

**Tracing enhanced oil recovery signatures in casing gases from the Lost Hills oil field  
using noble gases**

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**Abstract**

Enhanced oil recovery (EOR) and hydraulic fracturing practices are commonly used methods to improve hydrocarbon extraction efficiency; however, the environmental effects of such practices remain poorly understood. EOR is particularly prevalent in oil fields throughout California where water resources are in high demand and the disposal of large volumes of produced water may affect groundwater quality. Consequently, it is essential to better understand the fate of injected (EOR) fluids in California, and other subsurface petroleum systems, as well as any potential effect on nearby aquifer

systems. Noble gases can be used as tracers to understand hydrocarbon generation, migration, and storage conditions, as well as the relative proportions of oil and water present in the subsurface. In addition, a noble gas signature diagnostic of injected (EOR) fluids can be readily identified. We report noble gas isotope and concentration data in casing gases from oil production wells in the Lost Hills oil field, northwest of Bakersfield, California, and injectate gas data from the Fruitvale oil field, located within the city of Bakersfield. Casing and injectate gas data are used to: 1) establish pristine hydrocarbon noble-gas signatures and the processes controlling noble gas distributions, 2) characterize the noble gas signature of injectate fluids, 3) trace injectate fluids in the subsurface, and 4) construct a model to estimate EOR efficiency. Noble gas results range from pristine to significantly modified by EOR, and can be best explained using a solubility exchange model between oil and connate/formation fluids, followed by gas exsolution upon production. This model is sensitive to oil-water interaction during hydrocarbon expulsion, migration, and storage at reservoir conditions, as well as any subsequent modification by EOR.

## **1. Introduction**

Noble gases are excellent tracers of a variety of subsurface fluid flow processes due to their inert nature and distinct isotopic signatures. Noble gas isotopes and relative concentrations can be used to provide limits on the volumes of differently sourced fluids that have contributed to any particular system (Ballentine et al., 1991; 1996; 2002). Terrestrial reservoirs (i.e., atmospheric, crustal and mantle) have diagnostic noble gas isotopic compositions, and fluids derived from each reservoir can be differentiated. Noble gases from fluids in sedimentary basins have been used to successfully quantify physical exchange mechanisms between water, oil and gas phases in conventional and unconventional hydrocarbon systems (Ballentine et al., 2002; Hunt et al., 2012; Prinzhofer, 2013; Darrah et al., 2014; 2015; Wen et al., 2015; 2016; Barry et al., 2016; 2017; Byrne et al., 2018).

Noble gases are also useful tracers of fugitive natural gases detected in groundwater (Darrah et al., 2014; 2015; Wen et al., 2015; Harkness et al., 2017) and for understanding the effects of CO<sub>2</sub> injection and EOR (LaForce et al., 2014; Gyore et al., 2015; Sathaye et al., 2016). To date, the use of noble gas signatures as tracers of reservoir processes in oil dominated systems is not well-developed, and requires precise characterization of noble gas signatures of oil reservoir fluids. However, subsurface hydrocarbon systems are intrinsically complex, marked by multiphase interactions (e.g., Barry et al., 2017) and affected by a range of natural geologic and anthropogenic processes. As a result, determining proportions of gas, oil, and formation water is difficult using geophysical (e.g., seismic, wireline logs, continuous core) and traditional geochemical techniques (Batzle and Wang, 1992; Schoell, 1980). For example, geochemical variability within a hydrocarbon system is controlled by the timing and origin of hydrocarbon charge, secondary fluid migration, and the relative distribution of regional trapping structures. Further complicating these signatures is the superimposed effect of EOR techniques, including steam or CO<sub>2</sub> flooding, and unconventional techniques such as hydraulic fracturing (e.g., Darrah et al., 2014).

In this study, a novel geochemical approach is used to differentiate between source and secondary (EOR) geochemical features in an actively producing hydrocarbon system. This distinction is critical, as it provides the necessary baseline values for fingerprinting any fugitive hydrocarbons that might be identified in nearby aquifer systems, which are valuable natural resources as a source of water supply for drinking water and agricultural irrigation (Mount et al., 2014). A primary goal of the California State Water Resource Control Board's (SWRCB) Oil and Gas Regional Groundwater Monitoring Program (RMP) is to determine if any fluid exchange has occurred between regions affected by oil and gas activities and the surrounding groundwater (Taylor et al., 2014). The RMP will investigate approximately 100 onshore oil fields, thus providing information to resource management agencies about relative risks to groundwater. This study utilizes noble gas concentrations and isotope ratios in order to understand the

88 extent of exchange between oil and formation waters. Our models describe how noble  
89 gas concentrations and elemental ratios evolve as 1) natural formation waters (i.e., air  
90 saturated water [ASW]) interact and exchange with oil in a pristine system, and 2) what  
91 the secondary effects of human-induced EOR are on the noble gases within a subsurface  
92 petroleum system.

## 94 **2. Lost Hills Production History**

95 Oil was discovered in the Lost Hills area in 1910, which is located in the southwestern  
96 portion of San Joaquin Valley, approximately 220 km northeast of Los Angeles and 70  
97 km northwest of Bakersfield (Long et al., 2015), California, USA. The Lost Hills field  
98 currently produces oil, with no separate hydrocarbon gas phase (or 'gas-cap') being  
99 encountered. Details about geologic and depositional history are provided in the  
100 Supplementary Material.

101  
102 Oil and gas operators commonly inject water and steam into subsurface petroleum-  
103 bearing rocks (e.g., Tulare Formation, Etchegoin Formation) throughout the San Joaquin  
104 Valley in order to promote secondary recovery of hydrocarbons and to maintain  
105 adequate fluid pressures in petroleum reservoirs. Injected fluids commonly consist of  
106 briny produced, surface or well water. The Tulare Formation was extensively  
107 steamflooded, beginning in 1964. The Diatomite, within the Monterey Formation, was  
108 waterflooded, hydraulically fractured, and subjected to CO<sub>2</sub> injection, with the earliest  
109 hydraulic fracturing occurring in 1962 (Land, 1984). All wells where injection has  
110 occurred (since 1977) in the Lost Hills field are shown in Figure 1a. Tracers injected into  
111 the Belridge Diatomite of the Monterey Formation have been shown to breakthrough to  
112 the oil producing wells in less than one month, which may indicate a naturally fractured  
113 reservoir (Zhou et al., 2002). The total number of injection wells (within a 100 and 500  
114 meter radius of the sample well) has been tabulated in Table 1 along with aggregate  
115 injection volumes. Supplementary Figures S3-S7 show the injection history of Lost Hills  
116 (since 1977) and the proximity to wells sampled in this study. Digitized records of

injection volumes per well prior to 1977 were not available, but for the entire Lost Hills field, injection occurring during 1952-1976 accounted for about 2% of total historic injections.

Formation water at Lost Hills has salinity between 0.54-0.60 M NaCl and an average reservoir temperature of 57°C. All oil gravities (expressed as API) are homogeneous at approximately 23.1, apart from sample 'LH-05', which has an API of 21.5. All temperature, salinity and API data are taken from (n=7) well-logs in the Lost Hills area. Heavier oil occurs in the (overlying) Tulare Formation (12-18 API).

### **3. Methods**

#### **3.1 Sample Collection**

In October 2016 (n=10) gas samples were collected: seven oil wells (and a duplicate) in the Lost Hills field were sampled, and two injection sites in the Fruitvale field. We assume that all injection fluids are air-dominated, as they are all stored in surface tanks, which are not sealed effectively enough to prevent interaction with the atmosphere.

Samples were collected from well casing and injectate gases as part of the SWRCB's RMP. Casing gases are pressurized gases that exsolve from produced fluids within the wellhead, whereas injectate gases are those that are re-injected as part of the EOR process. All Lost Hills samples are derived from oil systems and the gas samples collected were producing 'casing gases'. Notably, Lost Hills casing gas samples were collected from both wells that have been recently subjected to EOR, and wells that have been producing oil without enhanced recovery methods. In contrast, Fruitvale samples were collected from the gas phase of co-mingled injectate fluids held in storage tanks, directly from different oil wells. The goal of collecting injectate gas samples was to determine the noble gas signature of fluids and gases associated with EOR. Samples from each site are geochemically distinct from one another, and represent co-mingling of injectate fluids with air additions, likely from leaking surface storage tanks.

Gases from each location were collected in copper tubes for noble gases using standard techniques (Burnard, 2013). Necessary precautions were taken during sampling to ensure sample integrity. For example, noble gas samples were collected in 3/8" diameter, refrigeration-grade copper tubes, which were connected using hose clamps and reinforced polytetrafluoroethylene tubing to a 2-stage pressure regulator attached directly to the wellhead. The regulator was used to step down the pressure from the wellstream to 1-2 bars. The copper tubes were then flushed with the produced gas to avoid air contamination and the gas outlet was submerged in a bucket of water to prevent any backflow of air into the system. After ten minutes of flushing the Cu-tubes were sealed with stainless steel clamps (Weiss, 1968).

