in some cases,
1231 OU models were statistically favored by model comparison, but the resulting
1232 estimates of the α parameter implied evolutionary dynamics that were only
1233 trivially different from BM. This observation is consistent with the finding of Ho
1234 & Ané (2014), that low values the constraint parameter α (i.e. values that
1235 imply a long phylogenetic half-life relative to the total tree height, and
1236 therefore Brownian motion-like dynamics) were likely to be over-estimated, or
1237 falsely found to be different from zero. However, examination of the α
1238 parameter estimates (and the corresponding phylogenetic half-lives) of our
1239 model fits indicated dynamics substantially different from BM, a conclusion
1240 that is supported by the very large AICc differences we almost always found
1241 between OU and BM or trend models.

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1242 Shifts in the optima of our OU models cause transient, strongly
1243 directional evolution bridging order-of-magnitude body size transitions (i.e.
1244 large changes in θ). Because the statistical power to detect regime shifts
1245 depends more strongly on the size of change in θ between regimes than on
1246 counts of taxa analysed (Ho & Ané 2014), it is possible that regime shifts with
1247 smaller changes of θ exist, but were not detected. Therefore, our results
1248 provide a summary of only the larger events in the evolution of dinosaur body
1249 size. The distribution of dinosaur body sizes can be described as the
1250 distribution of originations of descendent regimes from ancestor regimes
1251 through time, the direction of body size change involved in regime
1252 originations, and the pattern of regime extinctions through time. We examined
1253 these patterns to explain body size diversification in Dinosauria. In doing so,
1254 we made a distinction between regime shifts that led to clades containing
1255 multiple taxa, which often represent long-lived clades (Fig. 14A,B: large
1256 symbols), and ‘singleton’ regimes containing only single taxa, which may
1257 represent transient instances of gigantism or dwarfism as an evolutionary
1258 response to local environmental conditions (Fig. 14A,B: small symbols).
1259 Originations giving rise to larger-bodied descendant regimes are more
1260 frequent (11 occurrences) than those giving rise to smaller-bodied descendant
1261 regimes (six occurrences) (Fig. 14A: large symbols representing originations
1262 of multi-taxon regimes). This is common in ornithischians, and especially in
1263 sauropodomorphs (for which transitions to small-bodied multi-taxon regimes
1264 are absent). By contrast, increases are approximately as common as
1265 decreases in pre-Albian theropods, a pattern that is entirely due to the

1266 occurrence of successive, stepwise decreases on the lineage leading to birds
 1267 (Figs 9–10).

1268 *Constraint-breaking pulses of body size evolution generated most*
 1269 *dinosaurian body size disparity.* Ornstein-Uhlenbeck models describe
 1270 constrained patterns of evolution, in which the total disparity obtainable within
 1271 a regime is limited. For dinosaur body mass, the expected asymptotic trait
 1272 variances ($\sigma^2/2\alpha$; Hansen 1997) of individual regimes are typically much less
 1273 than one order of magnitude (Dataset S3: taking maximum values in \log_{10} kg
 1274 of 0.579 [Sauropoda], 0.132 [Thyreophora], 0.406 [Marginocephalia], 0.188
 1275 [Ornithopoda], and 0.278 [Theropoda: Coelurosauria]; all results excluding the
 1276 primitive dinosaur regimes). There are a few exceptions to this, among
 1277 ‘prosauropods’ (non-sauropodan sauropodomorphs) for some trees, and for the
 1278 primitive dinosaur regime on trees of Sauropodomorpha and Theropoda
 1279 (however, the regimes characterizing these relatively short-lived grades may
 1280 not have persisted for sufficiently long to constrain model fits accurate).

1281 The expected asymptotic trait variance is an estimate of the disparity
 1282 that could have accumulated under a single macroevolutionary regime, and
 1283 our confidence in these estimates is boosted by the observation that they are
 1284 similar to the observed clade disparities (e.g. Fig. 13A for sauropods).
 1285 Expected asymptotic trait variances in our models are small compared to
 1286 differences between trait optima (θ) for successive regimes (Fig. 14A–B:
 1287 typically 1–2 orders of magnitude [$1\text{--}2 \log_{10}$ kg]). This pattern may largely
 1288 explain the goodness of fit of multi-regime OU models, and indicates that
 1289 most of the variance of dinosaur body sizes that evolved during the Mesozoic
 1290 resulted from stepwise shifts in the ‘optimal’ body sizes of groups of lineages,

1291 and not from the gradual expansion of lineages within regimes to wider trait
1292 spaces. In other words: large, constraint-breaking body size changes occur on
1293 a minority of branches corresponding to the transitions between
1294 macroevolutionary regimes. These branches are associated with high
1295 magnitude shifts in trait values, as documented in previous work (Benson *et*
1296 *al.* 2014a: as high rates of body size change at nodes). And this overall
1297 pattern is also consistent with that inferred from patterns of body size change
1298 at different taxonomic ranks among Phanerozoic marine invertebrates (Heim
1299 *et al.* 2015).

1300 Transitions between macroevolutionary regimes caused rapid,
1301 directional trait evolution, under which most dinosaurian body size disparity
1302 was accumulated. For example, sauropodomorphs explored roughly three
1303 orders of magnitude of body size variation due to a minimum of two regime
1304 shifts that occurred during less than 40 million years of the Triassic (Fig. 7).
1305 These gave rise to high body size disparity among Triassic and Early Jurassic
1306 sauropodomorphs, which was comparable to that of all dinosaurs during those
1307 intervals (Fig. 13A). During their subsequent evolution, which was dominated
1308 by occupancy of a single regime, sauropodomorphs explored little more than
1309 a single order of magnitude in body size (Figs 7,13A). In an even more
1310 extreme example, tetanuran theropods explored more than three orders of
1311 magnitude of body size variation though a minimum of three regime shifts
1312 lasting less than 10 million years during the Middle Jurassic (Figs 9–10). This
1313 culminated in apparently the greatest body size disparity of (non-pygostylian)
1314 theropod evolutionary history during the Middle Jurassic (Fig. 13B). Also
1315 during the Early–Middle Jurassic, expansions over two orders of magnitude

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3 1316 resulted from three regime shifts among ornithischian dinosaurs (Fig. 8; in
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5 1317 Thyreophora, Marginocephalia and Iguanodontia). The hypothesis that
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7 1318 dinosaur body size disparity resulted from gradual dispersal of lineages
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9 1319 through trait space is easy to reject: even variable-rate Brownian motion
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11 1320 models provide substantially worse fit to the data than do multi-peak OU
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13 1321 models.
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16 1322 *The adaptive landscape and suboptimal or unrealised regions of*
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18 1323 *dinosaurian body size space.* The pulsed pattern of body size evolution
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20 1324 documented here is consistent with the hypothesis of quantum evolution, or
21
22 1325 the existence of a heterogeneous macroevolutionary adaptive landscape
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24 1326 (Simpson 1953; Hansen 2013). Under this paradigm, transitions between
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26 1327 macroevolutionary adaptive zones or peaks involve rapid, directional
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28 1328 evolutionary treks across apparently unstable regions of trait space. Given
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30 1329 that the correspondence between body size and ecological niches is coarse, it
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32 1330 is clear that the peaks represented by our macroevolutionary regimes are
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34 1331 broad entities. We therefore view them as comparable to the ‘maximum
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36 1332 adaptive zones’ of Stanley (1973) rather than as fitness functions that are
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38 1333 exactly shared among species (see also Hansen 1997, 2013 for a discussion
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40 1334 of the difficulties scaling microevolutionary models to macroevolutionary
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42 1335 observations). Furthermore, the apparent timescales of transitions between
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44 1336 regimes span several million years, a long duration compared to
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46 1337 microevolutionary timescales that are measured in generations. In general,
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48 1338 macroevolutionary changes occur at slower rates than predicted from simple
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50 1339 extrapolation of microevolutionary changes to longer timescales (e.g. Uyeda
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52 1340 *et al.* 2011). Hansen (2013) has described a model in which microevolutionary

1341 changes result from rapid attraction to adaptive peaks, whereas
1342 macroevolutionary changes reflect shifts in the positions of those peaks in the
1343 adaptive landscape due to changes in the abiotic and biotic environments. In
1344 this context, our study provides information on the topology of the
1345 macroevolutionary adaptive landscape for dinosaurs.

1346 There may be ultimate limits to the maximum and minimum body sizes
1347 attainable by terrestrial vertebrates. Dinosaurs seem to have approached both
1348 during their evolutionary history, attaining maximum body masses as small as
1349 2 g in birds, and as large as 40 tonnes (Mazetta *et al.* 2004; Lacovara *et al.*
1350 2014; Bates *et al.* 2015) and perhaps 90 tonnes (*Argentinosaurus*, herein) in
1351 giant sauropods (Table 4). The regimes relevant to these extremely large and
1352 extremely small body sizes are both unique, having each evolved only once
1353 among Mesozoic dinosaurs (Fig. 14D,F). This contrasts with the more
1354 frequent attainment of less extreme body size regimes in other groups of
1355 dinosaurs (Fig. 14C–F). One possible explanation is that extreme small and
1356 large body sizes both require fundamental structural reorganization. If so, then
1357 terrestrial tetrapods significantly outside of the body size range seen in
1358 dinosaurs will rarely, or never, evolve.

1359 An apparently underexploited region of trait space between 30 and 300
1360 kg ($1.5\text{--}2.5 \log_{10}$ kg; grey rectangles in Figs 14–15) was represented by the
1361 trait optimum of only one regime, meaning that it was attained only once
1362 during dinosaur evolution, in early tetanuran or coelurosaurian theropods
1363 (Figs 9–10). In fact, many coelurosaur species of this size are known, so the
1364 central portion of dinosaurian body size space is densely populated among
1365 theropods (Figs 14–15). Sauropodomorpha and Ornithischia do contain taxa

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3 1366 within this size range. However these taxa belong either to lineages making
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5 1367 rapid transitions from smaller- to larger-body sized regimes, or are at the
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7 1368 margins of the body size distributions for their clades. This is most evident in
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9 1369 the body mass frequency distributions for Jurassic–Early Cretaceous
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11 1370 ornithischians (Fig. 15E,H,L) and Triassic sauropodomorphs (Fig. 15A). It is
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13 1371 less evident during other times because large-bodied ornithischians are not
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15 1372 known from the Triassic (Fig. 15B) and small-bodied sauropodomorphs are
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17 1373 absent by the Middle Jurassic (Fig. 15G). It is likely that the proportion of
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19 1374 dinosaur species at smaller body sizes has been underestimated due to
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21 1375 taphonomic biases that limit the preservation potential of individuals weighing
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23 1376 less than 60 kg in some deposits (Brown *et al.* 2013). Nevertheless, our data
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25 1377 confirm the presence of a bimodal or trimodal body mass distribution in
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27 1378 dinosaurs: fitting a set of mixture models (McLachlan & Basford 1988) using
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29 1379 the R package mclust (Fraley *et al.* 2012) finds support for models positing
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31 1380 that the distribution of log-transformed dinosaur body sizes is composed of a
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33 1381 mixture of either two or three normally distributed sub-distributions. These
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35 1382 models receive similar levels of support from Bayesian information criteria
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37 1383 (BIC), and are substantially better ($\Delta\text{BIC} > 100$) than a model with a single,
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39 1384 normal distribution for dinosaur body size. A bimodal body size distribution
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41 1385 was also proposed by some previous studies (Codron *et al.* 2012; O’Gorman
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43 1386 and Hone 2012), and this is explained here by the bimodal distribution of body
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45 1387 size optima for macroevolutionary regimes (Fig. 14).
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52 1388 Previous studies (Codron *et al.* 2012; O’Gorman & Hone 2012) invoked
53
54 1389 the non-uniformitarian life history mode of dinosaurs (e.g. Janis & Carrano
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56 1390 1992; Varicchio 2011) to explain bimodality of dinosaur body sizes, arguing
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1391 that oviparity and high fecundity resulted in an abundance of medium-sized
1392 individuals representing juveniles of large-bodied dinosaur species in
1393 Mesozoic ecosystems, reducing the fitness of medium-sized adults. This
1394 hypothesis is somewhat consistent with our models, although the theropod
1395 clade Coelurosauria included high species diversity at intermediate body
1396 masses (Figs 9–10, 14E, 15Q: Late Cretaceous Theropoda), which is not
1397 consistent with ecological exclusion. Furthermore, a bimodal pattern of body
1398 size evolution is also seen in Cenozoic mammals, in which lineages are either
1399 constrained by attraction to small, plesiomorphic body masses or otherwise
1400 attracted to large, derived body masses, with an apparent lack of optima at
1401 intermediate sizes (0.4 kg–8 kg; Alroy 1999). The different numerical values of
1402 dinosaurian and mammalian body size optima suggest that these patterns are
1403 analogous, rather than strictly equivalent, and may have different causes.
1404 Nevertheless, this presence of bimodality in mammalian body size evolution
1405 suggests that distinct life history factors need not be invoked to explain the
1406 same pattern in dinosaurs. Instead, bimodal body size dynamics have
1407 characterized terrestrial vertebrates since early in the Mesozoic. The
1408 microevolutionary explanation of this pattern is currently unknown.
1409 Furthermore, some of the best-sampled dinosaur assemblages are biased
1410 against preservation of smaller taxa (Brown *et al.* 2013) and detailed
1411 appraisals of the effect of this bias on body size distributions and
1412 macroevolutionary patterns, such as those obtained here, have yet to be
1413 carried out.

1414 *Niche-filling (early burst) or constrained models of trait evolution?*

1415 Rapid increases in disparity are observed early in the fossil records of many

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3 1416 animal groups (e.g. Foote 1997a; Hughes *et al.* 2013). A similar pattern
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5 1417 occurs for the body mass disparity of non-pygostylian dinosaurs, and in the
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7 1418 major subclades Sauropodomorpha (which peaked in the Triassic),
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9 1419 Ornithischia (which peaked in the Late Jurassic) and non-pygostylian
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11 1420 Theropoda (which peaked in the Middle Jurassic) (Fig. 13). A key question
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13 1421 that can be addressed by phylogenetic studies of trait evolution such as ours
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15 1422 is whether these patterns result from high, but decelerating, early rates of
16
17 1423 evolution (the 'early burst' model; Harmon *et al.* 2010) or from constraints on
18
19 1424 the range of phenotypes attainable by a clade. Under the constraint
20
21 1425 hypothesis, trait space becomes rapidly saturated (e.g. Slater 2013; Oyston *et*
22
23 1426 *al.* 2015) even if evolutionary rates do not decelerate. The early burst model
24
25 1427 was specifically formulated to test Simpson's niche-filling hypothesis of
26
27 1428 adaptive radiation (Simpson 1954; Harmon *et al.* 2008). This hypothesis
28
29 1429 suggests a link between rates of evolution and the availability of ecological
30
31 1430 niches, which could be made available by key innovations that provide access
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33 1431 to new adaptive zones, or by other factors such as mass extinctions, or
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35 1432 geographic dispersal (reviewed by Etienne & Haegeman 2012).

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38 1433 A 'constrained' pattern, as represented by Ornstein-Uhlenbeck (OU)
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40 1434 models, is evident from our results, and is consistent with the widespread
41
42 1435 existence of constraints on body size evolution within animal groups (Heim *et*
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44 1436 *al.* 2015: marine animals). Nevertheless, pulses of body size evolution that
45
46 1437 cause increases in phenotypic disparity are associated with
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48 1438 macroevolutionary regime shifts, and these are not homogeneously
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50 1439 distributed through time. In fact, sauropodomorph and ornithischian regime
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52 1440 shifts are concentrated in the first half of the evolutionary history of dinosaurs
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(Fig. 14B), in the Triassic (sauropodomorphs) and Triassic–Jurassic (ornithischians; especially the Early/Middle Jurassic following a major extinction event at the Triassic/Jurassic boundary; e.g. Raup & Sepkoski 1982; Brusatte *et al.* 2008). Therefore, the temporal distribution of macroevolutionary pulses seen in non-theropod dinosaurs is arguably consistent with a broad scale niche-filling model of dinosaurian body size evolution (i.e. adaptive radiation). Under this hypothesis, opportunities (i.e. access to unfilled niches) resulted from fundamental key innovations of the dinosaur body plan in the Triassic, and from extinctions among of potential competitors around the Triassic/Jurassic boundary.

