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•UNIVERSITY OF OXFORD•

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•DEPARTMENT OF ENGINEERING SCIENCE•

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# UV-C Short-Range Wireless

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# Quantum Key Distribution

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## 1 **Abstract**

2       The rapid advancement of quantum technology increases the  
3 security risks of conventional encryption schemes, such as the  
4 Rivest-Shamir-Adleman (RSA) algorithm. Quantum key distribution  
5 (QKD) has emerged as a secure cryptographic solution, using  
6 Heisenberg's uncertainty principle and the no-cloning theorem to  
7 ensure unconditional security in key exchange. Significant progress  
8 has been made in fibre-based and satellite QKD systems, achieving  
9 long transmission distances and high secure key rates (SKR).  
10 However, short-range wireless QKD remains largely unexplored,  
11 primarily due to severe ambient light interference, which degrades  
12 system performance and limits practical deployment.

13       This thesis reports a short-range wireless QKD operating in the  
14 ultraviolet-C (UV-C) spectrum, utilizing the solar-blind region to  
15 mitigate ambient light interference. A high-sensitivity UV-C receiver  
16 was developed using a cooled silicon photomultiplier (SiPM),  
17 enabling a 1 Mbit/s optical wireless communication (OWC) link.  
18 Subsequently, ultralow photon flux links were demonstrated using a  
19 UV-C micro-LED and a single-photon detector (SPD), achieving a  
20 maximum data rate of 120 Mbit/s with a minimum transmitted  
21 photon number per bit of 7.7. Finally, the world's first wireless QKD  
22 system operating in the UV-C region was implemented, inferring an  
23 SKR exceeding 1 Mbit/s. Additionally, the impact of LED-based  
24 ambient light and sunlight interference was analysed, and  
25 mitigation strategies were proposed to enhance system robustness.

26       Overall, this work establishes the feasibility of UV-C wireless QKD  
27 and provides a foundation for future improvements in ambient light

1 rejection and real-world deployment.

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## 1 **Publications**

- 2 [1] **F. Liu**, J. Farmer, A. Schreier, G. Faulkner, H. Chun, W.  
3 Matthews, Z. Wang and D. O'Brien, "Ultra-sensitive UV solar-  
4 blind optical wireless communications with an SiPM," *Optics*  
5 *Letters*, vol. 48, no. 20, pp. 5387-5390, 2023/10/15 2023.
- 6 [2] **F. Liu**, J. Farmer, G. Faulkner, Z. Wang, J. Wang, D. O'Brien, E.  
7 Xie, J. Hill, J. Herrnsdorf, J. McKendry and M. Dawson, "10 Mbit/s  
8 UV Solar-Blind OWC at 30 Photons Per Bit," in *2024 Conference*  
9 *on Lasers and Electro-Optics (CLEO)*, 5-10 May 2024, pp. 1-2.
- 10 [3] **F. Liu**, J. Farmer, G. Faulkner, J. McKendry, E. Xie, Q. Zheng, Z.  
11 Wang, J. Wang, J. Hill, J. Herrnsdorf, M. Dawson and D. O'Brien,  
12 "Ultralow Photon Flux OWC Links Using UV-C Single-Photon  
13 Detection," *Laser & Photonics Reviews*, vol. n/a, no. n/a, p.  
14 2401804.
- 15 [4] **F. Liu** et al., "UV-C Wireless Quantum Key Distribution," *Nature*  
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# 1 **Chapter 1:**

## 2 **Introduction**

### 3 **1.1 OVERVIEW**

4       With the increasing demand for wireless capacity, optical  
5 wireless communication (OWC) is regarded as a promising  
6 technology for providing wireless access services in future wireless  
7 communication systems because of its broad bandwidth, low latency  
8 and lack of regulatory requirements [1-3]. In comparison to the 1  
9 Tbit/s peak data rate target for next-generation wireless systems  
10 (6G) [4], OWC systems with data rates of Tbit/s have already been  
11 demonstrated, with peak values reaching 100 Tbit/s [5-8]. However,  
12 with the rapid development of quantum computing, encryption and  
13 data security are increasing concerns for future secure wireless  
14 communications [9, 10]. In a typical communication system,  
15 traditional asymmetric encryption algorithms, such as the Rivest-  
16 Shamir-Adleman (RSA) algorithm, elliptic-curve cryptography (ECC)  
17 and Diffie-Hellman (DH) key exchange, are used to provide security  
18 [11-13]. Asymmetric encryption algorithms are based on the  
19 principle that solving certain mathematical problems, including large  
20 integer factorisation and discrete logarithms, is computationally  
21 infeasible for classical computers. However, these traditional  
22 encryption algorithms may not remain secure in the long term,  
23 given the development of quantum computing [14-19]. This is  
24 because Shor's algorithm provides a method for quantum  
25 computers to efficiently solve the large integer factorisation problem  
26 and the discrete logarithm problem [20]. Classical algorithms

1 require exponential time to solve these problems as the key size  
2 increases, but quantum computers running Shor's algorithm can  
3 solve them in polynomial time. This means that a quantum  
4 computer with sufficient scale could derive the private key from the  
5 public key, effectively breaking the encryption. As a result, the data  
6 encrypted or signed with these schemes could be intercepted by  
7 quantum-capable eavesdroppers.

8 To prevent information interception against quantum computing,  
9 quantum cryptography offers a new paradigm for secure  
10 communications even in the presence of quantum computers [21]. It  
11 primarily includes quantum key distribution (QKD) and post-  
12 quantum cryptography (PQC), two complementary approaches that  
13 aim to ensure long-term security in the quantum era [21-27]. QKD  
14 provides theoretically full security to generate secure keys and  
15 encrypt data based on fundamental quantum mechanics, including  
16 Heisenberg's uncertainty principle and the no-cloning theorem [21-  
17 24]. Heisenberg's uncertainty principle states that accurately  
18 measuring a pair of conjugate properties, such as the position and  
19 momentum of a photon, simultaneously is impossible [28, 29]. This  
20 is because measuring one property of one particle inevitably  
21 perturbs it and thus makes the measurement of another property  
22 inaccurate. The non-cloning theorem states that exactly replicating  
23 arbitrary unknown quantum states is impossible [30-34]. Thus, in a  
24 QKD system, the eavesdropper (Eve) can only attempt to eavesdrop  
25 and measure the quantum states directly (she cannot accurately  
26 measure and replicate states) from the transmitter (Alice) and  
27 transmit these states to the receiver (Bob). Thus, eavesdropping will  
28 change the quantum states, allowing the transmitter and receiver to  
29 detect eavesdropping by a change in the link error rate. PQC resists

1 attacks from quantum computers based on mathematical problems  
2 believed to be hard even for quantum computers, such as lattice-,  
3 code-, hash-, and multivariate-based problems [35-38]. Unlike QKD,  
4 PQC can be deployed using conventional hardware without requiring  
5 quantum channels, making it more accessible to integrate into  
6 existing communication networks. PQC is especially suitable for  
7 applications such as digital signatures and public-key encryption  
8 and is expected to operate with QKD to create robust and flexible  
9 quantum-secure communication systems.

10 This thesis focuses on the development of QKD. The first QKD  
11 protocol was the BB84 protocol which was proposed by C. H.  
12 Bennett and G. Brassard in 1984 [23]. In the past approximately 4  
13 decades, different types of QKD protocols have been developed,  
14 including Discrete-Variable QKD (DV-QKD), Coherent-One Way QKD  
15 (COW-QKD), Continuous-Variable QKD (CV-QKD), Measurement-  
16 Device-Independent QKD (MDI-QKD) and Twin-Field QKD (TF-QKD)  
17 [23, 39-41]. For QKD experimental demonstrations, a maximum  
18 distance of more than 1000 km and a secure key rate (SKR) of over  
19 110 Mbit/s have been achieved [42, 43]. However, previous QKD  
20 research has mainly focused on fibre-based or satellite systems [42-  
21 45], and there are a very limited number of short-range wireless  
22 QKD demonstrations [46-50].

23 Short-range wireless QKD has the potential to provide  
24 encryption services for wireless information in mobile devices in the  
25 future, such as smartphones and wireless payments [47]. However,  
26 short-range wireless QKD faces the challenge of severe ambient  
27 light interference, mainly due to sunlight [46-50]. According to the  
28 sunlight spectrum near the ground [51], there exists a “solar-blind”  
29 region in the ultraviolet-C (UV-C) band (less than 280 nm). The

1 sunlight irradiance in this region is extremely low because of  
2 absorption by the ozone layer in the atmosphere. Therefore, a short-  
3 range wireless QKD system operating in daylight in the UV-C region  
4 is feasible.

5 However, available devices operating in the UV-C region are  
6 very limited, especially high-bandwidth UV-C transmitters and a  
7 high-sensitivity UV-C receiver. To address the challenges from  
8 individual components to the complete QKD system, this thesis  
9 details the development of the UV-C transmitter and receiver, and a  
10 demonstration of a UV-C wireless QKD system. The layout of this  
11 thesis is outlined below.

## 12 **1.2 LAYOUT OF THE THESIS**

13 In Chapter 2, the theory and background of the QKD are  
14 provided. The structure and scheme of QKD systems are presented.  
15 Then, different QKD protocols are listed and compared, followed by  
16 summarized wireless QKD demonstrations. Additionally, the  
17 potential of operating wireless QKD in the UV-C region is analysed.

18 In Chapter 3, ultra-sensitive UV solar-blind OWC links based on a  
19 commercial UV-C LED and a high-sensitivity receiver are reported.  
20 The receiver was designed based on a cooled silicon photomultiplier  
21 (SiPM). A 1 Mbit/s UV-C OWC link was demonstrated, with an  
22 average received photon number per bit of approximately 40. The  
23 successful design of the high-sensitivity receiver paves the way for  
24 the receiver design for the final UV-C wireless QKD systems. Results  
25 are published in [52] (listed below).

26 **F. Liu**, J. Farmer, A. Schreier, G. Faulkner, H. Chun, W. Matthews, Z.  
27 Wang and D. O'Brien, "Ultra-sensitive UV solar-blind optical wireless

1 communications with an SiPM," *Optics Letters*, vol. 48, no. 20, pp.  
2 5387-5390, 2023/10/15 2023.

3 In Chapter 4, ultralow photon flux links are reported. These are  
4 based on a UV-C micro-LED and a UV-C single-photon detector  
5 (SPD). The UV-C SPD achieved a lower dark count rate (DCR)  
6 compared to the receiver described in Chapter 3, due to a lower  
7 temperature and a smaller active area. A maximum data rate of 120  
8 Mbit/s and a minimum transmitted average photon number per bit  
9 of 7.7 was demonstrated. Moreover, a 500 MHz repetition rate  
10 single-photon link was demonstrated, showing the potential of  
11 building a UV-C wireless QKD. Results are reported in [53, 54] (listed  
12 below):

13 **F. Liu**, J. Farmer, G. Faulkner, Z. Wang, J. Wang, D. O'Brien, E. Xie, J.  
14 Hill, J. Herrnsdorf, J. McKendry and M. Dawson, "10 Mbit/s UV Solar-  
15 Blind OWC at 30 Photons Per Bit," in *2024 Conference on Lasers and*  
16 *Electro-Optics (CLEO)*, 5-10 May 2024, pp. 1-2.

17 **F. Liu**, J. Farmer, G. Faulkner, J. McKendry, E. Xie, Q. Zheng, Z.  
18 Wang, J. Wang, J. Hill, J. Herrnsdorf, M. Dawson and D. O'Brien,  
19 "Ultralow Photon Flux OWC Links Using UV-C Single-Photon  
20 Detection," *Laser & Photonics Reviews*, vol. n/a, no. n/a, p. 2401804.

21 Chapter 5 describes the design and demonstration of a UV-C  
22 wireless B92 QKD system, which to the best of our knowledge is the  
23 first demonstration operating in the UV-C region. The modelling of  
24 the system is presented and then details of the demonstration are  
25 reported. The system was operated with an inferred SKR of over 1  
26 Mbit/s. Details of strategies to reject ambient light from LEDs and  
27 sunlight are proposed, together with a detailed analysis. The  
28 corresponding publications are in preparation and listed below:

---

1 **F. Liu** et al., "UV-C Wireless Quantum Key Distribution," *Nature*  
2 *Communications* (Submitted).

3 In Chapter 6, the contributions of this thesis are summarized,  
4 and conclusions are drawn. Future work on the potential  
5 improvement of ultralow photon flux links and QKD links is  
6 discussed.

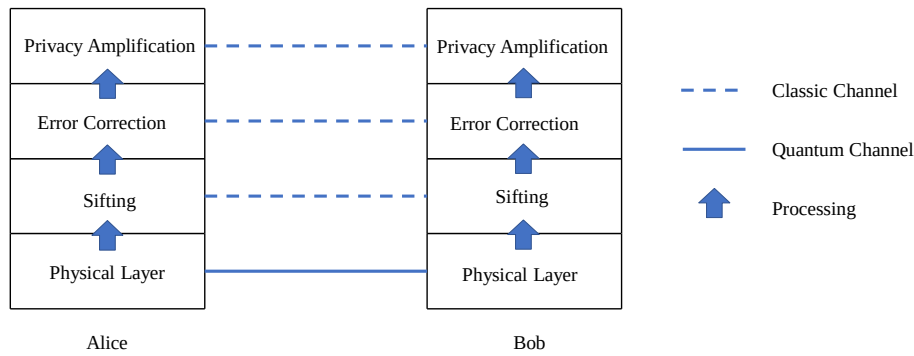
# 1 **Chapter 2:**

## 2 **Literature Review**

3 This Chapter begins by introducing the architecture and  
4 operation of QKD systems. Then, different protocols are analysed  
5 and compared, highlighting their respective advantages and  
6 disadvantages in terms of performance and implementation. State-  
7 of-the-art wireless QKD demonstrations are reviewed and  
8 summarized, identifying UV-C regions that remain underexplored.  
9 Finally, the potential of pursuing a UV-C wireless QKD system is  
10 discussed.

### 11 **2.1 QKD ARCHITECTURE**

12 A block diagram of a QKD scheme, consisting of a transmitter  
13 (Alice), a receiver (Bob) and channels between them, is illustrated in  
14 Fig. 2.1. The goal of a QKD system is to generate secure keys that  
15 can be used to encrypt information data. Photons are sent from  
16 Alice to Bob in the physical layer. Sifted keys are then obtained by  
17 matching compatible transmitted and received photons. Error  
18 correction removes errors in the sifted keys, and privacy  
19 amplification eliminates the information intercepted by an  
20 eavesdropper (Eve). A detailed description of the process is  
21 provided in the following sections.



1  
2 Figure 2.1 QKD scheme block diagram

### 3 2.1.1 Physical Layer

4 Alice uses single-photon sources to generate and transmit  
5 photons. However, ideal single-photon sources are unavailable, and  
6 weak light sources serve as an alternative. A weak light source is  
7 typically implemented as a light source with an attenuator to reduce  
8 the average number of emitted photons in each qubit to less than  
9 one. The transmitted photons are randomly encoded in one of  
10 several bases. The type of basis varies across different QKD  
11 systems, and the use of polarization, phase, or orbital angular  
12 momentum (OAM) has all been demonstrated [23, 55, 56]. Bob  
13 implements a corresponding detection channel for each basis. One  
14 of these bases is randomly selected to decode the received photons  
15 using SPDs. An SPD-based receiver typically consists of a high-  
16 sensitivity photodetector and subsequent amplification.

17 The channel between Alice and Bob can be classified into either  
18 a guided (predominantly fibre-based) or free-space channel. The  
19 channel loss in fibre-based QKD comes from material absorption,  
20 bending, fibre splices and fibre connections, while geometrical loss  
21 and atmospheric attenuation are the primary contributors in free  
22 space QKD [57]. Beyond channel loss, fibre-based systems are also  
23 affected by noise sources, mainly spontaneous Raman scattering

1 (SRS) induced by co-propagating classical signals [58, 59]. SRS is  
2 particularly prominent in wavelength-division multiplexed systems  
3 where quantum and classical signals share the same fibre. Free-  
4 space systems, on the other hand, experience noise from ambient  
5 light, which increases background noise at the receiver, as well as  
6 beam wandering and wavefront distortion caused by atmospheric  
7 turbulence [60-62]. Thus, fibre-based QKD provides stable and  
8 environmentally isolated links suitable for metropolitan networks,  
9 while free-space QKD offers greater flexibility for infrastructure-  
10 limited environments, such as ground-to-satellite or mobile links.

### 11 **2.1.2 Sifting**

12 Bob sends his chosen decoding bases to Alice, who compares  
13 them with the encoding bases. Alice then informs Bob of the  
14 incompatible bases, and they discard the corresponding qubits. The  
15 remaining qubits with compatible bases are referred to as the sifted  
16 keys.