### 3.2 Noble Gas Results

Noble gas analyses were performed at the University of Oxford using standard techniques (details in Supplementary Material). Lost Hills helium ( $^4\text{He}$ ) concentrations range from  $0.1$  to  $7.5 \times 10^{-6} \text{ cm}^3\text{STP/cm}^3$ , where STP is defined as Standard Temperature and Pressure. Samples are strongly radiogenic and have helium isotope ( $^3\text{He}/^4\text{He}$ ) values ranging between  $0.018 \pm 0.001$  and  $0.060 \pm 0.003 R_A$  (Table 2), reported relative to the air value  $R_A = 1.38 \times 10^{-6}$  (air =  $1R_A$ ; Clarke et al., 1976). Fruitvale helium ( $^4\text{He}$ ) concentrations range from  $2.9$  to  $3.2 \times 10^{-6} \text{ cm}^3\text{STP/cm}^3$ . A single helium isotope ( $^3\text{He}/^4\text{He}$ ) value of  $0.125 \pm 0.006 R_A$  was determined for the 'FR-03' sample. All noble gas data are reported in Table 2 along with  $1\sigma$  uncertainties.

Lost Hills neon ( $^{20}\text{Ne}$ ) concentrations range from  $1.85$  to  $24.0 \times 10^{-9} \text{ cm}^3\text{STP/cm}^3$ . Measured  $^{20}\text{Ne}/^{22}\text{Ne}$  ratios vary between  $9.38 \pm 0.005$  and  $10.17 \pm 0.005$ ; deviations from the atmospheric  $^{20}\text{Ne}/^{22}\text{Ne}$  value of  $9.8$  are likely due to a combination of nucleogenic  $^{22}\text{Ne}$  production and mass fractionation effects. Measured  $^{21}\text{Ne}/^{22}\text{Ne}$  values vary between  $0.0283 \pm 0.0001$  and  $0.0305 \pm 0.0002$ . Similarly,  $^{21}\text{Ne}/^{22}\text{Ne}$  isotopic ratios in excess of the air value ( $0.0290$ ) can be attributed to mass fractionation and

nucleogenic crustal  $^{21}\text{Ne}$  addition. Fruitvale neon ( $^{20}\text{Ne}$ ) concentrations are higher and range from 79.7 to  $606 \times 10^{-9} \text{ cm}^3\text{STP/cm}^3$ . Measured  $^{20}\text{Ne}/^{22}\text{Ne}$  and  $^{21}\text{Ne}/^{22}\text{Ne}$  ratios are nearly identical to air.

Lost Hills argon ( $^{36}\text{Ar}$ ) concentrations range from 0.02 to  $0.09 \times 10^{-6} \text{ cm}^3\text{STP/cm}^3$  and reveal significant deviations from the atmospheric  $^{40}\text{Ar}/^{36}\text{Ar}$  value of 298.56 (Lee et al., 2006), with measured  $^{40}\text{Ar}/^{36}\text{Ar}$  values varying between  $299 \pm 1.50$  and  $334 \pm 1.67$ . Fruitvale argon ( $^{36}\text{Ar}$ ) concentrations range from 0.19 to  $1.22 \times 10^{-6} \text{ cm}^3\text{STP/cm}^3$  and show smaller deviations from air, with measured  $^{40}\text{Ar}/^{36}\text{Ar}$  values varying between  $295 \pm 1.5$  and  $311 \pm 1.6$ . Both areas show a resolvable radiogenic  $^{40}\text{Ar}$  contribution within the petroleum system. Argon isotopes ( $^{40}\text{Ar}/^{36}\text{Ar}$ ) inversely correlate with  $^{20}\text{Ne}$  concentrations, suggesting that the source of the  $^{20}\text{Ne}$  is also a source of air-like argon isotopes, particularly in the Fruitvale samples. Measured  $^{38}\text{Ar}/^{36}\text{Ar}$  values are predominantly air-like (0.1885; Lee et al., 2006; Burnard, 2013) and vary between  $0.189 \pm 0.003$  and  $0.197 \pm 0.003$ .

Lost Hills krypton ( $^{84}\text{Kr}$ ) concentrations range from 1.42 to  $6.27 \times 10^{-9} \text{ cm}^3\text{STP/cm}^3$  and from 4.59 to  $26.5 \times 10^{-9} \text{ cm}^3\text{STP/cm}^3$  in Fruitvale samples. Lost Hills samples overlap and extend above air values (0.303; Aregbe et al., 1996) with respect to  $^{86}\text{Kr}/^{84}\text{Kr}$ , ranging between  $0.29 \pm 0.001$  and  $0.39 \pm 0.002$ , whereas Fruitvale samples range from  $0.31 \pm 0.001$  to  $0.49 \pm 0.001$ , plotting above the air-value, most likely due to mass fractionation processes.

Lost Hills xenon ( $^{132}\text{Xe}$ ) concentrations range from 0.73 to  $2.12 \times 10^{-9} \text{ cm}^3\text{STP/cm}^3$  (Table 2). Lost Hills  $^{132}\text{Xe}/^{130}\text{Xe}$  values range between  $6.56 \pm 0.03$  and  $6.58 \pm 0.03$ , and are distinctly lower than air values ( $^{132}\text{Xe}/^{130}\text{Xe}_{\text{air}} = 6.60705$ ; Pepin, 2000). Fruitvale xenon ( $^{132}\text{Xe}$ ) concentrations range from 1.73 to  $2.02 \times 10^{-9} \text{ cm}^3\text{STP/cm}^3$ . Fruitvale  $^{132}\text{Xe}/^{130}\text{Xe}$  values bracket air ratios, ranging between  $6.57 \pm 0.03$  and  $6.67 \pm 0.03$ .

## 4. Discussion

### 4.1 Sample integrity and classification

Before interpreting and discussing the variability in gas samples collected from the Lost Hills and Fruitvale oil fields, it is first necessary to classify sample types. Only injectate gas samples (n=2) were collected from Fruitvale, and these are broadly similar to air-like noble gases (i.e.,  $^{20}\text{Ne}/^{36}\text{Ar}$ ). These gases represent headspace gases in closed surface storage tanks, however there is a strong atmospheric signal, likely due to air leaks in the tanks. Casing gases (n=7) were collected from Lost Hills; these gases are exsolved from produced fluids (i.e., oil and water) as it is pumped to the surface, and noble gas isotope ratios are distinct from air. Lost Hills samples are sourced from either the Belridge Diatomite member or the deeper Cahn member of the Monterey Formation (consisting of the Brown and Antelope shales; stratigraphic cross section – Figure S2).

Cahn gas samples (LH-10 and LH-09) have lower non-radiogenic noble gas concentrations (e.g.,  $^{20}\text{Ne}$ ,  $^{36}\text{Ar}$ ) and relative ratios (e.g.,  $^{20}\text{Ne}/^{36}\text{Ar}$ ), which plot in the range expected for solubility partitioning between gas and oil (Figure 3a and 3b), and are thus considered ‘pristine’ or unaffected by air-like injection. In contrast, Diatomite samples have higher non-radiogenic noble gas concentrations, relative elemental ratios (above the water solubility range), and exhibit air-influenced noble gas isotope ratios (Table 2), all of which are attributed to the injection of air-saturated fluids related to recent EOR activities, which are much more prevalent in the vicinity of Diatomite samples. In Table 1, we show the number of injection wells (since 1977) within an area with a 100m, and a 500m radius from each sampled well. This table also provides the median injection depth, and the total volume of injectate from the injection wells identified. The Diatomite sample (LH-04) has the most pristine noble gas signature amongst all Diatomite samples, and is in fact the oldest well in production (since 1990; DOGGR, 1998). Well ‘LH-04’ is also located approximately 1km to the north of the densest injection area (Figure 1a) and has had no injection since 1977 from wells within



an area with a 100m radius, suggesting that injection fluids do not traverse such distances.