 In contrast, theropod dinosaurs underwent a large number of high magnitude body size changes in the Cretaceous (Benson *et al.* 2014a), to the extent that post-Aptian macroevolutionary regimes of theropods could not be characterized in the present study, and will likely not be characterized until we obtain a significantly greater sample of their fossil diversity. In illustration of the intensity with which Cretaceous theropods explored body size space, just one clade, pygostylian birds made their first appearance in the Early Cretaceous (Zhang *et al.* 2000, 2008; Wang *et al.* 2015), and explored a body size range of approximately 4.5 orders of magnitude by the end of the Cretaceous (Fig. 14,15N,R; Tables 2, 5), over a time interval lasting only one-third of the total evolutionary history of Dinosauria (Fig. 13C). Cretaceous pygostylians range from 15 g in taxa such as *Qiliania* and *Iberomesornis* to 190 kg in *Gargantuavis* (Buffetaut & Le Loeuff 1998). This range is only slightly less than the range of body masses achieved by non-avian dinosaurs through the entire Mesozoic (maximum 5.5 orders of magnitude: from 434 g in

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3 1466 *Microraptor*, the smallest certainly non-juvenile non-avian dinosaur up to a
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5 1467 maximal estimate of 95,000 kg in *Argentinosaurus*; smaller masses are
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7 1468 recorded in some alvarezsauroids: *Parvicursor* [153 g] and *Ceratonykus* [262
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9 1469 g], but we cannot confirm that these specimens represent adults; Table 3).
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11 1470 Given the expansive pattern of body size evolution seen in Cretaceous
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13 1471 pygostylians, and the high frequency of large evolutionary body size changes
14
15 1472 among other Cretaceous theropods (Benson *et al.* 2014a), the pattern of body
16
17 1473 size evolution seen in theropod dinosaurs is not consistent with any
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19 1474 conception of the 'early burst' model. Nevertheless, it could be reconciled with
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21 1475 niche-filling as a mechanism of evolutionary diversification if some groups are
22
23 1476 able to maintain high levels of biological versatility (Vermeij 1973; =
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25 1477 'evolvability') that allow the continued discovery of new niches, whilst others
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27 1478 do not (Benson *et al.* 2014a).
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34 1480 **CONCLUSIONS**

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38 1482 (1) We infer scaling relationships among measurements of dinosaur limb
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40 1483 bones to estimate the adult body masses of 526 dinosaur species.
41
42 1484 (2) Scaling relationships among the hindlimb elements of bipedal dinosaurs,
43
44 1485 and among the forelimb and hindlimb elements of quadrupedal dinosaurs are
45
46 1486 typically conserved within groups such as Ornithischia, Sauropodomorpha
47
48 1487 and Theropoda (bipedal) and Sauropoda, Thyreophora and Ceratopsidae
49
50 1488 (quadrupedal), as they do not show strong phylogenetic signal. This suggests
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52 1489 that locomotor design is functionally constrained within clades with distinct
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3 1490 body plans, but that it varies among those clades, resulting from shifts in large
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5 1491 limb proportions at just a few nodes on the dinosaurian tree.
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7 1492 (3) Trend-based models, which describe strict conceptions of Cope's rule,
8
9 1493 provide poor explanations of dinosaurian body mass evolution in almost all
10
11 1494 cases. We therefore reject the widespread presence of multi-lineage
12
13 1495 directional trends of body size evolution in dinosaurs.
14
15 1496 (4) Instead, dinosaurian body mass can be explained by multi-regime
16
17 1497 Ornstein-Uhlenbeck models, which describe constrained evolution around
18
19 1498 macroevolutionary 'optima' or adaptive zones (Stanley 1972; Hansen 2013).
20
21 1499 Rare shifts in these optima cause lineages to undergo rapid excursions
22
23 1500 through trait space characterized as 'quantum evolution' by Simpson (1954),
24
25 1501 and under which most of the disparity of dinosaurian body masses
26
27 1502 accumulated.
28
29 1503 (5) Our analyses reject two recent hypotheses regarding miniaturization on
30
31 1504 the bird stem lineage. (i) Theropod dinosaurs did not undergo sustained
32
33 1505 evolutionary body size reductions from the Triassic to the Late Jurassic.
34
35 1506 Support for this hypothesis resulted from overestimation of the ancestral
36
37 1507 theropod body mass. (ii) Pygostylian birds, principally in the range of 15–300
38
39 1508 g do not share a body size optimum with other paravians (minimum body
40
41 1509 masses ~ 400 g; maximum: 600 kg; Tables 3, 5).
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43 1510 (6) Dinosaur body mass optima are unevenly distributed among clades and
44
45 1511 through time. Most sauropodomorph regime shifts occurred in the Triassic,
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47 1512 most ornithischian regime shifts occurred in the Triassic–Jurassic, and
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49 1513 theropods continued to undergo regime shifts for their entire evolutionary
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51 1514 history. This pattern might be explained by the distribution of opportunity
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1515 resulting from key innovations during the early evolutionary history of
1516 Dinosauria, extinctions of potentially competing taxa at the Triassic/Jurassic
1517 boundary, and potentially from greater evolutionary versatility in theropods
1518 than in other dinosaurs.

1519 (7) Body mass optima are unevenly distributed in trait space. Shifts to larger
1520 body size are more frequent than to smaller size in ornithischians and
1521 sauropodomorphs, and the distribution of body size optima is bimodal, with a
1522 sparsely populated intermediate region of trait space around 30–300 kg. This
1523 explains previous observations of bimodality in dinosaur body mass
1524 distributions. However, as bimodality is also present in mammalian body size
1525 evolution, the occurrence of bimodal dinosaur body mass evolution may not
1526 require an explanation that invokes specific attributes of dinosaur life history.

1527 (8) Dinosaur body size disparity can be explained predominantly by
1528 constrained (Ornstein-Uhlenbeck) models of evolution. This suggests that the
1529 early attainment of maximal disparity seen in many groups of animals may
1530 result from constraints on the range of phenotypes achievable by a clade.

1531

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1541
1542 **SUPPORTING INFORMATION**
1543
1544 Additional Supporting Information can be found in the online version of this
1545 article. Data for this study are available in the Dryad Digital Repository:
1546 <http://dx.doi.org/10.5061/dryad.1t3r4>. [please note that the data for this paper
1547 are not yet published and this temporary link should not be shared without the
1548 express permission of the author].
1549
1550 Appendix S1. Methods and results for dinosaur body mass estimates,
1551 including Tables S1–S12.
1552
1553 Appendix S2. Summary figures from SURFACE analyses for Triassic–
1554 Jurassic Dinosauria, Sauropodomorpha, Ornithischia and Triassic–Aptian
1555 Theropoda, ordered by AICc or pBIC score.
1556
1557 Appendix S3. Diagrams comparing the pBIC scores of SURFACE regime
1558 configurations from analysis of timescaled phylogenies mapped to the same
1559 phylogenies timescaled using other methods.
1560
1561 Appendix S4. Diagrams illustrating trajectories of stepwise-AICc and
1562 stepwise-pBIC searches following the SUFACE algorithm and resulting
1563 regime configurations.

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5 1565 Appendix S5. Phenograms illustrating regime configurations mapped to
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7 1566 phylogenies evolving in body size space through time.
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12 1568 Dataset S1. Measurements, metadata and body mass estimates for
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14 1569 Dinosauria.
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19 1571 Dataset S2. AICc scores for trend-based model fits compared to Brownian
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21 1572 motion.
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26 1574 Dataset S3. AICc scores and model fits for OU and other models
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28 1575 implemented in Ouwie.
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33 1577 Dataset S4. Phylogenetic trees used in our analyses, in Newick format.
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38 1579 Scripts S1. Scripts used in R version 3.3.3 to conduct analyses in the current
39
40 1580 paper including estimation of standard errors of variables estimated through
41
42 1581 regression, fitting models of trends that can shift over time or across nodes,
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44 1582 and versions of surfaceForward and surfaceBackward (Ingram & Mahler
45
46 1583 2013) that use pBIC (Khabbazzian et al. 2016), rather than AICc.
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49 1584

50 1585 **REFERENCES**
51

52
53 1586

54
55 1587 AKAIKE, H. 1974. A new look at the statistical model identification. *IEEE*
56
57 1588 *Transactions on Automatic Control*, **19**, 716–723.
58
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56
57
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59
60

1589 ALEXANDER, R. M. 1998. All-time giants: the largest animals and their
1590 problems. *Palaeontology*, **41**, 1231–1245.

1591 ALROY, J. A. 1999. Cope's rule and the dynamics of body mass evolution in
1592 North American fossil mammals. *Science*, **280**, 731–734.

1593 ANDERSON, J. F., HALL-MARTIN, A. and RUSSELL, D. A. 1985. Long-bone
1594 circumference and weight in mammals, birds and dinosaurs. *Journal of*
1595 *the Zoological Society of London A*, **207**, 53–61.

1596 ANÉ, C. 2008. Analysis of comparative data with hierarchical autocorrelation.
1597 *Annals of Applied Systematics*, **2**, 1078–1102.

1598 APALDETTI, C., POL, D. and YATES, A. 2013. The postcranial anatomy of
1599 *Coloradisaurus brevis* (Dinosauria: Sauropodomorpha) from the Late
1600 Triassic of Argentina and its phylogenetic implications. *Palaeontology*,
1601 **56**, 277–301.

1602 BAKER, J., MEADE, A., PAGEL, M. and VENDITTI, C. 2015. Adaptive
1603 evolution toward larger size in mammals. *Proceedings of the National*
1604 *Academy of Sciences of the USA*, **112**, 5093–5098.

1605 BAKKER, R. T. 1971. Ecology of the brontosaurus. *Nature*, **229**, 172–174.

1606 BAPST, D. W. 2012. paleotree: an R package for paleontological and
1607 phylogenetic analyses of evolution. *Methods in Ecology and Evolution*, **3**,
1608 803–807.

1609 — 2013. A stochastic rate-calibrated method for time-scaling phylogenies of
1610 fossil taxa. *Methods in Ecology and Evolution*, **4**, 724–733.

1611 — 2014a. Preparing paleontological datasets for phylogenetic comparative
1612 methods. 515–544. In GARAMSZEGLI, L. Z. (ed.). *Modern Phylogenetic*

- 1613 *Comparative Methods and Their Application in Evolutionary Biology*.
1614 Springer: Heidelberg.
- 1615 — 2014*b*. Assessing the effect of time-scaling methods on phylogeny-based
1616 analyses in the fossil record. *Paleobiology*, **40**, 331–351.
- 1617 — and HOPKINS S. 2017. Comparing *cal3* with other a posteriori time-scaling
1618 approaches in a case study with the pterocephaliid trilobites.
1619 *Paleobiology* **43**, 49–67.
- 1620 BATES, K. T., FALKINGHAM, P. L., MACAULAY, S., BRASSEY, C. and
1621 MAIDMENT, S. C. R. 2015. Downsizing a giant: re-evaluating
1622 *Dreadnoughtus* body mass. *Biology Letters*, **11**, 20150215.
- 1623 BEAULIEAU, J. M., JHUWUENG, D. –C., BOETTIGER, C. and O'MEARA,
1624 BC. 2012. Modelling stabilizing selection: expanding the Ornstein–
1625 Uhlenbeck model of adaptive evolution. *Evolution*, **66**, 2369–2383.
- 1626 BENSON, R. B. J. and CHOINIERE, J. N. 2013. Rates of dinosaur limb
1627 evolution provide evidence for exceptional radiation in Mesozoic birds.
1628 *Proceedings of the Royal Society B*, **280**, 20131780.
- 1629 —, BUTLER, R. J., CARRANO, M. T. and O'CONNOR, P. M. 2012. Air-filled
1630 postcranial bones in theropod dinosaurs: physiological implications and
1631 the 'reptile'-bird transition. *Biological Reviews*, **87**, 168–193.
- 1632 —, CAMPIONE, N. E., CARRANO, M. T., MANNION, P. D., SULLIVAN, C.,
1633 UPCHURCH, P. and EVANS, D. C. 2014*a*. Rates of dinosaur body mass
1634 evolution indicate 170 million years of sustained ecological innovation on
1635 the avian stem lineage. *PLOS Biology*, **12**(5), 1001853.

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43
44
45
46
47
48
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51
52
53
54
55
56
57
58
59
60

1636 —, FRIGOT, R. A., GOSWAMI, A., ANDRES, B. and BUTLER, R. J. 2014*b*.
1637 Competition and constraint drove Cope's rule in the evolution of giant
1638 flying reptiles. *Nature Communications*, **5**, 3567.
1639 —, HUNT, G., CARRANO, M. T. and CAMPIONE, N. 2017. Data from: Cope's
1640 rule and the adaptive landscape of dinosaur body size evolution. *Dryad*
1641 *Digital Repository*. doi.org/10.5061/dryad.1t3r4.
1642 BENTON, M. J., CSIKI, Z., GRIGORESCU, D., REDELSTORFF, R.,
1643 SANDER, P. M., STEIN, K. and WEISHAMPEL, D. B. 2010. Dinosaurs
1644 and the island rule: The dwarfed dinosaurs from Haţeg Island.
1645 *Palaeogeography, Palaeoclimatology, Palaeoecology*, **293**, 438–454.
1646 BONAPARTE, J. F. and CORIA, R. A. 1993. Un Nuevo y gigantesco
1647 sauropodo titanosaurio de la Formacion Rio Limay (Albiano-
1648 Ceonmaniano) de la Provincia del Neuquen, Argentina. *Ameghiniana*,
1649 **30**, 271–282.
1650 — and PUMARES, J. A. 1995. Notas sobre el primer craneo de *Riojasaurus*
1651 *incertus* (Dinosauria, Prosauropoda, Melanorosauridae) del Triasico
1652 Superios de La Rioja, Argentina. *Ameghiniana*, **32**, 341–349.
1653 BONNAN, M. F. 2003. The evolution of manus shape in sauropod dinosaurs:
1654 implications for functional morphology, forelimb orientation, and
1655 phylogeny. *Journal of Vertebrate Paleontology*, **23**, 595–613.
1656 BROWN, J. H. 1995 *Macroecology*. University of Chicago Press, Chicago,
1657 269 pp.
1658 — and MAURER, B. A. 1986. Body size, ecological dominance, and Cope's
1659 Rule. *Nature*, **324**, 248–250.

- 1660 BROWN, C. M., EVANS, D. C., CAMPIONE, N. E., O'BRIEN, L. J. and
1661 EBERTH, D. A. 2013. Evidence for taphonomic size bias in the Dinosaur
1662 Park Formation (Campanian, Alberta), a model Mesozoic terrestrial
1663 alluvial-paralic system. *Palaeogeography, Palaeoclimatology,*
1664 *Palaeoecology*, **372**, 108–122.
- 1665 BRUSATTE, S. L., BENTON, M. J., RUTA, M. and LLOYD, G. T. 2008.
1666 Superiority, competition, and opportunism in the evolutionary radiation of
1667 dinosaurs. *Science*, **321**, 1485–1487.
- 1668 BUFFETAUT, E., LE LOEUFF, J., MECHIN, P. and MECHIN-SALESSY, A.
1669 1995. A large French Cretaceous bird. *Nature*, **377**, 110.
- 1670 BURNES, G. P., DIAMOND, J. and FLANNERY, T. 2001. Dinosaurs,
1671 dragons, and dwarfs: the evolution of maximal body size. *Proceedings of*
1672 *the National Academy of Sciences of the USA*, **25**, 14518–14523.
- 1673 BURNHAM, K. P. and ANDERSON, D. R. 2004 *Model selection and*
1674 *multimodel inference*. Springer, New York, 488 pp.
- 1675 —, — and HUYVAERT, K. P. 2011. AIC model selection and multimodel
1676 inference in behavioral ecology: some background, observations, and
1677 comparisons. *Behavioural Ecology and Sociobiology*, **65**, 23–35.
- 1678 BUTLER, M. A. and KING, A. A. 2004. Phylogenetic comparative analysis: a
1679 modelling approach for adaptive radiation. *American Naturalist*, **164**,
1680 683–695.
- 1681 BUTLER, R. J., GALTON, P. M., PORRO, L. B., CHIAPPE, L. M.,
1682 HENDERSON, D. M. and ERICKSON, G. M. 2009. Lower limits of
1683 ornithischian dinosaur body size inferred from a new Upper Jurassic

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43
44
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46
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49
50
51
52
53
54
55
56
57
58
59
60

1684 heterodontosaurid from North America. *Proceedings of the Royal*
1685 *Society B*, **277**, 375–381.

1686 CAMPBELL, K. E. and MARCUS, L. 1992. The relationships of hindlimb bone
1687 dimensions to body weight in birds. *Natural History Museum of Los*
1688 *Angeles County Science Series*, **36**, 395–412.

1689 CAMPIONE, N. E. and EVANS, D. C. 2012. A universal scaling relationship
1690 between body mass and proximal limb bone dimensions in quadrupedal
1691 terrestrial tetrapods. *BMC Biology*, **10**, 1–21.

1692 —, EVANS, D. C., BROWN, C. M. and CARRANO, M. T. 2014. Body mass
1693 estimation in non-avian bipeds using a theoretical conversion to
1694 quadrupedal stylopodial proportions. *Methods in Ecology and Evolution*,
1695 **5**, 913–923.

1696 CANTALAPIEDRA J. L., PRADO, J. L., FERNÁNDEZ M. H. and ALBERDI, M.
1697 T. 2017. Decoupled ecomorphological evolution and diversification in
1698 Neogene–Quaternary horses. *Science*, **355**, 627–230.

1699 CARBALLIDO, J. L., POL, D., OTERO, A., CERDA, I. A., SALGADO, L.,
1700 GARRIDO, A. C., RAMEZANI, J., CÚENO, N. R. and KRAUSSE, J. M. A
1701 new giant titanosaur sheds light on body mass evolution among
1702 sauropod dinosaurs. *Proceedings of the Royal Society B*, **284**,
1703 20171219.

1704 CARPENTER, K. 2006. Biggest of the big: a critical re-evaluation of the
1705 mega-sauropod *Amphicoelias fragillimus*. 131–138, In FOSTER, J. R.
1706 and LUCAS, S. G., (eds). *Paleontology and Geology of the Upper*
1707 *Jurassic Morrison Formation. New Mexico Museum of Natural History*
1708 *and Science Bulletin* 36.

- 1
2
3 1709 CARRANO, M. T. 1998 Locomotion in non-avian dinosaurs: integrating data
4
5 1710 from hindlimb kinematics, in vivo strains, and bone morphology.
6
7 1711 *Paleobiology*, **34**, 450–469.
8
9
10 1712 — 2001 Implications of limb bone scaling, curvature and eccentricity in
11
12 1713 mammals and non-avian dinosaurs. *Journal of Zoology*, **254**, 41–55.
13
14 1714 — 2006. Body-size evolution in the Dinosauria. 225–268. In CARRANO, M.
15
16 1715 T., GAUDIN, T. J., BLOB, R. W. and WIBLE, J. R. (eds). *Amniote*
17
18 1716 *Paleobiology*. University of Chicago Press, Chicago.
19
20
21 1717 — 2007. The appendicular skeleton of *Majungasaurus crenatissimus*
22
23 1718 (Theropoda: Abelisauridae) from the Late Cretaceous of Madagascar.
24
25 1719 *Journal of Vertebrate Paleontology*, **27**(suppl. 2), 163–179.
26
27 1720 CHINNERY, B. 2004. Morphometric analysis of evolutionary trends in the
28
29 1721 ceratopsian postcranial skeleton. *Journal of Vertebrate Paleontology*, **24**,
30
31 1722 591–609.
32
33
34 1723 — and HORNER J. R. 2007. A new neoceratopsian dinosaur linking North
35
36 1724 American and Asian taxa. *Journal of Vertebrate Paleontology*, **27**, 625–
37
38 1725 641.
39
40 1726 CLOSE, R. A., FRIEDMAN M., LLOYD, G. T. and BENSON, R. B. J. 2015.
41
42 1727 Evidence for a Mid-Jurassic Adaptive Radiation in Mammals. *Current*
43
44 1728 *Biology*, **25**, 2137–2142.
45
46
47 1729 CODRON, D., CARBONE, C., MÜLLER, D. W. H. and CLAUSS, M. 2012.
48
49 1730 Ontogenetic niche shifts in dinosaurs influenced size, diversity and
50
51 1731 extinction in terrestrial vertebrates. *Biology Letters*, **8**, 620–623.
52
53
54 1732 COLBERT, E. H. 1962. The weights of dinosaurs. *American Museum*
55
56 1733 *Novitates*, **2076**, 1–16.
57
58
59
60

1734 COOPER, M. R. A reassessment of *Vulcanodon karibaensis* Raath
1735 (Dinosauria: Saurischia) and the origin of Sauropoda. *Palaeontologia*
1736 *Africana*, **25**, 203–231.