### 17 **2.1.3 Error Correction**

18 The sifted keys still have residual errors because of the  
19 imperfection in data transmission, detection and processing. Thus,  
20 error correction, also known as information reconciliation in QKD, is  
21 carried out to make sure that the sifted keys in Alice and Bob are the  
22 same [63, 64]. Unlike conventional communications, where error  
23 correction aims to recover the information from noisy data, QKD  
24 transmits random numbers rather than live data. As a result, QKD  
25 error correction focuses on reconciling minor discrepancies between  
26 the two parties' sifted keys, rather than reconstructing the  
27 information data. Low-density parity code check (LDPC) codes are  
28 typically applied with a small quantum bit error rate ( $QBER \leq 5\%$ ),

1 as their performance theoretically approaches the Shannon limit  
2 [65]. When the QBER is relatively high ( $QBER > 5\%$ ), alternative  
3 methods, such as segmented error correction and cascade  
4 protocols, can be employed, or the corresponding sifted keys can be  
5 discarded to ensure security [66, 67]. Removing these errors does  
6 not affect the overall randomness of keys, ensuring their security.  
7 The keys after error correction are termed corrected keys.

#### 8 **2.1.4 Privacy Amplification**

9       Guaranteeing that Alice and Bob have the same keys is not  
10 enough, because the influence of the eavesdropper may still exist.  
11 However, Alice and Bob cannot identify which qubits have been  
12 intercepted by Eve in the corrected keys. Privacy amplification  
13 addresses this issue by reducing Eve's knowledge to a negligible  
14 level using universal hash functions [68-74]. By applying a universal  
15 hash function, corrected keys with and without eavesdropping will  
16 produce entirely different outputs, due to the sensitivity of the hash  
17 function. Thus, the parts possibly known by Eve are effectively  
18 removed. This means that Alice and Bob compress the key length to  
19 ensure that any partial knowledge Eve may have is eliminated.

#### 20 **2.1.5 Attacks**

21       The number of photons (emitted by a weak light source) in each  
22 qubit follows a Poisson distribution. As a result, a small proportion of  
23 qubits contain more than one photon, which makes a QKD system  
24 more vulnerable to attacks, such as a photon-number-splitting (PNS)  
25 attack [75-77]. In a PNS attack, Eve allows one photon to pass  
26 through to Bob and intercepts all other photons in each qubit. One  
27 method to reduce the impact of PNS attacks is using decoy states  
28 [78]. Eavesdropping on the communication link can be discovered

1 by a rising QBER [78, 79].

## 2 **2.2 QKD PROTOCOLS**

### 3 **2.2.1 DV-QKD Protocols**

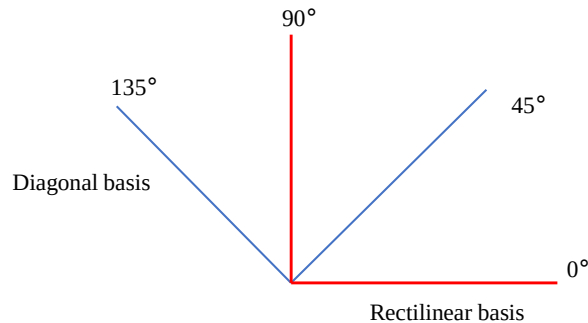
4 This section introduces different DV-QKD protocols, including  
5 BB84, B92, six-state, SARG04, KMB09, T12 and S13, along with their  
6 respective advantages and disadvantages. BB84 is the most widely  
7 studied and implemented, offering strong security but lower  
8 efficiency in noisy environments. B92 simplifies the protocol by  
9 using fewer states, but it is more sensitive to noise. Six-State and  
10 SARG04 enhance security compared to BB84 but at the cost of  
11 increased hardware complexity and lower efficiency. KMB09 and T12  
12 introduce advanced techniques to improve noise resistance, in cost  
13 of requiring more complex implementations. Additionally, T12 offers  
14 improved transmission efficiency.

#### 15 2.2.1.1 BB84 Protocol

16 The BB84 protocol is the first QKD protocol, which was proposed  
17 in 1984 by C. H. Bennett and G. Brassard [23]. BB84 is  
18 straightforward to implement, with many experimental validations  
19 [43, 80-87]. However, BB84 has drawbacks in high-noise and long-  
20 distance applications, where error rate increases and key generation  
21 efficiency decreases.

22 In this protocol, Alice first prepares two random bit strings  $a$  and  
23  $b$ . The first string  $a$  represents the bit values (0 or 1) that Alice  
24 wants to send to Bob. The second string  $b$  is applied to encode the  
25 bit in two different bases (the  $X$ -basis and  $Y$ -basis), and each basis  
26 contains two states representing 0 and 1, respectively. The  
27 description of these bases is illustrated in Fig. 2.2. The rectilinear

1 basis includes a horizontal state ( $0^\circ$ ) and a vertical state ( $90^\circ$ ), and  
2 the diagonal basis contains a diagonal state ( $45^\circ$ ) and an anti-  
3 diagonal state ( $135^\circ$ ).



4

5 **Figure 2.2 Bases used in BB84**

6 The encoded bits are called qubits and are transmitted to Bob  
7 via quantum channels. Bob chooses the bases randomly to decode  
8 the received qubits and obtains a measured string. Then, the sifting,  
9 error correction and privacy amplification are performed to generate  
10 the final secure keys. The process of BB84 can be summarized as  
11 follows:

12 • Alice chooses random bases to encode a string of generated bits  
13 to create qubits. Then, the qubits are sent to Bob via the quantum  
14 channel.

15 • Bob receives qubits and applies a random basis to measure each  
16 qubit. The measurement result is called the raw key.

17 • Bob sends his chosen bases to Alice, and Alice compares them  
18 with the bases she used. Then, Alice tells Bob which of the bases  
19 he used are incorrect. The corresponding bits are discarded, and  
20 the others are reserved. The reserved bits are known as the sifted  
21 key.

22 • Bob sends part of his sifted keys to Alice, and Alice calculates

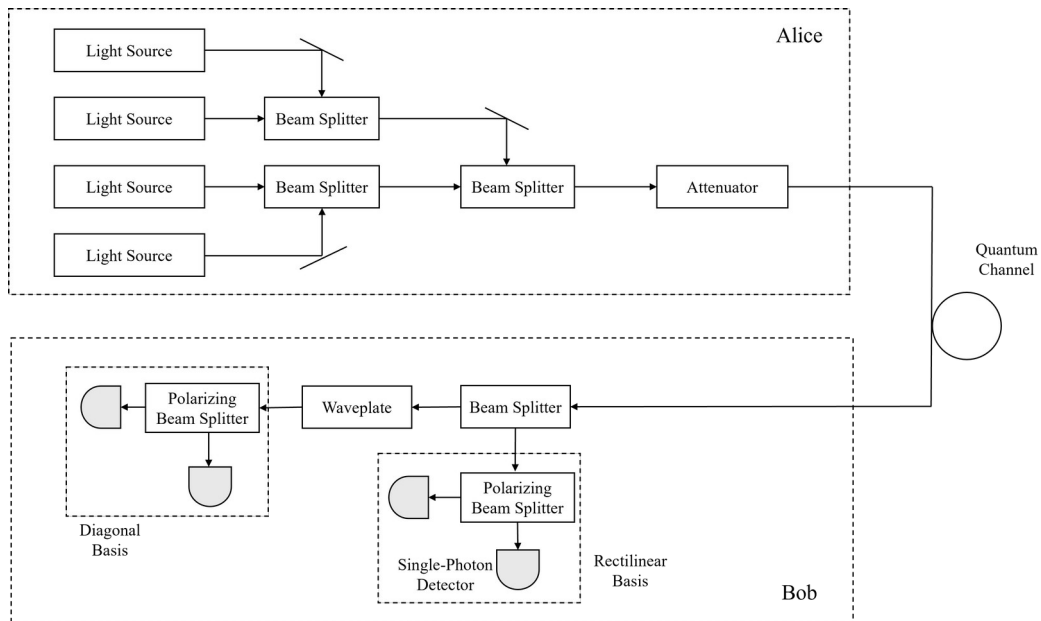
1 the error rate, which is called QBER. Then, error correction,  
2 typically using an LDPC code, is carried out to remove the errors in  
3 the sifted keys. The generated bits are called the corrected keys.

4 • Based on the calculated QBER, privacy amplification is  
5 undertaken using universal hash functions to remove the  
6 eavesdropped information by Eve and obtain final secure keys.

7 A block diagram of a typical polarization-based BB84 QKD  
8 system is shown in Fig. 2.3 [88]. There are four light sources in Alice,  
9 which generate photons polarized at  $0^\circ$ ,  $45^\circ$ ,  $90^\circ$  and  $135^\circ$ ,  
10 respectively. Only one light source is randomly chosen to be  
11 triggered at the same time. Then, the generated photons are  
12 combined by three beam splitters (BSs). An attenuator is installed to  
13 reduce the number of photons in each qubit to less than one. The  
14 photons emitted from Alice travel through the quantum channel to  
15 Bob. A BS randomly allocates received photons to two polarizing  
16 beam splitters (PBSs), with one waveplate in front of one of these  
17 two PBSs. Then, photons are detected by four SPDs (each SPD  
18 detects photons in one of four polarization states) behind PBSs.

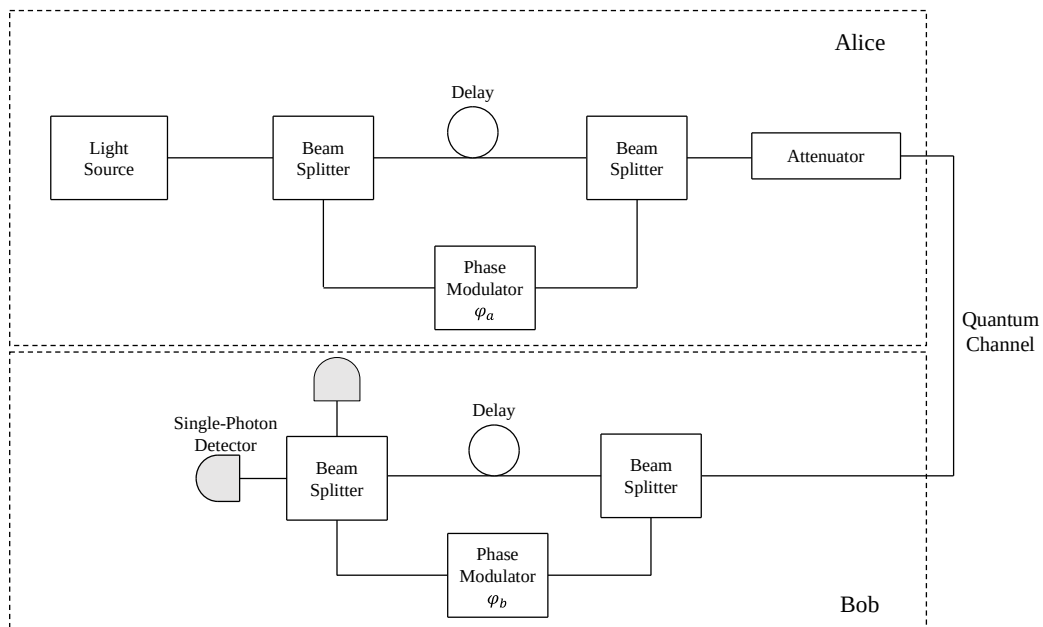
19 Apart from the polarization-based BB84 system, a phase-based  
20 BB84 system is an alternative way to realize the protocol in practice  
21 [89, 90]. A corresponding block diagram is shown in Fig. 2.4. In  
22 Alice, a BS splits light from a source into two paths: a short arm with  
23 a delay module and a long arm with a phase modulator  $\varphi_a$ . The light  
24 through two paths is combined by a BS and attenuated for  
25 transmission. In Bob, by using a BS, the photons passing Alice's long  
26 arm go into Bob's short (L-S route) arm and vice versa (S-L route).  
27 The delay modules in Alice and Bob ensure that the two routes are  
28 the same length. The photons sent by Alice are detected by only

- 1 one of the two SPDs with a probability based on the phase
- 2 difference  $\Delta \varphi = \varphi_a - \varphi_b$ .



3

4 Figure 2.3 Block diagram of a typical polarization-based BB84 QKD  
5 system



6

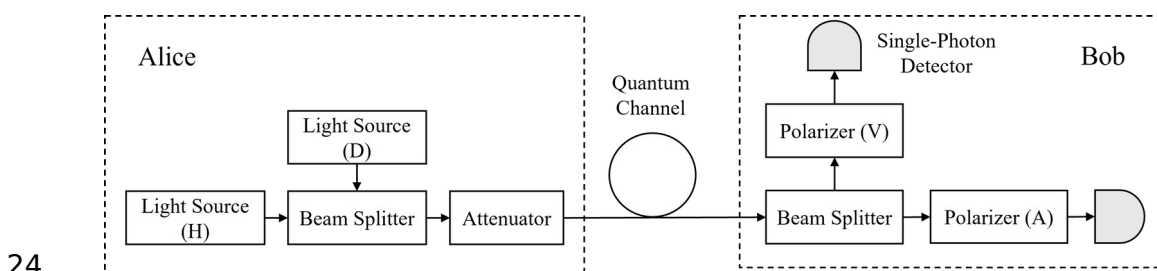
7 Figure 2.4 Block diagram of a typical phase-based BB84 QKD  
8 system

9 2.2.1.2 B92 Protocol

2  
3

1 In 1992, Bennett found that only two nonorthogonal states were  
 2 enough for quantum communications, rather than using four states  
 3 in the BB84 protocol [90]. As a result, B92 is more resource-efficient,  
 4 requiring fewer components such as half light sources and half  
 5 SPDs, though the key generation rate is reduced by approximately  
 6 half. Moreover, the impact of noise in B92 is greater, which makes it  
 7 harder to implement in practice. This is because B92 relies on  
 8 unambiguous state measurement, where each SPD detection  
 9 directly corresponds to a key bit. In noisy channels, the  
 10 nonorthogonal states become harder to distinguish, thereby  
 11 increasing the error rate. By contrast, BB84 benefits from basis  
 12 comparison and redundancy in state choices, which enhances its  
 13 noise resistance. Thus, B92 is more sensitive to noise, leading to  
 14 higher QBER in practice.

15 A block diagram of a typical polarization-based B92 QKD system  
 16 is shown in Fig. 2.5. In Alice, there are only two light sources,  
 17 generating photons polarized at  $0^\circ$  (H, horizontal) and  $45^\circ$  (D,  
 18 diagonal), respectively. Then, the generated photons are combined  
 19 by a BS, and an attenuator reduces the number of photons in each  
 20 qubit to less than one. In Bob, a BS randomly allocates received  
 21 photons to two polarizers, selecting photons with specific  
 22 polarizations (V, vertical or A, anti-diagonal). Then, photons are  
 23 detected by two SPDs.



1 Figure 2.5 Block diagram of a polarization-based typical B92 QKD  
 2 system

3 The working principle of the B92 protocol is shown in Table 2.1.  
 4 Alice first randomly chooses one of the horizontal (“0”) and diagonal  
 5 (“1”) states to generate a qubit sequence. Then, the generated  
 6 qubits are sent to Bob via the quantum channel. To measure one  
 7 qubit, Bob randomly chooses a state (vertical or anti-diagonal).  
 8 When the qubit is received and the measurement result is  $(\uparrow)$ ,  
 9 Bob knows the qubit sent by Alice is  $\rightarrow$  ( ). Sifting is omitted and  
 10 Bob keeps all received bits.

11 Table 2.1 Working principle of the B92 protocol

Bit sequences in Alice	0	1	1	0	0	1
States in Alice	$\rightarrow$			$\rightarrow$	$\rightarrow$	
States in Bob	$\uparrow$	$\uparrow$	$\uparrow$			
Reserved or not	N	Y (50%)	N (50%)	Y (50%)	N (50%)	N
Measurement result		1		0		

12 2.2.1.3 Six-State Protocol

13 The six-state protocol improves upon BB84 by using six states  
 14 for encoding and decoding [91, 92]. Compared to BB84, the two  
 15 additional bases are a right-circular state and a left-circular state.  
 16 This protocol provides better noise resistance and reduces the  
 17 alignment complexity[46]. However, the key generation rate is  
 18 lower.

19 2.2.1.4 SARG04 Protocol

20 The SARG04 protocol is another version of the BB84 protocol  
 21 [77, 93, 94], which provides better robustness against channel loss

1 and eavesdropping and can be compatible with BB84 hardware.  
2 However, SARG04 is more complex in terms of protocol design and  
3 implementation.