Two samples were collected from the Cahn member: 'LH-10' (collected in duplicate; Table 2), which has been in production since 2001 and 'LH-09', which has been in production since 1981 (DOGGR, 1998). Notably, these wells also do not have any injection wells located within a 100m radius (Table 1; Figure S6). The duplicate analyses (LH-10) are highly reproducible (within 1%), which provides a high level of confidence in the integrity of the sampling, transport and storage procedure (Table 2; supplemental info S3.1). The relatively unmodified Cahn samples are best explained by a closed system equilibrium exchange model (4.3.1), which is used to determine the extent of hydrocarbon interactions with fluids in the reservoir and/or during migration from source to reservoir. In contrast, Diatomite wells have a higher density of injection wells within a 100m radius (Figure S4) and a higher cumulative injection volume (Table 1) in that radius. As a result, Diatomite samples are considerably 'less pristine', marked by excess air-derived noble gases. We surmise that these excess noble gas signals are the result of surface fluid injection; however, they can be corrected and volumetric ratios of water to oil can be calculated. In section 4.3.3, we quantify the extent of EOR using a radiogenic  $^4\text{He}$  model. Determining initial noble gas values for pristine and EOR affected regions of each formation is imperative for interpreting the extent of fluid exchange between oil fields and nearby aquifers.

All gases are strongly radiogenic with respect to helium isotopes; however, Fruitvale injectate samples have higher  $^3\text{He}/^4\text{He}$  ratios, closer to the ratio of air (Figure 2), which is consistent with a larger air contribution. EOR samples generally have higher concentrations of air-derived noble gases than pristine samples and more air-like elemental ratios (e.g.,  $^{20}\text{Ne}/^{36}\text{Ar}$ ; Figure 3), both of which are consistent with air-entrainment during re-injection. Lost Hills gases show a range from pristine crustal signatures – indicative of closed-system exchange with formation fluids – to slightly air-

contaminated and/or diluted signatures in the EOR region. In 4.5 we present a radiogenic model that explains the disparate  $^4\text{He}$  contents in Lost Hills samples by evoking dilution associated with recent EOR activity.

#### **4.2.1 Original non-radiogenic noble gas inventory**

In order to quantify the extent of interaction between an oil phase and formation waters in the subsurface (i.e., determine subsurface water-oil ratios), we must first constrain the original noble gas inventory of ASW in deep meteoric water. As with previous solubility models (e.g., Ballentine et al., 1991; Ballentine et al., 1996; Zhou et al., 2005; Barry et al., 2016; Barry et al., 2017; Tolstikhin et al., 2017), we assume that air-derived noble gases ( $^{20}\text{Ne}$ ,  $^{36}\text{Ar}$ ,  $^{84}\text{Kr}$  and  $^{132}\text{Xe}$ ) are originally input into the subsurface, dissolved in groundwater (i.e., ASW), during aquifer recharge. It is assumed that once groundwater recharge has occurred, the atmospheric noble gas characteristics of these fluids are conservative, as they are isolated from further additions of light ( $^{20}\text{Ne}$ ,  $^{36}\text{Ar}$ ) air-derived noble gases, which do not have significant subsurface sources. In contrast, addition of radiogenic and/or mantle-derived noble gases to deep meteoric waters will result in a distinct isotope signal. Initial noble gas solubility in water is a function of recharge conditions (temperature, salinity and recharge elevation), but generally increases with mass ( $\text{Ne} < \text{Ar} < \text{Kr} < \text{Xe}$ ). Therefore, the initial ASW noble gas inventory in formation water can be approximated using solubility coefficients ( $K_{\text{GW}}$ ) at a given temperature and salinity (Crovetto et al., 1982; Smith and Kennedy, 1983). In Lost Hills we assume 15°C recharge of freshwater at sea-level elevation; notably if different recharge conditions are used to calculate the initial noble gas inventory, this will have only a minor effect on noble gas reconstructions (see section 4.2.2).

#### **4.2.2 Reconstructing noble gas concentrations in oil**

During petroleum production, oil is pumped from the reservoir to the surface, resulting in decompression and eventual gas exsolution. The gas-oil-ratio (GOR) is typically

measured at the wellhead by combining the volume of gas produced with the volume of oil produced. The volumes are typically reported by the oil industry as Standard Cubic Feet per Barrel of Oil (SCF/BBL). We convert these to SI units ( $\text{m}^3$  of gas/ $\text{m}^3$  of oil), using eq. 1:

$$\text{GOR}_{\text{metric}} = \frac{(\text{GOR}_{\text{imperial}})(0.028316)}{0.1589873} \quad (1)$$

The GORs vary as a function of time and production history and some Lost Hills wells do not currently produce gas, and thus we use a time-integrated average GOR value from the Diatomite member to reconstruct the concentration of noble gases in the oil phase.

The concentration of a given noble gas, originally in the oil (in units of mol/Kg oil) phase at Reservoir Temperature and Pressure (RTP) (i.e., prior to production and exsolution), is estimated using the following equation:

$$[C_i^{\text{oil}}] = \frac{[C_i^{\text{gas}}](\text{GOR}_{\text{metric}})}{(\text{SG})(22.4)} \quad (2)$$

Where  $[C_i^{\text{gas}}]$  is the gas concentration of noble gas  $i$  in the gas phase, in units of  $\text{cm}^3\text{STP}/\text{cm}^3$  (Table 2). This calculation assumes one mole of gas occupies 22.4 L at STP. The specific gravity (SG) of oil is in units of  $\text{g}/\text{cm}^3$  and is a function of oil gravity: typically expressed in American Petroleum Institute (API) gravity units (i.e., inverse of density). The specific gravity of a given oil can be calculated using eq. 3 (Petrosky and Farshad, 1993):

$$\text{SG} = \frac{141.5}{\text{API} + 131.5} \quad (3)$$

In Table 2 the measured gas concentrations for all air-derived noble gas species  $[C_{i(\text{meas})}^g]$  (units are  $\text{cm}^3\text{STP}/\text{cm}^3$ ) are provided. In Table 3 measured API, SG,

metric GOR, and calculated concentrations in oil [ $C_i^{oil}$ ] (units are mol/Kg oil) are given along with solubility coefficients ( $K_{wo}^M$ ), calculated using Kharaka and Specht, 1988 at RTP (Lost Hills = 57°C and 0.57 M NaCl). Importantly, the calculation in equation 2 assumes equilibrium between the produced fluids and gases; if equilibrium is not reached then estimates of initial concentration in oil will be skewed. We also assume that gas concentrations are representative of a time-integrated average or production conditions.

Importantly, we use an average integrated GOR (=107.6) from Diatomite wells in order to model the Lost Hills data. We compiled production data at each well sampled in Lost Hills, and took an average value of all Diatomite sourced samples. Notably, there are differences in GOR between Diatomite and Cahn sourced wells, however due to the fact that all wells have a common API, we assume a common GOR in order to do reconstructions (Eq. 2). Determining an accurate GOR is critical for converting measured gas concentrations into concentrations in the reservoir. It is evident from the variation of the wellhead GOR values (measured during sample collection) that they are not always representative of virgin GOR conditions. Furthermore, historical well-log data demonstrate that GORs vary significantly over sustained periods of production. For example, measured GORs vary by nearly 4 orders of magnitude in the Lost Hills field, whereas measured noble gas concentrations vary by just 1 order of magnitude within the same fields (Figure 4). The fact that we observe relatively little variability in noble gas concentrations, suggests that using an average integrated GOR value is most appropriate.

#### **4.3.1 Assessing pristine samples with an equilibrium model**

As oil is produced and migrates from source to reservoir, noble gases are partitioned from deep formation connate/formation waters into the oil phase ( $K_{wo}$ ) in a predictable manner, based on empirically-derived solubility coefficients (Kharaka and Specht, 1988). Predicted concentrations are fixed by several parameters, including: 1) the initial noble

gas inventory in ASW ( $[C_i^{asw}]$ ), 2) temperature and salinity conditions where phase exchange occurs (reservoir conditions) in the hydrocarbon system, and 3) density of oil (API). Recharge conditions (i.e., freshwater at 15°C) are assumed for the first set of parameters. Temperature, salinity and API are constrained for each sample by measured values at the wellhead during sample collection. For Lost Hills we use 57°C and 0.57 M NaCl, which represents an average between Diatomite and Cahn samples. Oil density was used on a well-by-well basis, however most Lost Hills wells are homogeneous (Table 3).

Using an estimate of the initial noble gas content of undegassed oil ( $[C_i^{oil}]$ ) in the reservoir (calculated in Eq. 2; Table 3), we constructed a solubility model (after Ballentine, 1996; Barry et al., 2016), which predicts the relative volume of water that is required to have exchanged with an oil phase (assuming equilibrium conditions), in order to produce the observed noble gas concentrations in casing gases. Eq. 4 is modified from Barry et al., 2016:

$$\frac{V_o}{V_w} = \frac{[C_i^{asw}]}{[C_i^{oil}]} - \left[ \frac{(K_{WO}^M)}{(SG)(K_{GW}^M)} \right] \quad (4)$$

We note that  $(K_{GW}^M)$  is the reservoir temperature and salinity specific Henry's constant (Crovetto et al., 1982; Smith and Kennedy, 1983),  $K_{WO}^M$  is the reservoir temperature and salinity specific Henry's constant at a given specific gravity (Kharaka and Specht, 1988).