1737 COOPER, N., THOMAS, G. H., VENDITTI, C., MEADE, A. and
1738 FRECKLETON, R. P. 2016. A cautionary note on the use of Ornstein
1739 Uhlenbeck models in macroevolutionary studies. *Biological Journal of*
1740 *the Linnean Society*, **118**, 64–77.

1741 DAVIS, A. M. and BETANCUR-R, R. 2017. Widespread ecomorphological
1742 convergence in multiple fish families spanning the marine-freshwater
1743 interface. *Proceedings of the Royal Society B*, **284**, 20170565.

1744 DE SOUZA, L. M. and SANTUCCI, R. M. 2014. Body size in
1745 Titanosauriformes (Sauropoda, Macronaria). *Journal of Evolutionary*
1746 *Biology*, **27**, 2001–2012.

1747 DEL HOYO, J., ELLIOTT, A. and SARGATAL, J. 1999. *Handbook of the Birds*
1748 *of the World*, volume 5: Barn Owls to Hummingbirds. Lynx Edicions,
1749 Barcelona, 759 pp.

1750 ERICKSON, G. M. 2014. On dinosaur growth. *Annual Review of Earth and*
1751 *Planetary Sciences*, **42**, 675–697.

1752 —, CURRIE, P. J., INOUE, B. D. and WINN, A. A. 2006. Tyrannosaur life
1753 tables: The first look at non-avian dinosaur population biology. *Science*,
1754 **313**, 213–217.

1755 —, RAUHUT, O. W. M., ZHOU, Z. –H., TURNER, A. H., INOUE, B. D., HU,
1756 D. –Y, NORELL, M. A. 2009a Was dinosaurian physiology inherited by
1757 birds? Reconciling slow growth in *Archaeopteryx*. *PLOS ONE*, **4**, e7390.

- 1758 —, MAKOVICKY, P. J., INOUE, B. D., ZHOU, C. –F. and GAO, K. 2009b.
1759 Life table for *Psittacosaurus lujiatunensis*: The first glimpse into
1760 ornithischian dinosaur population biology. *Anatomical Record*, **292**,
1761 1514–1521.
- 1762 —, CURRIE, P. J., INOUE, B. D., and WINN, A. A. 2010. A revised Life table
1763 and survivorship curve for *Albertosaurus sarcophagus* based on the Dry
1764 Island mass death assemblage. *Canadian Journal of Earth Sciences*, **47**,
1765 1269–1275.
- 1766 ETIENNE, R. S. and HAEGEMANN, B. 2012. A conceptual and statistical
1767 framework for adaptive radiations with a key role for diversity
1768 dependence. *American Naturalist*, **180**, E75–E89.
- 1769 FELSENSTEIN, J. 1973. Maximum-likelihood estimation of evolutionary trees
1770 from continuous characters. *American Journal of Human Genetics*, **25**,
1771 471–492.
- 1772 — 1985. Phylogenies and the comparative method. *American Naturalist*, **125**,
1773 1–15.
- 1774 FOOTE, M. 1997a. The evolution of morphological diversity. *Annual Review*
1775 *of Ecology and Systematics*, **28**, 129–152.
- 1776 — 1997b. Estimating taxonomic durations and preservation probability.
1777 *Paleobiology*, **23**, 278–300.
- 1778 FRALEY, C., RAFTERY, A. E., MURPHY, T. B. and SCRUCICA, L. 2012.
1779 mclust Version 4 for R: Normal Mixture Modeling for Model-Based
1780 Clustering, Classification, and Density Estimation Technical Report No.
1781 597, Department of Statistics, University of Washington.

1782 GALTON, P. M. 1970. The posture of hadrosaurian dinosaurs. *Journal of*
1783 *Paleontology*, **44**, 464–473.

1784 — 1974. Notes on *Thescelosaurus*, a conservative ornithischian dinosaur
1785 from the Upper Cretaceous of North America, with comments on
1786 ornithopod classification. *Journal of Paleontology*, **48**, 1048–1067.

1787 GARLAND, T. JR. and IVES, A. R. 2001. Using the past to predict the
1788 present: confidence intervals for regression equations in phylogenetic
1789 comparative methods. *American Naturalist*, **155**, 346–364.

1790 GRILLO, O. and DELCOURT, R. 2017. Allometry and body length of
1791 abelisauroid theropods: *Pycnonemosaurus nevesi* is the new king.
1792 *Cretaceous Research*, **69**, 71–89.

1793 HANSEN, T. F. 1997. Stabilizing selection and the comparative analysis of
1794 adaptation. *Evolution*, **51**, 1341–1351.

1795 — 2013. Adaptive landscapes and macroevolutionary dynamics. 205–226. *In*
1796 SVENSSON E. and CALSBEEK R. (eds). *The adaptive landscape in*
1797 *evolutionary biology*. Oxford University Press, Oxford.

1798 — and MARTINS EP. 1996. Translating between microevolutionary process
1799 and macroevolutionary patterns: the correlation structure of interspecific
1800 data. *Evolution*, **50**, 1404–1417.

1801 HARMON, L. J., LOSOS, J. B., DAVIES, T. J., GILLESPIE, R. G.,
1802 GITTLEMAN, J. L., JENNINGS, W. B., KOZAK, K. H., MCPEEK, M. A.
1803 MORENO-ROARK, F., NEAR, T. J., PURVIS, A., RICKLEFS, R. E.,
1804 SCHLUTER, D., SCHULTE, J. A. II, SEEHAUSEN, O., SIDLAUSKAS, B.
1805 L., TORRES-CARAVAJAL, O., WEIR, J. T. and MOOERS, A. O. 2010.

- 1806 Early bursts of body size and shape evolution are rare in comparative
1807 data. *Evolution*, **64**, 2385–2396.
- 1808 HEATH, T. A., HUELSENBECK, J. P. and STADLER, T. 2014. The fossilized
1809 birth-death process for coherent calibration of divergence-time
1810 estimates. *Proceedings of the National Academy of Sciences of the*
1811 *USA*. **111**, E2957–E2966.
- 1812 HEDMAN, M. 2010. Constraints on clade ages from fossil outgroups.
1813 *Paleobiology*, **36**, 16–31.
- 1814 HEIM, N. A., KNOPE, M. L., SCHAAL, E. K., WANG, S. C. and PAYNE, J. L.
1815 2015. Cope's rule in the evolution of marine animals. *Science*, **347**, 867–
1816 870.
- 1817 HO, L. S. T. and ANÉ, C. 2014. Intrinsic inference difficulties for trait
1818 evolution with Ornstein-Uhlenbeck models. *Methods in Ecology and*
1819 *Evolution*, **5**, 1133–1146.
- 1820 HONE, D. W. E. and BENTON, M. J. 2005. The evolution of large size: how
1821 does Cope's Rule work? *Trends in Ecology and Evolution*, **20**, 4–6.
- 1822 — KEESEY T. M., PISANI D. and PURVIS A. 2005. Macroevolutionary trends
1823 in the Dinosauria: Cope's rule. *Journal of Evolutionary Biology*, **18**, 587–
1824 595.
- 1825 HOPKINS M. J. and SMITH A. B. 2015. Dynamic evolutionary change in post-
1826 Paleozoic echinoids and the importance of scale when interpreting
1827 changes in rates of evolution. *Proceedings of the National Academy of*
1828 *Sciences of the U.S.A.*, **112**, 3758–3763.

1829 HUGHES, M., GERBER, S. and WILLS, M. A. 2013. Clades reach highest
1830 morphological disparity early in their evolution. *Proceedings of the*
1831 *National Academy of Sciences of the USA*, **110**, 13875–13879.

1832 HUNN, C. A. and UPCHURCH, P. 2001. The importance of time/space in
1833 diagnosing the causality of phylogenetic events: towards a
1834 ‘chronobiogeographical’ paradigm? *Systematic Biology*, **50**, 1–17.

1835 HUNT, G. 2006. Fitting and comparing models of phyletic evolution: random
1836 walks and beyond. *Paleobiology*, **32**, 578–601.

1837 — 2008. Evolutionary patterns within fossil lineages: model-based
1838 ASSESSMENT of modes, rates, punctuations and process. 117–131. *In*
1839 BAMBACH, R. K. and KELLEY, P. H. (eds). *From Evolution to*
1840 *Geobiology: Research Questions Driving Paleontology at the Start of a*
1841 *New Century*. The Paleontological Society.

1842 — 2012. Measuring rates of phenotypic evolution and the inseparability of
1843 tempo and mode. *Paleobiology*, **38**, 351–373.

1844 — and CARRANO, M. T. 2010. Models and methods for analyzing phenotypic
1845 evolution in lineages and clades. *Special Papers of the Paleontological*
1846 *Society*, **16**, 245–269.

1847 — and ROY, K. 2006. Climate change, body size evolution, and Cope's Rule
1848 in deep-sea ostracodes. *Proceedings of the National Academy of*
1849 *Sciences of the USA*, **103**, 1347–1352.

1850 INGRAM, T. and MAHLER, D. L. 2013. SURFACE: detecting convergent
1851 evolution from comparative data by fitting Ornstein-Uhlenbeck models
1852 with stepwise Akaike Information Criterion. *Methods in Ecology and*
1853 *Evolution*, **4**, 416–425.

- 1
2
3 1854 IVES, A. R., MIDFORD, P. E. and GARLAND, T. JR. 2007. Within-species
4
5 1855 variation and measurement error in phylogenetic comparative methods.
6
7 1856 *Systematic Biology*, **56**, 252–270.
8
9
10 1857 JANIS, C. M. and CARRANO, M. T. 1992. Scaling of reproductive turnover in
11
12 1858 archosaurs and mammals: why are large terrestrial animals so rare?
13
14 1859 *Acta Zoologica Fennica*, **28**, 201–216.
15
16 1860 KELLNER, A. W. A. and CAMPOS, D. A. 2002. On a theropod dinosaur
17
18 1861 (Abelisauria) from the continental Cretaceous of Brazil. *Arquivos do*
19
20 1862 *Museu Nacional Rio de Janeiro*, **60**, 163–170.
21
22
23 1863 KHABBAZIAN M., KRIEBEL R., ROHE K. and ANÉ C. 2016. Fast and
24
25 1864 accurate detection of evolutionary shifts in Ornstein-Uhlenbeck models.
26
27 1865 *Methods in Ecology and Evolution*, **7**, 811–824.
28
29
30 1866 KINGSOLVER, J. G. and PFENNIG, D. W. 2004. Individual-level selection as
31
32 1867 a cause of Cope's rule of phyletic size increase. *Evolution*, **58**, 1608–
33
34 1868 1612.
35
36 1869 LACOVARA, K. J., IBIRICU, L. M., LAMANNA, M. C., POOLE, J. C.,
37
38 1870 SCHROETER, E. R., ULLMANN, P. V., VOEGELE, K. K., BOLES, Z. M.,
39
40 1871 EGERTON, V. M., HARRIS, J. D., MARTÍNEZ, R. D. and NOVAS, F. E.
41
42 1872 2014. A gigantic, exceptionally complete titanosaurian sauropod
43
44 1873 dinosaur from southern Patagonia, Argentina. *Scientific Reports*, **4**,
45
46 1874 6196.
47
48
49 1875 LAPIEDRA, O., SOL, D., CARRANZA, S. and BEAULIEU, J. M. 2013.
50
51 1876 Behavioural changes and the adaptive diversification of pigeons and
52
53 1877 doves. *Proceedings of the Royal Society B*, **280**, 20122893.
54
55
56
57
58
59
60

1
2
3
4
5
6
7
8
9
10
11
12
13
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42
43
44
45
46
47
48
49
50
51
52
53
54
55
56
57
58
59
60

1878 LAURIN, M. 2004. The evolution of body size, Cope's rule and the origin of
1879 amniotes. *Systematic Biology*, **53**, 594–622.

1880 LEE, A. H. and WERNING, S. 2008. Sexual maturity in growing dinosaurs
1881 does not fit reptilian growth models. *Proceedings of the National*
1882 *Academy of Sciences of the USA*, **105**, 582–587.

1883 LEE, M. S., CAU, A., NAISH, D. and DYKE, G. J. 2014. Sustained
1884 miniaturization and anatomical innovation in the dinosaurian ancestors of
1885 birds. *Science*, **345**, 562–566.

1886 LLOYD, G. T., BAPST, D. W., FRIEDMAN, M. and DAVIS, K. E. 2016.
1887 Probabilistic divergence time estimation without branch lengths: dating
1888 the origins of dinosaurs, avian flight and crown birds. *Biology Letters*, **12**,
1889 20160609.

1890 LÜ, J. –C., XU, L., JIA, S. –H. and JI, Q. 2009. A new gigantic sauropod
1891 dinosaur from the Cretaceous of Ruyang, Henan, China. *Geological*
1892 *Bulletin of China*, **28**, 1–10.

1893 MAHLER, D. L. and INGRAM, T. 2014. Phylogenetic comparative methods for
1894 studying clade-wide convergence. 425–450. In GARAMSZEGLI, L. Z.
1895 (ed.). *Modern Phylogenetic Comparative Methods and Their Application*
1896 *in Evolutionary Biology*. Springer-Verlag, Berlin.

1897 MAIDMENT, S. C. R., LINTON, D. H., UPCHURCH, P. and BARRETT, P. M.
1898 2012. Limb-bone scaling indicates diverse stance and gait in
1899 quadrupedal ornithischian dinosaurs. *PLOS One*, **7**, e36904.

1900 MARTINS, E. P. and HANSEN, T. F. 1997. Phylogenies and the comparative
1901 method: a general approach to incorporating phylogenetic information

- 1902 into the analysis of interspecific data. *American Naturalist*, **149**, 646–
- 1903 667.
- 1904 MARYANSKA, T. and OSMÓLSKA, H. 1975. Protoceratopsidae (Dinosauria)
- 1905 of Asia. *Palaeontologia Polonica*, **33**, 133–181.
- 1906 MATLEY, C. A. 1924. Note on an armoured dinosaur from the Lameta beds of
- 1907 Jubbulpore. *Records of the Geological Survey of India*, **55**, 105–109.
- 1908 MAZZETTA, G. V., CHRISTIANSEN, P. and FARIÑA, R. A. 2004. Giants and
- 1909 bizarres: body size of some southern South American Cretaceous
- 1910 dinosaurs. *Historical Biology*, **2004**, 1–13.
- 1911 MCSHEA, D. W. 1994. Mechanisms of large-scale evolutionary trends.
- 1912 *Evolution*, **48**, 1747–1763.
- 1913 MCLACHLAN, G. J. and BASFORD, K. E. 1988. *Mixture models: Inference*
- 1914 *and applications to clustering*. Marcel Dekker, New York, 253 pp.
- 1915 MOTANI, R. and SCHMITZ, L. 2011. Phylogenetic versus functional signals in
- 1916 the evolution of form-function relationships in terrestrial vision. *Evolution*,
- 1917 **65**, 2245–2257.
- 1918 NORMAN, D. B. 1980. On the ornithischian dinosaur *Iguanodon*
- 1919 *bernissatensis* from the Lower Cretaceous of Bernissart (Belgium).
- 1920 *Mémoires de l'Institut Royal des Sciences Naturelles de Belgique*, **178**,
- 1921 1–103.
- 1922 — 1986. On the anatomy of *Iguanodon atherfieldensis* (Ornithischia:
- 1923 Ornithopoda). *Bulletin de l'Institut Royal des Sciences Naturelle de*
- 1924 *Belgique: Sciences de la Terre*, **56**, 281–372.
- 1925 NOVAS, F. E., EZCURRA, M. D., AGNOLÍN, F. L., POL, D. and ORTÍZ, R.
- 1926 2012. New Patagonian Cretaceous theropod sheds light about the early

1
2
3 1927 radiation of Coelurosauria. *Revista del Museo Argentino de Ciencias*
4
5 1928 *Naturales, nueva serie*, **14**, 57–81.
6
7 — SALGADO, L., SUÁREZ, M., AGNOLÍN, F. L., EZCURRA, M. N. D.,
8
9 1930 CHIMENTO, N. S. R., DE LA CRUZ, R., ISASI, M. P., VARGAS, A. O.
10
11 1931 and RUBILAR-ROGERS, D. 2015. An enigmatic plant-eating theropod
12
13 1932 from the Late Jurassic period of Chile. *Nature*, **522**, 331–334.
14
15 1933 O'CONNOR, P. M. 2009 Evolution of archosaurian body plans: skeletal
16
17 1934 adaptations of an air-sac-based breathing apparatus in birds and other
18
19 1935 archosaurs. *Journal of Experimental Zoology*, **311A**, 629–646.
20
21 1936 O'CONNOR, J., CHIAPPE, L. M. and BELL, A. 2011. *Pre-modern birds: avian*
22
23 1937 *divergences in the Mesozoic*. 39–116. In DYKE, G. and KAISER, G.
24
25 1938 (eds). *Living Dinosaurs: the Evolutionary History of Modern Birds*. Wiley-
26
27 1939 Blackwell, Oxford.
28
29 1940 O'GORMAN, E. J. and HONE, D. W. E. 2012. Body size distribution of the
30
31 1941 dinosaurs. *PLOS ONE*, **7**(12), e51925.
32
33 1942 O'MEARA, B. C., ANÉ, C., SANDERSON, M. J. and WAINWRIGHT, P. C.
34
35 1943 2006. Testing for different rates of continuous trait evolution using
36
37 1944 likelihood. *Evolution*, **60**, 922–933.
38
39 1945 ŐSI, A., PRONDVAI, E., BUTLER, R. and WEISHAMPEL, D. B. 2012.
40
41 1946 Phylogeny, histology and inferred body size evolution in a new
42
43 1947 rhabdodontid dinosaur from the Late Cretaceous of Hungary. *PLOS*
44
45 1948 *ONE*, **7**(9), e44318.
46
47 1949 OYSTON, J. W., HUGHES, M., WAGNER, P. J., GERBER, S. and WILLS, M.
48
49 1950 A. 2015. What limits the morphological disparity of clades? *Interface*
50
51 1951 *Focus* **5**, 20150042.
52
53
54
55
56
57
58
59
60

