

1     **Quantifying the impact of fire events on dust emission potential from partially**  
2                                   **vegetated dunes in the southwest Kalahari**

3

4

5     **Abstract**

6     The removal of stabilising vegetation from sand dunes by fire has been widely  
7     linked to increases in aeolian sediment transport and dune movement. However,  
8     substantial gaps exist in our knowledge of whether burned dunes in arid  
9     environments have the potential to emit dust. To explore relationships between fire  
10    and dust emission on partially vegetated sand dunes in the Namibian Kalahari  
11    Desert, 180 measurements of wind erosion threshold and dust flux were carried out  
12    using a portable wind tunnel (Portable In-Situ Wind Erosion Laboratory or PI-  
13    SWERL). Data were analysed to compare erosion thresholds and dust emission flux  
14    on adjacent burned and unburned sites. The data suggest that both burned and  
15    unburned dune crests, flanks, and interdunes have a low potential for dust emission.  
16    Whilst there was no significant difference in dust emission flux between burned and  
17    unburned control surfaces (Kruskal-Wallis,  $p > 0.05$ ), there was evidence of  
18    significantly higher erosion thresholds on burned surfaces (T-test,  $p < 0.01$ ). Where  
19    the surface had been disturbed, resulting in the removal of the typically present  
20    biological soil crusts (biocrust), our data suggest that dust emission fluxes are, on  
21    average, 8-13 times higher those of undisturbed surfaces. The analysis reveals that  
22    even when burned and devoid of vegetation, the Kalahari linear dune system is  
23    sediment-availability limited. This finding indicates the importance of ground  
24    surface characteristics, such as biocrusts, in preventing dust emission from the  
25    Kalahari dune field.

26

27    *Key words:* Dust; Fire; Kalahari; PI-SWERL; Sand dunes

28    Highlights:

- 29     • Experimental results from the first wind tunnel study of burned dunes are  
30       presented
- 31     • Fire does not have a significant intensifying effect on dust emission
- 32     • Threshold shear velocities increased on burned areas compared to unburned  
33       surfaces
- 34     • Surfaces where biological soil crusts had been removed produced large dust  
35       fluxes

## 36 1. Introduction

37 Accurately quantifying mineral dust emissions from diverse landscapes is crucial for  
38 understanding atmospheric dust loading and its climatic and environmental impacts  
39 (Schepanski, 2018; Field et al., 2010; Kok et al., 2023; Gill, 2018). However, current  
40 atmospheric models are not in agreement on the sources and emission flux of  
41 existing atmospheric dust (Kok et al., 2021b; Chappell et al., 2023; Mahowald et al.,  
42 2024; Kim et al., 2024). Models rely on field studies to quantify surface controls on  
43 emissions and satellite data to estimate the amount of dust in the atmosphere (Webb  
44 and Strong, 2011; Haustein et al., 2015; Klose et al., 2019). There are several  
45 uncertainties in the data on dust flux and source locations used in the atmospheric  
46 models including: low-resolution remote sensing imagery often fails to detect  
47 emission sources (Urban et al., 2018; von Holdt et al., 2019); challenges in  
48 characterising surface properties of key emission sources (Webb and Strong, 2011);  
49 and the miscalculation of coarse dust in the atmosphere (Adebisi et al., 2023).  
50 Another key uncertainty in dust and atmospheric modelling is the accurate  
51 quantification of low-concentration or infrequent emission dust sources, which are  
52 difficult to capture due to their irregularity and the limitations of remote sensing  
53 technologies (Urban et al., 2018; Sweeney et al., 2023; Wagenbrenner et al., 2017;  
54 Okin et al., 2011).

55

56 The entrainment of dust from the Earth's surface is a complex interaction of factors  
57 that influence the erosivity of the wind and/or the erodibility of the surface (Webb et  
58 al., 2021; Wiggs et al., 2022). For dust particles to be entrained, wind speed must  
59 exceed a critical shear velocity threshold, though this threshold is influenced by

60 numerous surface characteristics (Kok et al., 2012; Reynolds et al., 2007; Field et al.,  
61 2010; von Holdt et al., 2019). These characteristics are highly variable in time and  
62 space and include sediment size, surface moisture, presence of a surface crust (both  
63 biological and physical), particle abrasion, surface roughness, agglomeration of soil  
64 grains, surface geochemistry, and vegetation cover (Dickey et al., 2023; Sweeney et  
65 al., 2011; Langston and McKenna Neuman, 2005; Engelstaedter et al., 2003; Nield et  
66 al., 2011). Particle size is especially important in dust emission, as it dictates both the  
67 force required for entrainment and the potential distance particles can travel (Kok et  
68 al., 2012). Consequently, low-lying landforms filled with fine sediment, such as  
69 paleolake basins and ephemeral river or lake beds, represent some of the largest dust  
70 sources globally (Prospero et al., 2002; Bullard et al., 2011; Ginoux et al., 2012;  
71 Parajuli and Zender, 2017).

72

73 Sand dunes are not typically considered to be significant sources of atmospheric dust  
74 as they generally contain less than 5 % silt and clay (Prospero et al., 2002). Vegetated  
75 dunes are often excluded from dust emission studies because plants act to stabilise  
76 the surface, preventing sediment movement and abrasive interactions. Plant  
77 architecture increases surface roughness and provides physical sheltering, but also  
78 traps allochthonous dust particles and allows a finer fraction of sediment to build up  
79 (Suter-Burri et al., 2013; Gonzales et al., 2018). However, there is evidence to suggest  
80 that such stabilised dunes can yield larger fluxes of dust in comparison to active  
81 dunes where the surfaces have suffered recent disturbance (such as grazing,  
82 drought, or fire; Bullard et al, 2011; Sweeney et al., 2023).

83

84 It has been hypothesised that if vegetated dunes were to lose their vegetation cover  
85 and become more geomorphologically dynamic (e.g., Wiggs et al., 1995), they could  
86 potentially become large sources of dust into the future (Sweeney et al., 2023;  
87 Bhattachan et al., 2012, 2022; Pye, 1989). One of the key mechanisms by which  
88 vegetated desert dunes may become denuded of vegetation and act as low  
89 concentration emission sources of dust over short timescales is fire (Bullard et al.,  
90 2008; McGowan and Clark, 2008; Yu and Ginoux, 2022).

91

92 Sediment in many vegetated dune systems around the globe is coated in iron oxide-  
93 rich clay which can rapidly be abraded to emit fine grains (Bullard et al., 2004;  
94 Bullard and White, 2005; Swet et al., 2019). For example, abrasion studies using  
95 sediment from the vegetated dunes of the Simpson Desert in Australia found that  
96 the iron-rich clay coating of sand grains can be rapidly abraded into smaller particles  
97 (Baddock et al., 2013; Bullard and White, 2005). The Kalahari Desert in southern  
98 Africa also contains a fine fraction within its sediment that has high concentrations  
99 of soluble iron in its fine material ( $< 45 \mu\text{m}$ ; Bhattachan et al., 2012). The iron oxide is  
100 noteworthy as, if deposited in the marine system, it can have a significant influence  
101 on ocean productivity (Jickells et al., 2005; Bhattachan et al., 2015; Dansie et al., 2022).

102

103 Burned surfaces in other desert landscapes often exhibit higher emissivity rates due  
104 to volatilised fatty acids from vegetation, which condense onto nearby sediment and  
105 form a hydrophobic coating on grains. The coating reduces interparticle cohesion by  
106 altering the contact angle of the grains (Ravi et al., 2006; Sankey et al., 2012). In the  
107 arid shrubland of the USA, Sankey et al. (2012) found that burned surfaces emitted

108 three times more sediment than unburned areas. Yet, there have been no field  
109 studies investigating the emissivity of burned vegetated dune fields. To understand  
110 the complexities of the impact of fire on dune dust emissions , it is crucial to examine  
111 the fire–aeolian erosion feedback mechanisms on dune environments.

112

113 There is some evidence for post-fire dust emissions from sand dune systems. For  
114 example, Yu and Ginoux (2022) used Moderate Resolution Imaging  
115 Spectroradiometer (MODIS) data (deep blue aerosol product and burned area  
116 product) and found that at the global scale wildfires lead to enhanced dust  
117 emissions. This included dust emissions from vegetated dune systems in Australia,  
118 the Sahel, and Argentina, but there are notable gaps in emission from the southwest  
119 Kalahari linear dune system (Yu and Ginoux, 2022). Further, Bullard et al. (2008) and  
120 McGowan and Clarke (2008) used remote sensing methods to detect dust plumes  
121 originating from fire scars in the Simpson Desert in Australia. But emission potential  
122 from the southwest Kalahari dune field has been poorly constrained. The Kalahari is  
123 ecologically and geomorphologically similar to the Simpson Desert in Australia  
124 (Buckley, 1981) and burns frequently (Andela et al., 2019). The Kalahari is often  
125 suggested as being a source of dust if it were to become spontaneously de-vegetated  
126 (Bhattachan et al., 2012). Yet, no research has investigated post-fire dust emissions in  
127 the region.