Thus, we use Eq. 4 in order to calculate  $\frac{V_o}{V_w}$  at reservoir conditions.

#### 4.3.2 Model sensitivity: implications for exchange and migration

We note that calculated  $\frac{V_o}{V_w}$  values are highly dependent on input assumptions (i.e., recharge and reservoir temperature and salinity conditions as well as measured API gravity). In the Lost Hills (Cahn and Diatomite) system, reservoir temperatures range between 44-69 °C, whereas salinity in the reservoirs spans a relatively narrow range of

0.54-0.60 M NaCl, both based on down-borehole data. Using an average reservoir temperature of 57°C and salinity of 0.57 M NaCl we calculate  $\frac{V_o}{V_w}$  at reservoir conditions using Eq. 4, and values are given in Table 1.

Cahn  $\frac{V_o}{V_w}$  values range between 0.17-0.36, which is in good agreement with the range in  $\frac{V_o}{V_w}$  (0.33-0.37) calculated indirectly by integrating the oil to water volume ratio in all Cahn wells, using historical well-log data (Table 1). The agreement between these two independent estimates suggests that this geochemical approach is well calibrated. Importantly, this method enables down borehole volumetric estimates to be made in the absence of well-log data.

The model was also run for Diatomite samples and calculated a significantly lower range in  $\frac{V_o}{V_w}$  (0.01-0.11), which also overlaps with integrated well-log data estimates (0.05-0.19) from the Diatomite samples. Notably, none of these values are corrected for the effect of EOR, however details of the correction are provided in 4.3.3 and the corrected values are tabulated in Table 1. The integrated Diatomite estimates ( $\frac{V_o}{V_w}$  from 0.05-0.19) are significantly lower than the Cahn estimates, suggesting greater water-oil interaction and partitioning. We suggest that the difference in  $\frac{V_o}{V_w}$  is directly related to migration distance, and thus the  $\frac{V_o}{V_w}$  can be used as a semi-quantitative ‘secondary-migration odometer’. Importantly, the range for the Diatomite samples encompasses the solubility model-derived  $\frac{V_o}{V_w}$  of 0.05-0.11, for  $^{20}\text{Ne}$  and  $^{36}\text{Ar}$  estimates respectively. Again, the agreement between the two independent estimates lends credence to the geochemical approach. Outputs from sample ‘LH-08’ clearly suggest that this particular sample has been isolated and subjected to much less water than the other Diatomite samples. This is likely the result of geologic compartmentalization, whereby portions of the reservoir have been segregated and experience a different amount of contact with extraneous

fluids. In this instance, the sample would be subjected to significantly less water contact, and thus would retain a much higher  $\frac{V_o}{V_w}$ .

In contrast, a longer migration pathway would result in more interaction with formation water, and thus a lower  $\frac{V_o}{V_w}$ . In view of the possible range in  $\frac{V_o}{V_w}$  that could explain the Lost Hills dataset; we have conducted a data inversion in order to determine the best-fit  $\frac{V_o}{V_w}$  value for Cahn and Diatomite samples respectively. Using a data inversion technique we fit variables in the conceptual model (Eq. 4) to measured data by minimizing the square of the residuals between model-derived (i.e., predicted)  $\frac{V_o}{V_w}$  and measured  $\frac{V_o}{V_w}$  values (Figure 5). Specific gravity, initial noble gas concentrations in ASW and partition coefficients are fixed (see Table 3), based on assumed recharge conditions. Neglecting helium – which is dominated by radiogenic contributions, and heavy noble gases which are disproportionately modified by excess heavy noble gases (section 4.4) – two observables remain:  $^{20}\text{Ne}$  and  $^{36}\text{Ar}$  concentrations in oil, which allows us to solve using data inversion for  $\frac{V_o}{V_w}$  (e.g., Ballentine and Hall 1999). The best-fit  $\frac{V_o}{V_w}$  can thus be inverted for both  $^{20}\text{Ne}$  and  $^{36}\text{Ar}$  concentrations in oil separately, providing a range in possible  $\frac{V_o}{V_w}$  values.

In Figure 5 we show measured noble gas concentration in oils vs. model predicted values, at a predicted  $\frac{V_o}{V_w}$  value of 0.322, based on a best fit for Ne and Ar using the inverse technique described above. At RTP conditions there is very good agreement for both Cahn samples, whereas the Diatomite samples have measured/modeled values  $>1$ , and show a marked enrichment in Ne and Ar, likely due to EOR (see section 4.3.3 for discussion). The disparity is the direct result of EOR, the effects of which are discussed in detail in section 4.3.3. Sample ‘LH-08’ is a clear outlier, plotting anomalously below all other samples, suggesting that this sample has evolved differently than the other samples – perhaps due to geological isolation. Heavy noble gas enrichment (only Kr is

shown) relative to the light noble gases remains pervasive in all samples; notably this is a phenomenon that has been observed previously and is discussed in section 4.4.

Agreement between predicted and measured  $\frac{V_o}{V_w}$  demonstrate that the solubility model presented here (Eq. 4) is highly effective at identifying oil-water interaction in the subsurface. The findings also have important implications for the mass balance of the system, i.e., the observation that Cahn samples have higher  $\frac{V_o}{V_w}$  (Table 1) implies less water contact, and thus a relatively short migration pathway, as is observed in Diatomite samples. This observation is consistent with the low permeability and high porosity nature of the Lost Hills hydrocarbon system, as such systems would be expected to limit the amount of fluid migration.

The model presented here is a zero-dimensional solubility exchange and mixing model that explains noble gas variations in Lost Hills casing gases. We use measured gas concentrations to determine  $\frac{V_o}{V_w}$  and surmise that differences between Cahn and Diatomite samples are the result of different migration pathways and EOR contributions. Previously, Sathaye et al., 2016 used a reactive transport model to demonstrate noble gas fractionation during subsurface gas migration, showing a characteristic pattern of initial co-enrichment of noble gases from both water and gas phases in banks at the gas front, followed by a depletion of the dissolved noble gas. This is an important observation, however, this enrichment would not be preserved if the leading edge passes through the system and, for example, is produced at the surface. We suggest that the Lost Hills system is unlikely to preserve the leading edge of the injected water as these are large volumes injected over time relative to the reservoir volume and the respective phases should approach the solubility controlled equilibrium conditions that we model. If equilibrium conditions were not reached, the Sathaye et al., 2016 model would provide an alternative explanation for the observed enrichments in EOR modified samples. Differentiating between these two processes is difficult to test using the available data set, but should be carefully considered. However, we argue that



despite the source of the noble gas enrichment (air-entrainment during injection or noble gas banking) we can effectively correct back to solubility conditions and calculate relative differences in  $\frac{V_o}{V_w}$ .

#### **4.3.3 Assessing EOR using air derived noble gases**

Fruitvale data provide direct evidence that injectate fluids are geochemically similar to air-like signatures (Figures 3), as they have orders of magnitude higher air-derived noble gas concentrations (i.e.,  $^{20}\text{Ne}$ ,  $^{36}\text{Ar}$ ) and air-like isotope ratios (Table 2). As a result, we can differentiate between pristine hydrocarbon signatures – that can be explained by an equilibrium exchange model and samples that have been recently subjected to other methods of EOR, and thus have higher measured noble gas concentrations than predictions would indicate (Figure 5). The equilibrium model (Eq. 4) assumes that an oil phase was in direct contact with a water phase, and that equilibrium was reached at reservoir conditions. The ‘1<sup>st</sup> bubble’ value (Figures 6a, 6b) represents the point at which there is an infinitesimally small volume of oil partitioning or ‘stripping’ noble gases from a significantly larger water volume, during initial charging of the reservoir, and importantly, no excess air entrainment from EOR is present. As the amount of oil relative to water increases, the system evolves along the line from infinitely small  $\frac{V_o}{V_w}$  towards infinitely large  $\frac{V_o}{V_w}$ . As the  $\frac{V_o}{V_w}$  increases, the noble gas concentration in oil decreases due to dilution. In Figure 6a, we show the predicted evolution of the oil phase at different  $\frac{V_o}{V_w}$ , combined with three different predicted EOR trajectories. The predicted EOR trajectories represent modeled two-component exchange between formation water and an oil phase (i.e., no gas-cap), followed by air-like addition assumed to be associated with recent EOR activities. Clearly, Diatomite samples have a more pronounced EOR signature (Figure 6), characterized by a range in  $^{20}\text{Ne}/^{36}\text{Ar}$  well above the range predicted by simple solubility exchange. The convergence of the predicted EOR trajectory (colored lines) and the  $\frac{V_o}{V_w}$  line (black) represents the initial pristine conditions which evolve along the EOR trajectories after the addition of different

amounts of air. Notably, the  $\frac{V_o}{V_w}$  evolves over the lifetime of the hydrocarbon system, whereas the effect of EOR is only superimposed upon those pristine gas signatures. As a result, the point at which the two predictive trends converge, the pristine  $\frac{V_o}{V_w}$  can be estimated. This is referred to as the EOR corrected  $\frac{V_o}{V_w}$  value and oil to water ratios following correction for the EOR component are tabulated in Table 1 (denoted with an asterisk). In Figures 6a,b there are distinct, mixing trajectories for deeper Cahn samples and another for Diatomite samples. There is also a single outlying Diatomite sample 'LH-08' that falls on a completely different trend, perhaps due to geologic isolation.