- 1952 PAGEL, M. 1999. Inferring the historical patterns of biological evolution.
1953 *Nature*, **401**, 877–884.
- 1954 — 2002. Modelling the evolution of continuously varying characters on
1955 phylogenetic trees: the case of hominid cranial capacity. 269–286. *In*
1956 MACLEOD, N. and FOREY, P. L. (eds). *Morphology, Shape and*
1957 *Phylogeny*. Taylor and Francis, London.
- 1958 PARADIS, E., CLAUDE, J. and STRIMMER, K. 2004. APE: analyses of
1959 phylogenetics and evolution in R language. *Bioinformatics*, **20**, 289–290.
- 1960 PINHEIRO, J., BATES, D., DEBROY, S., SARKAR, D. and R CORE TEAM.
1961 2014. nlme: Linear and Nonlinear Mixed Effects Models. R package
1962 version 3.1-117, <URL: <http://CRAN.R-project.org/package=nlme>>.
- 1963 PUTTICK, M. N., THOMAS, G. H. and BENTON, M. J. 2014. High rates of
1964 evolution preceded the origin of birds. *Evolution*, **68**, 1497–1510.
- 1965 R CORE TEAM. 2017. R: A language and environment for statistical
1966 computing. R Foundation for Statistical Computing, Vienna, Austria. URL
1967 <http://www.R-project.org/>.
- 1968 RAUP, D. M. and SEPKOSKI, J. J. 1976. Mass extinctions in the marine fossil
1969 record. *Science*, **215**, 1501–1503.
- 1970 REVELL, L. J. 2012. phytools: An R package for phylogenetic comparative
1971 biology (and other things). *Methods in Ecology and Evolution*, **3**, 217–
1972 223.
- 1973 SAARINEN, J. J., BOYER, A. G., BROWN, J. H., COSTA, D. P., ERNEST, S.
1974 K. M., EVANS, A. R., FORTELIUS, M., GITTLEMAN, J. L., HAMILTON,
1975 M. J., HARDING, L. E., LINTULAAKSO, K., LYONS, S. K., OKIE, J. G.,
1976 SIBLY, R. M., STEPHENS, P. R. THEODOR, J. UHEN, M. D. and

1
2
3 1977 SMITH, F. A. 2014. Patterns of maximum body size evolution in
4
5 1978 Cenozoic land mammals: eco-evolutionary processes and abiotic
6
7 1979 forcing. *Proceedings of the Royal Society B*, **281**, 20132049.
8
9
10 1980 SANDER, P. M. and CLAUSS, M. 2008. Perspective: sauropod gigantism.
11
12 1981 *Science*, **322**, 200–201.
13
14 1982 —, MATEUS, O., LAVEN, T. and KNÖTSCHKE, N. 2006. Bone histology
15
16 1983 indicates insular dwarfism in a new Late Jurassic sauropod dinosaur.
17
18 1984 *Nature*, **441**, 739–741.
19
20
21 1985 —, CHRISTIAN, A., CLAUSS, M., FECHNER, R., GEE, C. T., GRIEBLER, E. -
22
23 1986 M., GUNGA, H. -C., HUMMEL, J., MALLISON, H., PERRY, S. F.,
24
25 1987 PREUSCHOTT, H., RAUHUT, O. W. M., REMES, K., TÜTKEN, T.,
26
27 1988 WINGS, O. and WITZEL, U. 2010. Biology of the sauropod dinosaurs:
28
29 1989 the evolution of gigantism. *Biological Reviews*, **86**, 117–155.
30
31
32 1990 SCHLUTER, D., PRICE, T., MOOERS, A. O. and LUDWIG, D. 1997.
33
34 1991 Likelihood of ancestor states in adaptive radiation. *Evolution*, **51**, 1699–
35
36 1992 1711.
37
38 1993 SERENO, P. C. 1997. The origin and evolution of dinosaurs. *Annual Reviews*
39
40 1994 *of Earth and Planetary Sciences*, **25**, 435–489.
41
42 1995 — 1999. The evolution of dinosaurs. *Science*, **284**, 2137–2147.
43
44
45 1996 SHEETS, H. D. and MITCHELL, C. E. 2001. Why the null matters: statistical
46
47 1997 tests, random walks and evolution. *Genetica*, **112-113**, 105–125.
48
49 1998 SILVESTRO, D., ANTONELLI, A., SALAMIN, N. and QUENTAL, T. B. 2015.
50
51
52 1999 The role of competition in evolutionary replacements of North American
53
54 2000 canids. *Proceedings of the National Academy of Sciences of the USA*,
55
56 2001 **112**, 8684–8689.
57
58
59
60

- 2002 SIMPSON. G. G. 1953. The Major Features of Evolution. Columbia University
2003 Press, New York, 434 pp.
- 2004 SLATER. G. J. 2013. Phylogenetic evidence for a shift in the mode of
2005 mammalian body size evolution at the Cretaceous-Palaeogene
2006 boundary. *Methods in Ecology and Evolution*, **4**, 734–744.
- 2007 — 2015. Iterative adaptive radiations of fossil canids show no evidence for
2008 diversity-dependent trait evolution. *Proceedings of the National Academy*
2009 of Sciences of the U.S.A., **112**, 4897–4902.
- 2010 —, HARMON. L. J. and ALFARO. M. E. 2012. Integrating fossils wit molecular
2011 phylogenies improves inference of trait evolution. *Evolution*, **66**, 3931–
2012 3944.
- 2013 —, GOLDBOGEN, J. A. and PYENSON, N. D. 2017. Independent evolution of
2014 baleen whale gigantism linked to Plio-Pleistocene ocean dynamics.
2015 *Proceedings of the Royal Society B*, **284**, 20170546.
- 2016 SMITH. F. A., BOYER. A. G., BROWN. J. H., COSTA. D. P., DAYAN. T.,
2017 ERNEST. S. K. M. EVANS. A. R., FORTELIUS. M., GITTLEMAN. J. L.,
2018 HAMILTON. M. J., HARDING. L. E., LINTULAAKSO. K., LYONS. S. K.,
2019 MCCAIN. C., OKIE. J. G., SAARINEN. J. J., SIBLY. R. M., STEPHENS.
2020 P. R., THEODOR. J. and UHEN. M. D. 2012. The evolution of maximum
2021 body size of terrestrial mammals. *Science*, **330**, 1216–1219.
- 2022 SOOKIAS, R. B., BUTLER, R. J. and BENSON, R. B. J. 2012. Rise of
2023 dinosaurs reveals major body-size transitions are driven by passive
2024 processes of trait evolution. *Proceedings of the Royal Society B*, **279**,
2025 2180–2187.
- 2026 STANLEY, S. M. 1973. An explanation for Cope's rule. *Evolution*, **27**, 1–26.

1
2
3
4
5
6
7
8
9
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11
12
13
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42
43
44
45
46
47
48
49
50
51
52
53
54
55
56
57
58
59
60

2027 STUBBS, T. L. and BENTON, M. J. 2016. Ecomorphological diversifications
2028 of Mesozoic marine reptiles, the roles of ecological opportunity and
2029 extinction. *Paleobiology*, **42**, 547–573.

2030 SUGIURA, N. 1978. Further analysis of the data by Akaike's information
2031 criterion and the finite corrections. *Communications in Statistics —*
2032 *Theory and Methods*, **7**, 13–26.

2033 TARVER, J. E. and DONOGHUE, P. C. J. 2011. The trouble with topology:
2034 phylogenies without fossils provide a revisionist perspective of
2035 evolutionary history in topological analyses of diversity. *Systematic*
2036 *Biology*, **60**, 700–712.

2037 TURNER, A. H., POL, D., CLARKE, J. A., ERICKSON, G. M. and NORELL,
2038 M. A. 2007. A basal dromaeosaurid and size evolution preceding avian
2039 flight. *Science*, **317**, 1378–1381.

2040 UPCHURCH, P., BARRETT, P. M. and GALTON, P. M. 2007. A phylogenetic
2041 analysis of basal sauropodomorph relationships: Implications for the
2042 origin of sauropod dinosaurs. *Special Papers in Palaeontology*, **77**, 57–
2043 90.

2044 —, HUNN C. A. 2002. “Time”: the neglected dimension in cladistic
2045 biogeography. *Geobios*, **35**, 277–286.

2046 UYEDA, J. C., HANSEN, T. F., ARNOLD, S. J. and PIENAAR, J. 2011. The
2047 million-year wait for macroevolutionary bursts. *Proceedings of the*
2048 *National Academy of Sciences of the USA*, **108**, 15908–15913.

2049 VAN VALKENBURGH, B., WANG, X. -M. and DAMUTH, J. 2004. Cope's rule,
2050 hypercarnivory, and extinction in North American canids. *Science*, **306**,
2051 101–104.

- 2052 VARICCHIO, D. J. 2011. A distinct dinosaur life history? *Historical Biology*,
2053 **23**, 91–107.
- 2054 VERMEIJ, G. J. 1973. Biological versatility and Earth history. *Proceedings of*
2055 *the National Academy of Sciences of the USA*, **70**, 1936–1938.
- 2056 WANG, M., ZHENG, X., O'CONNOR, J. K., LLOYD, G. T., WANG, X.,
2057 WANG, Y., ZHANG, X. and ZHOU, Z. 2015. The oldest record of
2058 Ornithuromorpha from the Early Cretaceous of China. *Nature*
2059 *Communications*, **6**, 6987.
- 2060 WERNER, J. and GRIEBELER, E. M. 2011. Reproductive biology and its
2061 impact on body size: comparative analysis of mammalian, avian and
2062 dinosaurian reproduction. *PLOS ONE*, **6**(12), e28442.
- 2063 WERNING, S. 2012. The ontogenetic osteohistology of *Tenontosaurus tilletti*.
2064 *PLOS One*, **7**, e33539.
- 2065 WOODRUFF, D. C. and FOSTER, J. R. 2014. The fragile legacy of
2066 *Amphicoelias fragillimus* (Dinosauria: Sauropoda; Morrison Formation –
2067 latest Jurassic). *Volumina Jurassica*, **12**, 211–220.
- 2068 YATES, A. M. 2007 The first complete skull of the Triassic dinosaur
2069 *Melanorosaurus* Haughton (Sauropodomorpha: Anchisauria). *Special*
2070 *Papers in Palaeontology*, **77**, 9–55.
- 2071 — and KITCHING, J. W. 2003. The earliest known sauropod dinosaur and the
2072 first steps towards sauropod locomotion. *Proceedings of the Royal*
2073 *Society B*, **270**, 1753–1758.
- 2074 —, BONNAN, M. F., NEVELING, J., CHINSAMY, A. and BLACKBEARD, M.
2075 G. 2010. A new transitional sauropodomorph dinosaur from the Early

2076 Jurassic of South Africa and the evolution of sauropod feeding and
2077 quadrupedalism. *Proceedings of the Royal Society B*, **277**, 787–794.

2078 ZANNO, L. E. and MAKOVICKY, P. J. 2013. No evidence for directional
2079 evolution of body mass in herbivorous theropod dinosaurs. *Proceedings*
2080 *of the Royal Society B*, **280**, 20122526.

2081 ZHANG, F. and ZHOU, Z. 2000. A primitive enantiornithine bird and the origin
2082 of feathers. *Science*, **290**, 1955–1959.

2083 ZHANG, F., ZHOU, Z. and BENTON, M. J. 2008. A primitive confuciusornithid
2084 bird from China and its implications for early avian flight. *Science in*
2085 *China, series D — Earth Sciences*, **51**, 625–639.

2086 ZHAO, Q., BENTON, M. J., SULLIVAN, C., SANDER, P. M. and XU, X. 2013.
2087 Histology and postural change during the growth of the ceratopsian
2088 dinosaur *Psittacosaurus lujiatunensis*. *Nature Communications*, **4**, 2079.

2090 **FIGURE CAPTIONS**

2091 **Fig. 1.** Simulated behavior of macroevolutionary models under various
2092 parameter values. A, B, Brownian motion (BM) under high (A) and low (B)
2093 values of the Brownian variance parameter showing diffusive behavior and
2094 unbiased directionality of trait evolution along lineages; C, D, Ornstein-
2095 Uhlenbeck (OU) models showing bounded or constrained evolution under (C)
2096 a single regime in which the trait optimum (θ) equals the starting values (Z_0),
2097 and (D) in which a second regime originates from the first and has a trait
2098 optimum that is greater than the starting value; E, F, trend models originating
2099 at the base of the phylogeny (E) and descending from an ancestral OU

2100 regime (F). For this figure, a trend model with $\mu = 0.7$ was approximated using
 2101 an OU model with low α ($= 0.005$) and distant, unrealized θ ($= 150$).

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2103 **FIG. 2.** Simulated behavior of new trend-based models under various
 2104 parameter values. A, node-shift model, showing a change in the trend
 2105 parameter μ at a node on the phylogeny. B, time-shift model, showing a
 2106 change in the trend parameter μ at a time during the evolutionary history of a
 2107 group. Z_0 is the trait value at the root node and σ is the Brownian variance.
 2108 For this figure, a trend model with $\mu = 0.7$ was approximated using an OU
 2109 model with low α and distant, unrealized θ .

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2111 **FIG. 3.** Comparisons of AICc scores for SURFACE (multi-regime Ornstein-
 2112 Uhlenbeck), best trend-based, and Brownian motion (BM1) models. A,
 2113 Triassic–Jurassic Dinosauria; B, Sauropodomorpha; C, Ornithischia; and
 2114 Triassic–Aptian Theropoda. Models were fit across 20 phylogenies scaled to
 2115 time using the mbl1 algorithm, and results for each phylogeny are connected
 2116 by lines. Results based on other timescaling algorithms were essentially
 2117 identical. AICc scores for all trend-based models and BM1 models are given
 2118 in Dataset S2. AICc scores and visualizations of SURFACE models are given
 2119 in Appendix S2.

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2121 **FIG. 4.** Plots of $\log_{10}(\alpha)$ on AICc across 20 phylogenies for SURFACE model
 2122 fits showing bimodal distribution of model outcomes for most groups. A,
 2123 Triassic–Jurassic Dinosauria; B, Sauropodomorpha; C, Ornithischia; and
 2124 Triassic–Aptian Theropoda. Models were fit across 20 phylogenies scaled to

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2125 time using the mbl1 algorithm., Tree numbers correspond to those used
2126 throughout this paper, e.g. Appendix S2 and elsewhere.
2127
FIG. 5. Comparison of low-AICc/high- α class and high-AICc/low- α class
2128 SURFACE model fits for Triassic–Jurassic Dinosauria. A, low-AICc/high- α
2129 class of model based on tree 9; B, high-AICc/low- α class of model based on
2130 tree 10. The suboptimal model fit is characterized by a low value of α ($= 0.005$
2131 in (A) compared to 0.074 in (B)) and few distinct regimes. Red and blue
2132 lineages exhibit trend-like attraction to unrealized low (blue) or high (red) trait
2133 optima.
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FIG. 6. SURFACE stepwise-AICc model for phylogeny 12 of Triassic–Jurassic
2136 Dinosauria. Results for other phylogenies show little variation from this
2137 (except that described above; Figs 4–5), and are presented in Appendix S2.
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FIG. 7. SURFACE model fit and regime evolution through time for
2140 Sauropodomorpha. A, SURFACE stepwise-AICc model for phylogeny 17 of
2141 Sauropodomorpha. Results for other phylogenies show little variation from
2142 this and are presented in Appendix S2. B, Evolution of body size regimes in
2143 Sauropodomorpha simplified from (A) by collapsing each phylogenetically-
2144 independent multi-taxon regime to a single branch.
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FIG. 8. SURFACE model fit and regime evolution through time for
2147 Ornithischia. A, SURFACE stepwise-AICc model for phylogeny 5 of
2148 Ornithischia. Results for other phylogenies show little variation from this and

2150 are presented in Appendix S2. B, Evolution of body size regimes in
 2151 Ornithischia, simplified from (A) by collapsing each phylogenetically-
 2152 independent multi-taxon regime to a single branch.

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2154 **Fig. 9.** SURFACE model fit and regime evolution through time for Theropoda,
 2155 showing a regime configuration consistent with a large ancestral body size for
 2156 Tetanurae. A, SURFACE stepwise-AICc model for phylogeny 6 of Theropoda.
 2157 B, Evolution of body size regimes in Theropoda simplified from (A) by
 2158 collapsing each phylogenetically-independent multi-taxon regime to a single
 2159 branch.

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2161 **Fig. 10.** SURFACE model fit and regime evolution through time for
 2162 Theropoda, showing a regime configuration consistent with a smaller
 2163 ancestral body size for Tetanurae, with multiple independent origins of larger
 2164 body size (within Ceratosauria, Megalosauroida, and Allosauroida). A,
 2165 SURFACE stepwise-AICc model for phylogeny 3 of Theropoda. B, Evolution
 2166 of body size regimes in Theropoda simplified from (A) by collapsing each
 2167 phylogenetically-independent multi-taxon regime to a single branch.

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2169 **Fig. 11.** Comparisons of AICc scores for best OU-based models fit using
 2170 OUwie (Table 1; Dataset S3), best trend-based (Dataset S2), and Brownian
 2171 motion (BM1) models. A, Sauropodomorpha; B, Triassic–Aptian Theropoda;
 2172 C, thyreophoran ornithischians; D, marginocephalian ornithischians; and E,
 2173 ornithopod ornithischians. Taxonomic content of ornithischian subtrees
 2174 corresponds to Appendix S5. Models were fit across 20 phylogenies for

2175 Sauropodomorpha, and to 16 phylogenies representing the low-AICc/high- α
2176 cluster for Ornithischia (Fig. 4C). Two phylogenies for Sauropodomorpha
2177 (trees 2 and 8) returned nonsensical parameter estimates (defined in text) and
2178 were discarded. Results for each phylogeny are connected by lines. Trees
2179 supporting Brownian motion or trend-based models are indicated and
2180 explained in the text.

2181

2182 **FIG. 12.** Visualisation of best trend model results for trees on which trend
2183 models had the best AICc weights (Fig. 11). A, Trend-based model for tree 10
2184 of Theropoda. B, Trend-based model for tree 9 of ornithopod ornithischians.
2185 Although the entire trees of Theropoda and Ornithischia are shown, grey-
2186 shaded internodes and terminals were not included in this analysis, which
2187 focused on comparison with OUwie (taxonomic inclusion described in text).
2188 ‘Songlingorn.’ Is an abbreviation of Songlongornithidae, including *Yanornis*
2189 and *Yixianornis* in our tree.