4 The SARG04 protocol utilizes a different sifting procedure from  
5 the BB84 protocol. Instead of revealing the bases used to encode  
6 qubits, Alice selects one state from the rectilinear basis and one  
7 state from the diagonal basis and sends them to Bob. For Bob, his  
8 goal is to distinguish which state is the correct one. Based on his  
9 measurement, he checks which states his result is consistent with. If  
10 the result is consistent with both states, the bit will be discarded  
11 because Bob cannot know which state is used by Alice. If one of the  
12 states is inconsistent with Bob's measurement result, he will keep  
13 this bit and extract the correct sifting key. For example, Alice  
14 generates qubits in rectilinear basis, including a horizontal state  
15 ("0") and a vertical state ("1"), and a diagonal basis, including a  
16 diagonal state ("0") and an anti-diagonal state ("1"). Suppose that a  
17 horizontal state is sent from Alice. Then, she publicly announces two  
18 "0" states (horizontal state and diagonal state). Depending on the  
19 basis Bob applies to measure the qubit, two situations occur:

20 • If Bob uses the rectilinear basis to measure the received qubit, the  
21 measurement result must be 0. This result is obviously consistent  
22 with the transmitted state. However, if the transmitted state is the  
23 diagonal state, the measurement result can also be the horizontal  
24 state with a probability of 50 %. Therefore, Bob cannot know which  
25 state is sent by Alice. This bit is invalid and discarded.

26 • If Bob uses the diagonal basis to measure the received qubit, he  
27 can obtain a diagonal state or an anti-diagonal state, each with a  
28 probability of 50 %. Thus, if the outcome is the diagonal state, it can

1 be obtained by transmitting either the horizontal state or the  
 2 diagonal state. However, when the outcome is the anti-diagonal  
 3 state, Bob can be certain that the correct state is the horizontal  
 4 state because the anti-diagonal state can never be obtained from  
 5 the diagonal state. Therefore, Bob announces that this bit is valid  
 6 and extracts a sifting key.

#### 7 2.2.1.5 KMB09 Protocol

8 The KMB09 protocol builds on BB84 by introducing decoy states  
 9 to mitigate the impact of PNS attacks and improve performance in  
 10 high channel loss scenarios. However, the use of decoy states adds  
 11 complexity to implementation and post-processing [95].

#### 12 2.2.1.6 T12 Protocol

13 The T12 protocol is designed based on BB84 and adjusts the  
 14 usage rates of two bases [96]. Thus, the key generation rate is  
 15 improved, but the device calibration and system implementation are  
 16 more complex.

Table 2.2 A survey of DV-QKD protocols

Protoc ol	Year	Advantage	Disadvantage	Reference
BB84	1984	Easy to understand and widely implement	Limited performance with high-loss and long- distance	C. H. Bennett and G. Brassard [23]
B92	1992	Resource-efficient	Low key generation rate and sensitive to noise	C. H. Bennett [90]
Six- state	1998 &199 9	Robust against noise and mitigate alignment complexity	Low key generation rate	D. Bruß [91] & H. Bechmann- Pasquinucci and N. Gisin [92]

SARG04	2004	Robust against channel loss and eavesdropping, and compatible with BB84 hardware	Complex protocol design and implementation	V. Scarani, A. Acín, G. Ribordy and N. Gisin [77]
KMB09	2009	Mitigate PNS attacks and robust against path loss	Complex implementation and post-processing	M. Mubashir Khan, M. Murphy and A. Beige [95]
T12	2013	High key generation rate	Complex device calibration and system implementation	M. Lucamarini <i>et al.</i> [96]

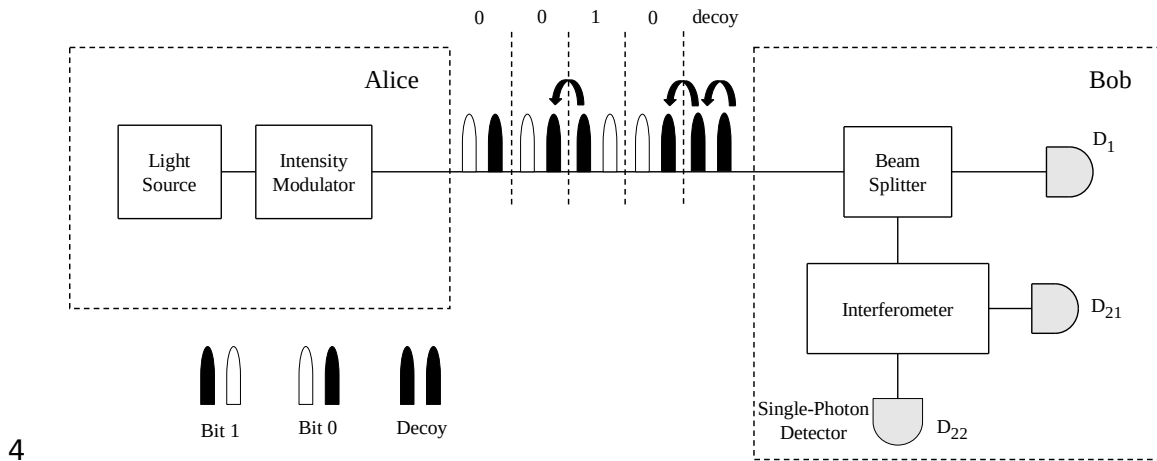
1 In the BB84 protocol, the number of qubits encoded by each  
2 basis is identical (50 % on the  $X$  basis and 50 % on the  $Y$  basis).  
3 Then, Bob chooses one of the two bases randomly to measure every  
4 qubit, so the possibility of getting the correct outcome is 50 %.  
5 Thus, only  $50\% \times 50\% + 50\% \times 50\% = 50\%$  of initial qubits can be  
6 converted into raw keys. However, the numbers of qubits encoded  
7 and decoded in two bases are different in the T12 protocol. For  
8 instance, the usage rates of  $Y$  basis and  $X$  basis in both Alice and  
9 Bob are 80 % and 20 %, respectively. The conversion ratio from  
10 initial qubits to raw keys is  $80\% \times 80\% + 20\% \times 20\% = 68\%$  which is 18 %  
11 higher than standard BB84 protocol. Further increase in the number  
12 of qubits in the  $Y$  basis can further increase the length of raw keys.  
13  $1/16$  of total qubits for the  $X$  basis and  $15/16$  for the  $Y$  basis were  
14 applied in [96] and nearly 88.3 % of initial qubits can be used to  
15 generate raw keys.

16 A summary of the above DV-QKD protocols is shown in Table 2.2.

### 17 2.2.2 COW-QKD Protocol

18 In 2004, D. Stucki, N. Brunner, N. Gisin, V. Scarani, and H.  
19 Zbinden proposed the COW-QKD protocol [97, 98] and a wireless  
20 QKD demonstration based on COW-QKD has been demonstrated

1 [48]. The COW-QKD protocol obtains secure keys based on  
 2 measuring the arrival time of qubits [99]. A typical block diagram of  
 3 the COW-QKD system is illustrated in Fig. 2.6.



6 Figure 2.6 A typical block diagram of a COW-QKD system

7 Alice prepares non-empty pulses with an average photon  
 8 number  $\mu$  or empty pulses by using a coherent light source and an  
 9 intensity modulator. Each bit "1" or "0" contains an empty pulse and  
 10 a non-empty pulse:  $|1_k\rangle = |\sqrt{\mu}\rangle_{2k-1} - |0\rangle_{2k}$ ,  $|0_k\rangle = |0\rangle_{2k-1} - |\sqrt{\mu}\rangle_{2k}$ , where  $k$  is  
 11 the bit number. Bob obtains the bit value by measuring the arrival  
 12 time of non-empty pulses. Most of the pulses go through the BS in  
 13 Bob to  $D_1$  to generate raw keys and the rest are reflected to the  
 14 interferometer to check the coherence both within a bit and across  
 15 bit separations. By adding decoy states, checking coherence can  
 16 detect a PNS attack from Eve.

### 17 2.2.3 CV-QKD Protocols

18 There are two kinds of CV-QKD protocols: the squeezed-state  
 19 Gaussian protocol [100] and the coherent-state Gaussian protocol  
 [39]. This section focuses on the coherent-state Gaussian protocol

1 because it is easier to implement and more robust against noise and  
2 loss, making it more suitable for practical applications.

3 The coherent-state CV-QKD protocol was proposed by F.  
4 Grosshans and P. Grangier in 2002 [39]. The protocol operates as  
5 follows:

6 • Alice generates two random numbers  $x_A$  and  $p_A$ . These both obey  
7 a Gaussian distribution with variance  $V_A N_0$  and mean 0, where  $V_A$  is  
8 an integer and  $N_0$  is the variance of the noise. Then,  $x_A$  and  $p_A$  are  
9 respectively applied to modulate the amplitude and phase of a  
10 coherent laser beam to generate a coherent state to be transmitted  
11 to Bob.

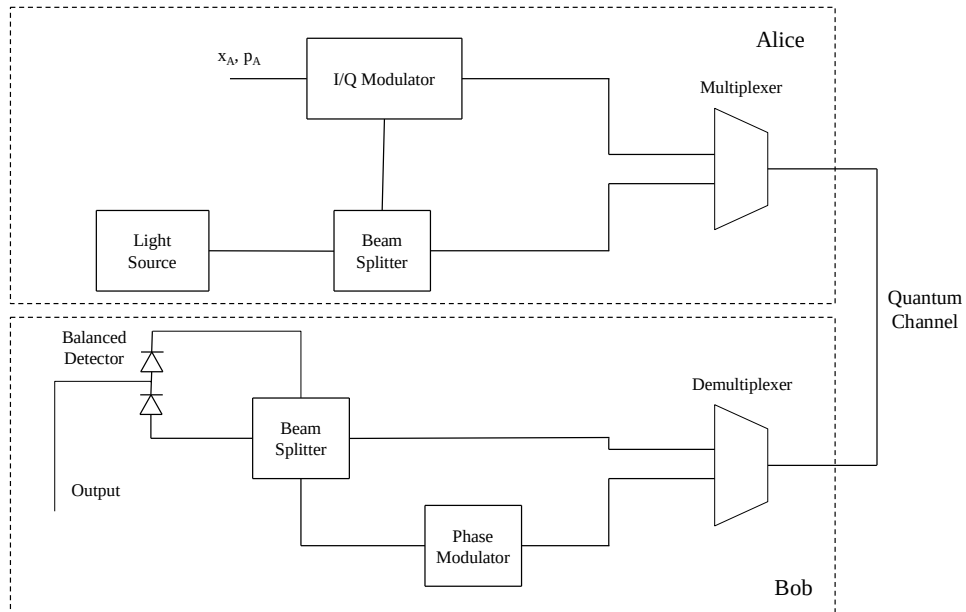
12 • Bob receives the coherent states and measures either  $X$  or  $P$   
13 randomly.  $X$  and  $P$  can be seen as bases for encoding and decoding  
14 the amplitude value  $x_A$  and the phase value  $p_A$ .

15 • Bob tells Alice about  $X$  or  $P$  that he measured. Alice compares the  
16 bases she used with Bob's and removes incompatible ones.

17 • Error correction and privacy amplification are carried out to  
18 further exclude the errors and Eve's knowledge to obtain final  
19 secure keys.

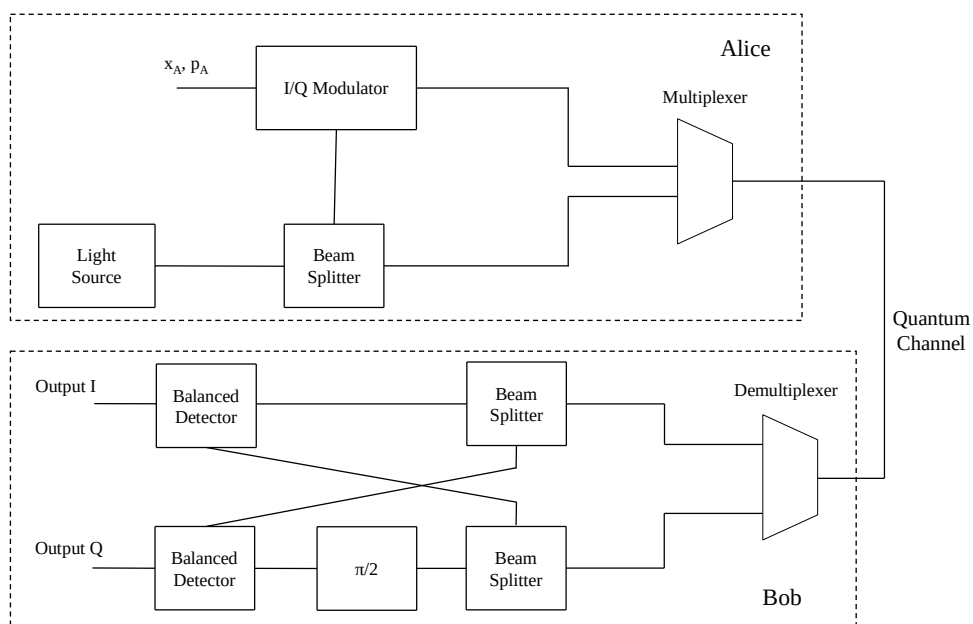
20 The implementation of coherent CV-QKD can be divided into two  
21 categories: homodyne and heterodyne [100-103]. Fig. 2.7 shows a  
22 block diagram for a real coherent CV-QKD system using the  
23 homodyne detection scheme [101]. In Alice, a weak coherent light  
24 source is used to generate initial coherent laser beams which are  
25 then passed through a BS. The optical power from each of the two  
26 paths of the BS is different. Only a small portion of light goes into  
27 the I/Q modulator and is modulated using Gaussian random

1 numbers  $x_A$  and  $p_A$ . Then, the light from the two paths is multiplexed  
 2 together to be transmitted to Bob. In Bob, the received light is  
 3 separated into two paths by a demultiplexer. The light in one path  
 4 passes through a phase modulator and combines with the light from  
 5 another path via a BS. A balanced detector is then used to measure  
 6 the amplitude or phase.



7

8 Figure 2.7 A block diagram for a coherent-state Gaussian protocol  
 9 using the homodyne detection



10

2  
3

1 Figure 2.8 A block diagram for a coherent-state Gaussian protocol  
2 using the heterodyne detection

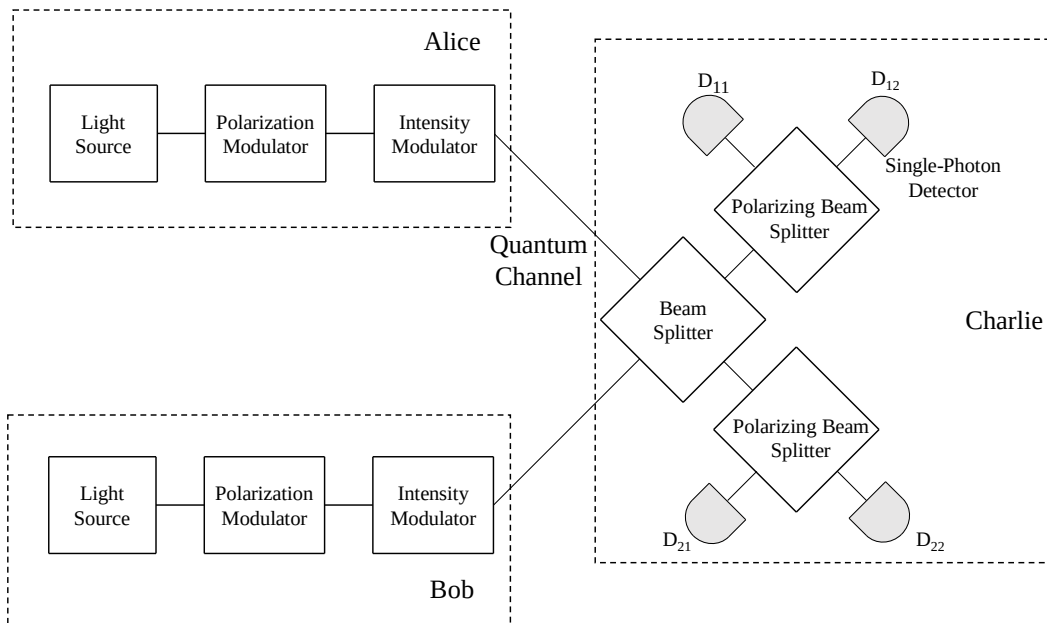
3 The heterodyne detection is another approach for coherent CV-  
4 QKD protocols [102]. A block diagram of the heterodyne system is  
5 shown in Fig. 2.8. The main difference when compared with the  
6 homodyne structure is the structure of the receiver that Bob uses. A  
7 90-degree hybrid is used to process received light and two balanced  
8 detectors are installed to measure both the in-phase and orthogonal  
9 components ( $X$  and  $P$  components). Therefore,  $x_A$  and  $p_A$  can both be  
10 obtained so that the secure key generation rate is increased  
11 compared to measuring one quadrature in homodyne detection.

#### 12 **2.2.4 MDI-QKD Protocol**

13 In 2012, the MDI-QKD protocol was proposed by H. Lo, M. Curty,  
14 and B. Qi, in order to eliminate detector side-channel attacks [104].  
15 An intermediate station, referred to as “Charlie”, is introduced  
16 between Alice and Bob as a receiver to collect and process the  
17 quantum information sent from Alice and Bob. The MDI-QKD  
18 protocol offers the advantage of achieving twice the communication  
19 distance compared to DV-QKD, COW-QKD and CV-QKD.