128

129 This research aims to fill the knowledge gap of whether the partially vegetated  
130 dunes of the southwest Kalahari have the potential to emit dust into the atmosphere  
131 because of fire destroying their protective vegetation cover and altering surface

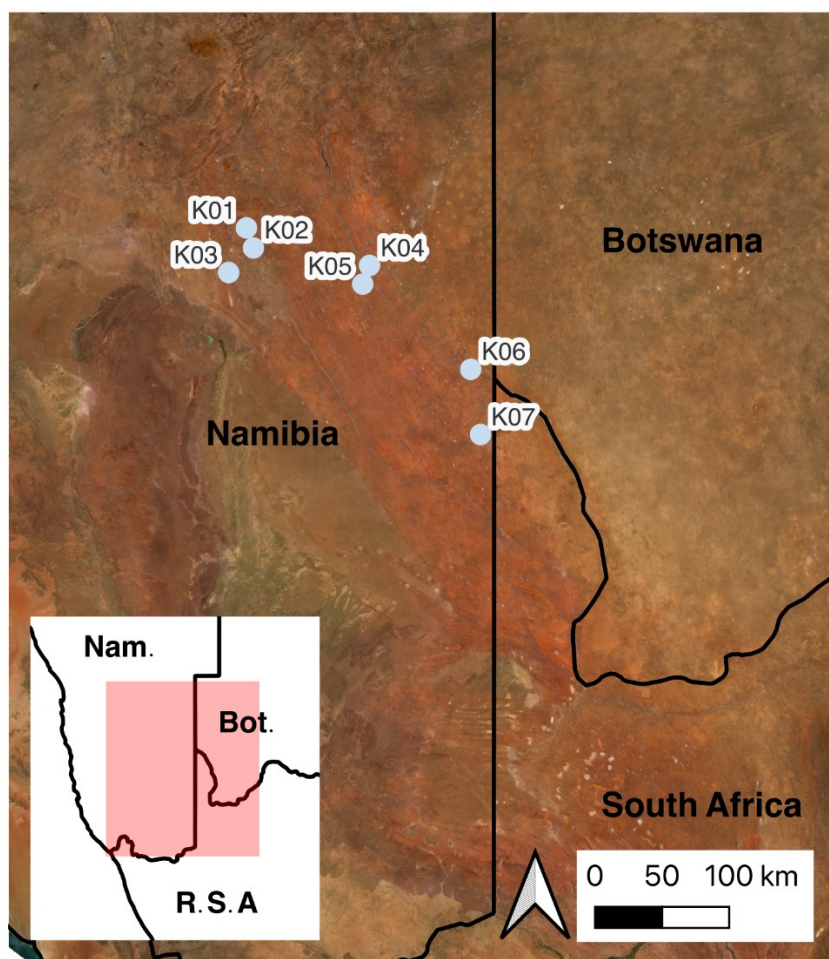
132 erodibility and emissivity. Dune surface sedimentology is examined to determine  
133 changes between burned and co-located unburned plots to establish if there is fine  
134 material typical of dust in the sediment texture. Next, the emission potential between  
135 burned and unburned plots is quantified using a portable wind tunnel. The research  
136 can be broken down into two aims which are to:

- 137 1. Establish whether the southwest Kalahari dunes have a sedimentary potential to  
138 emit dust.
- 139 2. Quantify whether burning of dune vegetation has a significant effect on dust  
140 emission.

## 141 2. Methods

### 142 2.1. Study area

143 The research was conducted on the vegetated linear dunes (Figure 1) in the arid  
144 southwest Kalahari in Namibia in September and October 2022. The dunes consist of  
145 unconsolidated aeolian sands of mainly quartz and feldspar (Garzanti et al., 2022).  
146 There is no statistical difference in grain sizes on the east or west flank of the dunes  
147 (Lancaster, 1986): therefore, sediments in this study were collected from either face  
148 of the dune. The dune crests are spaced 0.2 to 2 km apart with heights that range  
149 from 5 to 25 m (Thomas and Leason, 2005). In this study, a crest is the highest point  
150 of the dune, a flank is 3 - 5 m down the side from the crest of the dune, and an  
151 interdune is classed as the area where the slope of the dune has become undetectable  
152 by eye.



153  
 154 **Figure 1.** Locations of sites (K01-K07; see Table 1). Background imagery © 2023 Planet Labs  
 155 PBC (Planet, 2023).

156  
 157 The linear dune system lies within the southern Africa summer rainfall zone and has  
 158 a mean annual precipitation of 150 – 300 mm but, characteristic of dryland regions,  
 159 interannual variability is high at 50 % (Thomas and Leason, 2005; Bhattachan et al.,  
 160 2014). The rain generally falls between November and April; accordingly, September  
 161 and October represent the end of the dry season. Additionally, the end of the dry  
 162 season also coincides with the windiest months (Nchaba et al., 2017; Bullard et al.,  
 163 1996). In the 12 months preceding September 2022 when measurements were made,  
 164 the sites received an average of 201 mm precipitation.

165

166 The fire season peaks in September but can continue into the early wet season. Fires  
167 in the region are limited by vegetation, which provides both fuel and affects the  
168 connectivity of the burnable surface. Vegetation mass and extent is strongly linked to  
169 precipitation and to land use (Andela et al., 2019). Consequently, like precipitation,  
170 fires have a high interannual variability, and occur regularly after wetter La Niña  
171 periods which provides the moisture for biomass build up (Chen et al., 2017).

172

173 The typical vegetation in the study area is Kalahari Xeric Savanna (van Rooyen and  
174 van Rooyen, 1998), with varying surface coverage depending on land use history.  
175 The vegetation composition on the dunes is a mix of annual grasses (mainly  
176 *Schmidtia kalahariensis*), perennial grasses (e.g., *Stipagrotis amabilis*), prostrate creepers  
177 (e.g., *Aptosimum elongatum*), patches of shrubs (e.g., *Lycium hirsutum*) and  
178 sporadically populated trees (e.g., *Acacia haematoxylon*). All measurements were  
179 conducted on farmland which is either sheep, goat, or cattle grazed. Most of the  
180 unburned farmland had a high coverage of annual grasses.

181

182 Biological soil crusts (biocrusts) were observed at each site and in burned interdune  
183 plots. This observation aligns with previous work in the region noting that biocrusts  
184 cover between 11 to 95 % of the ground surface, with its high variability being  
185 attributed to disturbance history (Thomas and Dougill, 2007). Additional studies  
186 have observed that biocrust cover is ubiquitous across Kalahari surfaces (Elliott et  
187 al., 2014, 2024; Lan et al., 2021; Thomas and Dougill, 2006; Mager, 2010; Dougill and  
188 Thomas, 2004). However, the absence of surface discolouration reduced the field  
189 visibility of the biocrust, making it not possible to reliably quantify its presence. This

190 was due to the lack of surface discolouration in the early stages of biocrust  
191 development (Bullard et al., 2022; Dougill and Thomas, 2004). No physical crusts  
192 were observed at the experimental sites.

## 193 2.2. Experimental set-up

194 Seven burned sites, with different fire ages (Time Since Fire: TSF), were identified  
195 using remote sensing, followed by conversations with local landowners (Table 1 and  
196 Figure 1). Each burned site was co-located with a nearby unburned control site.  
197 Where possible the site pairs (burned and control) were close to each other on the  
198 same dune (typically within 1 km of each other), or on the adjacent interdune. The  
199 sites were then further divided into plots to investigate the effect of the dune  
200 morphological unit on emission potential with particular focus on the dune crest and  
201 interdunes (Figure 2). The different morphological positions of the plots on the dune  
202 surfaces are summarised in Table 1. The accessibility of the area burned limited the  
203 ability to record at all the dune topographies. From herein, 'site' refers to the broader  
204 burned area and the co-located control area and 'plot' refers to the experimental area  
205 within a site in a distinct dune morphological unit.

206

207 All the measurements took place on grazed-disturbed land. Trampling by the  
208 grazing animals is known to break-up the surface and disrupt biocrusts (Thomas  
209 and Dougill, 2007). To evaluate the significance of the biocrust on erodibility,  
210 surfaces at each plot were disturbed by removing the top two centimetres using a  
211 spade, removing the biocrust. Herein, the plots where the surface was removed are  
212 termed 'scraped plots'. Once the surface sediments, including any biocrust, were  
213 removed, additional PI-SWERL measurements were made (Figure 2). The surface

214 scraping was not intended to replicate any particular real-world disturbance but  
 215 instead was aimed at identifying whether biocrust presence impacted on sediment  
 216 availability. Subsequently, the maximum number of experimental plots at each site  
 217 was nine - the burned crest, flank, and interdune, with co-located control plots and  
 218 additional scraped interdune, flank, and crest plots.