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2191 **FIG. 13.** Disparity of non-pygostylian Dinosauria (shaded polygon) and
2192 dinosaurian subclades (symbols and dashed lines) through the Mesozoic. A,
2193 Sauropodomorpha (non-sauropodan sauropodomorphs) and Sauropoda; B,
2194 Ornithischia; C, Theropoda (non-pygostylian theropods) and Pygostylia.
2195 Analysis uses the following time bins: Middle Triassic, Late Triassic, Early
2196 Jurassic, Middle Jurassic, Late Jurassic, Early Cretaceous, Cenomanian–
2197 Santonian, Campanian, and Maastrichtian. ‘Disparity’ is the standard
2198 deviation of \log_{10} body mass for each clade, and error bars are standard
2199 interquartile ranges of this value from 1000 bootstrapping replicates. These

2200 plots were constructed using the full set of $N = 526$ specimens for which adult
2201 body masses were available.

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2203 **FIG. 14.** Distribution of shifts in optimal body size (θ) between regimes (A–B)
2204 and body mass frequency distributions for dinosaurs (unfilled bars = all
2205 dinosaurs; filled bars = dinosaur subgroups) throughout the Mesozoic (C–F).
2206 A–B regime shifts plotted against θ of the ancestral regime (A) and (B) time.
2207 Large circles represent clade-level regime shifts, and small circles represent
2208 ‘singleton’ regime shifts. The vertical dashed line in A indicates θ for the
2209 ancestral dinosaurian body size regime and the grey box indicates a sparsely
2210 populated region of trait space at intermediate body masses. Dashed vertical
2211 lines in B indicate period and epoch boundaries. Vertical arrows extending up
2212 and down from the x axis indicate the initiation of ‘trend-like’ dynamics on
2213 single branches that give rise to exceptionally large-bodied (up) or small-
2214 bodied (down) taxa. These regimes are characterized by high-magnitude,
2215 unrealised values of theta. Body size frequency distributions for C,
2216 Sauropodomorpha; D, Ornithischia; E, non-pygostylian Theropoda; and F,
2217 Pygostylia. Grey boxes indicates a sparsely-populated region of trait space at
2218 intermediate body masses. Dashed vertical lines indicate the approximate
2219 range (1 kg – 10,000 kg) of body masses classes seen among non-
2220 pygostylian, non-sauropod dinosaurs (with a few exceptions among small
2221 ornithischians, small non-pygostylian theropods, and large ornithischians).
2222 The smallest size classes are entirely occupied by pygostylian birds, and the
2223 largest size class is entirely occupied by sauropod dinosaurs. Panels C–F
2224 were constructed using the full set of $N = 526$ specimens for which adult body

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2225 masses were available. “Sauropodom.” Is an abbreviation of
2226 Sauropodomorpha.
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2228 **FIG. 15.** Histograms showing (\log_{10}) body mass distributions for clades of
2229 dinosaurs among Mesozoic intervals (coloured bars): A–C, Triassic; D–F,
2230 Early Jurassic, G–J, Middle–Late Jurassic; K–N, Early Cretaceous; and O–P,
2231 Late Cretaceous. The unfilled bars indicate the body mass distributions for all
2232 dinosaurs during each interval. The grey rectangle indicates the
2233 underpopulated ‘intermediate’ range of dinosaur body masses. Dashed lines
2234 bracket the approximate minimum (1 kg) and maximum (10,000 kg) body
2235 masses for most dinosaurs (especially Theropoda and Ornithischia). These
2236 plots were constructed using the full set of $N = 526$ specimens for which adult
2237 body masses were available. “Sauropodom.” Is an abbreviation of
2238 Sauropodomorpha.

2239 **TABLES**

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Model	Description	Regimes	Parameters varying among regimes
BM1	Brownian motion	Single regime	NA
BMS	Brownian motion	Multi-regime	σ^2
OU1	Ornstein-Uhlenbeck	Single regime	(Z_0)
OUM	Ornstein-Uhlenbeck	Multi-regime	$\theta, (Z_0)$
OUMV	Ornstein-Uhlenbeck	Multi-regime	$\theta, \sigma^2 (Z_0)$
OUMA	Ornstein-Uhlenbeck	Multi-regime	$\theta, \alpha (Z_0)$
OUMVA	Ornstein-Uhlenbeck	Multi-regime	$\theta, \sigma^2, \alpha (Z_0)$

2241 **Table 1.** List of Ornstein-Uhlenbeck (OU)-based models and Brownian-motion
2242 (BM) models compared in the current work. Z_0 is listed in parentheses
2243 because each model was tested both allowing Z_0 to be distinct from θ and
2244 constraining Z_0 to equal θ for the regime present at the root node of the
2245 phylogeny.

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	Group	Age	Mass estimate (kg)
Theropoda: Pygostylia			
<i>Iberomesornis romerali</i>	Enantiornithes	Early Cretaceous	0.013 (0.0053–0.033)
<i>Paraprotopteryx gracilis</i>	Enantiornithes	Early Cretaceous	0.016 (0.0064–0.040)
<i>Qiliania graffini</i>	Enantiornithes	Early Cretaceous	0.017 (0.0076–0.040)
<i>Huoshanornis huji</i>	Enantiornithes	Early Cretaceous	0.022 (0.0090–0.055)
<i>Cathayornis yandica</i>	Enantiornithes	Early Cretaceous	0.03 (0.013–0.067)
<i>Sinornis santensis</i>	Enantiornithes	Early Cretaceous	0.035 (0.015–0.080)
<i>Alexornis antecedens</i>	Enantiornithes	Early Cretaceous	0.037 (0.018–0.079)
<i>Concornis lacustris</i>	Enantiornithes	Early Cretaceous	0.039 (0.016–0.098)
Theropoda:Maniraptora			
<i>Parvicursor remotus*</i>	Alvarezsauroidea	Late Cretaceous	0.15 (0.073–0.32)
<i>Ceratonykus oculus*</i>	Alvarezsauroidea	Late Cretaceous	0.26 (0.11–0.65)
<i>Microaptor zhaoianus</i>	Dromaeosauridae	Early Cretaceous	0.43 (0.20–0.92)
<i>Epidexipteryx hui</i>	Scansoriopterygidae	Early Cretaceous	0.47 (0.21–1.1)
<i>Rahonavis ostromi</i>	Dromaeosauridae	Late Cretaceous	0.58 (0.31–1.1)
<i>Anchiornis huxleyi</i>	Troodontidae	Late Jurassic	0.58 (0.23–1.4)
<i>Xixianykus zhang</i>	Alvarezsauroidea	Late Cretaceous	0.74 (0.29–1.9)
<i>Mahakala omnogovae</i>	Dromaeosauridae	Late Cretaceous	0.76 (0.41–1.4)
Theropoda			
<i>Procompsognathus triassicus</i>	Theropoda	Triassic	1.1 (0.61–2.1)
<i>Sinosauroptryx prima</i>	Coelurosauria	Early Cretaceous	1.6 (0.65–4.1)
<i>Compsognathus longipes</i>	Coelurosauria	Late Jurassic	2.6 (1.0–6.5)
<i>Velocisaurus unicus</i>	Ceratosauria	Late Cretaceous	3.2 (1.3–7.8)
<i>Tawa hallae</i>	Theropoda	Triassic	3.2 (1.4–7.3)
<i>Hexing qingyi*</i>	Ornithomimosauria	Early Cretaceous	4.2 (1.5–12)
<i>Segisaurus halli</i>	Theropoda	Early Jurassic	5 (1.7–14)
<i>Eodromaeus murphi</i>	Theropoda	Triassic	7.1 (2.5–20)
Sauropodomorpha			
<i>Pampadromaeus barberenai</i>	Sauropodomorpha	Triassic	8.5 (4.6–16)
<i>Saturnalia tupiniquim</i>	Sauropodomorpha	Triassic	11 (5.7–20)

<i>Chromogisaurus novasi</i>	Sauropodomorpha	Triassic	13 (7.0–24)
<i>Plateosaurus engelhardti</i>	Sauropodomorpha	Triassic	920 (490–1700)
Ornithischia			
<i>Fruitadens haagarorum</i>	Heterodontosauridae	Late Jurassic	0.73 (0.39–1.4)
<i>Tianyulong confuciusi</i>	Heterodontosauridae	Late Jurassic	0.89 (0.42–1.9)
<i>Abriktosaurus consors</i>	Heterodontosauridae	Early Jurassic	1.4 (0.68–3.0)
<i>Psittacosaurus sinensis</i>	Ceratopsia	Early Cretaceous	4.6 (1.8–11)
<i>Yueosaurus tiantaiensis</i>	Thescelosauridae	Early Cretaceous	4.6 (2.2–9.6)
<i>Gongbusaurus wucaiwansensis</i>	Ornithischia	Late Jurassic	4.9 (2.6–9.2)
<i>Heterodontosaurus tucki</i>	Heterodontosauridae	Early Jurassic	5.2 (1.8–15)
<i>Xiaosaurus dashanpensis</i>	Ornithischia	Late Jurassic	5.3 (2.3–12)

Table 2. Body mass estimates for the smallest dinosaurs. Body mass estimates and their $\pm 95\%$ confidence intervals (in brackets) are given for major groups within Dinosauria, demonstrating the uniquely small body masses of Pygostylian birds. A complete set of mass estimates and their standard errors is given in Dataset S1.

			Mass estimate
	Group	Age	(kg)
Saurischia: ?Theropoda			
<i>Eoraptor lunensis</i>	Theropoda	Triassic: Carnian	17 (9.3–32)
<i>Herrerasaurus ischigualastensis</i>	Theropoda	Triassic: Carnian	270 (150–510)
Theropoda			
<i>Staurikosaurus pricei</i>	Theropoda	Triassic: Carnian	23 (12–43)
<i>Coelophysis bauri</i>	Theropoda	Triassic: Norian	9.8 (5.3–18)
<i>Gojirasaurus quayi</i>	Theropoda	Triassic: Norian	190 (45–820)
<i>Guaibasaurus candelariensis</i>	Theropoda	Triassic: Norian	33 (16–72)
<i>Procompsognathus triassicus</i>	Theropoda	Triassic: Norian	1.1 (0.61–2.1)
<i>Tawa hallae</i>	Theropoda	Triassic: Norian	3.2 (1.4–7.3)
<i>Liliensternus liliensterni</i>	Theropoda	Triassic: Rhaetian	84 (45–160)
Saurischia:			
?Sauropodomorpha			
<i>Eodromaeus murphi</i>	Theropoda	Triassic: Carnian	7.1 (2.5–20)
Sauropodomorpha			
<i>Chromogisaurus novasi</i>	Sauropodomorpha	Triassic: Carnian	13 (7.0–24)
<i>Pampadromaeus barberenai</i>	Sauropodomorpha	Triassic: Carnian	8.5 (4.6–16)
<i>Panphagia protos</i>	Sauropodomorpha	Triassic: Carnian	12 (3.2–49)
<i>Saturnalia tupiniquim</i>	Sauropodomorpha	Triassic: Carnian	11 (5.7–20)
Ornithischia			
<i>Pisanosaurus mertii</i>	Ornithischia	Triassic: Carnian	35 (17–76)
		Triassic:	
<i>Eocursor parvus</i>	Ornithischia	Norian/Rhaetian	4.2 (2.0–8.8)

2256 **Table 3. Body mass estimates for Triassic dinosaurs**, especially for the
2257 earliest Triassic dinosaurs (Carnian stage). Body mass estimates and their
2258 ±95% confidence intervals (in brackets) are given for major groups within
2259 Dinosauria, demonstrating the autapomorphically large body mass of

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3 2260 *Herrerasarus* compared to other Carnian dinosaurs (see also Benson *et al.*
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5 2261 2014a). A complete set of mass estimates and their standard errors is given in
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7 2262 Dataset S1.
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	Group	Age	Mass estimate (kg)
Theropoda: Maniraptora			
<i>Deinococheirus mirificus</i>	Ornithomimosauria	Late Cretaceous	7300 (3900-14000)
<i>Segnosaurus galbinensis</i>	Therizinosauria	Late Cretaceous	4600 (2000–11000)
<i>Suzhousaurus megatherioides</i>	Therizinosauria	Early Cretaceous	3100 (1500–6700)
<i>Gigantoraptor erlianensis</i>	Oviraptorosauria	Late Cretaceous	2000 (1100-3700)
<i>Nothronychus graffami</i>	Therizinosauria	Late Cretaceous	1200 (640–2200)
<i>Austroraptor cabazai</i>	Dromaeosauridae	Late Cretaceous	600 (260–1400)
<i>Beishanlong grandis</i>	Ornithomimosauria	Late Cretaceous	580 (250–1300)
<i>Nothronychus mckinleyi</i>	Therizinosauria	Late Cretaceous	570 (130–2500)
Theropoda			
<i>Tyrannosaurus rex</i>	Tyrannosauroidae	Late Cretaceous	7700 (4100-14000)
<i>Giganotosaurus carolinii</i> ¹	Allosauroidae	Late Cretaceous	6100 (3300-11000)
<i>Tyrannotitan chubutensis</i>	Allosauroidae	Early Cretaceous	5400 (2900-10000)
<i>Saurophaganax maximus</i>	Allosauroidae	Late Jurassic	3800 (2000-7000)
<i>Chilantaisaurus tashuikouensis</i>	Allosauroidae	Late Cretaceous	3600 (1900-6700)
<i>Lametasaurus indicus</i> ²	Ceratosauria	Late Cretaceous	3600 (1400–9200)
<i>Acrocanthosaurus atokensis</i>	Allosauroidae	Early Cretaceous	3500 (1900-6400)
<i>Carcharodontosaurus saharicus</i>	Allosauroidae	Late Cretaceous	3300 (1400–7600)
Sauropoda			
<i>Argentinosaurus huinculensis</i> ³	Titanosauria	Late Cretaceous	95000 (32000–280000)
<i>Dreadnoughtus schrani</i>	Titanosauria	Late Cretaceous	59000 (29000–120000)
<i>Brachiosaurus altithorax</i>	Macronaria	Late Jurassic	58000 (23000–150000)
<i>Ruyangosaurus giganteus</i>	Titanosauriformes	Late Cretaceous	54000 (19000–160000)
<i>Turiasaurus riodevensis</i>	Turiasauria	Late Jurassic	51000 (25000–100000)
<i>Paralititan stromeri</i>	Titanosauria	Late Cretaceous	51000 (17000–150000)
		Late Jurassic/	
<i>Losillasaurus giganteus</i>	Turiasauria	Early Cretaceous	47000 (17000–130000)
<i>Argyrosaurus superbus</i>	Titanosauria	Late Cretaceous	42000 (14000–120000)
Ornithischia			
<i>Shantungosaurus giganteus</i>	Hadrosauroidae	Late Cretaceous	17000 (8600–35000)

<i>Iguanodon seelyi</i>	Iguanodontia	Early Cretaceous	15000 (7600–31000)
<i>Triceratops horridus</i>	Ceratopsia	Late Cretaceous	13000 (6500–27000)
<i>Pentaceratops sternbergii</i>	Ceratopsia	Late Cretaceous	11000 (5300–22000)
<i>Torosaurus latus</i>	Ceratopsia	Late Cretaceous	9700 (1400–66000)
<i>Magnapaulia laticaudus</i>	Hadrosauroidea	Late Cretaceous	8600 (2000–38000)
<i>Iguanodon bernissartensis</i>	Iguanodontia	Early Cretaceous	8300 (4100–17000)
<i>Edmontosaurus regalis</i>	Hadrosauroidea	Late Cretaceous	7600 (3700–15000)

Table 4. Body mass estimates for the largest dinosaurs. Body mass estimates and their $\pm 95\%$ confidence intervals (in brackets) are given for major groups within Dinosauria, demonstrating the uniquely large body masses of sauropod dinosaurs. A complete set of mass estimates and their standard errors is given in Dataset S1. Notes: 1, The body mass of *Giganotosaurus* is given here, but the phenotypically similar taxon *Mapusaurus* also attained large body size. 2, *Lametasaurus indicus* has a highly robust tibia (Matley 1924; Carrano 2007) that provides our highest estimated mass among ceratosaurian theropods. However, *Pycnonemosaurus*, which was not included in our dataset, has comparable measurements of the tibia shaft and is likely of comparable body mass (Kellner and Campos 2002; Grillo and Delcourt 2017). 3, Our estimated body mass for *Argentinosaurus* is based on a referred femur housed at Museo de la Plata, La Plata, Argentina. A robust body mass estimate around 70 tonnes was reported using multiple methods for a nearly-complete skeleton of a related taxon, *Patagotitan*, shortly before publication of the current work (Carballido *et al.* 2017).

	Group	Age	Mass estimate (kg)
Pygostylia			
<i>Gargantuavis philoinos</i>	Avialae indet.	Late Cretaceous	190 (100–350)
<i>Hesperornis crassipes</i>	Ornithuromorpha	Late Cretaceous	25 (11–53)
<i>Baptornis varneri</i>	Ornithuromorpha	Late Cretaceous	11 (5.0–23)
<i>Didactylornis jii</i> (= <i>Sapeornis</i>)	Avialae	Early Cretaceous	2.9 (1.1–7.2)
<i>Hollanda luceria</i>	Ornithuromorpha	Late Cretaceous	2.8 (1.3–6.0)
<i>Patagopteryx deferrariisi</i>	Ornithuromorpha	Late Cretaceous	2.6 (1.2–5.5)
<i>Jeholornis prima</i>	Avialae	Early Cretaceous	2.4 (0.95–6.1)
<i>Vorona berivotrensis</i>	Ornithuromorpha	Late Cretaceous	1.8 (0.85–3.9)
Sauropoda			
<i>Magyarosaurus dacus</i>	Titanosauria	Late Cretaceous	750 (370–1500)
<i>Europasaurus holgeri</i>	Titanosauriformes	Late Jurassic	1050 (510–2100)
<i>Lirainosaurus astibae</i>	Titanosauria	Late Cretaceous	1800 (740–4400)
<i>Dystrophaeus viaemalae</i>	Diplodocoidea	Late Jurassic	3100 (1100–8400)
<i>Gondwanatitan faustoi</i>	Titanosauria	Late Cretaceous	3600 (1300–9700)

Table 5. Body mass estimates for the largest pygostylians and smallest sauropods. Body mass estimates and their $\pm 95\%$ confidence intervals (in brackets) are given. A complete set of mass estimates and their standard errors is given in Dataset S1. Body masses for the smallest pygostylians and largest sauropods are shown in Tables 3 and 4.