It is important to note that other factors could also affect the predicted EOR trajectories on Figures 6a,b. The solubility model assumes simple two-component exchange between deep meteoric formation water and an oil phase, followed by air addition associated with recent EOR activities. In Lost Hills, there is no evidence for a gas phase at depth, which would cause fractionation in the opposite direction (i.e., towards lower  $^{20}\text{Ne}/^{36}\text{Ar}$  values) as mixing with EOR fluids (Figures 6a,b). Instead, the modeling results indicate that gases only exsolved during production. This observation is also entirely consistent with geological constraints in that there is no ancillary evidence for gas-cap formation.

Diatomite  $\frac{V_o}{V_w}$  model values range from 0.01 to 0.11, whereas Cahn  $\frac{V_o}{V_w}$  model values range from 0.17 to 0.36. The disparity between estimates suggests that Diatomite samples have been exposed to significantly more water than the Cahn samples. Water could be introduced in several different ways. For example the disparity could be the result of: 1) the depth of the samples, whereby samples in the Diatomite may have a similar hydrocarbon source but have likely seen more water during migration, resulting in lower  $\frac{V_o}{V_w}$ ; 2) the relative permeabilities of the reservoirs, for example, oil in tight shales are not exposed to as much water as those in more porous Diatomite units above; 3) the relative porosities of the reservoirs, for example, Diatomite samples

higher porosity than the Cahn samples, so there would be more water within each bulk volume of Diatomite than of Cahn; 4) injection and human induced water pumping. Importantly, none of these explanations are mutually exclusive and the resulting signatures are likely the result of a combination of all factors. Importantly, when we correct  $\frac{V_o}{V_w}$  for EOR addition (Table 1) there is still a disparity, albeit less pronounced. This suggests that there is a natural difference between reservoirs, but the disparity is further exaggerated by addition of human induced water. Nonetheless, volumetric estimates appear robust, as they are supported by independently determined volumetric estimates from well production logs.

#### **4.4 Heavy noble gas enrichments**

Heavy noble gases (Kr and Xe) are preferentially adsorbed onto organic sediments, which, when liberated into the fluid system, provide an additional heavy noble gas isotopic signature (Torgersen and Kennedy, 1999; Zhou et al., 2005). Importantly, a pristine (or enriched) ASW signature is distinct from air-like compositions that have been identified in EOR injectate. There are clear heavy noble excesses in all Lost Hills samples relative to predictions from the solubility exchange model (Figures 3a,b), with the most pronounced enrichment occurring in the Cahn. As a result, solubility models – that utilize heavy noble gas concentrations – are unable to accurately predict volumetric oil-water ratios, and thus are not given in Table 1. Such heavy noble gas enrichments have previously been observed and discussed in detail (Torgersen and Kennedy, 1999; Zhou et al., 2005; Barry et al., 2016). In organic laden lithologies – such as shales and cherts – Kr and Xe preferentially absorb onto minerals relative to He, Ne and Ar (e.g., Podosek et al., 1981). Interestingly, the ratio of excess Kr/Xe is consistent (~3.1) throughout the entire Lost Hills system (i.e., in both Cahn and Diatomite wells – apart from well ‘LH-08’), varying by only 13%, even as total concentration of Xe varies by nearly 30%. In contrast, Ne varies by 95% due to the strong control of solubility partitioning. We surmise that as heavy noble gas enriched lithologies are buried to pressures and temperatures where hydrocarbons are generated, they release heavy

noble gases and enrich the deep meteoric formation waters (Hiyagon and Kennedy, 1992; Torgersen and Kennedy, 1999; Zhou et al., 2005). As oils form, this enriched noble gas signature partitions into the hydrocarbon phase. Alternatively, oil interaction with sediments may directly partition heavy noble gases into the oil phase, prior to interaction with the deep meteoric formation waters. These processes could operate sequentially or simultaneously and thus potentially complicate interpretation of measured hydrocarbon signatures. As is evident in Figure 5 and Table 3, heavy noble (Kr, Xe) gases are ubiquitously enriched in gases derived from oils and thus yield lower  $\frac{V_o}{V_w}$  than Ne and Ar derived estimates, however heavy noble gas  $\frac{V_o}{V_w}$  values are not considered robust and therefore are not given in Table 1.

#### **4.5 Assessing EOR efficiency using radiogenic helium**

Radiogenic noble gas isotopes (e.g.,  $^4\text{He}$ ,  $^{40}\text{Ar}$ ) are continuously produced in the Earth's crust and can yield temporal information about fluids. The concentration of radiogenic nuclides can be predicted by estimating flux rates in open fluid systems (Torgersen and Clarke, 1985; Torgersen and Ivey, 1985; Castro et al., 1998a; 1998b; Kulongoski et al., 2005; Zhou and Ballentine 2006; Torgersen 2010). In a closed hydrocarbon system, the well-constrained production rate of radiogenic helium can be combined with measured  $^4\text{He}$  concentrations to estimate isolation ages (Tolstikhin et al., 1996; 2017; Solomon et al., 1996; Cook et al., 1996; Ballentine et al., 2002; Holland et al., 2013; Barry et al., 2017). In such systems, parent radionuclide concentrations in source rocks are critical, as is the efficiency of radiogenic discharge from the rock into the fluid phase (Tolstikhin et al., 2010; 2017; Hunt et al., 2012; Lowenstern et al., 2014; Darrah et al. 2014; Barry et al., 2015).

These radiogenic nuclides are then transferred into the surrounding formation waters. The amount of ingrowth and subsequent exchange is a function of U and Th abundance in the source mineral, reservoir temperature, salinity, rock thickness, and porosity. When radiogenic laden waters come into contact with oil, the noble gases are

partitioned exactly the same way as atmospheric-derived noble gases in groundwater (e.g.,  $^{20}\text{Ne}$  and  $^{36}\text{Ar}$ ), according to known relative partitioning behavior (i.e.,  $K_i/K_j$ ).

Helium isotope results suggest that the Lost Hills oils are predominantly radiogenic, with He-isotope values ranging from 0.018-0.060  $R_A$  (Figure 2). Assuming binary mixing between crustal (0.007  $R_A$ ; Zhou and Ballentine, 2006) and sub-continental lithospheric mantle (SCLM) endmembers (6.1  $R_A$ ; Day et al., 2015), we calculate a crustal contribution between 99.1 to 99.8%. We choose SCLM as the mantle endmember as this is a continental setting, with no evidence of direct mantle contributions. If arc or canonical upper mantle values are assumed the calculation would not be significantly affected. This indicates that SCLM contributions to the system are typically less than 1%, consistent with observed mantle contributions near the San Andreas Fault (Kennedy et al., 1997), yet insignificant for modeling purposes. In Lost Hills, the two samples (i.e., 'LH-10' and 'LH-09') that tap the deeper Cahn member are the most radiogenic, with the lowest  $^3\text{He}/^4\text{He}$  and the highest  $^4\text{He}$  concentrations of all samples ( $>5.3 \times 10^{-6} \text{ cm}^3\text{STP}/\text{cm}^3$ ). In contrast, shallower samples from the Diatomite sequence have slightly less radiogenic signals, marked by broadly similar  $^3\text{He}/^4\text{He}$  values and lower  $^4\text{He}$  concentrations ( $< 2.9 \times 10^{-6} \text{ cm}^3\text{STP}/\text{cm}^3$ ). This disparity could be the result of the more U and Th rich nature of the deeper shales vs. the above-lying Diatomite sequences, which is supported by gamma log data. Alternatively, this could be due to the fact that older shales had more time to accumulate helium than the younger Diatomite. Despite the source of the  $^4\text{He}$  disparity, the observation that both Cahn samples – as well as a single Diatomite sample (LH-04) – have high  $^4\text{He}$  concentrations, helps establish the background radiogenic signal, prior to EOR. The lower  $^4\text{He}$  in the remaining Diatomite samples (including outlying sample LH-08), therefore suggests dilution has occurred, which is attributed to EOR-induced oil expulsion, which has effectively liberated noble gas barren oil, thus decreasing radiogenic  $^4\text{He}$  concentrations in oil by an order of magnitude or more. This observation provides strong ancillary support for the hypothesis that variations in the air-derived noble gases are the direct result of EOR, and

not simply a source difference between the deeper and shallower units. Due to moderately high helium concentrations in the EOR fluids themselves ( $\sim 3 \times 10^{-6}$  cm<sup>3</sup>STP/cm<sup>3</sup>), they cannot be easily evoked as the diluting agent and instead it must be the liberation of oil that in fact dilutes [<sup>4</sup>He] of oils in place.