219

220 **Table 1.** Location, date of fire, and plot position for each burned site

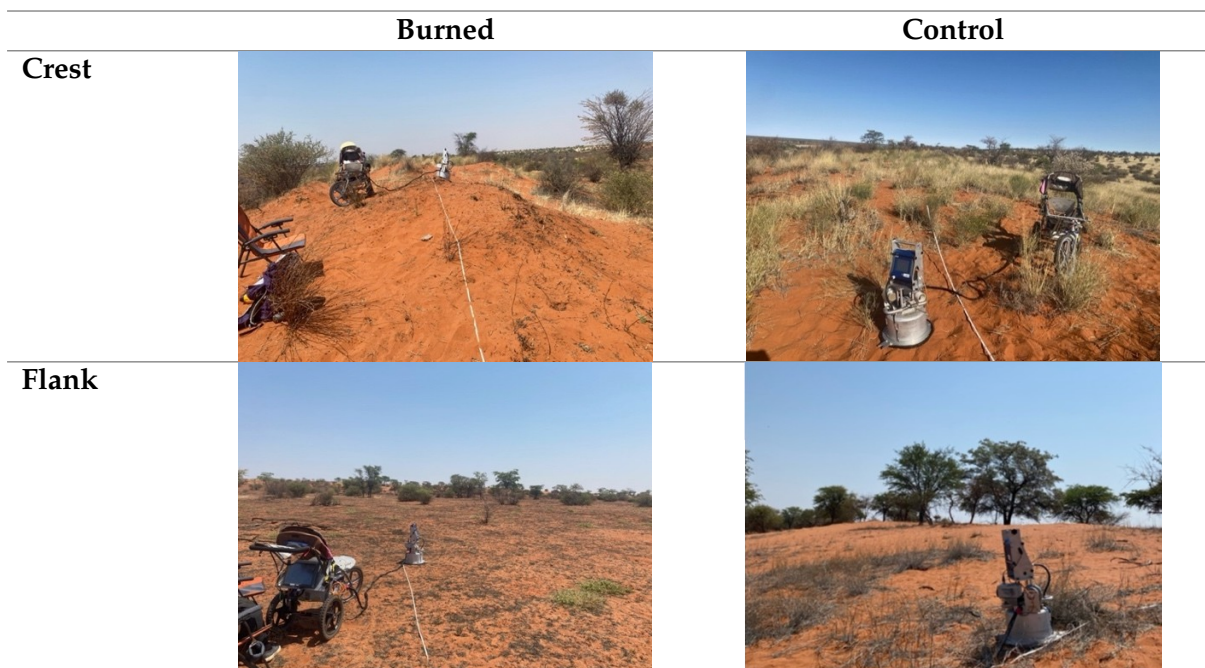
Site	Latitude	Longitude	Date of fire	TSF (months)	Dune morphology measured
K01	-23.722675	18.109997	01/09/2022	0 and 1	C*, F, ID
K02	-23.8606	18.1507	26/07/2022	1	C, ID
K03	-24.04225	17.98124	18/06/2022	4	C, ID
K04	-23.98248	19.01717	29/11/2021	10	C, F, ID
K05	-24.1159	18.99656	16/11/2021	10	ID
K06	-24.68887	19.82227	27/10/2020	25	C, ID
K07	-25.185	19.8947	28/02/2012	127	ID

*C – Crest, F – Flank, ID – Interdune*

*\*Control only*

221

222





223 **Figure 2.** Examples of PI-SWERL measurement locations on different topographies.

224

225 **2.3. Grain size distribution**

226 One surface sediment sample was taken per measurement plot (Table 1) for  
 227 laboratory analysis. Grain size distribution measurements for samples were  
 228 conducted using the wet (no pre-treatment) dispersal method with a Malvern  
 229 Mastersizer Hydro 2000MU laser diffraction particle size analyser. Before analysis,  
 230 samples were oven dried overnight and sieved to less than 2 mm and exposed to 10 s  
 231 of ultrasonic dissolution, with the ultrasound setting at 10  $\mu\text{m}$ . Samples were then  
 232 classified by type where clay is  $<2 \mu\text{m}$ , silt is  $2 \mu\text{m}$  to  $50 \mu\text{m}$ , and sand is  $50 \mu\text{m}$  to  
 233  $2000 \mu\text{m}$ . In addition, to classify grains that have the sedimentary potential to be  
 234 emitted as dust at the surface, sediment was further classed into particles which are  
 235 inhalable under  $10 \mu\text{m}$  and super-coarse dust particles under  $62.5 \mu\text{m}$ . These sizes  
 236 were chosen as  $< 10 \mu\text{m}$  aligns with the measurements made with the PI-SWERL and  
 237  $< 62.5 \mu\text{m}$  is the size bracket identified by Adebisi et al. (2023) as currently  
 238 underestimated in the atmosphere. Kruskal-Wallis chi-squared tests were used to

239 assess any statistical differences between the different morphological units where  
240 measurements were taken on the dune (i.e., crest, flank, and interdune) and grain  
241 size distribution.

242

## 243 2.4. PI-SWERL measurements

244 Emission flux was measured using the Portable In-Situ Wind Erosion Laboratory  
245 (PI-SWERL ®; DustQuant LLC, Las Vegas, NV, USA), now a widely used technique  
246 (Bacon et al., 2011; Sweeney and Mason, 2013; Vos et al., 2021; King et al., 2011;  
247 Walker et al., 2023; Kolesar et al., 2022; von Holdt et al., 2019; Wang et al., 2025). The  
248 instrument (as seen in Figure 2) consists of a flat, annular ring inside a closed  
249 chamber rotating about 6.5 cm above the ground which exerts shear stress on the  
250 surface and entrains particles. The emitted particles are measured by a DustTrak ® II  
251 model 8530 nephelometer to measure the PM<sub>10</sub> concentration of dust in the chamber  
252 per second in mg m<sup>-3</sup>. Fresh air is pumped into the chamber at a rate of  
253 approximately 0.1 m<sup>3</sup> s<sup>-1</sup>. The PI-SWERL was preprogrammed to run at a variety of  
254 revolutions per minute (RPM) which translate into specific friction velocities ( $u_*$  m s<sup>-1</sup>).  
255 The relationship between RPM and  $u_*$  is determined by surface roughness and  
256 defined as:

257

$$258 \quad u_{i,eff}(RPM) = C_1 \cdot \alpha^4 \cdot RPM^{C_2/\alpha}$$

259

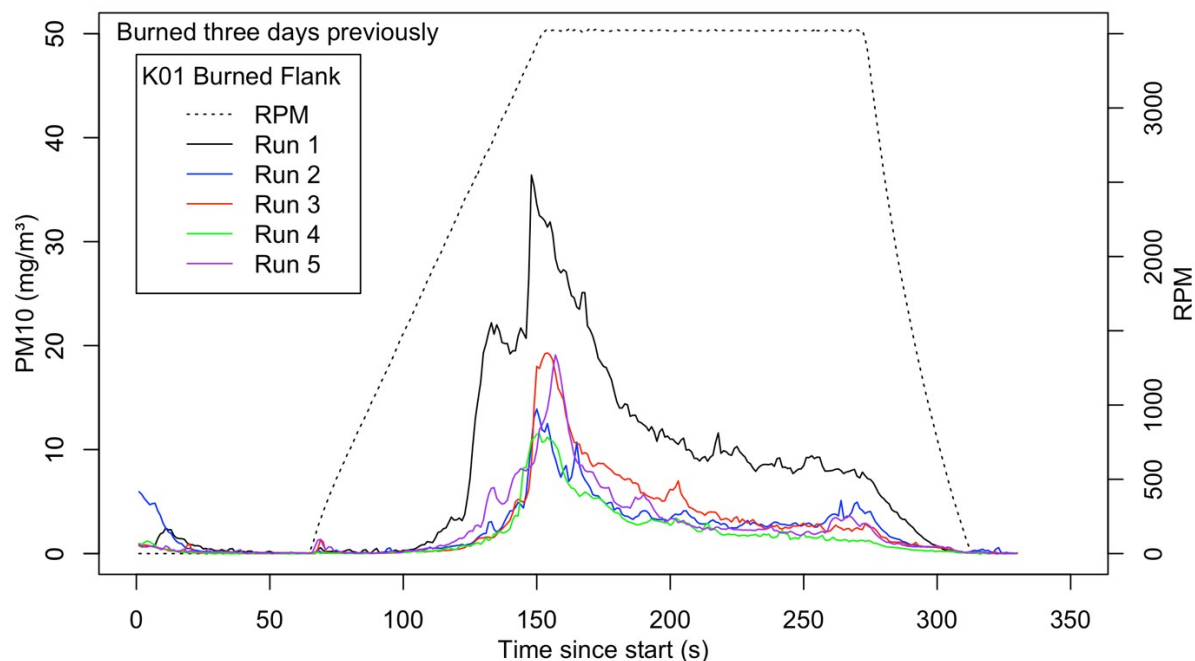
**Equation 1**

260 where,  $C_1$  and  $C_2$  are constants 0.000683 and 0.832 respectively (Etyemezian et al.,  
261 2014).  $\alpha$  is a value calculated based on surface roughness using the look-up table in  
262 Hartshorn et al. (2023). It was deemed that all surfaces consisted of loose

263 unconsolidated sand which was less than 2.5 % gravel cover, therefore, the  $\alpha$  value  
264 of 0.98 was chosen for all sites, equating to a peak  $u_*$  value of  $0.64 \text{ m s}^{-1}$  at 3500 RPM.  
265 This  $u_*$  value was chosen to account for high wind velocity events previously  
266 identified in the dune field by Wiggs et al. (1994).