20 A typical block diagram of an MDI-QKD system is shown in Fig.  
21 2.9. Alice and Bob both generate qubits and send them to Charlie  
22 simultaneously. The structures of Alice and Bob are the same. Each  
23 has a weak coherent light source, a polarization modulator and an  
24 intensity modulator. Charlie collects and measures the Bell state of  
25 qubits from Alice and Bob simultaneously using one BS, two PBSs  
26 and four SPDs. Only when two specific SPDs are triggered, the Bell  
27 state measurement is successful. The Bell state  $\psi$  is obtained when  
28  $D_{11}$  and  $D_{22}$  or  $D_{21}$  and  $D_{12}$  are triggered. When  $D_{11}$  and  $D_{12}$  or  $D_{21}$  and

1  $D_{22}$  are clicked, the measured Bell state is  $\psi^-$ . Then, Charlie  
 2 announces the measurement results publicly. Alice and Bob keep the  
 3 corresponding successful bits and discard the rest. In addition,  
 4 either Alice or Bob needs to flip her or his data to guarantee the bit  
 5 strings in Alice and Bob are the same.

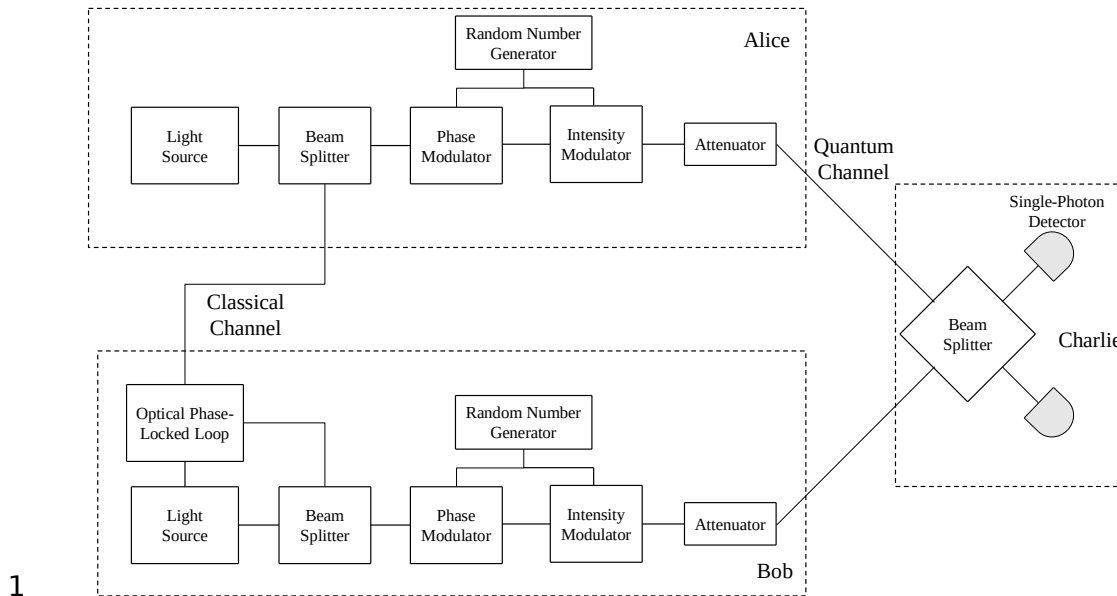


6

7 Figure 2.9 A typical block diagram of an MDI-QKD system

8 **2.2.5 TF-QKD Protocol**

9 The TF-QKD protocol was developed based on the MDI-QKD by  
 10 M. Lucamarini, Z. L. Yuan, J. F. Dynes and A. J. Shields [41]. The  
 11 communication distance is further improved compared to MDI-QKD  
 12 because it measures single-photon interference instead of two-  
 13 photon interference. A block diagram of a typical TF-QKD system is  
 14 illustrated in Fig. 2.10.



2 Figure 2.10 A typical block diagram of a TF-QKD system

3 Alice or Bob randomly chooses the bit value  $\alpha_i \in \{0, 1\}$ , the basis

4 value  $\beta_i \in \left\{0, \frac{\pi}{2}\right\}$  and the global phase value  $\gamma_i \in \mathcal{U}$ , where  $i$  is  $A$  or  $B$ ,

5 representing Alice or Bob, respectively. The distribution of the global

6 phase value is uniform between  $0$  and  $2\pi$ . The coherent state  $|\sqrt{\mu_i} e^{j\varphi_i}\rangle$

7 can be generated by intensity and phase modulation, where  $\mu_i$  is the

8 intensity and  $\varphi_i = \alpha_i \pi + \beta_i + \gamma_i$ . The intensity  $\mu_i$  can have different values

9 used to implement the decoy-state strategy. Then, the modulated

10 quantum information from both Alice and Bob is sent to Charlie.

11 Charlie measures the output of the beam splitter and publicly

12 announces the measurement results. Alice and Bob also announce

13 their intensity values  $\mu_i$ , basis values  $\beta_i$  and global phase values  $\gamma_i$ . If

14 the phase difference ( $|\varphi_A - \varphi_B|$ ) announced by Charlie is less than a

15 threshold  $\Delta$ , these two phases are called “twins” and the

16 corresponding qubits in Alice and Bob are kept. For example,  $\Delta$  is set

1 to be  $\frac{\pi}{16}$ . When the phase difference is  $\frac{\pi}{20}$ , the corresponding qubits  
 2 are kept. If the phase difference is  $\frac{\pi}{3}$ , the corresponding qubits are  
 3 discarded. When the phase difference is  $\frac{21\pi}{20}$ , the corresponding bits  
 4 are kept. This is because  $|\varphi_A - \varphi_B| - \pi < \Delta$ . It means that  $\beta_A = \beta_B$  and  
 5  $|\gamma_A - \gamma_B| < \Delta$ . In addition, if the phase difference is 0, the corresponding  
 6 bit values can be directly seen as the raw key. When the phase  
 7 difference is  $\pi$ , Bob should flip his bit values to obtain the raw key.  
 8 Finally, error correction and privacy amplification are operated to  
 9 get the final secure key.

## 10 2.3 WIRELESS QKD DEMONSTRATIONS

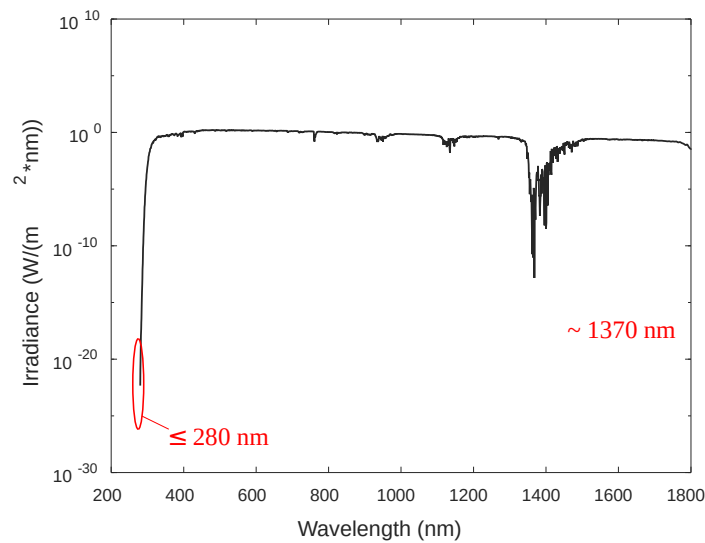
11 There have been a substantial number of QKD demonstrations  
 12 focused on fibre-based QKD systems [42, 43, 80-87, 101, 105-118].  
 13 The maximum SKR achieved is more than 100 Mbit/s using BB84 at  
 14 a distance of approximately 10 km [43] and the longest  
 15 transmission distance is more than 1000 km based on TF-QKD [42].  
 16 However, there are very few demonstrations of short-range wireless  
 17 QKD. These are summarized in Table 2.3 [46-50, 119-121]. The  
 18 maximum SKR achieved so far reaches 0.322 Mbit/s working at 1550  
 19 nm.

20 Table 2.3 Summary of short-range wireless QKD demonstrations

Year	Authors	Protocol	Wavelength	Link length	SKR
2017	H. Chun et al. [46]	DV	658nm	0.5 m	30kbit/s

202 1	D. Lowndes et al. [47]	DV	650nm	N/A	20kbit/s
202 1	M. Avesani et al. [119]	DV	1550nm	145 m	30kbit/s
202 2	A. T. Castillo et al. [48]	COW	850nm	1 m	6.4kbit/s
202 3	X. Zheng et al. [49]	CV	1550nm	1 m	322kbit/s
202 3	A. T. Castillo et al. [50]	COW	850nm	N/A	50kit/s
202 4	F. Honz et al. [121]	BB84	1400nm	63m	<3.7kbit/s

1 A key factor causing limited wireless QKD demonstrations is  
2 ambient light interference. The ambient light within the optical  
3 bandwidth of the wireless QKD system is detected by the receiver,  
4 potentially dominating the detected signal. The strongest source of  
5 ambient light is sunlight. Fig. 2.11 shows the near-ground sunlight  
6 irradiance according to ASTM 1.5 Spectra [51]. There are several  
7 potential wavelengths for short-range wireless QKD. Wavelengths  
8 around 1370 nm and in the UV-C region of the optical spectrum  
9 (less than 280 nm) are of particular interest because of the  
10 extremely low sunlight irradiance. The irradiance at 1370 nm and  
11 280 nm is  $1 \times 10^{-11} W/(m^2 \cdot nm)$  and  $5 \times 10^{-23} W/(m^2 \cdot nm)$  while over  
12  $0.8 W/(m^2 \cdot nm)$  for the visible light. A wireless QKD system around  
13 1370 nm has been investigated [60], but the establishment of a  
14 wireless QKD system in the UV-C region is an issue to be explored.



1

2

Figure 2.11 Near-ground sunlight irradiance

3

## 2.4 POTENTIAL OF UV-C SHORT-RANGE WIRELESS QKD

4 The lack of sunlight interference in the UV-C region provides the  
5 potential to develop a wireless QKD system, but the available light  
6 sources and photodetectors in the UV-C region are very limited. At  
7 the transmitter, the options for UV-C light sources are limited. There  
8 is no commercial laser diode (LD) operating below 300 nm.  
9 However, commercial UV-C LEDs are available with an operating  
10 wavelength even below 250 nm. The bandwidth of commercial LEDs  
11 is limited, however, typically these are around 10 MHz [122, 123].  
12 One potential option is to use UV-C micro-LEDs, which have been  
13 fabricated with a peak wavelength of  $\leq 280$  nm and bandwidth  
14 exceeding 100 MHz [124, 125]. Therefore, UV-C micro-LEDs are a  
15 good candidate for a UV-C wireless QKD transmitter. At the receiver,  
16 there are very few UV-C SPDs reported [126-132]. Moreover, each of  
17 these reported UV-C SPD has obvious disadvantages, such as  
18 requiring an ultralow working temperature (4 K) or having a high  
19 DCR ( $\geq 500$  kcps). One potential option is to use an SiPM, which is

2

3

1 designed for photon counting. In the UV-C region, SiPMs typically  
2 have a usable photon detection efficiency (PDE) of  $\sim 15\%$ , a DCR of  
3  $\sim 50 \text{kcps/mm}^2$  and a gain of  $1 \times 10^6$  [133]. Therefore, an SiPM is  
4 potentially a good choice for a UV-C wireless QKD receiver, but the  
5 DCR is too high at room temperature.

6 In addition, the B92 protocol is selected because of its resource-  
7 efficient characteristics. The final goal of the work reported in this  
8 thesis is to develop a short-range wireless B92 QKD system  
9 operating in the UV-C region, utilizing UV-C micro-LEDs and SiPMs.

## 10 **2.5 CONCLUSIONS**

11 This Chapter has reviewed different QKD protocols and state-of-  
12 the-art wireless QKD research. The typical structure of a QKD  
13 system was presented, including the physical layer, sifting, error  
14 correction and privacy amplification. A comparison of different QKD  
15 protocols was introduced, including DV-QKD, COW-QKD, CV-QKD,  
16 MDI-QKD and TF-QKD, followed by the summary of state-of-art  
17 wireless QKD demonstrations. Finally, the potential of building a UV-  
18 C short-range wireless QKD system was discussed.

19 A QKD system requires ultralow photon flux links with an  
20 average photon number per bit of less than 1. In the next Chapter,  
21 the development of a low photon flux link is discussed. This serves  
22 as the initial work leading to a QKD link.

## 1 Chapter 3:

# 2 Ultra-Sensitive UV Solar-Blind OWC with 3 an SiPM

### 4 3.1 INTRODUCTION

5 OWC is a promising technology to provide high-quality wireless  
6 access service because of its high bandwidth, low latency,  
7 robustness to electromagnetic interference and lack of regulatory  
8 requirements [1, 3, 134]. A typical OWC system uses a light source  
9 to transmit and a photodetector to receive optical signals.  
10 Transmitters either use Light Emitting Diodes (LEDs) or Lasers.  
11 Receivers consist of a photodetector, amplification and subsequent  
12 decoding and decision-making.

13 There are typically three bands used for OWC: Infrared (IR) band  
14 (760 nm - 2000 nm), visible band (390 nm - 760 nm) and ultraviolet  
15 (UV) band (10 nm - 390 nm). According to the sunlight spectrum, to  
16 avoid sunlight interference, there is a “solar-blind” region within the  
17 UV band whose wavelength is from 200 nm to 280 nm [51]. Sunlight  
18 within this spectrum is absorbed by the ozone, resulting in an  
19 extremely low irradiance near the ground. Thus, this solar-blind  
20 spectrum is a potential solution for both indoor and outdoor OWC.  
21 Several devices for solar-blind communications have been  
22 developed [135-141] and links operating at the Gbit/s level have  
23 been demonstrated [142-144].

24 However, compared to visible and IR light, the safe irradiance

1 level of UV light is much lower. According to eye safety standards,  
2 the irradiance of light between 200 - 400 nm should be no more  
3 than  $1mW/m^2$  [145]. Recent UV solar-blind OWC link demonstrations  
4 use received power levels of microwatts, corresponding to  $\sim 10^4$   
5 photons per bit [146, 147]. For these demonstrations, the  
6 corresponding beam irradiance (considering the detected power  
7 divided by the receiver collection area) is estimated to be more than  
8  $10mW/m^2$ , exceeding the safe irradiance level at the receiver. To the  
9 best of our knowledge, there was no demonstration showing a UV  
10 solar-blind OWC system, using a high-sensitivity receiver, that  
11 meets this eye safety criterion. Thus, to build OWC in the UV solar-  
12 blind region, improving receiver sensitivity to allow low transmission  
13 power is a key step.

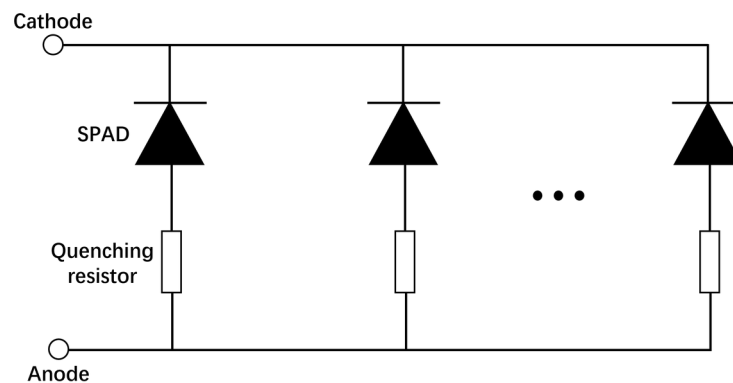
14 Several materials are promising candidates for developing high-  
15 sensitivity detectors in the UV solar-blind band, including  
16 Molybdenum Silicide (MoSi), SiC, Gallium nitride (GaN) and Silicon  
17 (Si) [126-132]. A MoSi superconducting nanowire single-photon  
18 detector (SNSPD) reported in [128] has a PDE exceeding 70% and a  
19 DCR of approximately 0.25 count per second (cps) but is limited in  
20 its applicability due to the low operating temperature of 4K. In the  
21 case of SiC and GaN single-photon avalanche diodes (SPAD), the  
22 DCR is over  $1Mcps/mm^2$  at room temperature with a PDE typically at  
23 around 10 % [129]. Among all Si-based photodetectors, the SiPM is  
24 specifically designed for photon counting [148-152]. Compared with  
25 the above three materials, typical SiPMs have a low DCR ( $\sim$   
26  $50kcps/mm^2$ ), can have a large active area ( $\geq 1mm^2$ ) and a high  
27 enough PDE ( $\sim 15\%$ ) for UV detection at room temperature [153].  
28 Thus, an SiPM has the potential to form the basis of a high-  
29 sensitivity receiver working in this UV solar-blind region.

1 This Chapter presents a UV solar-blind OWC scheme based on  
2 such an ultra-sensitive receiver. First, a modelling of UV OWC links is  
3 reported. Then, a cooled UV-sensitive SiPM is described. This is  
4 combined with a commercial 275 nm LED used as a transmitter.  
5 Results from several different link demonstrations are then  
6 reported.