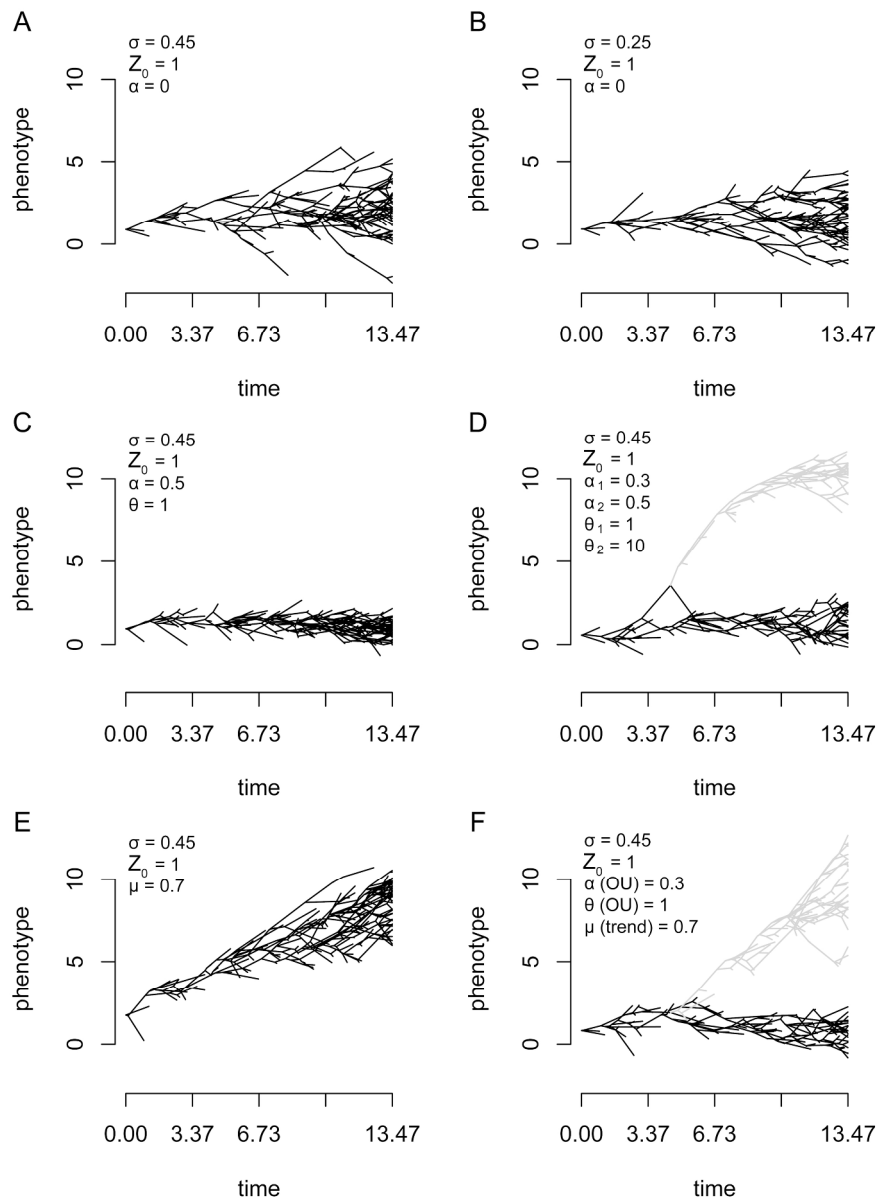


FIG. 1. Simulated behavior of macroevolutionary models under various parameter values. A, B, Brownian motion (BM) under high (A) and low (B) values of the Brownian variance parameter showing diffusive behavior and unbiased directionality of trait evolution along lineages; C, D, Ornstein-Uhlenbeck (OU) models showing bounded or constrained evolution under (C) a single regime in which the trait optimum (θ) equals the starting values (Z_0), and (D) in which a second regime originates from the first and has a trait optimum that is greater than the starting value; E, F, trend models originating at the base of the phylogeny (E) and descending from an ancestral OU regime (F). For this figure, a trend model with $\mu = 0.7$ was approximated using an OU model with low α ($= 0.005$) and distant, unrealized θ ($= 150$).

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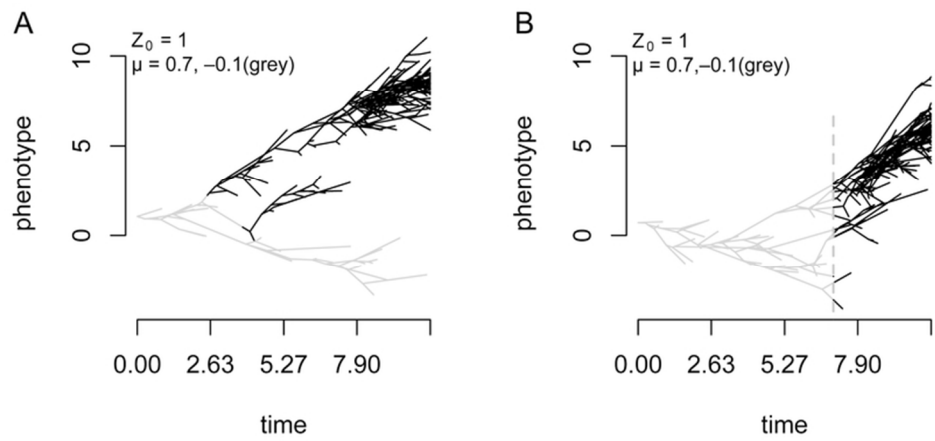


FIG. 2. Simulated behavior of new trend-based models under various parameter values. A, node-shift model, showing a change in the trend parameter μ at a node on the phylogeny. B, time-shift model, showing a change in the trend parameter μ at a time during the evolutionary history of a group. Z_0 is the trait value at the root node and σ is the Brownian variance. For this figure, a trend model with $\mu = 0.7$ was approximated using an OU model with low α and distant, unrealized θ .

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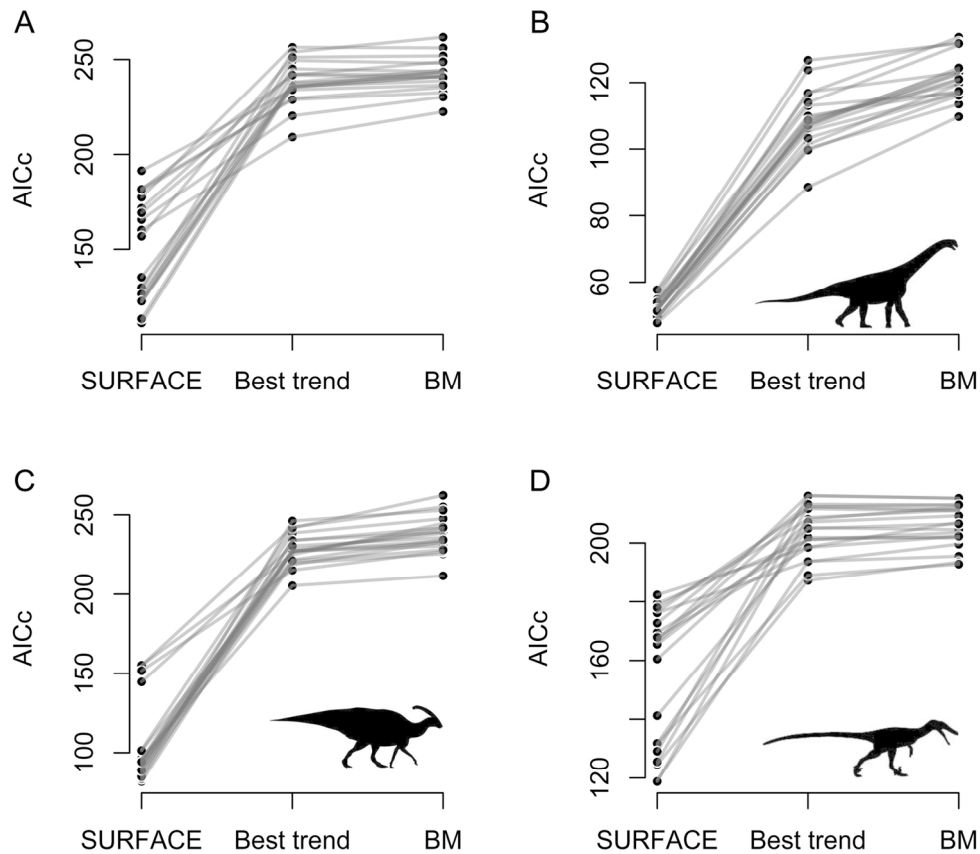


FIG. 3. Comparisons of AICc scores for SURFACE (multi-regime Ornstein-Uhlenbeck), best trend-based, and Brownian motion (BM1) models. A, Triassic-Jurassic Dinosauria; B, Sauropodomorpha; C, Ornithischia; and Triassic-Aptian Theropoda. Models were fit across 20 phylogenies scaled to time using the mbl1 algorithm, and results for each phylogeny are connected by lines. Results based on other timescaling algorithms were essentially identical. AICc scores for all trend-based models and BM1 models are given in Dataset S2. AICc scores and visualizations of SURFACE models are given in Appendix S2.

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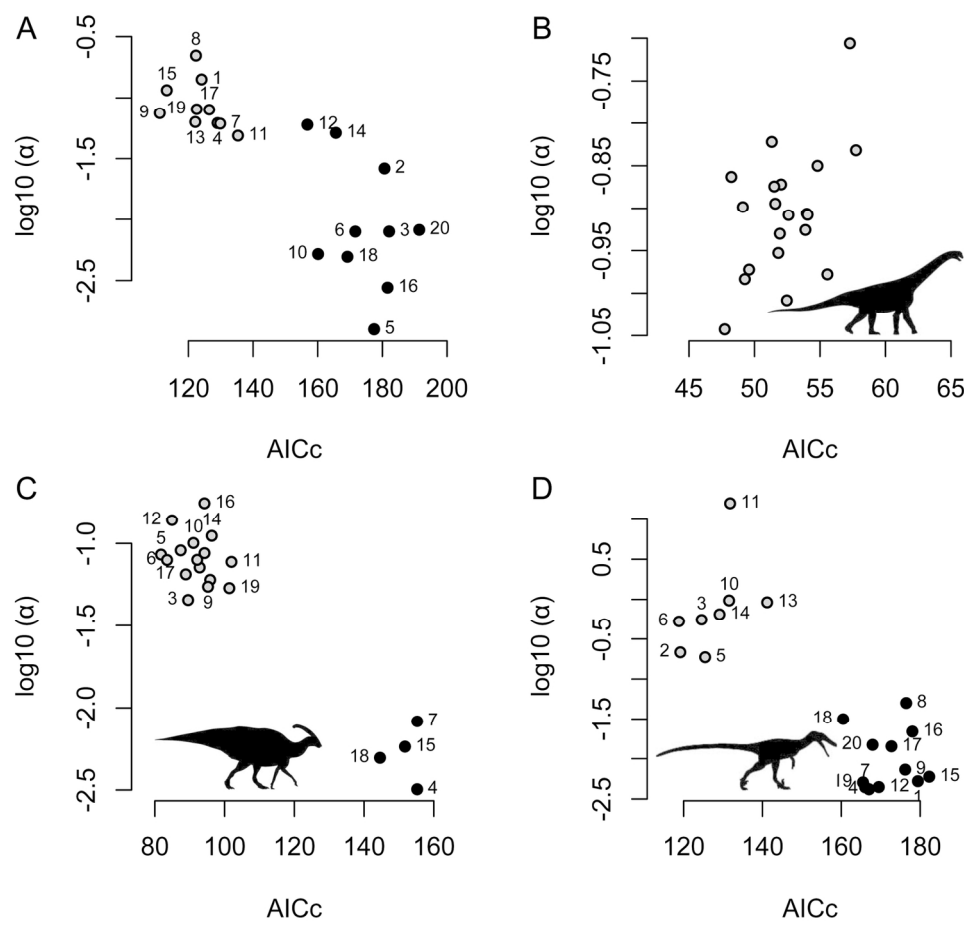


FIG. 4. Plots of $\log_{10}(\alpha)$ on AICc across 20 phylogenies for SURFACE model fits showing bimodal distribution of model outcomes for most groups. A, Triassic–Jurassic Dinosauria; B, Sauropodomorpha; C, Ornithischia; and Triassic–Aptian Theropoda. Models were fit across 20 phylogenies scaled to time using the mbl1 algorithm., Tree numbers correspond to those used throughout this paper, e.g. Appendix S2 and elsewhere.

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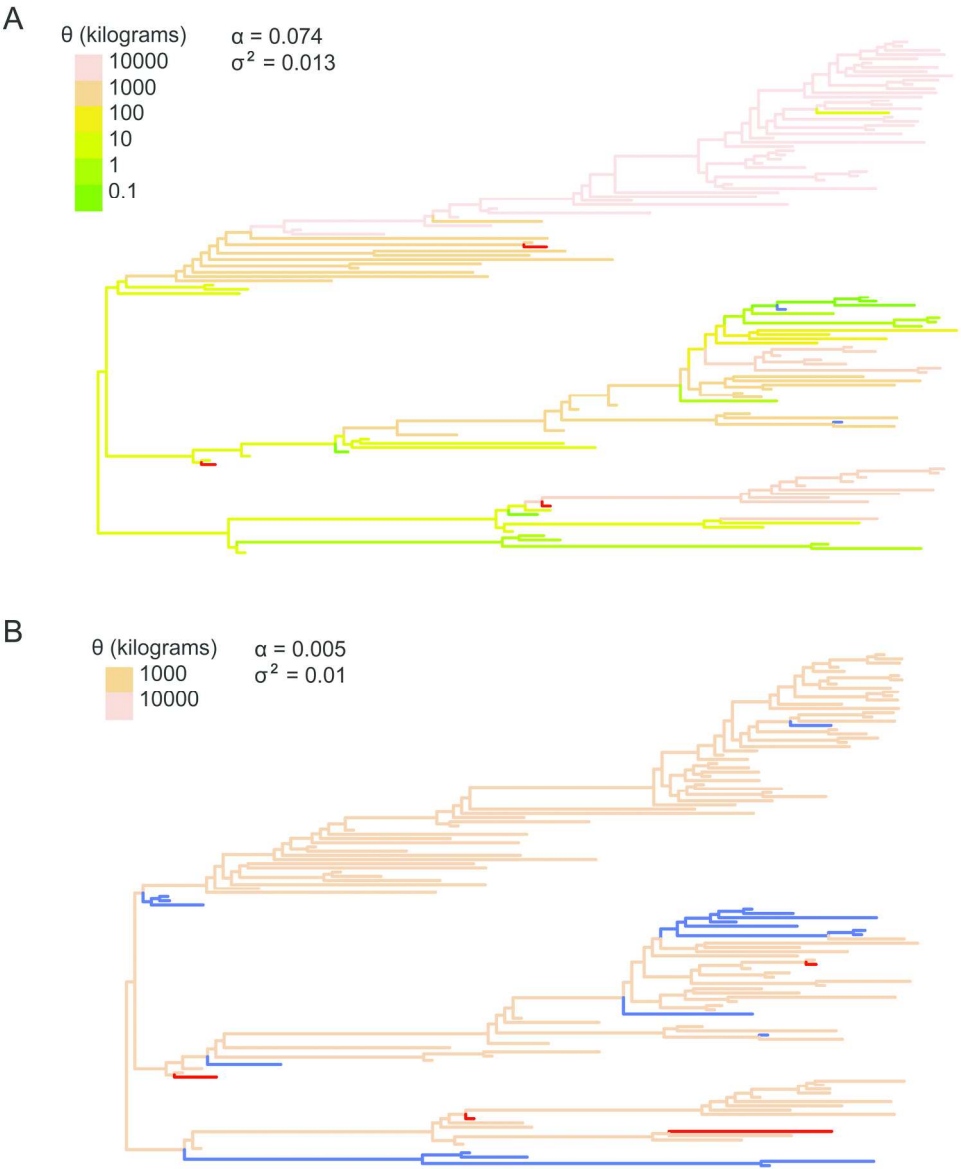


FIG. 5. Comparison of low-AICc/high- α class and high-AICc/low- α class SURFACE model fits for Triassic–Jurassic Dinosauria. A, low-AICc/high- α class of model based on tree 9; B, high-AICc/low- α class of model based on tree 10. The suboptimal model fit is characterized by a low value of α ($= 0.005$ in (A) compared to 0.074 in (B)) and few distinct regimes. Red and blue lineages exhibit trend-like attraction to unrealized low (blue) or high (red) trait optima.

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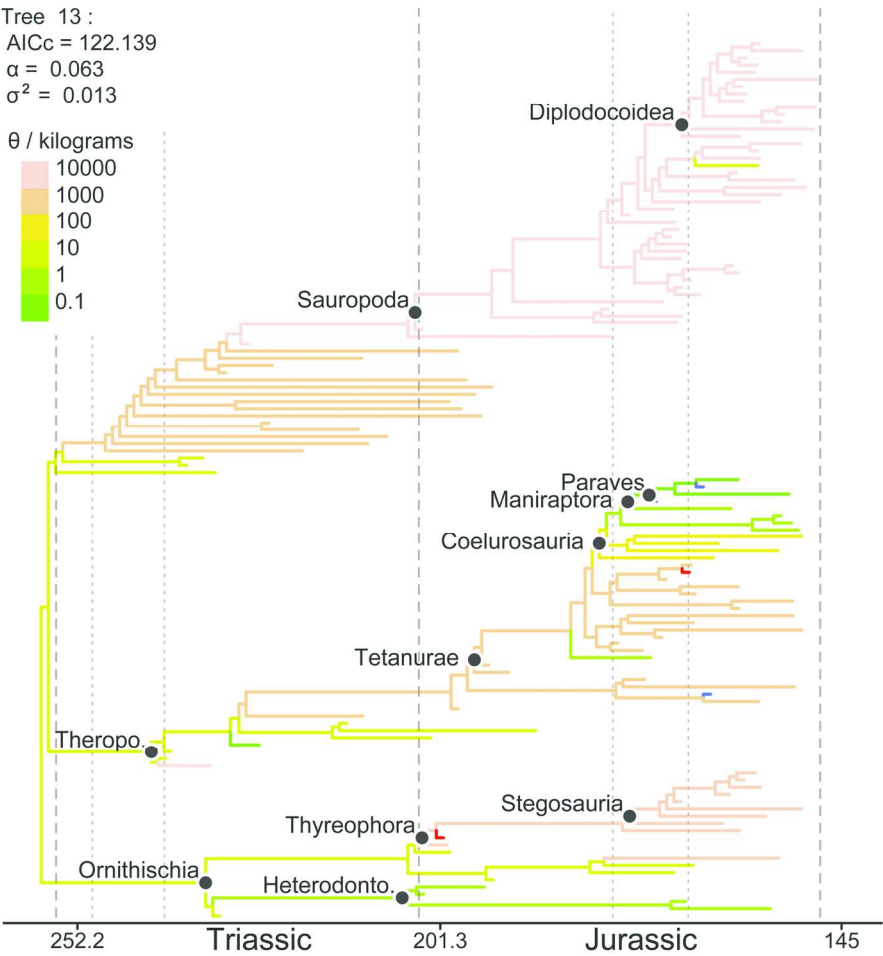


FIG. 6. SURFACE stepwise-AICc model for phylogeny 12 of Triassic–Jurassic Dinosauria. Results for other phylogenies show little variation from this (except that described above; Figs 4–5), and are presented in Appendix S2.