267

268 The rotation of the flat ring was increased at a constant rate (ramp test). The ramp  
269 test was chosen as it enables the determination of (a) if the plot is capable of emitting  
270 dust, (b) enables an easy identification of the erosion threshold, and (c) any changes  
271 in dust emission flux whilst at a constant rate. At the start of each test, a 60 s clean air  
272 run was undertaken to flush out any remaining particles in the chamber (Figure 3).  
273 The RPM was then increased from 0 to 3500 RPM over 90 s, held at 3500 RPM for 120  
274 seconds, and then reduced to 0 RPM together with a clean air flush for the final 60  
275 seconds of the test (Figure 3). To reduce contamination between each run, the inside  
276 of the chamber was brushed down to remove any remaining visible sediment and  
277 after concluding measurements at each plot, the PI-SWERL was run at max RPM for  
278 three minutes.



279

280 **Figure 3.** Example runs of a PI-SWERL at each measurement plot, displaying the  
 281 programming used for each measurement.

282

283 Many previous experiments with a PI-SWERL have utilised a ‘step’ test, where the  
 284 flux of dust is measured at different  $u_*$  to assess differing wind velocities (e.g., Bacon  
 285 et al., 2011; Cui et al., 2019; Sankey et al., 2012). As the aim of this study is to assess  
 286 the potential for the surface to be emissive, and if there is any change in thresholds  
 287 with different disturbances only one maximum velocity is used. This maximum RPM  
 288 was maintained for 90 s (Figure 3), which allows us to calculate flux and identify  
 289 controls on emission.

290

291 Threshold of emission ( $u_{*T}$ ) was determined using the same process as Cui et al.  
 292 (2019) and van Leeuwen et al. (2021) where the first recording when the  $PM_{10}$   
 293 concentration began to consistently increase for ten or more seconds is determined  
 294 as the threshold. The threshold RPM was then converted into  $u_*$  using the Equation  
 295 1. The  $PM_{10}$  flux was calculated using the equation:

296

$$E_i = \frac{\sum_{begin,i}^{end,i} C \cdot F}{(t_{end,i} - t_{begin,i}) \cdot A_{eff}}$$

297

298

Equation 2

299

300 where  $C$  is the measured  $PM_{10}$  concentration ( $mg\ m^{-3}$ ),  $F$  is the airflow rate ( $L\ s^{-1}$ ),  $t$  is  
 301 the duration (s) of step  $i$ , and  $A_{eff}$  is the effective area ( $m^2$ ) of the PI-SWERL ( $0.035\ m^2$ ;  
 302 Sweeney et al., 2011; Etyemezian et al., 2014). The flux was used to compare emission  
 303 fluxes to other measurements and PI-SWERL results.

304

305 Within each plot, a ten metre transect was set out and a PI-SWERL measurement  
 306 was taken every two metres along the transect. At each plot the biocrust was  
 307 removed at a plot away from the transect for further PI-SWERL measurements to be  
 308 conducted. One to six measurements were performed at each plot to get a  
 309 representative characterisation of the surface (Table 2). A smaller number of  
 310 measurements were performed at a plot if the peak concentrations of dust flux in the  
 311 initial measurements were consistent (within a tolerance of  $5\ mg\ m^{-3}$ ). If there was  
 312 grass and creeper coverage, to ensure that the vegetation did not interfere with the  
 313 rotating blade, vegetation was trimmed using shears to surface level (Figure 2),  
 314 taking care not to disturb the surface. In total, 180 Pi-SWERL measurements were  
 315 conducted at seven sites during September and October 2022 (Table 2).

316

317 **Table 2.** Measurements made at each plot. C denotes a control plot and B denotes a plot  
 318 which has burned. Scraped surfaces are where the top 2 cm of the surface was removed. K01  
 319 was visited twice, three days after burning and then 1 month after burning.

	K01 (3 Days)	K01 (1 month)	K0 2	K0 3	K0 4	K0 5	K0 6	K0 7
Interdune (C)	5		5	5	5	5	5	4
Interdune (B)	5	5	5	5	5	3	5	5
Flank (C)	5				5			
Flank (B)	5	5			5			
Crest (C)	3		5	5	5		6	
Crest (B)			5	5	5		4	
Scrape -Interdune	3	2	3	4	6		6	1
Scrape -Flank	6							
Scrape -Crest							3	

320

321 The PM<sub>10</sub> threshold and flux data were analysed to determine statistical differences  
 322 based on dune morphological unit (dune morphological unit, i.e., crest, flank, and  
 323 interdune) and status (burned or control). The choice of statistical test depended on  
 324 the data distribution and number of groups compared. Kruskal-Wallis chi-squared  
 325 tests were used to assess differences in emission flux across dune morphological  
 326 units and status (i.e., control, burned, or scraped), while Welch Two Sample t-tests  
 327 were applied to threshold data for the same factors. Additionally, for variations in  
 328 measurements taken at different dune morphological units, Kruskal-Wallis chi-  
 329 squared tests were used for emission flux, and Analysis of Variance (ANOVA) was  
 330 applied to threshold data.

331

### 332 3. Results

#### 333 3.1. Sediment Size distribution

334 88.4 % of the samples have a soil texture of sand, the remaining 11.6 % is classified as  
 335 sandy loam. Only 5.8 % contained any clay sized (< 2 µm) material and 92.8 %

336 containing silt sized (2 - 50  $\mu\text{m}$ ) material. Differences can be observed when  
337 assessing the topographical location of the sediments. 30 % of crest plots contained  
338 sediment smaller than 10  $\mu\text{m}$ , whilst all the interdune plots have a higher percentage  
339 of sediment under 10  $\mu\text{m}$  (with a maximum content of 8.9 % at K01 burned  
340 interdune plot).

341

342 Sediment finer than 62.5  $\mu\text{m}$  is typically considered to be able to be suspended in the  
343 atmosphere (Adebiyi et al., 2023). The proportion of sediment under 62.5  $\mu\text{m}$   
344 differed significantly with the different dune morphological unit (Kruskal-Wallis  $p <$   
345 0.0001). Interdunes had the largest mean proportion of sediment  $< 62.5 \mu\text{m}$  at 9.8 %,   
346 whilst the flanks had 4.3 % and the crests had 1.7 % on average (Table 3).

347

348

349 **Table 3.** Grain size data for samples from the study plots.

		% of sample																			
		Clay				Silt				Sand				< 10 µm				< 62.5 µm			
		$\bar{x}$	Max.	Min.	SD	$\bar{x}$	Max.	Min.	SD	$\bar{x}$	Max.	Min.	SD	$\bar{x}$	Max.	Min.	SD	$\bar{x}$	Max.	Min.	SD
	C	0.0	0.0	0.0	0.0	2.6	7.3	1.3	1.9	97.4	98.7	92.7	1.9	0.6	2.0	0.0	0.6	3.2	9.5	1.3	2.6
	F	0.0	0.0	0.0	0.0	2.6	3.2	1.4	1.5	97.4	96.8	95.2	1.5	1.4	1.7	0.3	0.6	4.2	6.9	2.0	2.0
<b>Burned</b>	ID	0.0	0.1	0.0	0.0	9.2	25.3	1.6	6.9	90.7	98.4	74.7	6.9	2.9	8.9	0.2	2.3	11.4	30.0	2.1	8.2
	C	0.0	0.0	0.0	0.0	0.4	0.8	0.0	0.3	99.7	100.0	99.2	0.3	0.0	0.0	0.0	0.0	0.6	1.4	0.0	0.4
	F	0.0	0.0	0.0	0.0	3.6	7.0	0.0	2.8	96.5	100.0	93.0	2.8	1.2	2.8	0.0	1.1	4.4	8.2	0.0	3.4
<b>Control</b>	ID	0.0	0.0	0.0	0.0	6.0	11.9	1.0	3.0	94.0	99.0	88.1	3.0	1.8	3.4	0.1	1.0	7.6	14.7	1.4	3.7

*C- Crest, F- Flank, and ID - Interdune*

350

351 **3.2. The impact of dune topography on emission threshold and**  
 352 **flux**

353 There was great variance in both  $u_{*T}$  and emission flux with the location of the plots  
 354 on the dune (Table 4). The dune morphological unit is highly significant for both the  
 355  $u_{*T}$  (ANOVA,  $p < 0.01$ ) and the flux of dust (Kruskal-Wallis test,  $p < 0.01$ ). Of the  
 356 undisturbed surfaces, interdunes had the largest peaks in emission, with the  
 357 unburned interdune at K07 recording a  $PM_{10}$  peak concentration of  $65.5 \text{ mg m}^{-3}$  (with  
 358 a flux of  $1.892 \text{ mg m}^{-2} \text{ s}^{-1}$ ). Yet, on average the flanks of the dunes have the largest  
 359 flux of dust (Figure 4b). However, scraped surfaces consistently saw the highest  
 360 fluxes of dust (Figure 4), recording a peak concentration of  $173 \text{ mg m}^{-3}$  (with a flux of  
 361  $6.305 \text{ mg m}^{-2} \text{ s}^{-1}$ ) observed at K02 scraped interdune.