## 7 **3.2 MODELLING OF UV WIRELESS COMMUNICATION LINKS**

### 8 **3.2.1 SiPM Operation**

9 A key component providing high sensitivity in detecting ultralow  
10 received power signals is an SiPM. An SiPM is an array of Si-based  
11 single-photon avalanche diodes (SPADs), as shown in Fig. 3.1 [154].  
12 Each SPAD operates in Geiger mode, due to the device design and  
13 the high bias voltage that reverse biases each diode.



14

15 Figure 3.1 Structure of SPADs in an SiPM [154]

16 The SPADs in an SiPM are wired in parallel, and the photocurrent  
17 from SPADs are therefore summed together. For each SPAD, a  
18 detected photon causes avalanche multiplication and, in turn,  
19 generates a current pulse flowing through a quenching resistor. As  
20 the avalanche current increases, the quenching resistor develops a  
21 voltage drop, which effectively reduces the bias across the SPAD

1 below its breakdown voltage. This causes the SPAD to exit the  
2 Geiger mode, thereby quenching the avalanche. As the avalanche  
3 current dissipates, the bias is gradually restored, and the SPAD  
4 becomes ready for the next photon detection. Dark Counts are  
5 spontaneously generated current pulses that occur without a  
6 detected photon and ultimately limit the sensitivity of the SiPM.  
7 When the number of detected photons and dark counts are low, the  
8 distribution of pulses follows a Poisson distribution [155].

### 9 3.2.2 System Model and Simulation

10 The UV wireless communication system model used in this  
11 section is a simple line-of-sight (LOS) model, including a transmitter,  
12 a receiver and an LOS wireless channel (see Fig. 3.2). The  
13 transmitter is assumed to be a commercial light source, such as a  
14 UV LED, transmitting OOK signals. The receiver is assumed to be a  
15 photon counting detector, such as an SiPM.



17 Figure 3.2 UV wireless communication model

18 The following theoretical analysis and simulation focus on the  
19 operation of the receiver and the estimation of the Bit Error Rate  
20 (BER) of the system. This is a function of the number of detected  
21 signal counts (caused by incident photons) and errors due to dark  
22 counts and the incorrect transmission of bits due to the Poisson  
23 statistics of the source. It is assumed that the transmitter sends  
24 photons as short pulses, and that the receiver has a detector that is

1 sensitive over a detection window that is matched to the width of  
2 the transmitted pulse. This ensures that the receiver has a  
3 maximum received signal-to-noise ratio as it is not sensitive when  
4 there is no transmitted signal.

5 The probability of detecting dark counts is:

$$6 \quad P_d = 1 - \frac{(DCR \cdot t_d)^0}{0!} e^{-DCR \cdot t_d} = 1 - e^{-DCR \cdot t_d} \quad (3.1)$$

7 where  $t_d$  is the time duration of the detection window. The  
8 probability of transmitting a bit "1" when no photons are detected  
9 is:

$$10 \quad P_n = \frac{\mu^0}{0!} e^{-\mu} = e^{-\mu} \quad (3.2)$$

11 Where  $\mu$  is the average number of detected photons per bit,  
12 which can be obtained from:

$$13 \quad \mu = PDE \cdot P / (h \cdot c / \lambda) \quad (3.3)$$

14 Where  $PDE$  and  $P$  are the photon detection efficiency and  
15 received power of the SiPM, respectively. Thus, the total BER is:

$$16 \quad BER = \frac{1}{2} P_d + P_n - P_d \cdot P_n \quad (3.4)$$

17 The factor 1/2 originates from the 50% probability of dark counts  
18 occurring in bits "1", which will not cause an error.  $P_d \cdot P_n$  is the  
19 probability that bits "1" are correctly detected due to dark counts  
20 instead of detected photons.

21 Simulation was conducted to examine the effect of DCR and

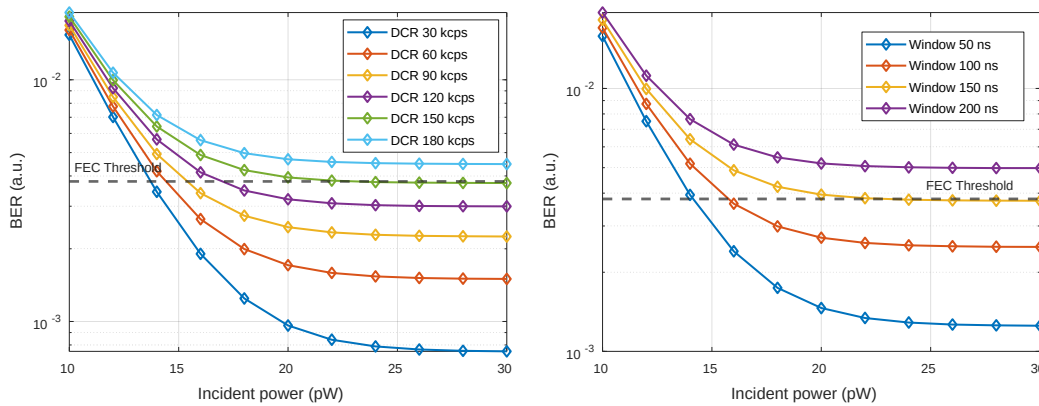
1 detection window on BER. The parameters and corresponding values  
 2 are listed in Table 3.1. At the transmitter, 1 Mbit/s OOK signals are  
 3 emitted with a wavelength of 280 nm. At the receiver, the received  
 4 power is varied, changing from 10 pW to 30 pW. The detection  
 5 window varies from 50 ns to 200 ns. The DCR changes from 30 kcps  
 6 to 180 kcps, and the PDE is assumed to be 15% for 280 nm. These  
 7 are typical values for an SiPM.

8 Table 3.1. Simulation parameters and values

Parameters	Values
Wavelength $\lambda$	280 nm
Data rate	1 Mbit/s
Received power	10 pW ~ 30 pW
DCR	30 kcps ~ 180 kcps
Detection window $t_d$	50 ns ~ 200 ns
PDE	15%

9 Fig. 3.3 shows a plot of BER versus the received power, with  
 10 varying DCR and detection window size. The detection window and  
 11 DCR are fixed at 50 ns and 50 kcps in Fig. 3.3 (a) and Fig. 3.3 (b)  
 12 respectively. The figure shows a forward error correction (FEC)  
 13 threshold of  $3.8 \times 10^{-3}$ , at which bit error-free can be achieved with a  
 14 7% coding overhead [125]. The results show the BER decreases with  
 15 the rise of received power. This is because the probability of  
 16 transmitting a bit "1" when no photons are detected is reduced  
 17 according to Eq. (3.2). When the received power is over 25 pW,  
 18 there is little change in BER. This is because the probability of  
 19 transmitted bits "1" without detected photons is very small, and the

1 unchangeable error of transmitted bits “0” detected as “1” due to  
2 DCR dominates the total BER, according to Eq. (3.1). Moreover, the  
3 product of DCR and detection window should not exceed  $7500 \text{ kcps} \cdot \text{ns}$ ,  
4 or the BER will be over the FEC threshold.



5

6

7 Figure 3.3 BER versus received power, DCR and detection window

8 (a) with a detection window of 50 ns (b) with a DCR of 50 kcps

9 Therefore, with a receiver having a DCR of less than 50 kcps and  
10 a detection window of lower than 100 ns, modelling indicates that a  
11 practical UV wireless communication system can be established.

12 The following sections describe the design of the transmitter  
13 and receiver for a link that will meet these requirements.

### 14 3.3 TRANSMITTER CHARACTERIZATION

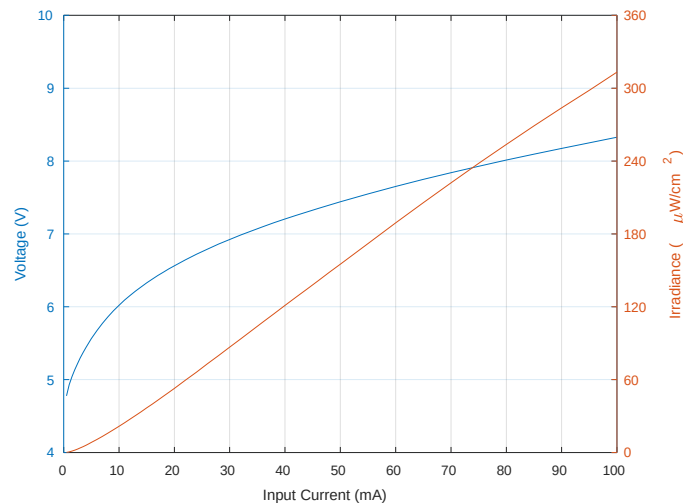
15 To build a UV solar-blind OWC link, a transmitter should have a  
16 central wavelength  $< 280 \text{ nm}$ , with as large a modulation bandwidth  
17 as possible. Because of the lack of commercial LDs operating below  
18  $300 \text{ nm}$ , a UV LED is the only option for the light source.

19 The Thorlabs LED275J was selected because the LEDs from  
20 different manufacturers are roughly the same. According to the

3

4

1 datasheet, the central wavelength is 275 nm and the FWHM is 12  
2 nm. The LED is packaged with an integrated ball lens to control  
3 emission from the device and the field-of-view (FOV) is 15°.  
4 Additionally, the maximum output optical power is approximately 1  
5 mW with a drive current of 100 mA.

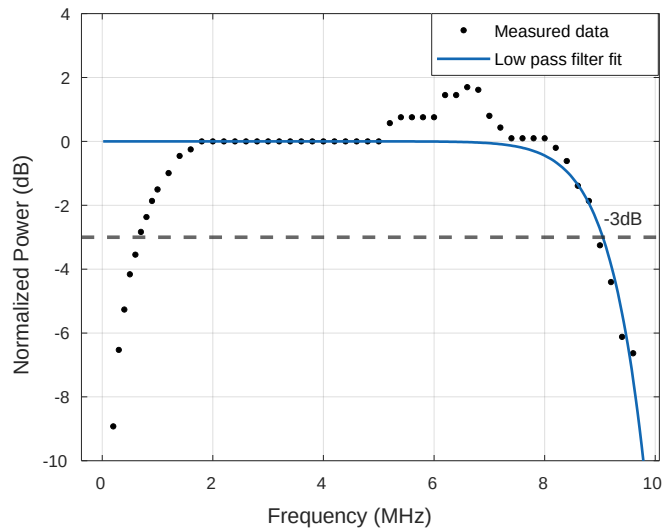


6

7 Figure 3.4 Plot of I-V and I-I for Thorlabs LED275J LED

8 In addition to the above characteristics the current-voltage-  
9 irradiance (I-V-I) curves and bandwidth need to be determined. An  
10 experiment to measure the I-V-I curves was undertaken. A source  
11 meter (KEITHLEY 2636B) was used to drive the LED and measure  
12 the voltage across it. The optical power emitted by the LED was  
13 measured by a power meter (Newport 843-R) with a UV-enhanced  
14 silicon (Si) head (818-UV/DB). The corresponding irradiance was  
15 calculated by dividing the measured power by the detection area of  
16 the Si head. The LED's output power was not measured due to its  
17 relatively large divergence angle and the limited detection area of  
18 the Si head. Even when the LED and Si head are placed nearby, the  
19 Si head is unable to collect all the emitted power from the LED.  
20 Thus, the LED and the Si head were aligned with a separation of

1 approximately 10 cm, in order to match the transmitter-to-receiver  
2 distance used in the OOK link experiment in this Chapter (see  
3 section 3.5). The measured I-V-I curves are illustrated in Fig. 3.4.  
4 The result shows the turn-on voltage is approximately 5 V.



5

6 Figure 3.5 Normalized signal power versus frequency

7 The bandwidth was then measured. DC bias provided by a laser  
8 driver (Thorlabs LDC205C) and an AC signal (sine wave) provided by  
9 an arbitrary waveform generator (AWG, AGILENT 81150A) were  
10 combined by a Bias Tee (Mini-Circuits ZFBT-6GW). The combined  
11 signal was used to drive the LED. Two convex lenses (LA4647-UV  
12 and LA4130-UV) were used to collimate the light from LED275j and  
13 focus light onto the active area of a SiC APD (GaN<sub>0</sub> Optoelectronics  
14 SPAD-ABC-S) detector. According to the SiC APD's datasheet, the  
15 rise time is ns-level, so the corresponding bandwidth is at least over  
16 35 MHz ( $\text{Bandwidth} \approx 0.35 / \text{rise time}$ ). Moreover, the PDE at 275 nm, the  
17 diameter of the active area and the gain are 10%, 120  $\mu\text{m}$  and 100,  
18 respectively. The APD was used in combination with a trans-  
19 impedance amplifier (TIA) MAX3665 with a gain of 8k $\Omega$  and a  
20 bandwidth of over 100 MHz. A small Printed Circuit Board (PCB) was

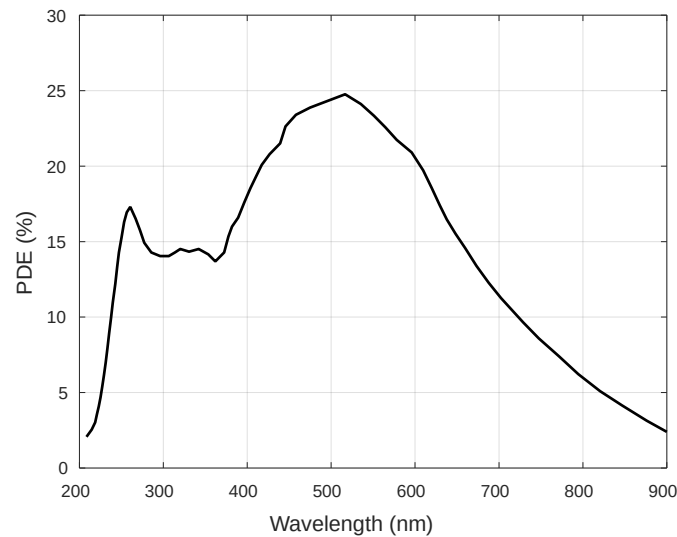
1 used to ensure a good connection between the TIA and APD. The  
2 amplified signal was then recorded on an oscilloscope (AGILENT  
3 MSO6104A).

4 The AWG frequency was varied between 100 kHz and 10 MHz,  
5 and the signal level on the Oscilloscope was recorded. This signal  
6 level is proportional to the optical power received by the APD. Fig.  
7 3.5 shows the frequency response of the link, showing an electrical-  
8 to-optical (E-O) bandwidth of approximately 9 MHz. Although it is  
9 not possible to accurately determine the bandwidth of the receiver  
10 the estimated bandwidth is much greater than 9MHz, indicating that  
11 this is the bandwidth of the LED device. The other notable feature in  
12 the figure is the lower frequency cut-off which is due to the  
13 response of the Bias Tee.

## 14 **3.4 RECEIVER DESIGN AND CHARACTERIZATION**

15 To build a UV solar-blind OWC link, an SiPM-based receiver  
16 should meet the following criteria: (1) enough PDE (2) high gain (3)  
17 wide bandwidth and (4) low DCR. Because of the lack of a suitable  
18 commercial detector module, a new receiver needs to be developed.  
19 The next section describes the choice of detector.

### 20 **3.4.1 SiPM Selection**



1

2 Figure 3.6 PDE of HAMAMATSU S13360-3025UQ-SPL

3 SiPMs with a range of different characteristics are available. A  
4 typical SiPM (HAMAMATSU S13360-3025UQ-SPL) is a feasible option.  
5 This SiPM has a 3x3 mm<sup>2</sup> active area and is designed for UV photon  
6 counting, with a window material of quartz glass, allowing UV light  
7 to pass through. The corresponding PDE curve is shown in Fig. 3.6.  
8 It shows the PDE at 280 nm is approximately 15 %.

9 The gain is significant for pulse detection because every pulse  
10 generated by a detected photon should be at least at the mV level  
11 to be observed by an oscilloscope. The gain of the S13360-3025UQ-SPL is  
12  $7 \times 10^5$ , which is significantly higher than the gains of 10 -100 for  
13 linear mode APDs. The rise time of the S13360-3025UQ-SPL is ns-  
14 level, corresponding to a bandwidth of at least tens of MHz. The fall  
15 time is on the order of tens of ns and is limited by the quenching  
16 time. This quenching time can be approximated by the RC time  
17 constant,  $T_f = R_{quench} \cdot C_{SPAD}$ , where  $R_{quench}$  is the quenching resistor and  
18  $C_{SPAD}$  is the capacitance of the SPAD.