The noble gas partitioning model presented in 4.3.1 indicates that Diatomite sample 'LH-04' has remained in production without EOR (Table 1) since the 1990s and thus has experienced the least secondary modification. As a result, we adopt a starting radiogenic <sup>4</sup>He starting composition of  $2.9 \times 10^{-6}$  cm<sup>3</sup>STP/cm<sup>3</sup> (i.e., the measured 'LH-04' value). By assuming simple dilution of <sup>4</sup>He-laden oil by mixing with <sup>4</sup>He-barren oil – liberated by injection – “EOR efficiency” can be estimated. EOR efficiency is thus defined as the amount of oil liberated by EOR, required to effectively dilute the noble gas signature of the oil in place. Using this approach, we estimate that EOR is between 48 and 97% effective in the Diatomite (Table 1). Not surprisingly, EOR efficiency is broadly inversely correlated with air-derived <sup>36</sup>Ar concentrations in gases, however the lack of a strong correlation suggests that each well has experienced a slightly different evolution due to local rock conditions (i.e., porosity, permeability), which also exert a control overall EOR efficiency. For example, sample 'LH-08', considered to be geologically compartmentalized, has the highest  $\frac{V_o}{V_w}$  as well as the highest EOR efficiency (97%) of all wells measured. If baseline noble gas values can be established for each well, then a quantitative measure of EOR efficiency can be assessed at any well that is sampled in the future.

## 5. Summary

In this study we present noble gas isotope and concentration data from casing gases derived from the Lost Hills oil field in the San Joaquin basin, California, as well as injectate data from the Fruitvale oil field. These data help establish both noble gas baseline values and EOR efficiencies. We show that within the Lost Hills petroleum system, the Cahn member remains largely pristine whereas the Diatomite member is

EOR modified with respect to noble gases. Only with knowledge of both source and secondary (EOR) features can a proper mass balance of the crustal system (i.e., both hydrocarbon and hydrogeological systems) be established, which will be critical for identifying and tracing the origin of any potential fugitive gases identified in nearby aquifers. For example, based on these results, we would predict lower air-derived noble gases and higher radiogenic  $^4\text{He}$  concentrations in any fugitive fluids derived from the Cahn reservoir and higher air-derived noble gases and lower radiogenic  $^4\text{He}$  concentrations if fluids derived from Diatomite. This study represents a successful calibration of noble gas technique for determining the effect of EOR on casing gas samples.

In summary, we have:

- 1) Established pristine hydrocarbon noble gas signatures in the Lost Hills (Cahn and Diatomite) casing gas samples and identified the processes controlling observed signatures; Cahn member samples retain the most pristine signatures that can be explained by a 2-phase solubility model.
- 2) Characterized injectate signatures in Fruitvale injectate gases; these values inform our models as to typical injectate signatures, which are air-like in nature with respect to noble gas elemental ratios.
- 3) Demonstrated the effect of injectate signature and effect in EOR wells; Diatomite samples which are disproportionately modified by EOR are marked by more air-like noble gas signatures, and radiogenic isotopes that appear heavily diluted by EOR.
- 4) Used a model to estimate EOR efficiency; Diatomite gas data reveal variability in EOR efficiency (48-97%) within this reservoir.

## Figure Captions

**Fig. 1a)** Map showing the location of where casing gases were collected in the Lost Hills Oil Field. Diatomite casing gases are shown as red triangles apart from sample LH-08, which is shown as an orange triangle (as it plots distinctly in several subsequent figures). Cahn casing gases are shown as green circles. The location of injection (EOR) wells are also shown as (small) yellow circles. This labeling scheme is used in all subsequent figures.  $^{20}\text{Ne}/^{36}\text{Ar}$  values are shown in parentheses next to each well name; the highest  $^{20}\text{Ne}/^{36}\text{Ar}$  values were measured in the densest regions of injection, whereas the lowest  $^{20}\text{Ne}/^{36}\text{Ar}$  values have limited injection in their vicinity.

**Fig. 1b)** Map of Injectate gases collected (large green squares) in the Fruitvale Oil Field and their proximity to all other injection wells in the region (small yellow circles).  $^{20}\text{Ne}/^{36}\text{Ar}$  values are shown in parentheses next to each well name. Notably, the two injectate samples that we collected had high (air-like)  $^{20}\text{Ne}/^{36}\text{Ar}$  values.

**Fig. 2a)** Helium isotopes ( $^3\text{He}/^4\text{He}$ ) vs. helium concentrations ( $10^{-6} \text{ cm}^3/\text{STPcm}^3$ ) in casing gases. Helium isotopes are normalized to air (where air =  $1R_A$ ). All He-isotope values are strongly radiogenic, however Cahn samples (green circles) have systematically higher He concentrations vs. Diatomite samples (red and orange triangles). Diatomite samples were subjected to substantially more EOR and are considered to have been diluted by liberation of  $^4\text{He}$ -free oil during EOR.  $1\sigma$  uncertainties are estimated to be less than 5% and this error is represented by the size of the symbols.

**Fig. 2b)** Helium isotopes ( $^3\text{He}/^4\text{He}$ ) vs.  $^4\text{He}/^{20}\text{Ne}$  in casing gases. Cahn samples (green circles) have systematically higher  $^4\text{He}/^{20}\text{Ne}$  values vs. Diatomite samples (red and orange triangles). Diatomite samples were subjected to substantially more EOR and are considered to have been diluted by liberation of  $^4\text{He}$ -free oil during EOR, thus lowering



the  $^4\text{He}/^{20}\text{Ne}$ .  $1\sigma$  uncertainties are estimated to be less than 5% and this error is represented by the size of the symbols.

**Fig. 3a)** Non-radiogenic ( $^{20}\text{Ne}/^{36}\text{Ar}$ ) vs.  $^{36}\text{Ar}$  ( $10^{-6} \text{ cm}^3/\text{STPcm}^3$ ) concentrations in casing gases and injectate gases. Cahn samples (green circles) retain the lowest (most pristine)  $^{20}\text{Ne}/^{36}\text{Ar}$  values. Diatomite samples (red and orange triangles) show higher (more air-like)  $^{20}\text{Ne}/^{36}\text{Ar}$  values. Injectate samples (green squares) are even more air-like with respect to their  $^{20}\text{Ne}/^{36}\text{Ar}$  values. The air value is shown as a blue diamond.  $1\sigma$  uncertainties are represented by the size of the symbols.

**Fig. 3b)** Spider diagram showing the average elemental ratios of air-derived noble gases for Cahn, Diatomite, Injectate, ASW and Air. Notably, Cahn samples plot furthest from Air and closest to '1<sup>st</sup>-bubble' solubility values. Diatomite samples plot intermediate between Air and '1<sup>st</sup>-bubble' solubility values. Critically, the Injectate (EOR) samples plot nearly identically to Air, showing the strong air-like signature associated with EOR.

**Fig. 4)** Non-radiogenic  $^{20}\text{Ne}$  ( $10^{-9} \text{ cm}^3/\text{STPcm}^3$ ) concentrations in casing gases vs. integrated gas-to-oil ratio (GOR), calculated from historic well-log data.  $1\sigma$  uncertainties are represented by the size of the symbols.

**Fig. 5)** Measured concentrations of non-radiogenic noble gases ( $^{20}\text{Ne}$ ,  $^{36}\text{Ar}$ ,  $^{84}\text{Kr}$ ) in oil vs. predicted (i.e., calculated using Eq. 2) values in oil at reservoir temperature and salinity. Notably, Cahn samples are in good agreement with each other and cluster near unity for Ne and Ar, suggesting a coherent  $\frac{V_o}{V_w}$  can be calculated. In contrast, Diatomite samples (apart from LH-08) have much higher measured vs. predicted noble gas concentrations in oil, which we attribute to addition of EOR.  $1\sigma$  uncertainties are represented by the size of the symbols.