362

363

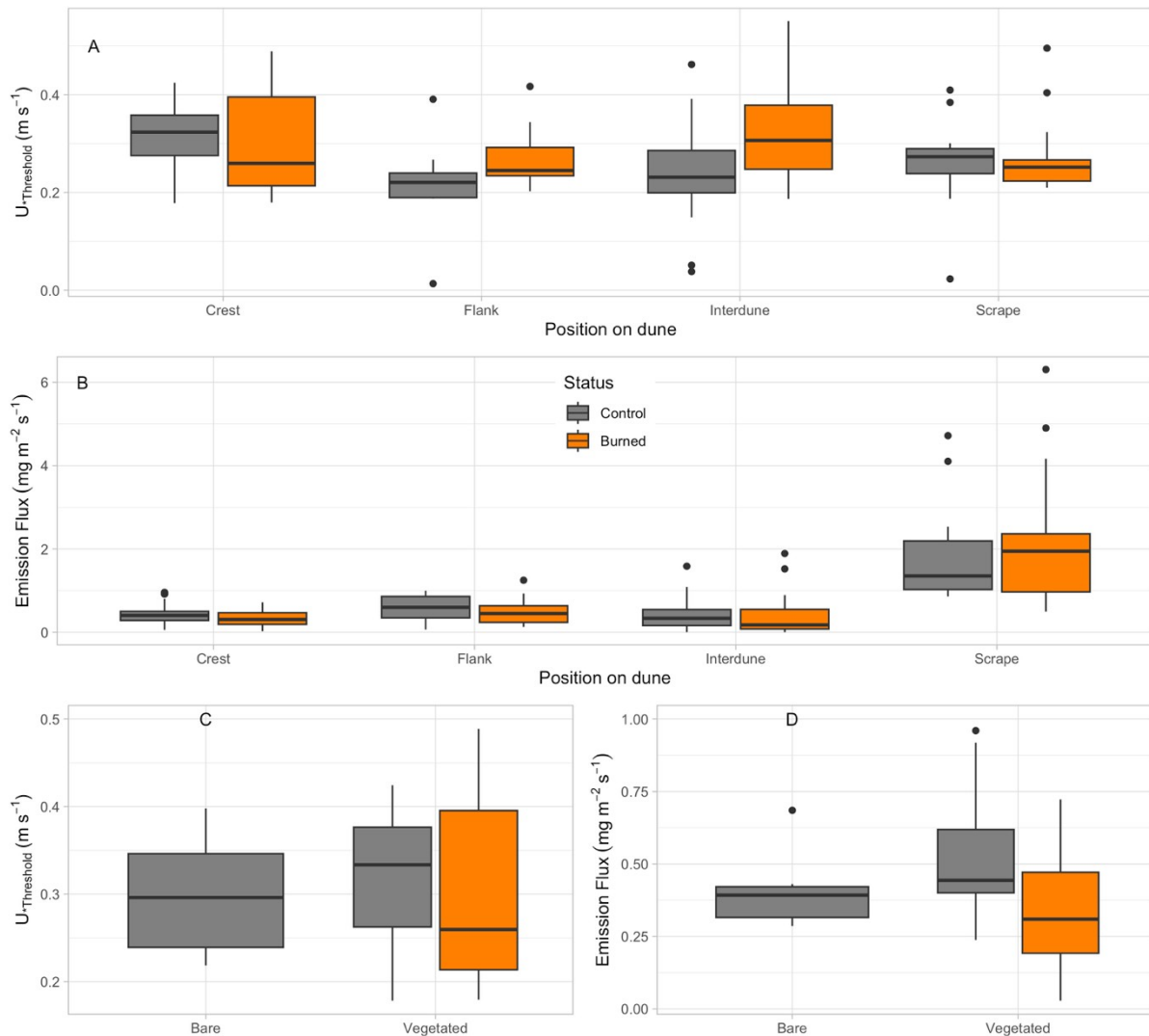
364 **Table 4.** Descriptive statistics of measured threshold velocities ( $\text{m s}^{-1}$ ) and emission flux ( $\text{mg}$   
 365  $\text{m}^{-2} \text{ s}^{-1}$ ) for each dune morphological unit and disturbance type.

Dune morph. unit	Disturbance	n	Threshold ( $\text{m s}^{-1}$ )				Flux at $0.64 \text{ m s}^{-1}$ ( $\text{mg m}^{-2} \text{ s}^{-1}$ )			
			$\bar{x}^*$	SD*	Min	Max	$\bar{x}^*$	SD*	Min	Max
ID	B	38	0.308	1.305	0.187	0.551	0.190	3.363	0.005	1.892
	C	34	0.221	1.610	0.038	0.462	0.281	3.061	0.005	1.587
	S	25	0.167	2.670	0.187	0.409	1.105	2.033	0.862	6.305
Flank	B	15	0.264	1.228	0.202	0.417	0.403	1.942	0.129	1.251
	C	10	0.175	2.525	0.013	0.391	0.447	2.397	0.069	0.999
	S	6	0.266	1.218	0.202	0.310	1.949	1.815	0.500	2.538
Crest	B	19	0.303	1.298	0.179	0.489	0.467	1.430	0.029	0.723
	C	24	0.282	1.414	0.178	0.424	0.253	2.428	0.238	0.960
	S	3	0.288	1.601	0.210	0.495	0.942	1.026	0.923	0.970

\*Geometric Mean and Standard Deviation

B = Burned, C = Control, and S = Scraped (biocrust removed)

366



367  
 368 **Figure 4.** Panels A and B:  $u_{*T}$  and emission flux for the crests, flanks, interdunes, and scraped  
 369 (biocrust removed) surfaces across all the measurement plots. Panels C and D:  $u_{*T}$  and  
 370 emission flux for the crest plots only subdivided into bare and vegetated crest plots. Burned  
 371 plots are in orange and control plots are displayed in grey.

372

### 373 3.3. Inter-plot comparisons

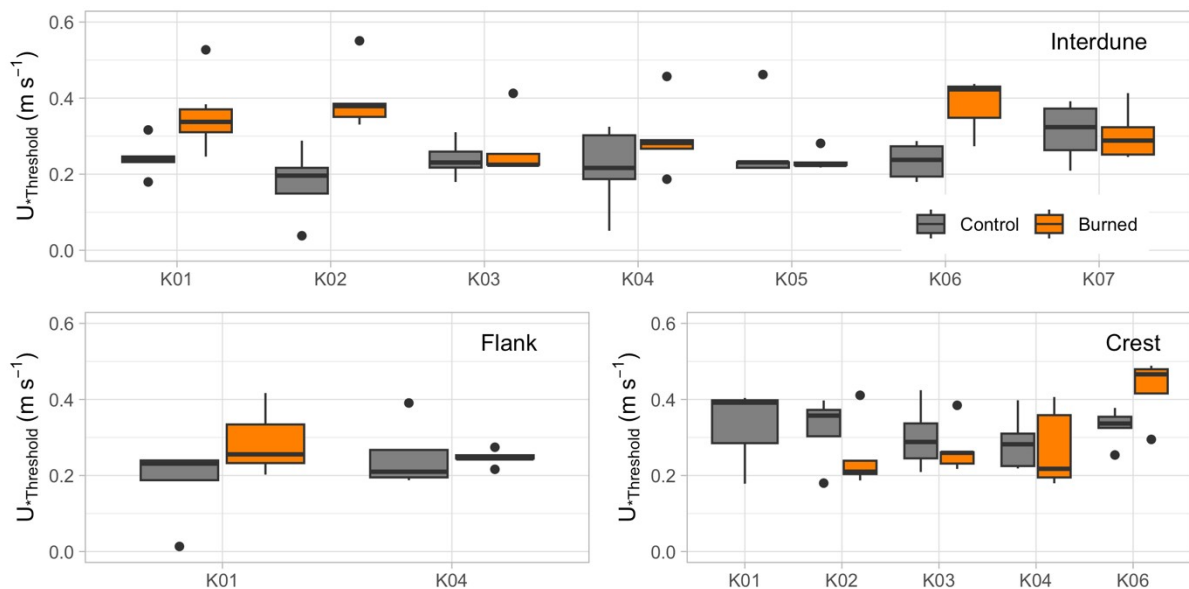
374 Each plot investigated has different surface properties that impact the ability to emit  
 375 dust, and there is a significant difference in the emission flux between each control  
 376 interdune plot (Kruskal-Wallis test,  $p < 0.01$ ). Emission fluxes from each  
 377 experimental plot are reported in Table 5. There was no significant difference in the  
 378 flux between the control crests and flanks between the sites. All control plots showed

379 similar  $u_{sT}$  values (Figure 5), which are not significantly different from one another  
 380 (ANOVA,  $p > 0.05$ ) in contrast to the emission flux.

381

382 The significant difference between each control plot made isolating the evolution of  
 383 dust emission with time since fire difficult. Further, there is no distinct pattern in  
 384 whether the burned or control plots had high emission fluxes or emission thresholds  
 385 with time since fire (Figure 5). Subsequently, there are no further investigations into  
 386 dust emission potential over time as burned areas heal.