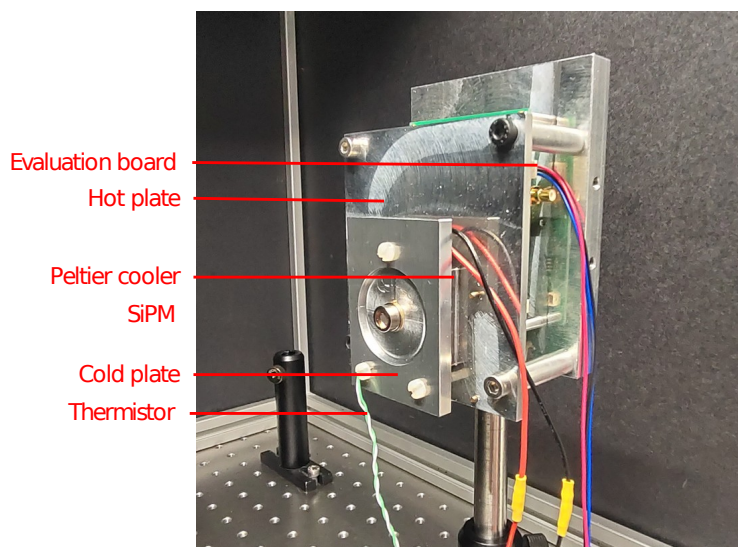
19 However, the DCR of S13360-3025UQ-SPL is over 400 kcps,

1 which is higher than the required value of 50 kcps according to the  
2 simulation in section 3.2. Thus, a method of cooling the detector is  
3 necessary to reduce the temperature to achieve a small enough  
4 DCR [153].

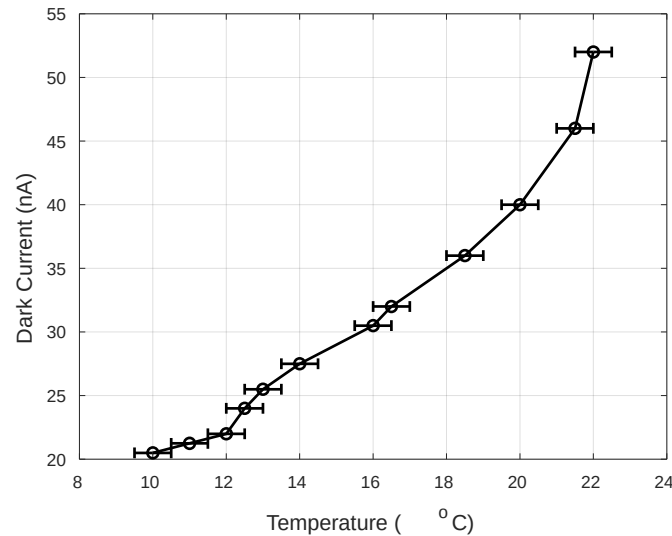
### 5 **3.4.2 Cooling Design**

6 A passive cooling design for the S13360-3025UQ-SPL using a  
7 Peltier cooler (Laird SH10-23-06-L1-W4.5) was developed and is  
8 shown in Fig 3.7.

9 The Peltier cooler has a centre hole and is sandwiched between  
10 a cold plate and a hot plate, both made of Al. A thermistor is  
11 connected to the cold plate to measure its temperature,  
12 representing the temperature of the SiPM. The SiPM was placed on  
13 the top of the cold plate, with its leads going through the Peltier  
14 cooler to connect to an evaluation board (HAMAMATSU C12332-02).  
15 The evaluation board was remotely controlled to provide the bias  
16 voltage to the SiPM. The dark current of the SiPM was measured by  
17 connecting the output of the evaluation board to a source meter  
18 (KEITHLEY 2636B).



1 Figure 3.7 Passive cooling design for S13360-3025UQ-SPL



2

3 Figure 3.8 Dark current versus temperature for S13360-3025UQ-SPL

4 The bias voltage was fixed at 56.86 V, which is the  
5 recommended value for the SiPM. The measured dark current  
6 versus temperature is shown in Fig. 3.9. (An error bar of 1 °C is  
7 added because the temperature values were measured as integers.)  
8 The results show the dark current decreases from 52 nA to 20 nA  
9 with the temperature reduced from 22 °C to 10 °C. A current of 20 nA  
10 corresponds to approximately 180 kcps, which is still too high to  
11 establish a UV solar-blind OWC link. The temperature cannot be  
12 further reduced, as the cooling capacity is limited by the heat  
13 dissipation capacity of the hot plate. However, cooling has proved to  
14 be effective in reducing DCR. A new cooling design is needed to  
15 further reduce the temperature of the SiPM.

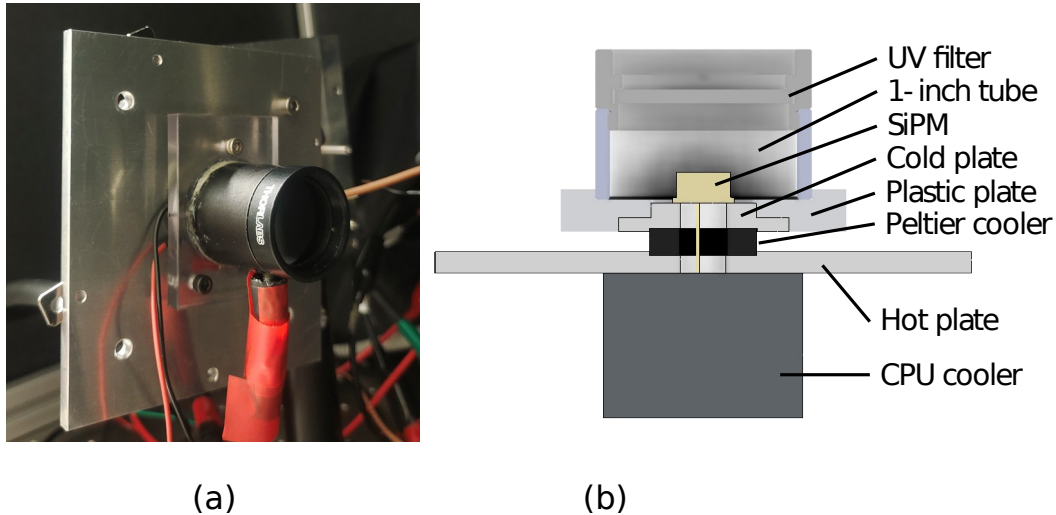
16 An alternative method to improve heat dissipation is using  
17 active liquid cooling instead of passive cooling. This is because the  
18 heat transfer coefficient of flowing liquid ( $\approx 500 W/(m^2 \cdot K)$ ) is much

1 higher than that of still air  $\approx 10 \text{ W}/(\text{m}^2 \cdot \text{K})$ . Thus, liquid cooling has the  
2 potential to efficiently dissipate the heat of the hot plate, which in  
3 turn allows the Peltier cooler to cool the SiPM to a low enough  
4 temperature.

5 The goal of the liquid cooling design is to achieve an SiPM  
6 temperature of  $-10 \text{ }^\circ\text{C}$ , ensuring a DCR of below 50 kcps (see section  
7 3.2.2). This is determined by an over 400 kcps of S13360-3025UQ-  
8 SPL at room temperature and a 10 dB reduction in DCR per  $30 \text{ }^\circ\text{C}$   
9 decrease [153]. There are two key considerations for the cooling  
10 design: (1) high-efficiency heat dissipation on the hot side and (2)  
11 overcoming the condensation on the SiPM's surface as the  
12 temperature of the SiPM falls below the dew point, which is typically  
13 around  $10 \text{ }^\circ\text{C}$ . A liquid CPU cooler (MASTER LIQUID LITE 240) was  
14 used to cool the hot plate. An image of the liquid cooling design is  
15 shown in Fig. 3.9 (a), and Fig. 9 (b) illustrates the sectional view.

16 The Peltier cooler (Laird SH10-23-06-L1-W4.5) is at the centre. At  
17 its cold side, there is a small Al plate (cold plate) held in place by an  
18 auxiliary plastic plate, thus reducing the overall thermal load that  
19 must be cooled. The SiPM is positioned on top of the cold plate and  
20 is held in place by an SM1 to TO-5 adaptor (Thorlabs S1LM05). To  
21 avoid condensation, the SM1 tube is glued on an auxiliary plastic  
22 plate with a UV filter (Thorlabs FGUV5M) on top to make a confined  
23 environment, and the air around the SiPM is pumped away by a  
24 vacuum pump (RS vacuum pump). The hot side of the Peltier and  
25 the CPU cooler are thermally connected through a large Al plate (hot  
26 plate). Graphene sheets are placed on the surface of adjacent  
27 components (SiPM, cold plate, Peltier cooler, hot plate and CPU  
28 cooler) to ensure efficient heat conduction. To connect the SiPM to

1 the PCB board, there is a hole in the middle of the cold plate, Peltier  
2 cooler and hot plate and a horizontal trench in the hot plate for the  
3 wires to pass through.



6 Figure 3.9 Liquid cooling design (a) Image (b) Sectional view

### 7 3.4.3 Amplification Circuit Design

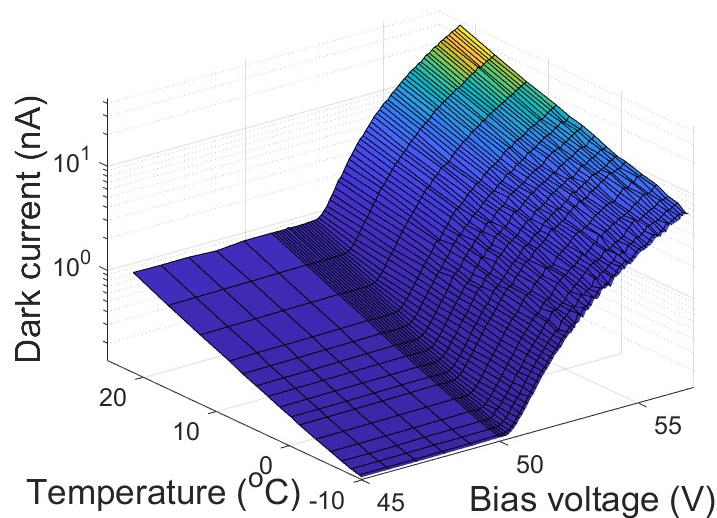
8 An amplification circuit using a TIA (ADA4817) was designed by  
9 James Farmer, a postdoctoral researcher in our group. The ADA4817  
10 has a gain bandwidth product (GBWP) of 410 MHz. The theoretical 3-  
11 dB bandwidth of this amplification circuit can be calculated as [156]:

$$12 \quad f_{3dB} \approx \sqrt{\frac{GBWP}{2\pi R_F C_i}} \quad (3.5)$$

13 where  $R_F$  is the resistance of the trans-impedance load and  $C_i$  is the  
14 total capacitance at the input of the TIA. The resistance of the trans-  
15 impedance load is 7500  $\Omega$ , and the total capacitance is dominated  
16 by the SiPM capacitance of approximately 320 pF. Thus, the  
17 theoretical 3-dB bandwidth of the receiver is calculated as 5.2 MHz.

### 18 3.4.4 Receiver Characterization

1 The cooling performance of the receiver was characterized by  
2 measuring the dark current as a function of temperature and bias  
3 voltage. A source meter (KEITHLEY 2636B) provided the bias voltage  
4 to the SiPM and measured the SiPM's output current as the dark  
5 current. The temperature of the SiPM is represented by the  
6 temperature of the cold plate, which was measured by a thermistor  
7 (TDK B57703M). The measured result is shown in Fig. 3.10. The SiPM  
8 was cooled from 21 °C to -10 °C. Further reducing the temperature  
9 is difficult because of the limitations of the Peltier cooler and CPU  
10 cooler.



11

12 Figure 3.10 Dark current versus temperature and bias voltage

13 The breakdown voltage of the detector decreases with  
14 decreasing temperatures by 54 mV/K. In order to keep a constant  
15 overvoltage, the bias voltage (breakdown voltage + overvoltage)  
16 must be adjusted according to the temperature setting. Thus, at -10  
17 °C the bias voltage was set to 55.2 V resulting in a measured dark  
18 current of 3.9 nA compared to 56.8 V and 42 nA at 21°C. The  
19 corresponding DCRs are 35 kcps and 375 kcps, respectively. These  
20 measurement values correspond to a reduction in the dark current

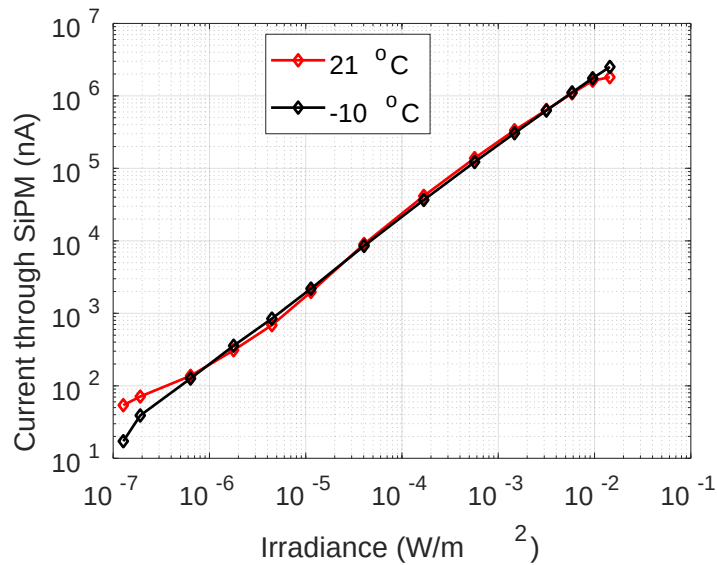
1 of about 10 dB with 30 K of cooling. This gradient agrees with  
2 previously reported values for SiPMs [153]. The temperature cannot  
3 be further reduced because of the limitation of the Peltier cooler and  
4 CPU cooler.

5 The gain of the SiPM can be obtained by:

$$6 \quad G = \frac{I_d}{DCR \cdot e_0} \quad (3.8)$$

7 where  $I_d$  is the dark current and  $e_0$  is the elementary charge. Thus,  
8 the gain of the SiPM is approximately  $7 \times 10^5$ .

9 The current from the SiPM versus irradiance from the LED275J  
10 was measured to obtain the PDE. The LED275J was driven by a laser  
11 driver (Thorlabs LDC205C), and the irradiance was measured by a  
12 power meter (Newport 843-R with a UV-enhanced Si head 818  
13 UV/DB). Then, the receiver replaced the power meter at the same  
14 position to obtain the photocurrent from the SiPM. The bias voltage  
15 was adjusted to recommended values with temperature change.  
16 The measured result is shown in Fig. 3.11, considering both 21 °C  
17 and -10 °C. The difference between the two curves at the beginning  
18 reflects different dark currents at these two temperatures.



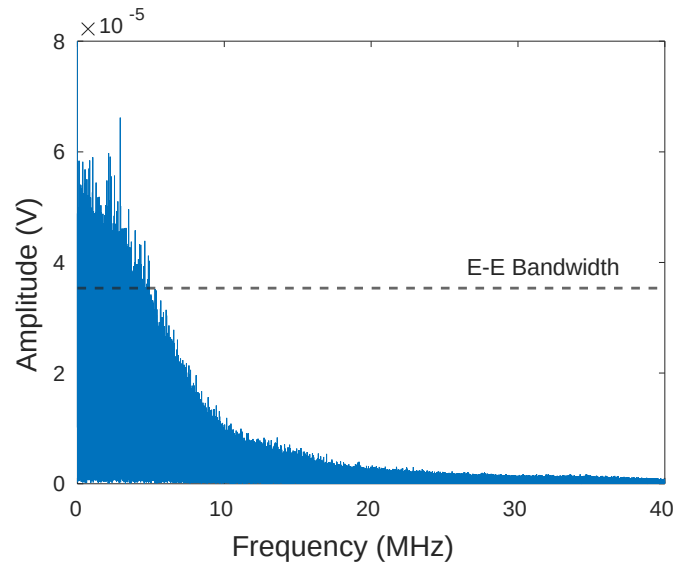
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2 Figure 3.11 Current from the SiPM versus irradiance from LED275J

3 As the plots for both temperatures are similar to each other, the  
 4 gain and the PDE of the SiPM can be considered constant. The PDE  
 5 can be obtained by:

6 
$$PDE = \frac{(I_t - I_d) \cdot h \cdot c}{G \cdot e_0 \cdot E \cdot A \cdot \lambda}$$
  
 (3.9)

7 where  $I_t$  is the total current from the SiPM, including photocurrent  
 8 and dark current.  $E$  is the irradiance and  $A$  is the detection area of  
 9 SiPM.  $h$  is Planck's constant and  $c$  is the speed of light.  $\lambda$  is the  
 10 operating wavelength. Thus, the PDE is calculated as 15.38 %.