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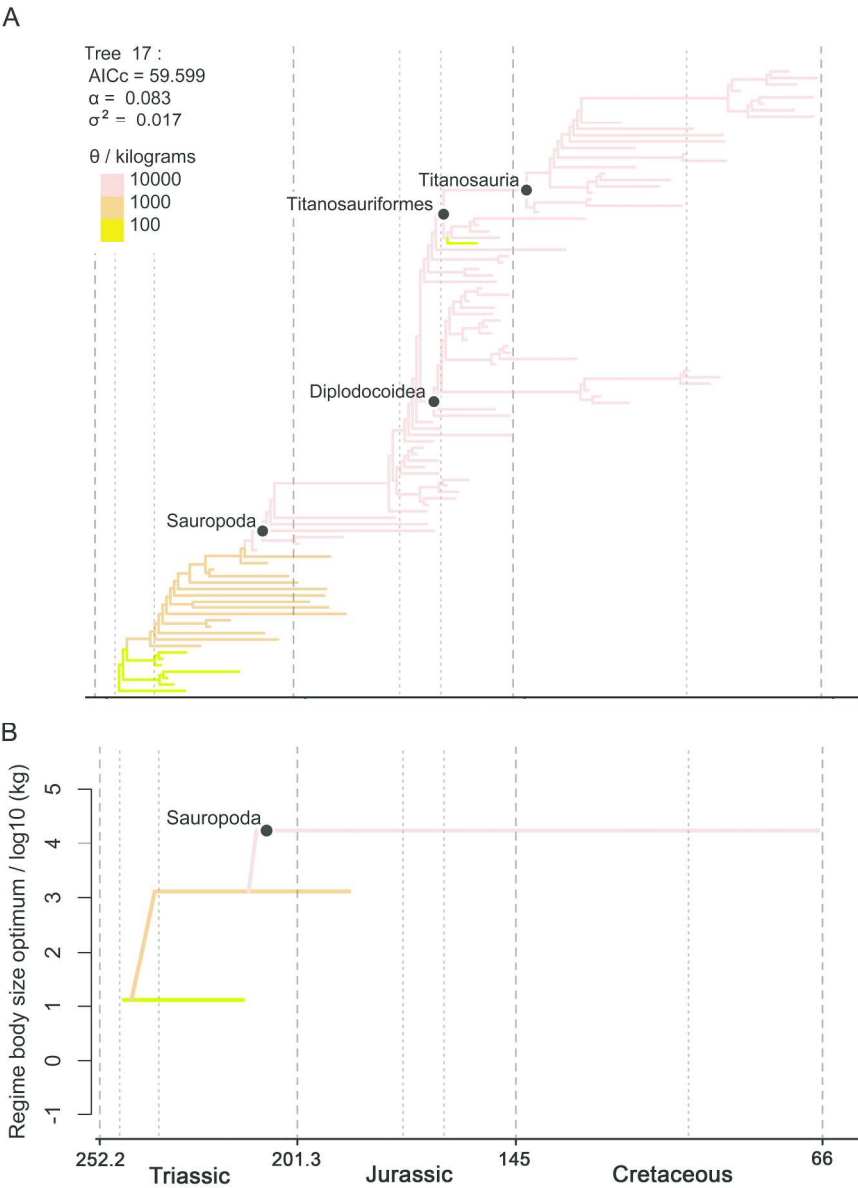


FIG. 7. SURFACE model fit and regime evolution through time for Sauropodomorpha. A, SURFACE stepwise-AICc model for phylogeny 17 of Sauropodomorpha. Results for other phylogenies show little variation from this and are presented in Appendix S2. B, Evolution of body size regimes in Sauropodomorpha simplified from (A) by collapsing each phylogenetically-independent multi-taxon regime to a single branch.

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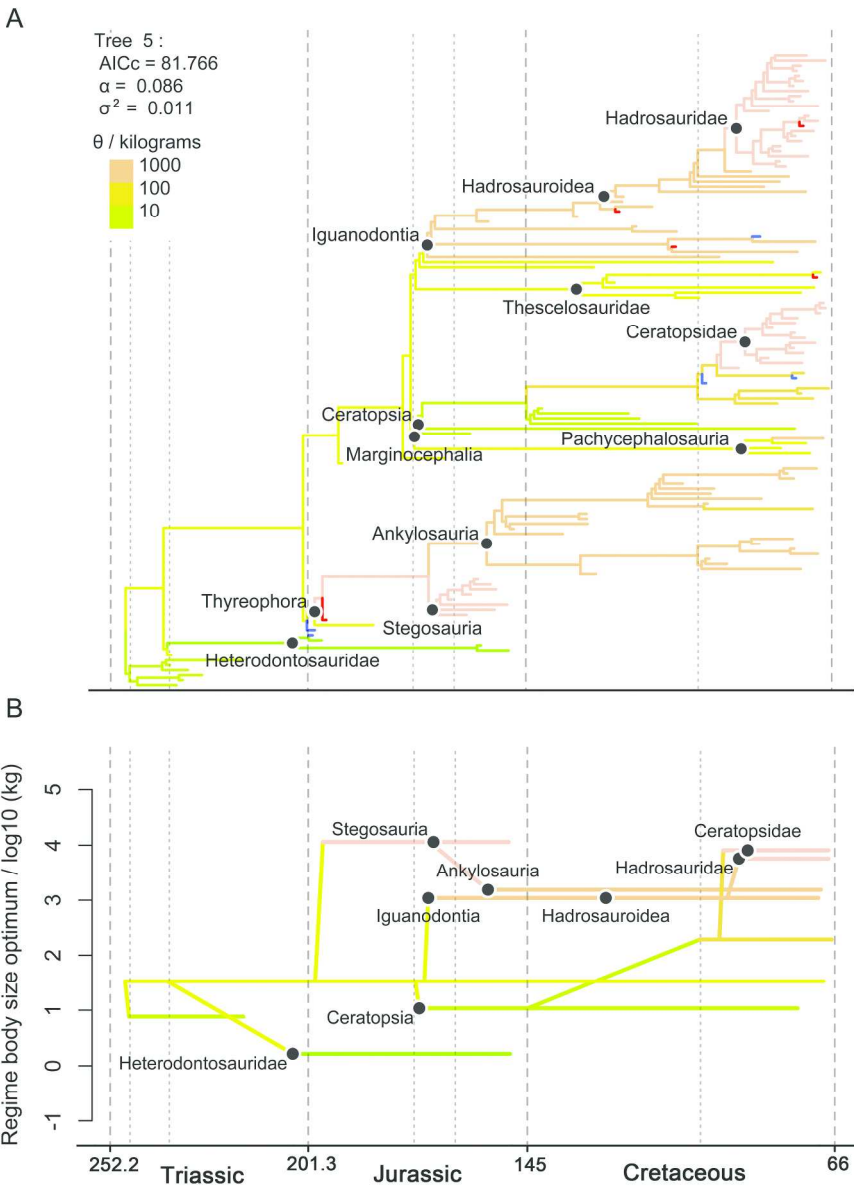


FIG. 8. SURFACE model fit and regime evolution through time for Ornithischia. A, SURFACE stepwise-AICc model for phylogeny 5 of Ornithischia. Results for other phylogenies show little variation from this and are presented in Appendix S2. B, Evolution of body size regimes in Ornithischia, simplified from (A) by collapsing each phylogenetically-independent multi-taxon regime to a single branch.

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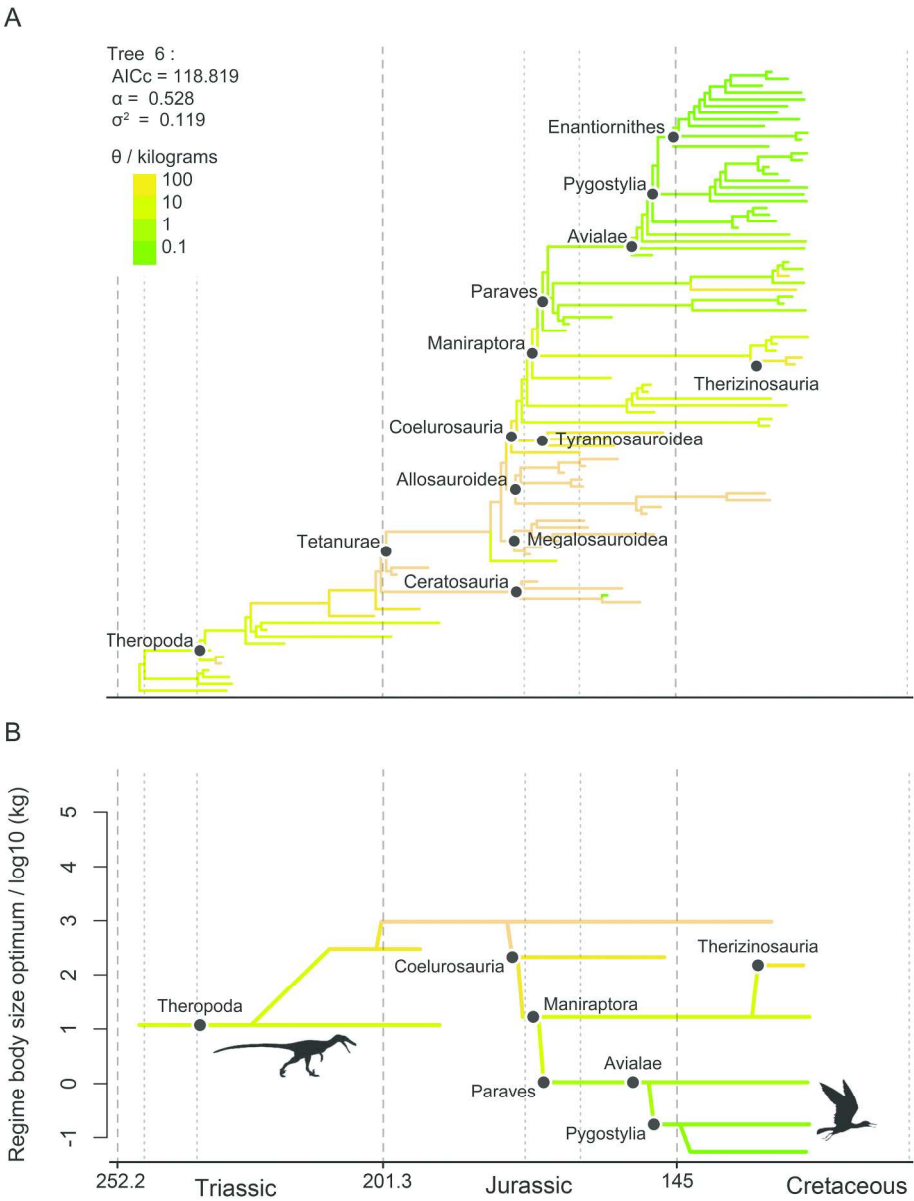


FIG. 9. SURFACE model fit and regime evolution through time for Theropoda, showing a regime configuration consistent with a large ancestral body size for Tetanurae. A, SURFACE stepwise-AICc model for phylogeny 6 of Theropoda. B, Evolution of body size regimes in Theropoda simplified from (A) by collapsing each phylogenetically-independent multi-taxon regime to a single branch.

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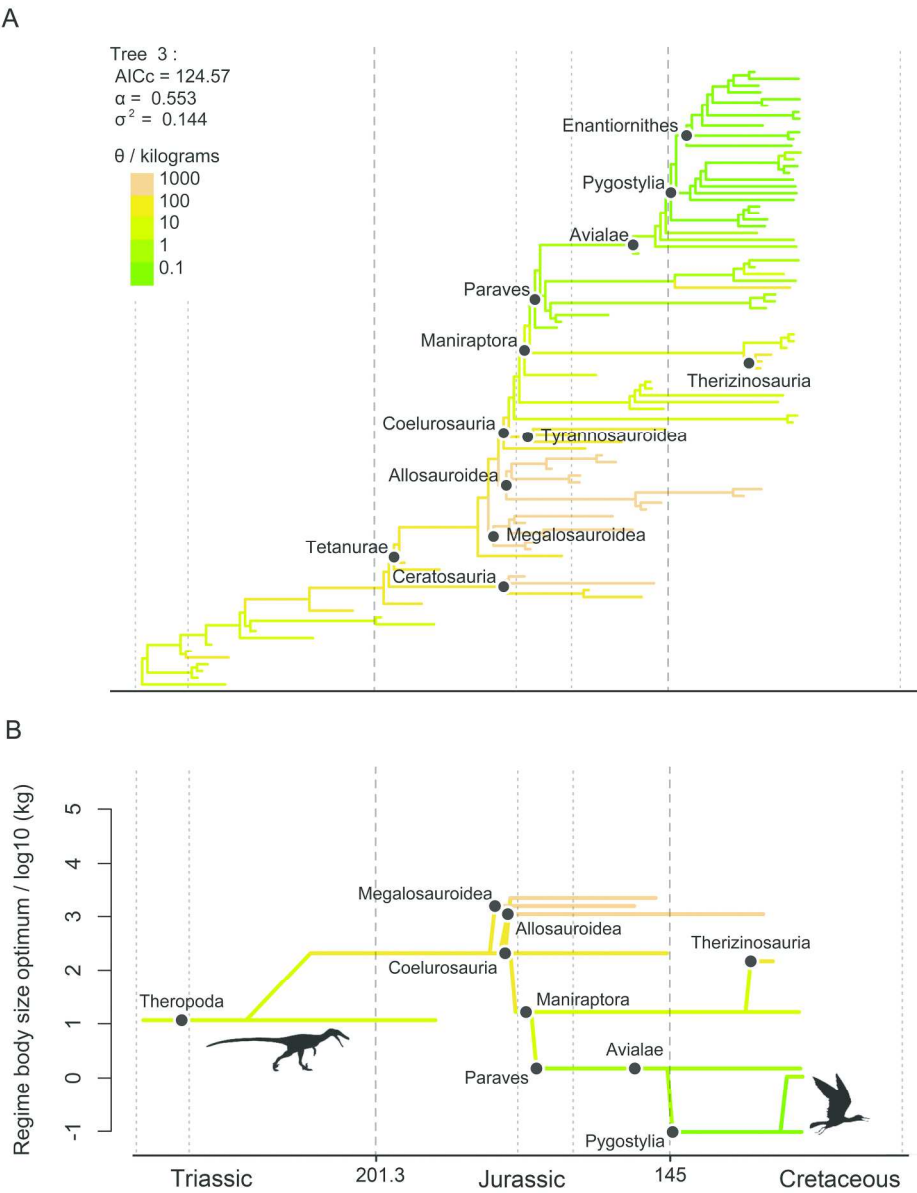


FIG. 10. SURFACE model fit and regime evolution through time for Theropoda, showing a regime configuration consistent with a smaller ancestral body size for Tetanurae, with multiple independent origins of larger body size (within Ceratosauria, Megalosauroidae, and Allosauroidae). A, SURFACE stepwise-AICc model for phylogeny 3 of Theropoda. B, Evolution of body size regimes in Theropoda simplified from (A) by collapsing each phylogenetically-independent multi-taxon regime to a single branch.

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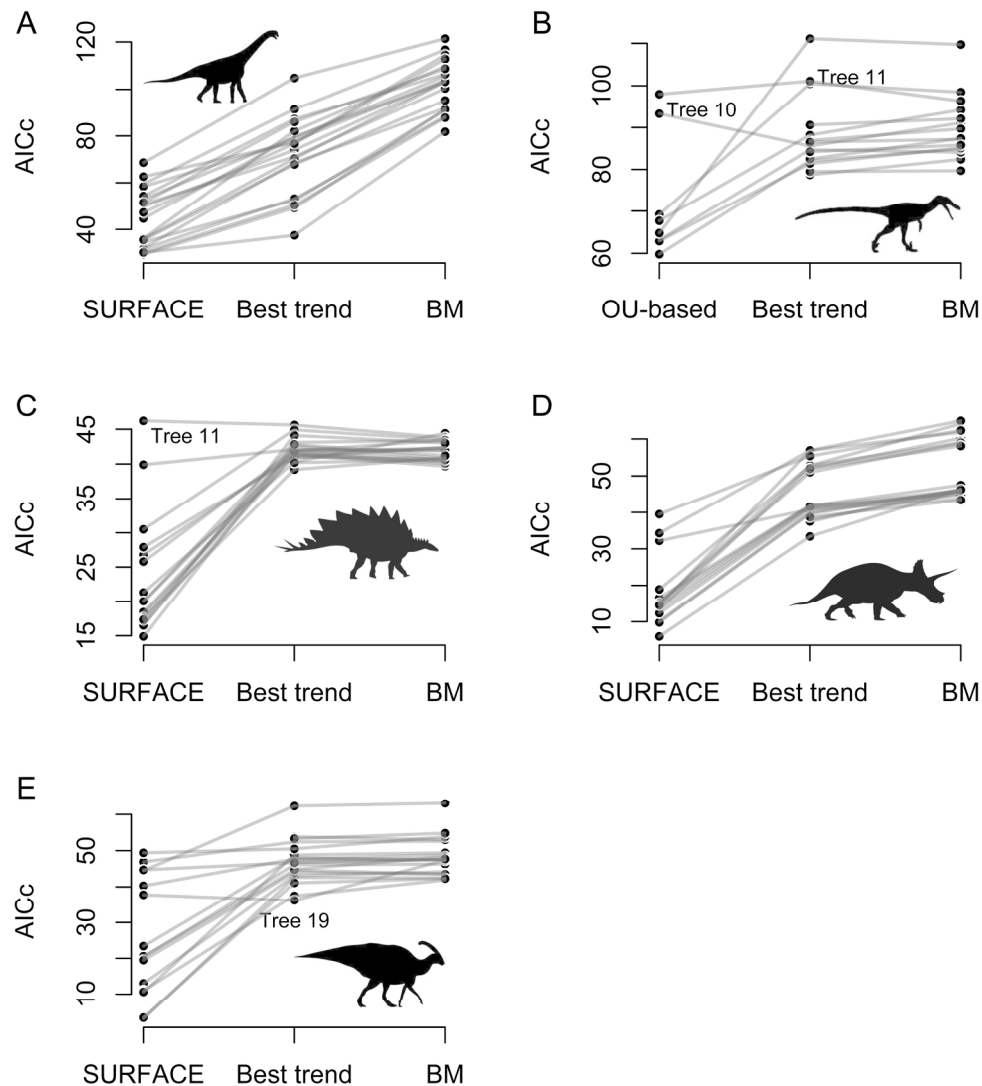


FIG. 11. Comparisons of AICc scores for best OU-based models fit using OUwie (Table 1; Dataset S3), best trend-based (Dataset S2), and Brownian motion (BM1) models. A, Sauropodomorpha; B, Triassic-Aptian Theropoda; C, thyreophoran ornithischians; D, marginocephalian ornithischians; and E, ornithopod ornithischians. Taxonomic content of ornithischian subtrees corresponds to Appendix S5. Models were fit across 20 phylogenies for Sauropodomorpha, and to 16 phylogenies representing the low-AICc/high- α cluster for Ornithischia (Fig. 4C). Two phylogenies for Sauropodomorpha (trees 2 and 8) returned nonsensical parameter estimates (defined in text) and were discarded. Results for each phylogeny are connected by lines. Trees supporting Brownian motion or trend-based models are indicated and explained in the text.