**Fig. 6a)** Non-radiogenic ( $^{20}\text{Ne}/^{36}\text{Ar}$ ) vs.  $^{20}\text{Ne}$  concentrations in oil (mol/kg oil) in casing gases. The black line represents the predicted solubility behavior as an oil system evolves from right to left. At the starting point,  $\frac{V_o}{V_w}$  = infinitely small (1<sup>st</sup> bubble) and as the oil interacts with more water (during migration) the predicted value moves towards the left, where  $\frac{V_o}{V_w}$  = infinitely large. Colored lines are mixing trajectories with air-like gases, which are closely linked to EOR Injectate gases. At the intersection of the solubility prediction and mixing with air, the air-corrected  $\frac{V_o}{V_w}$  is given. 1 $\sigma$  uncertainties are represented by the size of the symbols.

**Fig. 6b)** Non-radiogenic ( $^{20}\text{Ne}/^{36}\text{Ar}$ ) vs.  $^{36}\text{Ar}$  concentrations in oil (mol/kg oil) in casing gases. The black line represents the predicted solubility behavior as an oil system evolves from right to left. At the starting point,  $\frac{V_o}{V_w}$  = infinitely small (1<sup>st</sup> bubble) and as the oil interacts with more water (during migration) the predicted value moves towards the left, where  $\frac{V_o}{V_w}$  = infinitely large. Colored lines are mixing trajectories with air-like gases, which are closely linked to EOR Injectate gases. At the intersection of the solubility prediction and mixing with air, the air-corrected  $\frac{V_o}{V_w}$  is given. 1 $\sigma$  uncertainties are represented by the size of the symbols.

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#### References:

Aregbe, Y., Valkiers, S., Mayer, K. and De Bièvre, P., 1996. Comparative isotopic measurements on xenon and krypton. *International journal of mass spectrometry and ion processes*, 153(1), L1-L5.

Ballentine, C. J., O'nions, R. K., Oxburgh, E. R., Horvath, F., Deak, J., 1991. Rare gas constraints on hydrocarbon accumulation, crustal degassing and groundwater flow in the Pannonian Basin. *Earth and Planetary Science Letters*, 105(1), 229-246.

Ballentine, C.J., O'Nions, R.K., and Coleman, M.L., 1996. A Magnus opus: Helium, neon, and argon isotopes in a North Sea oilfield. *Geochimica et Cosmochimica Acta*, 60(5), 831-849.

Ballentine, C.J., and Hall, C.M., 1999, Determining paleotemperature and other variables by using an error-weighted, nonlinear inversion of noble gas concentrations in water: *Geochimica et Cosmochimica Acta*, 63, 2315–2336.

Ballentine, C.J., and Burnard, P.G., 2002. Production, release and transport of noble gases in the continental crust. *Reviews in mineralogy and geochemistry*, 47(1), 481-538.

Barry, P.H., Hilton, D.R., Day, J.M., Pernet-Fisher, J.F., Howarth, G.H., Magna, T., Agashev, A.M., Pokhilenko, N.P., Pokhilenko, L.N. Taylor, L.A., 2015. Helium isotopic evidence for modification of the cratonic lithosphere during the Permo-Triassic Siberian flood basalt event. *Lithos*, 216, 73-80.

769

770 Barry, P.H., Lawson, M., Meurer, W.P., Warr, O., Mabry, J.C., Byrne, D.J.,  
 771 and Ballentine, C.J., 2016. Noble gases solubility models of hydrocarbon charge  
 772 mechanisms in the Sleipner Vest methane field. *Geochimica et Cosmochimica Acta*, 194,  
 773 291–309, doi:10.1016/j.gca.2016.08.021.

774 Barry, P.H., Lawson, M., Meurer, W.P., Danabalan, D., Byrne, D.J.,  
 775 Mabry, J.C., Ballentine, C.J. 2017. Determining fluid migration and isolation times in  
 776 multiphase crustal domains using noble gases. *Geology*, 45(9), 775-778 doi:  
 777 10.1130/G38900.1.

778 Batzle, M., and Wang, Z. 1992. Seismic properties of pore fluids. *Geophysics*,  
 779 57(11), 1396-1408.

780

781 Burnard P., 2013. *The Noble Gases as Geochemical Tracers*. Springer.

782

783 Byrne, D.J., Barry, P.H., Lawson, M. and Ballentine, C.J., 2018. Noble gases in  
 784 conventional and unconventional petroleum systems. Geological Society, London,  
 785 Special Publications, 468(1), 127-149.

786

787 Castro, M. C., Jambon, A., Marsily, G., Schlosser, P., 1998a. Noble gases as  
 788 natural tracers of water circulation in the Paris Basin: 1. Measurements and discussion  
 789 of their origin and mechanisms of vertical transport in the basin. *Water Resources*  
 790 *Research*, 34(10), 2443-2466.

791

792 Castro, M. C., Goblet, P., Ledoux, E., Violette, S., de Marsily, G., 1998b. Noble  
 793 gases as natural tracers of water circulation in the Paris Basin. 2. Calibration of a  
 794 groundwater flow model using noble gas isotope data. *Water Resources Research*,  
 795 34(10), 2467-2483.

796

Clarke, W. B., Jenkins, W. J., Top, Z., 1976. Determination of tritium by mass spectrometric measurement of  $^3\text{He}$ . The international journal of applied radiation and isotopes, 27(9), 515-522.

Cook, P.G., Solomon, D.K., Sanford, W.E., Busenberg, E., Plummer, L.N., Poreda, R.J., 1996. Inferring shallow groundwater flow in saprolite and fractured rock using environmental tracers. Water Resources Research, 32(6), 1501-1509.

Crovetto, R., Fernández-Prini, R., Japas, M.L., 1982. Solubilities of inert gases and methane in  $\text{H}_2\text{O}$  and in  $\text{D}_2\text{O}$  in the temperature range of 300 to 600 K. The Journal of Chemical Physics 76, 1077–1086. doi:10.1063/1.443074.

Darrah, T.H., Vengosh, A., Jackson, R.B., Warner, N.R., Poreda, R.J., 2014. Noble gases identify the mechanisms of fugitive gas contamination in drinking-water wells overlying the Marcellus and Barnett Shales. Proceedings of the National Academy of Sciences, 111(39), 14076-14081.

Darrah, T. H., Jackson, R. B., Vengosh, A., Warner, N. R., and Poreda, R. J., 2015. Noble gases: a new technique for fugitive gas investigation in groundwater. Ground water, 53(1), 23-28.

Day, J.M., Barry, P.H., Hilton, D.R., Burgess, R., Pearson, D.G., Taylor, L.A., 2015. The helium flux from the continents and ubiquity of low- $^3\text{He}/^4\text{He}$  recycled crust and lithosphere. Geochimica et Cosmochimica Acta, 153, 116-133.

Division of Oil Gas and Geothermal Resources (DOGGR), 1998. California Oil and Gas Fields, Volume 1. 507.

825 Gannon, R.S., Saraceno, J.F., Kulongoski, J.T., Barry, P.H., Tyne, R.L., Teunis, J.A.,  
826 Kraus, T., Hansen, A.H., and Qi, S.L., 2018, Produced water chemistry data for the Lost  
827 Hills, Fruitvale, and North and South Belridge study areas, Southern San Joaquin Valley,  
828 California, 2016-17: U.S. Geological Survey Data Release.  
829

830 Györe, D., Stuart, F. M., Gilfillan, S. M., and Waldron, S., 2015. Tracing injected  
831 CO<sub>2</sub> in the Cranfield enhanced oil recovery field (MS, USA) using He, Ne and Ar  
832 isotopes. *International Journal of Greenhouse Gas Control*, 42, 554-561.  
833

834 Harkness, J.S., Darrah, T.H., Warner, N.R., Whyte, C.J., Moore, M.T., Millot, R.,  
835 Kloppmann, W., Jackson, R.B., Vengosh, A., 2017. The geochemistry of naturally  
836 occurring methane and saline groundwater in an area of unconventional shale gas  
837 development. *Geochimica et Cosmochimica Acta*, 208, 302-334.  
838

839 Hiyagon, H., and Kennedy, B. M., 1992. Noble gases in CH<sub>4</sub>-rich gas fields,  
840 Alberta, Canada. *Geochimica et Cosmochimica Acta*, 56(4), 1569-1589.  
841

842 Holland, G., Lollar, B.S., Li, L., Lacrampe-Couloume, G., Slater, G.F.,  
843 and Ballentine, C.J., 2013. Deep fracture fluids isolated in the crust since the  
844 Precambrian era: *Nature*, 497, 357–360, doi:10.1038/nature12127.  
845

846 Hunt, A.G., Darrah, T.H., and Poreda, R.J., 2012. Determining the source and  
847 genetic fingerprint of natural gases using noble gas geochemistry: A northern  
848 Appalachian Basin case study. *AAPG Bulletin*, 96(10), 1785-1811.  
849

850 Kennedy, B.M., Kharaka, Y.K., Evans, W.C., Ellwood, A., DePaolo, D.J., Thordsen,  
851 J., Ambats, G. and Mariner, R.H., 1997. Mantle fluids in the San Andreas fault system,  
852 California. *Science*, 278(5341), 1278-1281.  
853

854

855 Kharaka, Y.K., and Specht, D.J., 1988. The solubility of noble gases in crude oil at  
856 25–100 C. *Applied geochemistry*, 3(2), 137-144.