1

2 Figure 3.12 The frequency domain of dark pulses

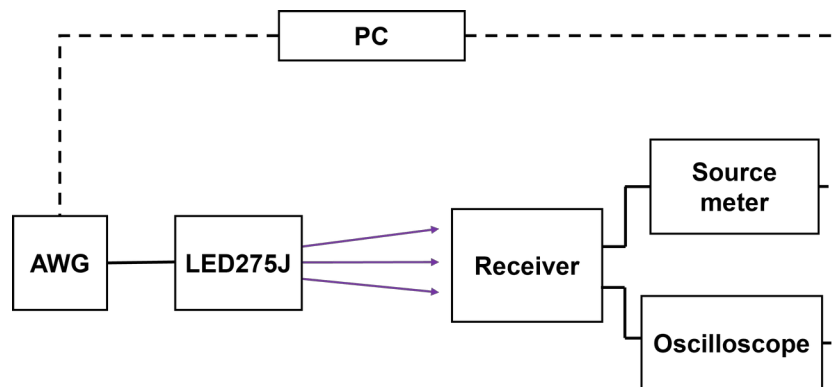
2

3 The bandwidth of this receiver can be estimated by Fourier  
4 transforming the output signal from the receiver due to dark counts.  
5 An individual dark count gives rise to a very short output pulse  
6 whose shape corresponds to the impulse response of the system so  
7 that the Fourier transform can be used to estimate bandwidth. The  
8 Fourier transform of a sequence of dark pulses recorded over 20 ms  
9 is shown in Fig. 3.12. The result shows the amplitude of around 0  
10 MHz is approximately 55  $\mu\text{V}$ . The receiver's electrical-to-electrical  
11 (E-E) bandwidth is estimated to be approximately 5 MHz,  
12 corresponding to an amplitude of  $55 \mu\text{V} / \sqrt{2} \approx 39 \mu\text{V}$ . The measured E-E  
13 bandwidth matches the theoretical bandwidth calculated in section  
14 3.4.3, with the limitation caused by the amplifier.

### 15 **3.5 ULTRA-SENSITIVE UV SOLAR-BLIND OWC LINKS**

16 A UV solar-blind OWC link was established to test the sensitivity  
17 of the receiver. The block diagram of the corresponding  
18 experimental setup is shown in Fig. 3.13. Initially, an On-Off-Key  
19 (OOK) link was tested.

1 A pseudorandom binary sequence (PRBS) OOK sequence with a  
2 length of  $2^{10} = 1024$  bits was generated by a PC. This sequence  
3 length exceeds that of the standard PRBS9 (511 bits), making it  
4 suitable for high-speed communication link testing. The generated  
5 sequence was then loaded into an arbitrary waveform generator  
6 (AWG, AGILENT 81150A) with a memory depth of 512k samples,  
7 which was used to drive LED275J. Light from the LED propagates to  
8 the receiver, which was placed 10 cm from the transmitter. A source  
9 meter (KEITHLEY 2636B) provided the bias voltage to the receiver.  
10 The OOK signal detected by the receiver was sent to an oscilloscope  
11 (AGILENT MSO6104A). The data captured by the oscilloscope was  
12 post-processed offline. In addition, the AWG, source meter and  
13 oscilloscope were remotely controlled by a PC.

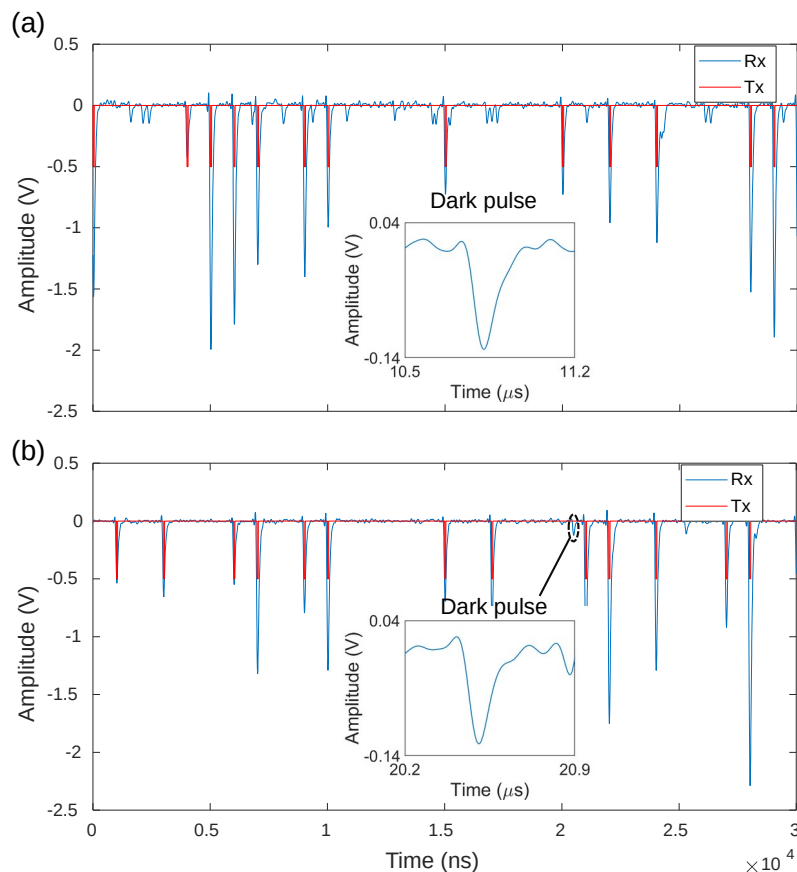


14

15 Figure 3.13 Block diagram of links' experimental setup

16 The link was characterized by transmitting a range of sequences  
17 at different data rates and different received power. Firstly, the  
18 received power was varied from 10 pW to 40 pW with a fixed data  
19 rate of 1 Mbit/s to measure the receiver's sensitivity. The duty cycle  
20 was set as 5%. Both 21 °C and -10 °C conditions were considered  
21 and the corresponding bias voltages were 56.8 V and 55.2 V,  
22 respectively. The overvoltage was constant, so the PDE and gain

1 were fixed at 15.38 % and  $7 \times 10^5$ , respectively. The detection window  
2 was set as 50 ns to match the duty cycle of 5%. The values of the  
3 detection window and the duty cycles are a compromise choice  
4 because they are a trade-off between BER and the transmitter's  
5 bandwidth. The data of the received signal was synchronized with  
6 the transmitted PRBS sequences by correlation in the time domain.

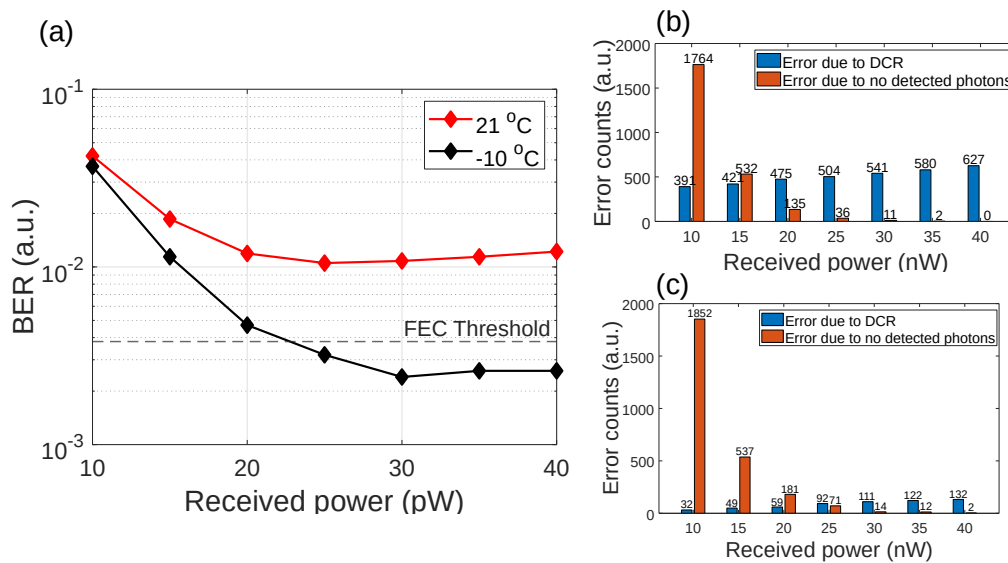


7

8 Figure 3.14 Synchronized transmitted sequences and received  
9 signal at (a) 21 °C (b) -10 °C

10 Fig. 3.14 shows the synchronized transmitted sequences and  
11 received signal during 30  $\mu$ s at both 21 °C and -10 °C. In the figure,  
12 the data rate shown is 1 Mbit/s, and the received power is 30 pW  
13 (corresponding to 41.5 photons per bit). In the transmitted sequence  
14 (red line), the value -0.5 V represents the transmitted bit "1". In the

1 received pulses (blue line), the amplitude of each “1” bit has  
 2 different amplitudes. This is because the number of detected  
 3 photons in each bit is different, following a Poisson distribution. The  
 4 zoomed single dark pulses in Fig. 3.14 (a) and (b) are nearly  
 5 identical, including the shape and amplitude, because of the  
 6 unchanged overvoltage. Moreover, the number of dark pulses is  
 7 significantly reduced from 21 °C to -10 °C.



8

9 Figure 3.15 (a) BER versus received power and temperature (b)  
 10 Error counts versus received power measured at 21 °C (c) Error  
 11 counts versus received power measured at -10 °C

12 The bit decision (whether “1” or “0”) depends on whether there  
 13 is any detected photon within the bit. This is achieved by comparing  
 14 the values of the received signal within the detection window to a  
 15 decision threshold. The decision threshold was set as -125 mV,  
 16 corresponding to approximately 90 % of a single pulse’s peak value  
 17 to ensure the counts from pulses with irregular shapes, without  
 18 losing generality. It is noted that the values of the received signal  
 19 are negative, due to the inverting operation of TIA. Thus, the bit is  
 20 determined to be “1” if the minimum value within the detection

1 window is below the decision threshold. The BER versus received  
2 power and temperature are plotted in Fig. 3.15 (a) for the data rate  
3 of 1 Mbit/s.

4 Each BER point is obtained by transmitting  $50 \times 2^{10}$  bits. The  
5 results show that the BER at  $-10\text{ }^{\circ}\text{C}$  is always lower than the BER at  
6  $21\text{ }^{\circ}\text{C}$  due to the reduction of DCR. The two BER curves each drop to  
7 a minimum value ( $2.4 \times 10^{-3}$  with 30 pW at  $-10\text{ }^{\circ}\text{C}$ ) because the error  
8 due to transmitted "1" detected as "0" decreases (referred to as the  
9 error due to no detected photons in the figure). This error occurs  
10 when there are no detected photons in a bit that should be "1", due  
11 to the Poisson distribution of photon numbers. The two BERs slightly  
12 increase when the received power exceeds 30 pW, due to the  
13 increase of afterpulses. Afterpulses are delayed pulses following  
14 pulses due to detected photons and dark counts, and are generated  
15 with a specific probability when a threshold current from SiPMs is  
16 exceeded [157]. Assuming the optimal power corresponds to a BER  
17 equal to the FEC threshold, the optimal received power is 23 pW.  
18 This corresponds to 32 photons per bit, approaching the quantum  
19 limit [158].

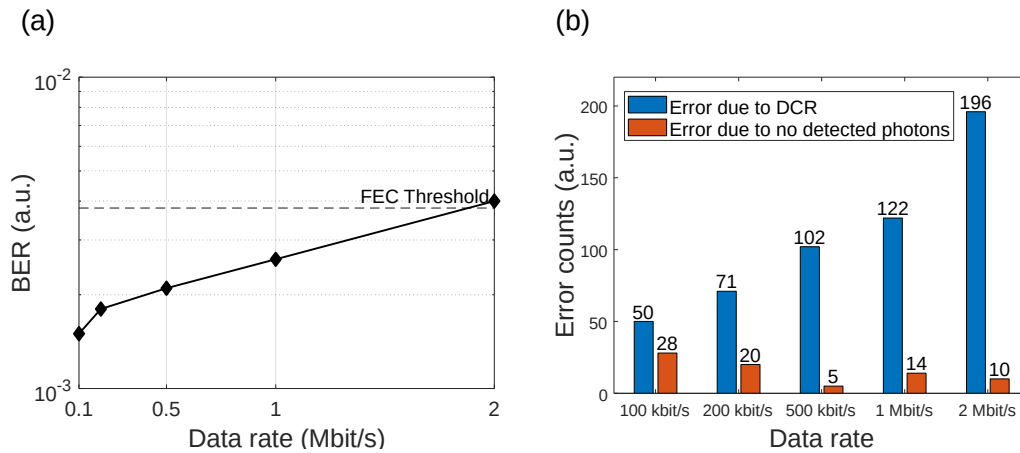
20 These experimental results in Fig. 3.15 (a) are worse than the  
21 simulation shown in Fig. 3.3. When the BER is equal to the FEC  
22 threshold, the received power in Fig. 3.3 is  $\sim 14$  pW while 23 pW in  
23 Fig. 3.15 (a). In order to investigate the reason for the difference  
24 between the simulation and experiment, the numbers of errors due  
25 to DCR and the error due to no detected photons versus received  
26 power are plotted in Fig. 3.15 (b) and (c). The reduction of error due  
27 to no detected photons with increasing received power is because of  
28 the increased number of average detected photons per bit. The  
29 error of DCR slightly increases with received power due to

1 afterpulses. Due to the bandwidth limitation of the detector, the  
2 time-broadened afterpulse falls into the detection window of the  
3 next bit. Thus, inter-symbol interference (ISI) occurs and a  
4 transmitted “0” can be detected as “1” due to the afterpulse from a  
5 previously transmitted bit. Higher received power increases the  
6 proportion of photon pulses exceeding the afterpulses threshold  
7 current and in turn, increases the number of afterpulses, resulting in  
8 higher BER.

9 Table 3.2 Duty cycle for different data rates

Data rate (Mbit/s)	0.1	0.2	0.5	1	2
Duty cycle (%)	0.5	1	2.5	5	10

10 Then, the bit rate was varied from 100 kbit/s to 2 Mbit/s with a  
11 fixed received power per bit (41.5 photons per bit). The  
12 corresponding duty cycles were varied to match the detection  
13 window of 50 ns, which are listed in Table 3.2. The BER versus data  
14 rate is shown in Fig. 3.16 (a). Each BER point is obtained by  
15 transmitting  $50 \times 2^{10}$  bits. The result shows a maximum data rate of 1  
16 Mbit/s with a BER of  $2.4 \times 10^{-3}$ . The corresponding numbers of two  
17 categories of errors are shown in Fig. 3.16 (b). It can be seen that  
18 the error due to DCR is the dominant error and increases with data  
19 rate resulting in the BER at 2 Mbit/s exceeding the FEC threshold.  
20 This is because increasing the data rate will increase the probability  
21 of detecting afterpulses and ISI is more severe at higher data rates  
22 because the response of the receiver is longer than the bit duration  
23 [159].



1  
2 Figure 3.16 (a) BER versus data rate (b) Error counts versus data  
3 rate

### 4 3.6 CONCLUSIONS

5 This Chapter details work on building ultra-sensitive UV solar  
6 blind OWC links, including light source characterization, receiver  
7 design and a link demonstration. Overall, a 1 Mbit/s OOK UV-C OWC  
8 communication link was demonstrated with a measured BER of  
9  $2.4 \times 10^{-3}$  and an average received photon number per bit of 41.5.  
10 This shows the potential for building ultralow photon flux links for a  
11 UV-C QKD system.

12 However, the receiver's sensitivity shown in this Chapter is not  
13 high enough to reach single-photon link level performance. The  
14 receiver design needs to be improved to further reduce the SiPM  
15 temperature for a lower DCR. The next Chapter will show an  
16 upgraded receiver design and ultralow photon flux links that are  
17 feasible for the final UV-C wireless QKD.

## 1 **Chapter 4:**

# 2 **Ultralow Photon Flux Links Using UV-C**

## 3 **Single-Photon Detection**

### 4 **4.1 INTRODUCTION**

5 Ultralow photon flux links have applications in such areas as  
6 deep-space communication and QKD [23, 160-162]. Building on the  
7 foundation established in Chapter 3, this Chapter will further reduce  
8 DCR to enhance SiPM-based receiver sensitivity to a single-photon  
9 sensitivity level.

10 In addition to the SPD, a high-speed transmitter is another  
11 significant element for building high data rate UV-C ultralow photon  
12 flux links. Due to the lack of a commercial UV-C LD and the low  
13 speed of commercial UV-C LEDs (with a typical bandwidth of  $\leq 10$   
14 MHz), micro-LEDs become an attractive option. Micro-LEDs have  
15 been used to implement high-speed data links [124, 125, 163, 164],  
16 and are available in the UV-C wavelength band of interest. In this  
17 region AlGaIn micro-LEDs with a diameter of around 100  $\mu\text{m}$  have  
18 been reported [125, 163]. They exhibit a peak wavelength of  $\leq 280$   
19 nm, an output power of  $\leq 1$  mW, and a bandwidth exceeding 100  
20 MHz [124, 125]. Over 1 Gbit/s OWC links using UV-C micro-LEDs  
21 have been reported and the maximum transmission distance has  
22 reached 100 m [146, 165, 166].