387



388

389 **Figure 5.**  $u_{sT}$  at each site on the interdune, flank, and crest plots.

390

391 **Table 5.** Emission flux for each experimental plot measured.

	Emission flux ( $\text{mg m}^{-2} \text{s}^{-1}$ )							
	(geometric mean $\pm$ geometric standard deviation)							
	K01 (3 days)	K01 (1 month)	K02	K03	K04	K05	K06	K07
<b>Interdune (C)</b>	$0.936 \pm 1.442$		$0.065 \pm 6.192$	$0.406 \pm 1.659$	$0.237 \pm 1.789$	$0.176 \pm 2.324$	$0.303 \pm 1.529$	$0.487 \pm 1.514$

<b>Interdune (B)</b>	0.108 ± 1.570	0.035 ± 3.249	0.221 ± 3.108	0.653 ± 1.235	0.197 ± 1.711	0.273 ± 1.897	0.065 ± 1.509	0.575 ± 3.336
<b>Flank (C)</b>	0.713 ± 1.340				0.280 ± 2.840			
<b>Flank (B)</b>	0.268 ± 1.691	0.322 ± 1.792			0.756 ± 1.489			
<b>Crest (C)</b>	0.893 ± 1.095		0.454 ± 1.318	0.449 ± 1.085	0.353 ± 1.210		0.451 ± 1.486	
<b>Crest (B)</b>			0.367 ± 2.290	0.384 ± 1.246	0.297 ± 1.495		0.078 ± 2.862	
<b>Scrape - Interdune</b>	2.245 ± 1.167	4.400 ± 1.104	5.007 ± 1.247	1.747 ± 2.021	1.732 ± 1.306		1.052 ± 1.119	1.872
<b>Scrape - Flank</b>	1.105 ± 2.033							
<b>Scrape - Crest</b>							0.942 ± 1.026	

392

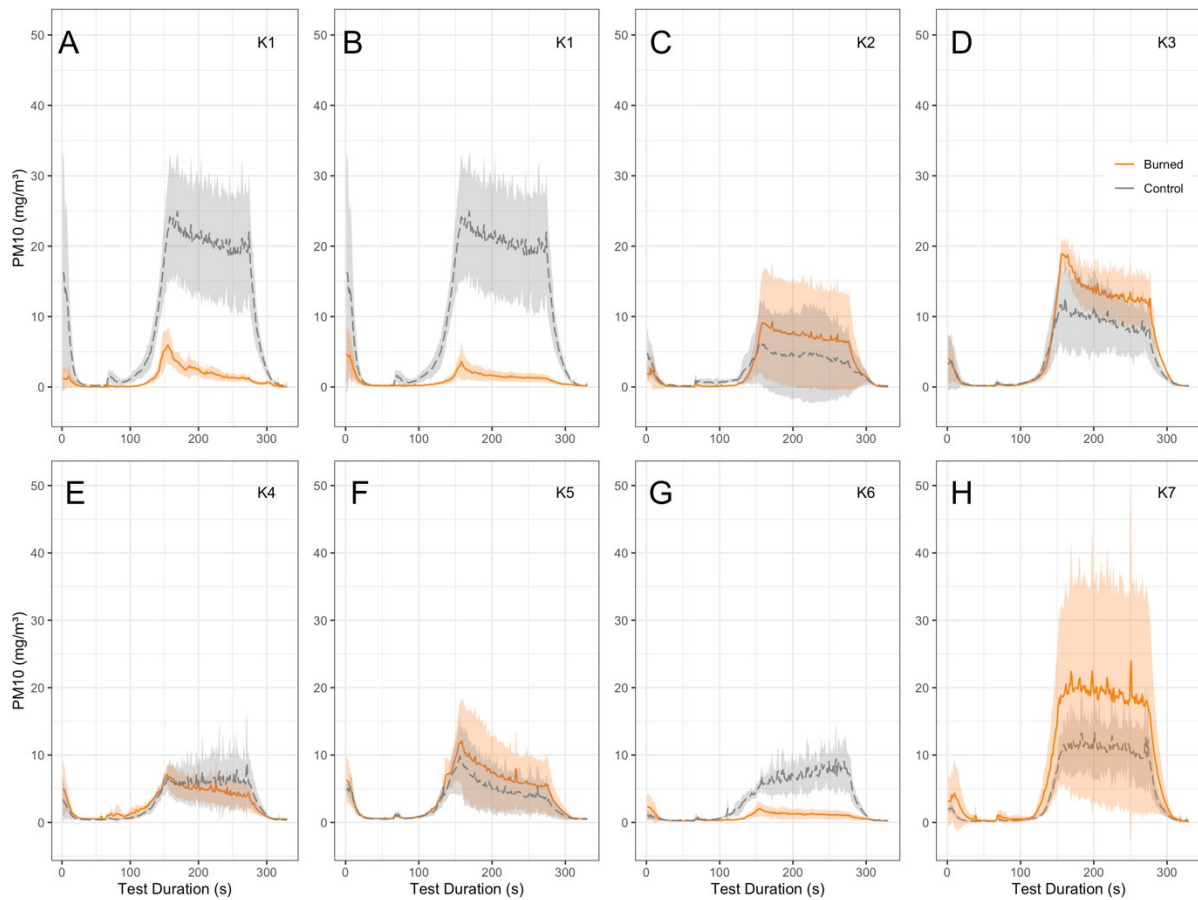
### 393 3.4. The effect of burning on emission threshold and flux

394 The magnitude of difference in dust flux between the burned and control plot varies  
 395 with site (Figure 5). In the interdunes, four plots had higher concentrations of dust in  
 396 the control plot and four had higher concentrations in the burned area (Figure 6).  
 397 The difference between the burned and control plots are minimal as four plots have  
 398 overlapping standard deviations (Figure 6 – panel C, D, E, H).

399

400 The burned and control surfaces had significantly different  $u_{*T}$  (T-test,  $p < 0.01$ ) on all  
 401 topographies, with burned sites having a higher threshold for erosion. In contrast,  
 402 burning does not have a significant impact on the emission flux on all topographies  
 403 (Kruskal-Wallis test,  $p > 0.05$ ). The largest average difference in the flux between the  
 404 burned and control plots at the same site was found at the interdune K01 where the

405 difference between the geometric means was  $0.901 \text{ mg m}^{-2} \text{ s}^{-1}$  with the control plot  
 406 having a larger dust flux.



407  
 408 **Figure 6.**  $\text{PM}_{10}$  emissions for the burned and control plots at each interdune plot. Mean  $\text{PM}_{10}$   
 409 ( $\text{mg m}^{-3}$ ) concentration is the solid line with one standard deviation shaded. The burned  
 410 plots are in orange and the control plots are in grey. In A the area was burned three days  
 411 before measurement and in panel B the plot was remeasured one month after burning, the  
 412 same control plot data was used for both measurements.

413  
 414 **3.5. The effect of biocrust removal on emission threshold and**  
 415 **flux**  
 416

417 Scraped surfaces where the biocrust was removed constantly reported the highest  
418 fluxes of dust (Table 5 and Figure 4b) but similar threshold for emissions (Table 4  
419 and Figure 4a). The ratio of geometric means for burned and scraped interdune plots  
420 was  $0.08 \pm 3.70$  and for the unburned and scraped interdune plots was  $0.13 \pm 1.74$   
421 (Table S1). These results indicate that the scraped (biocrust-removed) surfaces  
422 produced approximately 8 – 13 times higher dust fluxes than the burned and control  
423 surfaces. The largest difference between the control and scraped plots measured was  
424 at site K02 where the geometric mean of flux from the control plot was  $0.065 \text{ mg m}^{-2}$   
425  $\text{s}^{-1}$  whereas the biocrust-removed plot had a geometric mean flux of  $5.007 \text{ mg m}^{-2} \text{ s}^{-1}$ ,  
426 77 times greater than the control plot. The largest flux of dust from any one  
427 measurement was recorded on a biocrust-removed surface with a peak flux of  $6.305$   
428  $\text{mg m}^{-2} \text{ s}^{-1}$  at the burned interdune plot at site K02.

429

### 430 3.6. Summary of results

431 The Kalahari sediment has up to 9 % of grains by volume with a grainsize of  $< 10$   
432  $\mu\text{m}$ . This is particularly the case in the interdune areas. When tested using the PI-  
433 SWERL, all sites displayed some  $\text{PM}_{10}$  emission, but the threshold for erosion and  
434 emission flux varied with the surface status (control/burned/scraped), site, and the  
435 morphological unit of the dune. The dune crests had the lowest dust flux with an  
436 average emission of  $0.253 \text{ mg m}^{-2} \text{ s}^{-1}$  whereas the flanks showed the largest dust flux  
437 with an average emission of  $0.447 \text{ mg m}^{-2} \text{ s}^{-1}$ . The results of this study show that fine  
438 sediment is found in most of the experimental plots. Burning increases the erosion  
439 threshold of dune surfaces although this does not significantly impact the flux of

440 emitted dust (Table 4). Once the biocrust has been removed, dust flux can increase  
441 by up to 77 times compared to control and burned surfaces.