857

858 Kulongoski, J.T., Hilton, D.R., Izbicki, J.A. 2005. Source and movement of helium  
859 in the eastern Morongo groundwater basin: the influence of regional tectonics on  
860 crustal and mantle helium fluxes. *Geochimica et Cosmochimica Acta*, 69, 3857-3872.

861 <http://dx.doi.org/10.1016/j.gca.2005.03.001>.

862

863 LaForce, T., Ennis-King, J., Boreham, C., and Paterson, L., 2014. Residual CO<sub>2</sub>  
864 saturation estimate using noble gas tracers in a single-well field test: The CO<sub>2</sub> CRC  
865 Otway project. *International Journal of Greenhouse Gas Control*, 26, 9-21.

866

867 Land, P.E., 1984. Lost Hills Oil Field: Pub. No. TR32, California Department of  
868 Conservation Division of Oil, Gas, and Geothermal Resources, Sacramento, CA. 3-17.

869

870 Lee, J.Y., Marti, K., Severinghaus, J.P., Kawamura, K., Yoo, H.S., Lee, J.B. and Kim,  
871 J.S., 2006. A redetermination of the isotopic abundances of atmospheric Ar. *Geochimica*  
872 *et Cosmochimica Acta*, 70(17), 4507-4512.

873

874 Long, J.C., Feinstein, L.C., Birkholzer, J., Jordan, P., Houseworth, J., Dobson, P.F.,  
875 Heberger, M., Gautier, D.L. and Donald Gautier, L.L.C., 2015. An Independent Scientific  
876 Assessment of Well Stimulation in California. Volume III, Case Studies of Hydraulic  
877 Fracturing and Acid Stimulations in Select Regions: Offshore, Monterey Formation, Los  
878 Angeles Basin and San Joaquin Basin.

879

880 Lowenstern, J.B., Evans, W.C., Bergfeld, D., and Hunt, A.G., 2014. Prodigious  
881 degassing of a billion years of accumulated radiogenic helium at Yellowstone. *Nature*,  
882 506(7488), 355-358.

Mount, J., Freeman, E., Lund, J., 2014. Water Use in California. Public Policy Institute of California (PPIC).

Pepin, R.O., Palma, R.L. and Schlutter, D.J., 2000. Noble gases in interplanetary dust particles, I: The excess helium - 3 problem and estimates of the relative fluxes of solar wind and solar energetic particles in interplanetary space. *Meteoritics and Planetary Science*, 35(3), 495-504.

Petrosky Jr, G.E. and Farshad, F.F., 1993, January. Pressure-volume-temperature correlations for Gulf of Mexico crude oils. In SPE annual technical conference and exhibition. Society of Petroleum Engineers.

Podosek, F.A., Bernatowicz, T.J. and Kramer, F.E., 1981. Adsorption of xenon and krypton on shales. *Geochimica et Cosmochimica Acta*, 45(12), 2401-2415.

Prinzhofer, A., 2013. Noble Gases in Oil and Gas Accumulations. In *The Noble Gases as Geochemical Tracers* (225-247). Springer Berlin Heidelberg.

Schoell, M., 1980. The hydrogen and carbon isotopic composition of methane from natural gases of various origins. *Geochimica et Cosmochimica Acta*, 44(5), 649-661.

Smith, S.P., and Kennedy, B.M., 1983. The solubility of noble gases in water and in NaCl brine. *Geochimica et Cosmochimica Acta* 47, 503–515. doi:10.1016/0016-7037(83)90273-9.

Sathaye, K.J., Larson, T.E. and Hesse, M.A., 2016. Noble gas fractionation during subsurface gas migration. *Earth and Planetary Science Letters*, 450, 1-9.



Solomon, D.K., Hunt, A. and Poreda, R.J., 1996. Source of radiogenic helium 4 in shallow aquifers: Implications for dating young groundwater. *Water Resources Research*, 32(6), 1805-1813.

Taylor, K.A., Fram, M.S., Landon, M.K., Kulongoski, J.T., Faunt, C.C., 2014. Oil, Gas, and Groundwater Quality in California—a discussion of issues relevant to monitoring the effects of well stimulation at regional scales. Prepared in Cooperation with the California State Water Resources Control Board.  
[ftp://ftp.consrv.ca.gov/pub/oil/SB4EIR/docs/Taylor\\_et\\_al\\_2014.pdf](ftp://ftp.consrv.ca.gov/pub/oil/SB4EIR/docs/Taylor_et_al_2014.pdf).

Torgersen, T., and Clarke, W.B., 1985. Helium accumulation in groundwater, I: An evaluation of sources and the continental flux of crustal  $^4\text{He}$  in the Great Artesian Basin, Australia. *Geochimica et Cosmochimica Acta*, 49(5), 1211-1218.

Torgersen, T., and Ivey, G.N., 1985. Helium accumulation in groundwater. II: A model for the accumulation of the crustal  $^4\text{He}$  degassing flux. *Geochimica et Cosmochimica Acta*, 49(11), 2445-2452.

Torgersen, T., and Kennedy, B.M., 1999. Air-Xe enrichments in Elk Hills oil field gases: role of water in migration and storage. *Earth and Planetary Science Letters*, 167(3), 239-253.

Torgersen, T., 2010. Continental degassing flux of  $^4\text{He}$  and its variability. *Geochemistry, Geophysics, Geosystems*, 11(6).

Tolstikhin, I., Lehmann, B.E., Loosli, H.H., Gautschi, A., 1996. Helium and argon isotopes in rocks, minerals, and related ground waters: A case study in northern Switzerland. *Geochimica et Cosmochimica Acta*, 60(9), 1497-1514.

941 Tolstikhin, I., Kamensky, I., Tarakanov, S., Kramers, J., Pekala, M., Skiba, V.,  
 942 Gannibal, M., Novikov, D., 2010. Noble gas isotope sites and mobility in mafic rocks and  
 943 olivine. *Geochimica et Cosmochimica Acta*, 74(4), 1436-1447.  
 944

945 Tolstikhin, I., Ballentine, C.J., Polyak, B.G., Prasolov, E.M., Likvadze, O.E., 2017.  
 946 The noble gas isotope record of hydrocarbon field formation time scales. *Chemical*  
 947 *Geology*, 41, 141-152.  
 948

949 Weiss, R.F., 1968, December. Piggyback sampler for dissolved gas studies on  
 950 sealed water samples. In *Deep Sea Research and Oceanographic Abstracts* (Vol. 15, No.  
 951 6, 695-699). Elsevier.  
 952

953 Wen, T., Castro, M. C., Ellis, B. R., Hall, C. M., Lohmann, K. C., 2015. Assessing  
 954 compositional variability and migration of natural gas in the Antrim Shale in the  
 955 Michigan Basin using noble gas geochemistry. *Chemical Geology*, 417, 356-370.  
 956

957 Wen, T., Castro, M. C., Nicot, J. P., Hall, C. M., Larson, T., Mickler, P., and Darvari,  
 958 R., 2016. Methane Sources and Migration Mechanisms in Shallow Groundwaters in  
 959 Parker and Hood Counties, Texas: A Heavy Noble Gas Analysis. *Environmental science*  
 960 *and technology*, 50(21), 12012-12021.  
 961

962 Zhou, D., Friedman, F., Kamath, J., Morea, M., 2002. Understanding Waterflood  
 963 in Lost Hills Diatomite. *SPE*, 75142, 15-17.  
 964

965 Zhou, Z., Ballentine, C. J., Kipfer, R., Schoell, M., Thibodeaux, S., 2005. Noble gas  
 966 tracing of groundwater/coalbed methane interaction in the San Juan Basin, USA.  
 967 *Geochimica et Cosmochimica Acta*, 69(23), 5413-5428.  
 968

969           Zhou, Z., and Ballentine, C.J., 2006.  $^4\text{He}$  dating of groundwater associated with  
970   hydrocarbon reservoirs. *Chemical Geology*, 226(3), 309-327.