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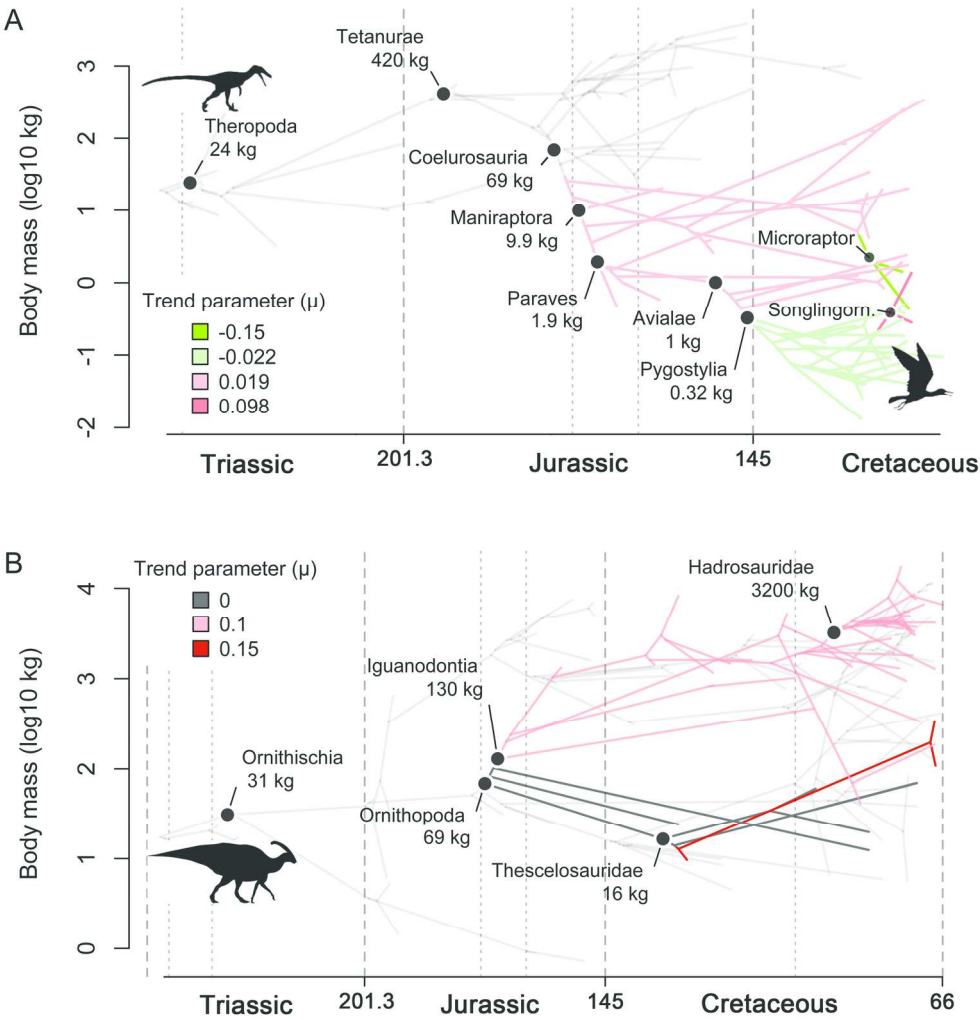


FIG. 12. Visualisation of best trend model results for trees on which trend models had the best AICc weights (Fig. 11). A, Trend-based model for tree 10 of Theropoda. B, Trend-based model for tree 9 of ornithopod ornithischians. Although the entire trees of Theropoda and Ornithischia are shown, grey-shaded internodes and terminals were not included in this analysis, which focused on comparison with OUwie (taxonomic inclusion described in text). ‘Songlingorn.’ Is an abbreviation of Songlongornithidae, including Yanornis and Yixianornis in our tree.

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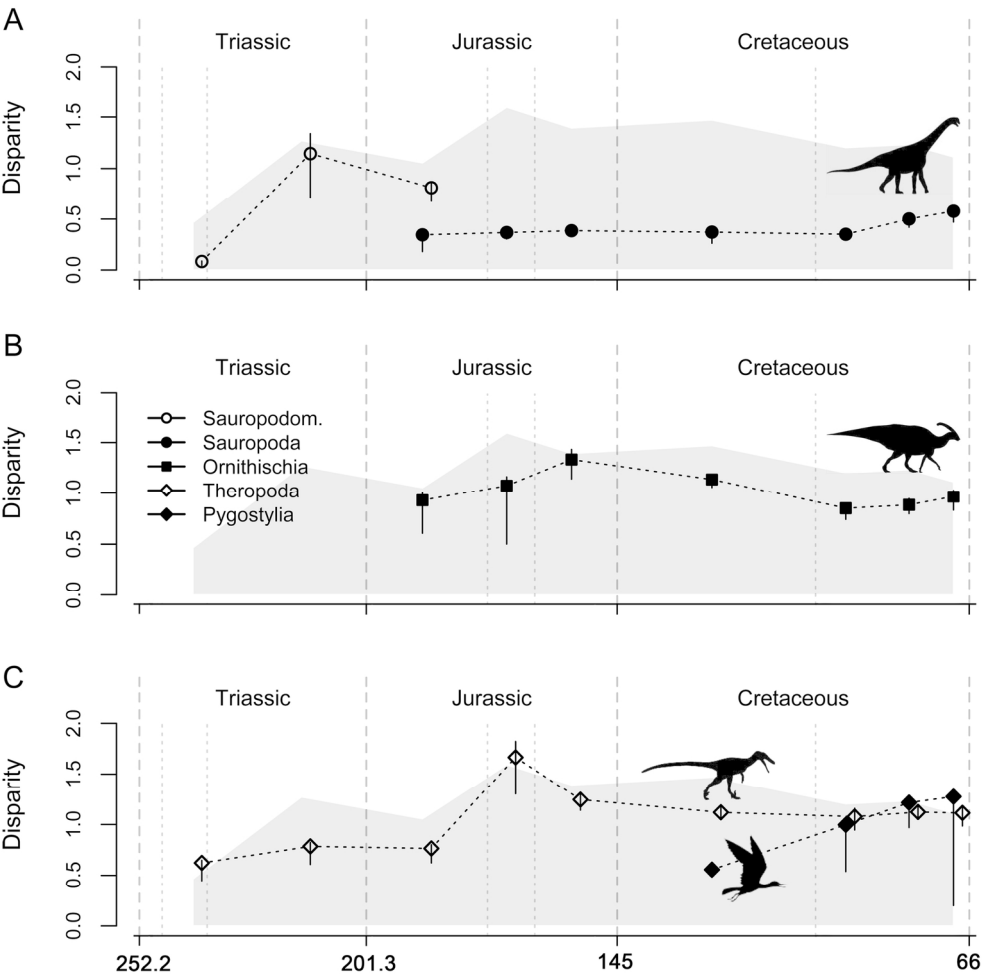


FIG. 13. Disparity of non-pygostylian Dinosauria (shaded polygon) and dinosaurian subclades (symbols and dashed lines) through the Mesozoic. A, Sauropodomorpha (non-sauropodan sauropodomorphs) and Sauropoda; B, Ornithischia; C, Theropoda (non-pygostylian theropods) and Pygostylia. Analysis uses the following time bins: Middle Triassic, Late Triassic, Early Jurassic, Middle Jurassic, Late Jurassic, Early Cretaceous, Cenomanian–Santonian, Campanian, and Maastrichtian. 'Disparity' is the standard deviation of log10 body mass for each clade, and error bars are standard interquartile ranges of this value from 1000 bootstrapping replicates. These plots were constructed using the full set of N = 526 specimens for which adult body masses were available.

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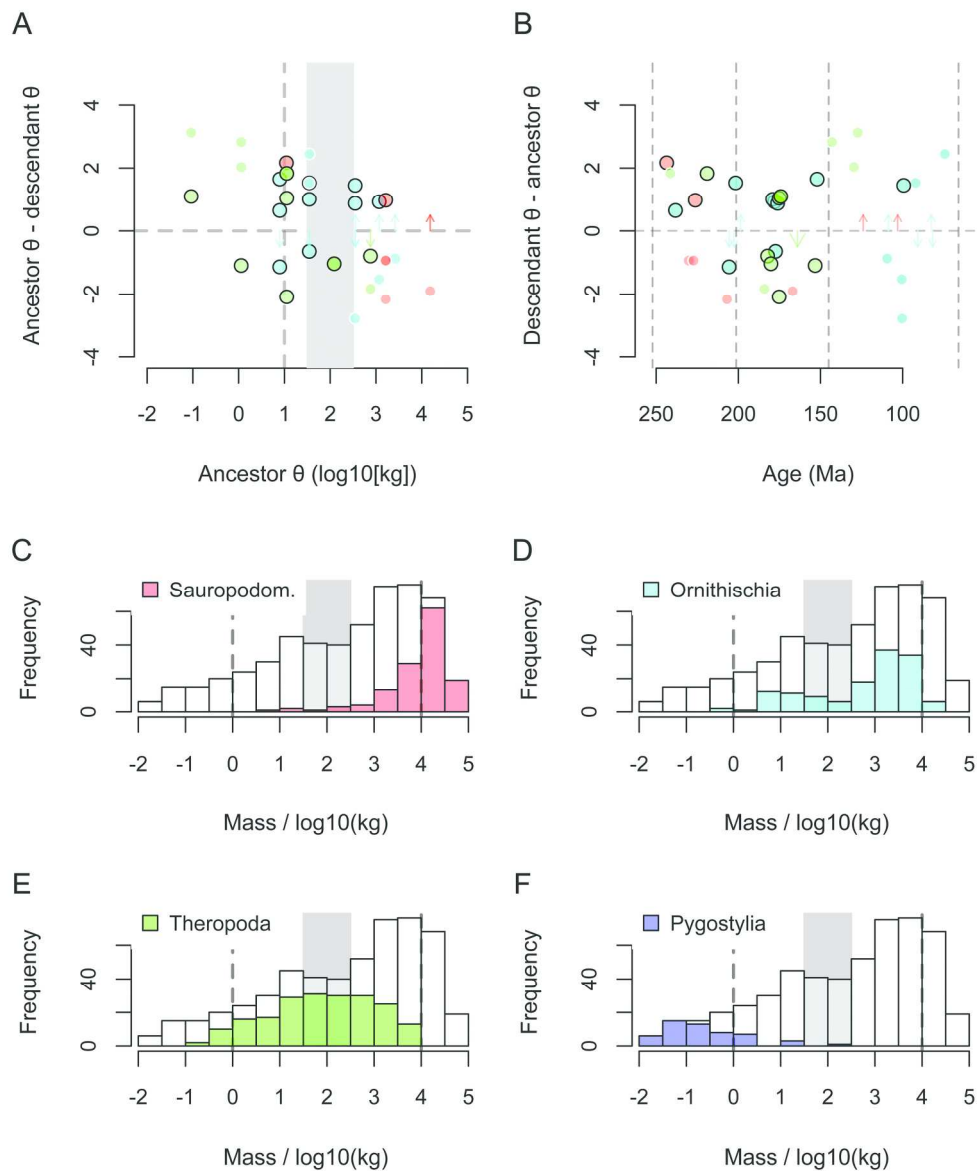


FIG. 14. Distribution of shifts in optimal body size (θ) between regimes (A–B) and body mass frequency distributions for dinosaurs (unfilled bars = all dinosaurs; filled bars = dinosaur subgroups) throughout the Mesozoic (C–F). A–B regime shifts plotted against θ of the ancestral regime (A) and (B) time. Large circles represent clade-level regime shifts, and small circles represent 'singleton' regime shifts. The vertical dashed line in A indicates θ for the ancestral dinosaurian body size regime and the grey box indicates a sparsely populated region of trait space at intermediate body masses. Dashed vertical lines in B indicate period and epoch boundaries. Vertical arrows extending up and down from the x axis indicate the initiation of 'trend-like' dynamics on single branches that give rise to exceptionally large-bodied (up) or small-bodied (down) taxa. These regimes are characterized by high-magnitude, unrealised values of theta. Body size frequency distributions for C, Sauropodomorpha; D, Ornithischia; E, non-pygostylian Theropoda; and F, Pygostylia. Grey boxes indicates a sparsely-populated region of trait space at intermediate body masses. Dashed vertical lines indicate the approximate range (1 kg – 10,000 kg) of body masses classes seen among non-pygostylian, non-sauropod dinosaurs (with a few exceptions among small ornithischians, small non-pygostylian theropods, and large ornithischians). The smallest size classes are entirely occupied by

pygostylian birds, and the largest size class is entirely occupied by sauropod dinosaurs. Panels C–F were constructed using the full set of $N = 526$ specimens for which adult body masses were available. "Sauropodom." Is an abbreviation of Sauropodomorpha.

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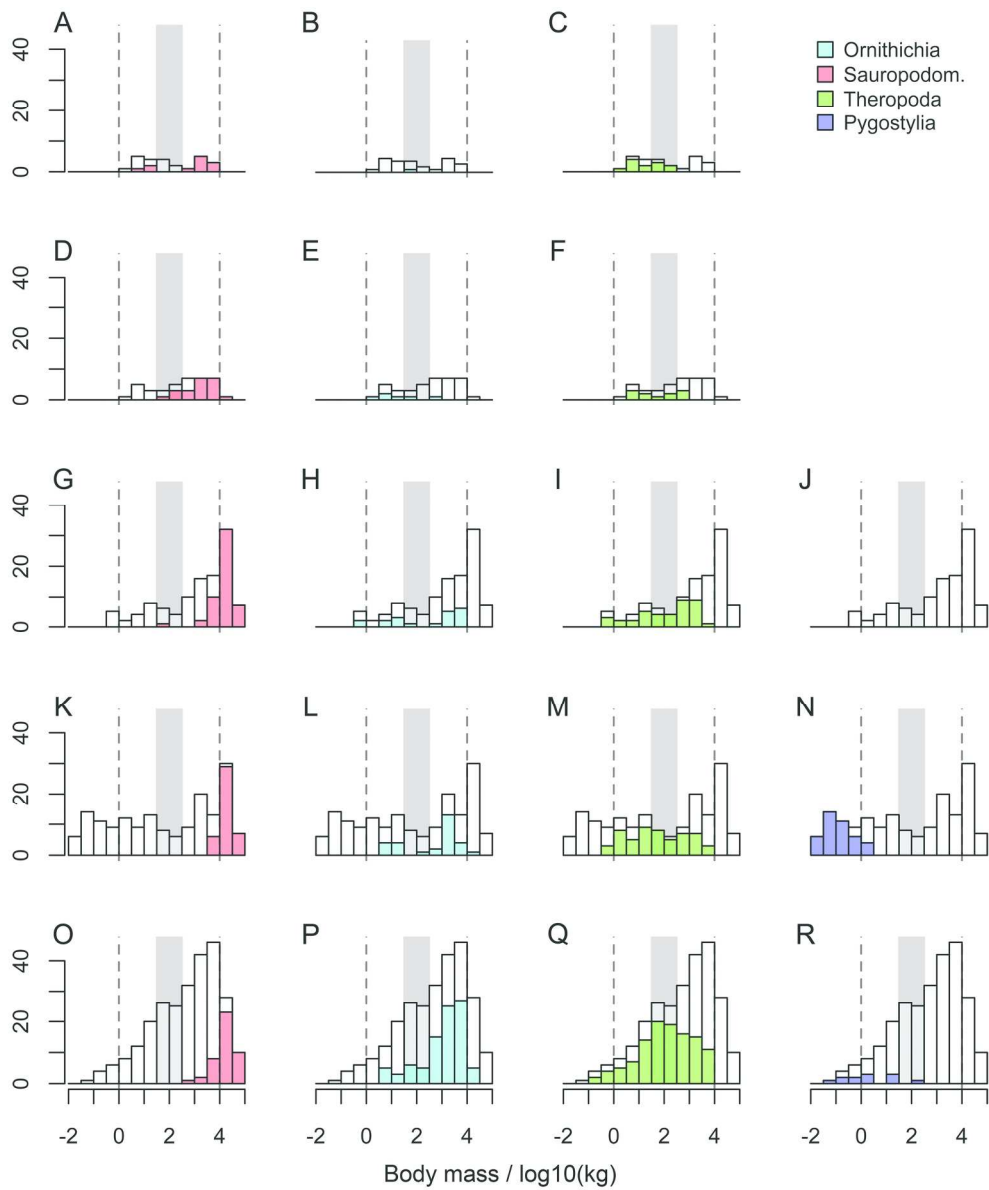


FIG. 15. Histograms showing (log10) body mass distributions for clades of dinosaurs among Mesozoic intervals (coloured bars): A–C, Triassic; D–F, Early Jurassic; G–J, Middle–Late Jurassic; K–N, Early Cretaceous; and O–P, Late Cretaceous. The unfilled bars indicate the body mass distributions for all dinosaurs during each interval. The grey rectangle indicates the underpopulated 'intermediate' range of dinosaur body masses. Dashed lines bracket the approximate minimum (1 kg) and maximum (10,000 kg) body masses for most dinosaurs (especially Theropoda and Ornithischia). These plots were constructed using the full set of $N = 526$ specimens for which adult body masses were available. "Sauropodom." Is an abbreviation of Sauropodomorpha.

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