442

## 443 4. Discussion

444 In this study, a wide range of variables (dune morphological unit, burning, and  
445 biocrust removal) were tested to assess their impact on dust emission from the  
446 southwest Kalahari dune system. Characterisation of emission potential of vegetated  
447 dune fields and factors that modulate the baseline emissivity are important to  
448 quantify for global emissivity/erodibility maps and models of dust flux, both now  
449 and into the future (Kok et al., 2021a). This study presents a broad dataset of  
450 potential PM<sub>10</sub> emissions and estimated threshold velocities from vegetated linear  
451 sand dunes and further explores within-dune landscape variations and different  
452 disturbance pressures. Overall, results indicate that potential emission of mineral  
453 dust is low across all sites in the southwest Kalahari linear dune system regardless of  
454 whether the dunes have burned or not. But disturbing the surface through the  
455 removal of biocrust results in higher emission fluxes (up to 77 times greater) of dust.

456

### 457 4.1. Fine sediments within dune compositions

458

459 Grain size analysis shows that there is fine material (< 62.5 µm) within the sediment  
460 composition (ranging from 0 % to 30 %) at most plots measured. Fine grain (< 62.5  
461 µm) presence indicates that if the dunes become de-vegetated the fine grains should  
462 be available to be eroded by the wind. This substantiates the observation by

463 Bhattachan et al. (2012) that the Kalahari has the potential to emit dust once de-  
 464 vegetated.

465

466 **4.2. Emission from the southwest Kalahari linear dunes**  
 467 **compared to other PI-SWERL tests**

468

469 **Table 6.** Measurements of dust flux from sand dunes in previous PI-SWERL studies. If the  
 470 study reports where on the dune measurements are made the topography is reported,  
 471 otherwise the terminology used in the study is used. Many studies did not report the  $\alpha$   
 472 value, if the value is not mentioned a hyphen is used in place. Where studies have multiple  
 473 measurements of the same surface at different  $u_*$  then the closest two  $u_*$  values to the current  
 474 study ( $0.64 \text{ m s}^{-1}$ ) are used.

Study	Topography	$u_*$ ( $\text{m s}^{-1}$ )	$\alpha$ value	Emission flux ( $\text{mg m}^{-2} \text{s}^{-1}$ )	Metric
Bacon et al., 2011	Crest	0.69	-	$4.21 \pm 4.72$	Geometric mean $\pm$ Standard Deviation
Cui et al., 2019	Interdune	0.55	0.96-0.90	$\sim 1$	Geometric mean
Dickey et al., 2023	Interdune	0.73	0.9	$2.30 \pm 0.81$	Geometric mean $\pm$ Standard Deviation
King et al., 2011	Interdune	0.56	-	$0.269 \pm 1.252$	Geometric mean $\pm$ Standard Deviation
King et al., 2011	Interdune	0.56	-	$0.279 \pm 1.511$	Geometric mean $\pm$ Standard Deviation
Kolesar et al., 2022	Active dune	-	-	$0.20 \pm 0.11$	Mean $\pm$ Standard Deviation
Sweeney et al., 2018	Coppice	0.7	0.98 - 0.92	$\sim 2$	Median
Sweeney et al., 2016	Stabilised	0.7	0.98 - 0.92	$<0.5$	Median
Sweeney et al., 2011	Dune	0.56	-	0.1443	Geometric mean
Sweeney et al., 2022	Vegetated	0.6	0.96	$0.356 \pm 0.241$	Geometric mean $\pm$ Standard Deviation
Sweeney et al., 2022	Interdune	0.6	0.88-0.96	$0.033 \pm 0.063$	Geometric mean $\pm$ Standard Deviation
Sweeney et al., 2022	Interdune	0.7	0.88-0.96	$0.070 \pm 0.193$	Geometric mean $\pm$ Standard Deviation
Sweeney et al., 2022	Active	0.6	0.96-0.98	$0.048 \pm 0.059$	Geometric mean $\pm$ Standard Deviation
von Holdt et al., 2019	Dune	-	-	$0.0640 \pm 1.199$	Geometric mean $\pm$ Standard Deviation
Wang et al., 2025	Active	0 – 0.69	-	0.25	Average
<b>This study</b>	<b>Interdune (undisturbed)</b>	<b>0.64</b>	<b>0.98</b>	<b><math>0.281 \pm 3.061</math></b>	<b>Geometric mean <math>\pm</math> Standard Deviation</b>
<b>This study</b>	<b>Crest</b>	<b>0.64</b>	<b>0.98</b>	<b><math>0.253 \pm 0.238</math></b>	<b>Geometric mean <math>\pm</math></b>

	(undisturbed)	Standard Deviation
475		
476	Comparison of the PI-SWERL data between studies must take account of the	
477	respective $u^*$ values used. Only one study used the same $u^*$ ( $0.64 \text{ m s}^{-1}$ ) as this study	
478	and used emission fluxes on a variety of different landscapes in the Salton Sea, USA	
479	(Dickey et al., 2023). However, all dune measurements by Dickey et al. (2023)	
480	reported higher emission flux measurements than both the burned and control plots	
481	in the current study. These higher fluxes included surfaces which typically would be	
482	expected to have lower flux such as sand with gravel lag (Dickey et al., 2023).	
483		
484	Previous PI-SWERL experiments have found that undisturbed bare sand dunes have	
485	a low emission potential when compared to landforms such as dry washes or playas	
486	(King et al., 2011; Sweeney et al., 2016) but higher emissions than many other	
487	surfaces such as desert pavements and alluvial fans (Dickey et al., 2023). The dust	
488	flux measured from the partially vegetated linear dunes in this study is frequently	
489	lower than those from dunes in other geographic locations. For example, Sweeney et	
490	al. (2023) found a $\text{PM}_{10}$ max of $18.99 \text{ mg m}^{-2} \text{ s}^{-1}$ from sand from a stabilised dune in	
491	Arizona. This is an order of magnitude higher than the maximum undisturbed flux	
492	of $1.89 \text{ mg m}^{-2} \text{ s}^{-1}$ which was recorded at K07 burned interdune plot, however	
493	Sweeney et al. (2023) exerted a much higher shear velocity of $0.82 \text{ m s}^{-1}$ on the	
494	surface compared to this study's $0.64 \text{ m s}^{-1}$ . Higher emission fluxes were also	
495	observed from dunes in the Salton Sea, USA where an average flux of $2.30 \text{ mg m}^{-2} \text{ s}^{-1}$	
496	was measured at a $u_*$ of $0.73 \text{ m s}^{-1}$ (Table 6; Dickey et al., 2023). Other studies have	
497	measured similar levels of emission flux to this study, for example in northern China	
498	Cui et al. (2019) at a slightly higher $u_*$ of $0.69 \text{ m s}^{-1}$ and in the southwest USA by King	

499 et al. (2011) at a lower  $u_*$  of  $0.56 \text{ m s}^{-1}$  (Table 6). The fluxes of dust from the scraped  
500 surfaces were of similar magnitude to highly emissive surfaces such as dry washes  
501 (King et al., 2011), smooth playas (Dickey et al., 2023) and Wadis (Cui et al., 2019)  
502 although the  $u^*$  values used in these studies ranged from  $0.55$  to  $0.73 \text{ m s}^{-1}$ . The  
503 similar fluxes indicates that if the biocrust is removed, emission fluxes can be similar  
504 to those reported from frequently dust producing topographies.

505

506 Notably, there is a wide range of  $\alpha$  values used to account for different surface  
507 roughness and to calculate  $u_*$  used in all PI-SWERL studies. Of the above-mentioned  
508 studies,  $\alpha$  values range from  $0.9 - 0.96$  on sand dune surfaces, which affect the  $u_*$   
509 value and cross-study comparison. These values were chosen based on a table in  
510 Etyemezian et al. (2014). The  $u_*$  value is sensitive to the choice in  $\alpha$  value. For  
511 example, the peak RPM of 3500 was used in this study with a resulting  $u_*$  value of  
512  $0.64 \text{ m s}^{-1}$  at an  $\alpha$  value of  $0.98$ , if the  $\alpha$  value is changed to  $0.9$  the  $u_*$  value becomes  
513  $0.85 \text{ m s}^{-1}$ . Recently a visual look up table was created by Hartshorn et al. (2023) and  
514 this table was used in the selection of  $0.98$  as the  $\alpha$  value for this study. The new  
515 look-up table should reduce the discrepancy in assigning  $\alpha$  values and thus  $u_*$   
516 estimates into the future.

517

### 518 4.3. Disturbance effects

519 Factors that are frequently cited as disturbance agents in the arid zone context  
520 include drought, fire, and grazing (e.g., Hesse and Simpson, 2006; Mayaud et al.,  
521 2016; Nield and Baas, 2008; Wasson and Nanninga, 1986). Drought and fire do not  
522 produce any mechanistic changes in sediment distribution at the surface, only

523 grazing reorganises the sediments through hoof action. In this study, we investigate  
524 two types of disturbance, biocrust-removal (scraped surfaces) and environmental  
525 (burned surfaces). In previous studies, disturbance has been an umbrella term to  
526 show that disturbed surfaces produce a higher flux of dust (Bacon et al., 2011; Cui et  
527 al., 2019).

528

529 To simulate disturbed biocrust surfaces in the Kalahari largely caused by ungulate  
530 hoof action (Thomas and Dougill, 2007) biocrusts were removed by scraping. In  
531 these modified plots fluxes were larger (geomean  $1.654 \text{ mg m}^{-2} \text{ s}^{-1}$ ) than undisturbed  
532 (geomean  $0.314 \text{ mg m}^{-2} \text{ s}^{-1}$ ) and fire disturbed (geomean  $0.240 \text{ mg m}^{-2} \text{ s}^{-1}$ ) plots. The  
533 scraped plots had dust fluxes similar to those of dust emission hotspots such as dry  
534 washes and playas (King et al., 2011) and other generalised disturbed surfaces (Cui  
535 et al., 2019). The difference between the biocrust-removed and fire or control  
536 measurements reveals the stabilising role of biocrusts on dune surfaces.

537

538 The survival of biocrusts post-fire is an explanation as to why there is no statistical  
539 difference in the potential emission flux between the burned and control plots. The  
540 results show that the protective effect of biocrust remains after burning. Biocrust  
541 toleration to fire has been found in the USA and Australia (Palmer et al., 2020), but at  
542 present there are no studies in southern Africa. Biocrusts in the dune field likely  
543 survive burning as the highly combustible vegetation in the Kalahari burns fast  
544 (Andela et al., 2019). The rapid combustion of vegetation reduces the amount of heat  
545 transferred to the surface sediment, allowing some biocrust to withstand fire (Palmer  
546 et al., 2020; Lentile et al., 2006). Even if the heat transfer is fatal to the biocrusts the

547 results from this study show biocrust do not lose their protective effects. In addition,  
548 the presence of biocrusts has previously been shown to limit PI-SWERL-generated  
549 dust flux (Vos et al., 2020; Fick et al., 2020) and globally biocrusts are thought to  
550 reduce dust emissions by around 60 % (Rodríguez-Caballero et al., 2022).

551

552 As noted in figure 4 and section 3.4 the threshold shear velocity ( $u_{*T}$ ) increases after  
553 burning. There is some uncertainty in why this can happen. Given biocrusts were  
554 noted at each site, including after burning, and the widespread nature of biocrusts in  
555 the southwest Kalahari it can be inferred that the biocrusts play a stabilising role in  
556 maintaining the  $u_{*T}$ . In addition, the higher  $u_{*T}$  in burned plots can result from several  
557 other factors, including the 'baking' of the ground surface during burning and  
558 infiltration induced sediment sorting.

559

560 Several studies have indicated that surface sediments may undergo thermal  
561 alteration or 'baking' following fire (Stavi et al., 2017; Agbeshie et al., 2022). This  
562 response is attributed to the presence of finer grain fractions in surface sediments,  
563 which are more susceptible to structural disruption and cementation at lower  
564 combustion temperatures (Agbeshie et al., 2022). However, this phenomenon  
565 remains insufficiently investigated in sandy soil systems. The possible 'baking' of the  
566 surface also explains the burned plot higher threshold friction velocity.

567

568 Similarly, infiltration-driven vertical grain-size sorting occurs in sandy soils, where  
569 finer grains accumulate deeper in the sediment profile and could account for the  
570 higher dust fluxes with mechanistic disturbance (Lei et al., 2017b, a). Again, these

571 processes are insufficiently investigated in sandy soil systems. The alteration of the  
572 surface through scraping removed any sorted, baked sediments, and biocrusts,  
573 leaving the finer sediment exposed. The removal of these factors together or in  
574 isolation may have driven the higher dust fluxes (Figure 4b), indicating that there is  
575 the potential for the southwest Kalahari to emit dust.

576

577 The higher  $u^*_{\tau}$  post-burn is in contrast to previous findings where PI-SWERL  
578 measurements in semi-arid shrubland has shown that desert surfaces have higher  
579 emissivity after fire (Sankey et al., 2011, 2012). The emissions recorded in Sankey et  
580 al., (2012) were concentrated in hotspots under shrubs and were likely due to burn-  
581 induced soil hydrophobicity under the shrub canopies (Ravi et al., 2007, 2006). The  
582 experimental design for the current study intentionally did not conduct wind tunnel  
583 tests under shrubs due to woody vegetation burning incompletely, increasing the  
584 surface roughness, and subsequently the shrubby areas would likely not emit dust  
585 (see Figure 7).

586



587  
588 **Figure 7.** An example of a burned formerly shrubby area three months after fire at K03. The  
589 remaining shrubs increase surface roughness and reduce the likelihood of the location  
590 emitting dust.

591

#### 592 4.4. The impact of the dune morphological unit on emission flux

593 The dust flux, erosion threshold, and grain size differ significantly with the dune  
594 morphological unit (Figures 4 and 5, Tables 3, 4, and 5), consequently results were  
595 divided into the morphological units of crest, flank, and interdune. Overall, the PI-  
596 SWERL emission flux results suggest that all morphological units have the potential  
597 of emitting low level dust emissions (Table 4). The minimum flux of dust recorded in  
598 this study was  $0.005 \text{ mg m}^{-2} \text{ s}^{-1}$  at the burned interdune of site K01 one month after  
599 burning (Table 4), demonstrating that, even though this is low, the surface can emit  
600 some  $\text{PM}_{10}$ . The interdunes had the largest range of emission flux, and the highest  
601 maximum flux of the dune morphological units ( $1.892 \text{ mg m}^{-2} \text{ s}^{-1}$  at the burned  
602 interdune K07 plot), but this was not reflected in the geometric mean where the  
603 emissions were mostly lower than the flanks and crests (Table 4). In previous work

604 in the Kalahari, Bhattachan et al. (2012) discounted the crests of the dunes from  
605 being dust-emitting surfaces. Instead, the results from this study suggests that both  
606 the flanks and crests have some capacity to emit dust and should be included in  
607 dune-field scale estimates of dust emission flux.

608

609 On the crests, there was a 1.8 times greater emission flux from the average burned  
610 ( $0.467 \text{ m s}^{-1}$ ) than the control ( $0.253 \text{ m s}^{-1}$ ) plots. The higher flux for emission on  
611 burned dune crests is likely an indicator of sediment size distribution rather than the  
612 effect of fire. In the Kalahari, many dune crests are not vegetated, if a crest burned it  
613 must have had vegetation cover sufficient to allow the spread of fire. Therefore, as  
614 vegetation traps allochthonous dust and allows some pedogenesis (Garcia-Pichel et  
615 al., 2016; Weber et al., 2022; Garzanti et al., 2022; Gonzales et al., 2018), the burned  
616 crest have some finer sediment fraction available to be eroded by the wind. Finer  
617 material in burned crests was reflected in our sediment size distributions as the  
618 burned crest plots had an average of 2.6 % silt content whereas the control crest plots  
619 had an average of 0.3 %. This finding substantiates previous work which has  
620 highlighted that vegetated dune fields have a higher potential to emit dust, if  
621 denuded of vegetation, than active dune fields (Bullard et al, 2011; Sweeney et al.,  
622 2023).

623

## 624 5. Conclusions

625 The aims of this study were to (1) establish whether the southwest Kalahari dunes  
626 have a sedimentary potential to emit dust, and (2) quantify whether the burning of  
627 dune vegetation has a significant effect on dust emission. Through utilising surface

628 sediment sampling and portable wind-tunnel experiments on co-located burned and  
629 unburned control plots, we present the first quantification of the impact of fire on the  
630 dust emission potential of partially vegetated dune systems. The principal findings  
631 from this study are:

- 632 1. Up to 9 % of grains by volume in Kalahari dune field surface sediments are  
633 smaller than 10  $\mu\text{m}$ , particularly in the interdune areas. This indicates the  
634 potential for dust emission from the southwest Kalahari dune field.
- 635 2. Burning increases the erosion threshold of dune surfaces although burning does  
636 not significantly impact the measured flux of emitted dust.
- 637 3. The stabilising role biological soil crusts is indicated by an average 8-13 times  
638 higher dust emission potential from scraped surfaces.

639

640 The findings from this study highlight the complex biogeomorphic interactions that  
641 govern arid desert dune systems. In the southwest Kalahari, vegetation is not the  
642 sole stabilising factor that prevents dust from being emitted from the dunes. Instead,  
643 biological soil crusts provide surface stabilisation in de-vegetated landscapes. This  
644 finding is critical to dune evolution and dust emission models. The role of biocrusts  
645 in suppressing desert dust emission worldwide is poorly quantified and future work  
646 should aim to enhance our understanding on dust-biocrust interactions. Further,  
647 other factors may govern post-fire dust emissions, such as infiltration sediment  
648 sorting and the 'baking' of sediments. These factors are poorly constrained and need  
649 to be investigated to quantify their role in dust emission from partially vegetated  
650 dune fields.

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655 **Data availability:** The raw PI-SWERL experimental data have been placed on an  
656 open-access repository and can be found at the following DOI:  
657 10.5281/zenodo.15470969. The grain size distribution data can be accessed upon  
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659

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