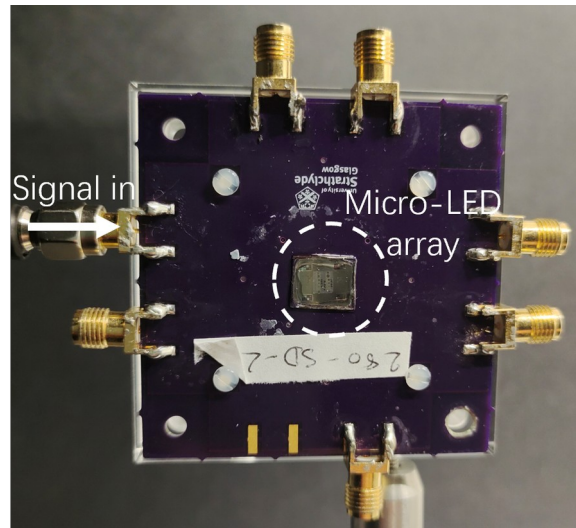
23 In this Chapter, a UV-C micro-LED with a bandwidth of over 400  
24 MHz and an enhanced SiPM-based UV-C SPD are introduced. Then,

1 ultralow photon flux links, including pulse position modulation (PPM)  
2 links and single-photon links are described.

### 3 **4.2 UV-C MICRO-LED CHARACTERIZATION**

4 As reported in Chapter 3, 1 Mbit/s is the maximum data rate  
5 that can be achieved for an OOK link [52], because of the low  
6 modulation bandwidth of the commercial UV-C LED (Thorlabs  
7 LED275J) used and the high DCR of the SiPM (HAMAMATSU S13360-  
8 3025UQ-SPL). In order to achieve a higher data rate link, a new  
9 high-bandwidth UV light source is a key goal. This section will  
10 introduce a high-bandwidth UV-C micro-LED.

11 UV-C micro-LEDs have been widely reported [124, 125, 163,  
12 164]. There are several potential materials for manufacturing these  
13 devices, including AlGa<sub>N</sub> [125, 163], ZnO [167], Ga<sub>2</sub>O<sub>3</sub> [168] and  
14 perovskite [169]. AlGa<sub>N</sub>-based UV-C LEDs have demonstrated the  
15 best performance. The emission wavelengths for these devices  
16 range from 200 nm to 400 nm, due to the bandgaps of Ga<sub>N</sub> and Al<sub>N</sub>  
17 of 3.4 eV and 6.2 eV respectively. By tuning the Al content within the  
18 AlGa<sub>N</sub>-based alloy material, the emission wavelength can be  
19 continuously adjusted within the UV range of wavelengths.



1

2 Figure 4.1. UV-C micro-LED array image

3 The AlGaN-based micro-LED used in this Chapter is a micro-LED  
4 array comprising three different diameters of micro-LEDs (100  $\mu\text{m}$ ,  
5 200  $\mu\text{m}$  and 300  $\mu\text{m}$ ). These micro-LEDs are manufactured at the  
6 Institute of Photonics, Department of Physics, SUPA, University of  
7 Strathclyde [125]. An image of the device is shown in Fig. 4.1. The  
8 three micro-LEDs are bonded at the central region of the PCB, with  
9 each micro-LED being connected to an externally mounted SMA  
10 port.

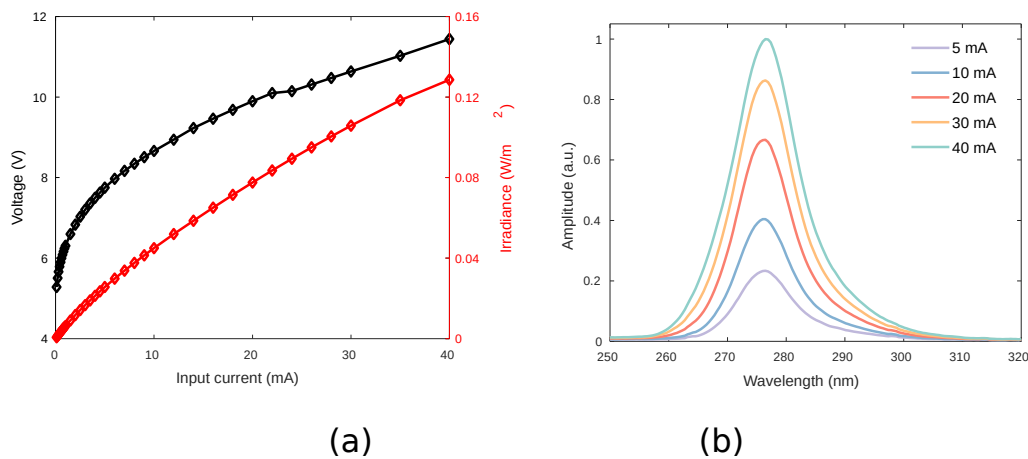
11 The 100  $\mu\text{m}$  micro-LED (the smallest available) is selected to  
12 pursue the highest data rate because the micro-LED's modulation  
13 bandwidth increases with decreasing diameter [163]. Note that all  
14 micro-LEDs mentioned in the rest of this Chapter refer to the 100  
15  $\mu\text{m}$  micro-LED.

16 An initial experiment to characterize these devices was  
17 undertaken. The micro-LED was driven by a source meter (KEITHLEY  
18 2636B), and the irradiance was measured by a power meter  
19 (Newport 843-R with a UV-enhanced Si head 818 UV/DB) to obtain  
20 the I-V-I curves. The divergence angle of the micro-LED is over 120°,

3

4

1 so the irradiance was measured instead of power (because the  
2 emitted power of the micro-LED is unable to be fully collected by the  
3 Si head even when they are placed nearby) [163]. The distance  
4 between the micro-LED and Si head is 6.5 cm. Fig. 4.2 (a) presents  
5 the I-V-I curves of the device. The result shows the turn-on voltage  
6 of the micro-LED is approximately 5 V.



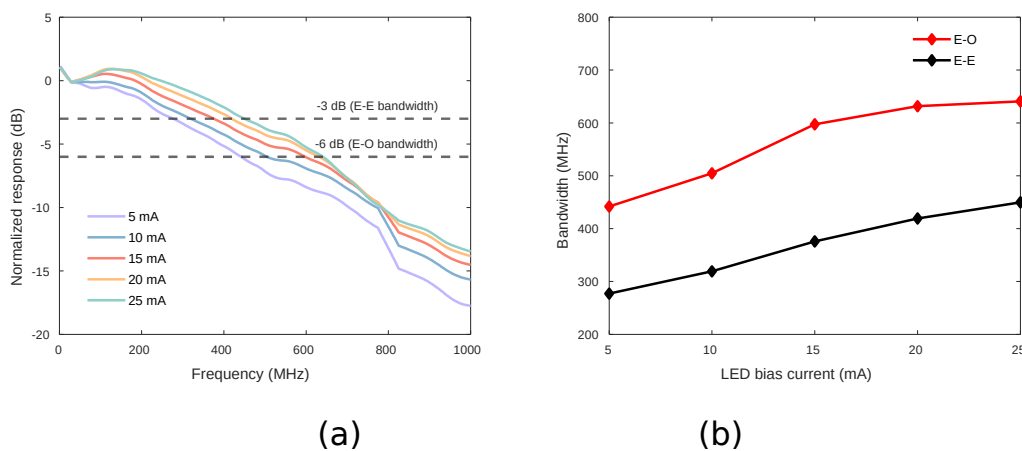
7  
8 (a) (b)  
9 Figure 4.2 (a) Plot of I-V and I-I for the micro-LED (b) Emission  
10 spectrum for a range of driving current from 5 mA to 40 mA

11 The emission spectrum of the micro-LED was measured by using  
12 an optical spectrum analyzer (Ocean Insight OFX00951). Fig. 4.2 (b)  
13 shows the spectrum versus wavelength with driving current varying  
14 from 5 mA to 40 mA. The spectrum covers both parts of the UV-B  
15 and UV-C regions and the FWHM is approximately 11 nm. The  
16 emitted peak wavelength is 276 nm.

17 The bandwidth of the micro-LED was then measured. Signals  
18 were generated by a vector network analyzer (VNA, Agilent  
19 Technologies FieldFox N9923A) and DC bias was provided by a  
20 power supply (ROHDE & SCHWARZ HMP4040). These were  
21 combined by a Bias Tee (Mini-Circuits ZFBT-6GW) and then used to  
22 drive the micro-LED. Two convex lenses (Thorlabs LA4052-UV)

1 collimated light from the micro-LED and focused it on the active  
 2 area of a 1 GHz bandwidth optical receiver (HAMAMATSU C5658 APD  
 3 module). The output from the receiver was then connected to the  
 4 VNA.

5 The 3-dB bandwidth was obtained by sweeping the S21 value  
 6 versus frequency. Fig. 4.3 (a) shows the normalized response of the  
 7 VNA's S21 value with the micro-LED bias current ranging from 5 mA  
 8 to 25 mA. The electrical-to-electrical (E-E) is defined as the  
 9 frequency at which there is a 3dB attenuation of electrical power  
 10 relative to a reference, and the (E-O) bandwidth is defined as the  
 11 frequency at which the optical power drops by 3dB relative to a  
 12 reference. As the VNA measures the E-E response the 3dB E-O  
 13 bandwidth corresponds to a 6dB drop in E-E response.



14

15

16

17

18

(a) Normalized response of micro-LED versus frequency (b) Measured E-E and E-O bandwidth versus LED bias current

19

20

21

22

The measured E-E and E-O bandwidth values versus LED bias current are shown in Fig. 4.3 (b). This shows that as the bias current increases from 5 mA to 25 mA, the E-E bandwidth increases from 277 MHz to 450 MHz, and the E-O bandwidth increases from 442

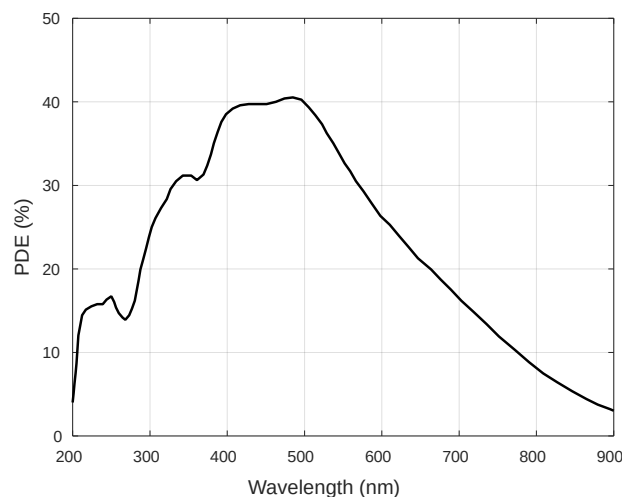
1 MHz to 641 MHz.

## 2 **4.3 UV-C SPD DESIGN AND CHARACTERIZATION**

### 3 **4.3.1 SiPM Selection**

4 The SiPM used in Chapter 3 has an active area of 3 mm x 3 mm,  
5 which has a DCR of approximately 400 kcps at room temperature.  
6 Reducing the temperature to about -10 °C, the measured DCR was  
7 still 35 kcps, which is higher than is required. There are two  
8 methods to further reduce the DCR: (1) reduce the active area of  
9 the SiPM and (2) further reduce the temperature of the SiPM.

10 A new SiPM (HAMAMATSU S13360-1350 SPL) with an active area  
11 of 1.3 mm x 1.3 mm was used for the enhanced SPD design. This  
12 SiPM is not usually used for UV-C applications but was custom-  
13 packaged in a TO-18 can with a quartz window by the manufacturer.  
14 This package has high transmission in the UV-C region. The reduced  
15 area of the device leads to a DCR of 90 kcps at room temperature  
16 according to the datasheet.



17

18

Figure 4.4 PDE of HAMAMATSU 13360-1350SPL

19

The device is available with a range of pixel sizes (25  $\mu\text{m}$ , 50  $\mu\text{m}$ )

3

4

1 and 75  $\mu\text{m}$ ), and a pixel size of 50  $\mu\text{m}$  was chosen as it provided a  
2 balanced trade-off between bandwidth and PDE. Typically, smaller  
3 pixels offer higher bandwidth but lower PDE, and 50  $\mu\text{m}$  pixels  
4 provide a good compromise between these two [154]. The PDE of  
5 the HAMAMATSU S13360-1350 SPL chip is shown in Fig. 4.4. The  
6 diagram shows the PDE at 280 nm is approximately 15 %.

### 7 **4.3.2 PCB Design**

8 A custom PCB was designed for the receiver. This has two  
9 functions: (1) supplying a bias voltage to the SiPM and (2)  
10 amplifying the signal generated by the SiPM sufficient for detection  
11 by an oscilloscope.

#### 12 4.3.2.1 Bias Voltage Supply

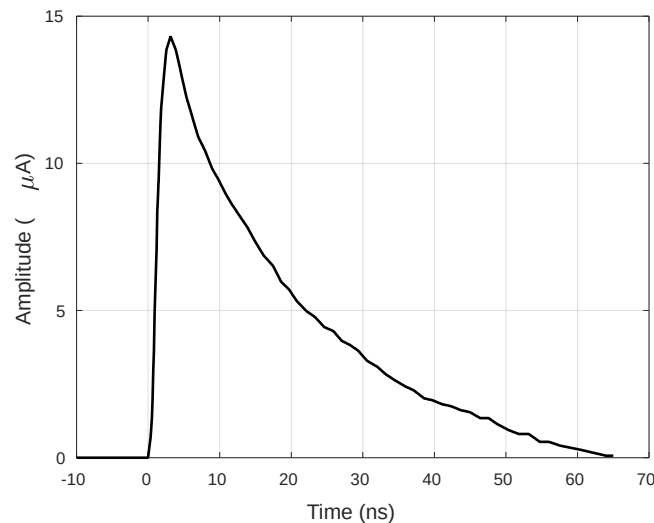
13 The bias voltage is supplied by an on-chip power supply module  
14 (HAMAMATSU C11204-01). The output voltage ranges from 20 V to  
15 90 V, which is sufficient for HAMAMATSU S13360 series SiPMs with a  
16 breakdown voltage between 50 V and 60 V. The output voltage is  
17 adjusted by software running on a connected PC.

#### 18 4.3.2.2 Signal Amplification

19 There are three requirements for signal amplification: (1)  
20 sufficient gain for observing pulses resulting from single-photon  
21 detection using an oscilloscope, (2) high bandwidth to prevent pulse  
22 distortion and (3) low noise.

23 The amplifier gain was determined based on the amplitude of a  
24 single-photon pulse. The data of a single pulse from a HAMAMATSU  
25 S13360-1350 SPL is not provided. Thus, the data is obtained from a  
26 HAMAMATSU 13360-3050 CS SiPM datasheet. This device has the  
27 same pixel size as the SiPM we are considering (so their single pulse

1 outputs should be the same). Fig. 4.5 shows the current of a single  
2 pulse versus time (taken from the datasheet). The peak amplitude is  
3 around 14.3  $\mu\text{A}$ . Considering the resolution of a typical oscilloscope  
4 (e.g. AGILENT MSO6104A) of 2 mV, an amplifier with a trans-  
5 impedance load of at least 150  $\Omega$  is needed to reach this minimum  
6 level.



7

8 Figure 4.5 Current of a single pulse versus time of an SiPM with 50  
9  $\mu\text{m}$  pixel

10 The bandwidth of the amplifier should be no less than the  
11 bandwidth of the SiPM. Based on the single-pulse shape shown in  
12 Fig. 4.5, the bandwidth of the SiPM can be calculated from the rise  
13 time of a single pulse by:

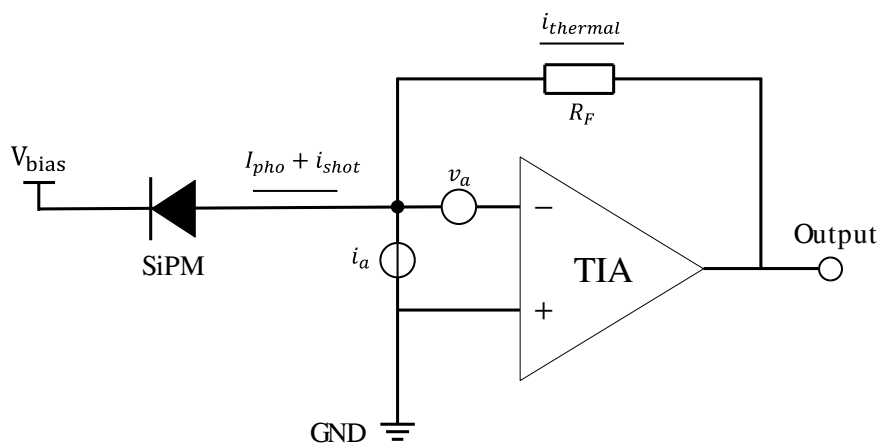
14 
$$BW_{SiPM} \approx 0.35 / \text{rise time} \quad (4.1)$$

15 With a rise time of 1.7 ns, the SiPM bandwidth is calculated as  
16 approximately 200 MHz. Thus, the bandwidth of the amplifier should  
17 be around this value. A TIA (Texas OPA855-Q1) was chosen because  
18 of its high GBWP of 8 GHz and low amplifier noise. The trans-  
19 impedance was chosen as 1500  $\Omega$  to achieve a sufficiently high

1 single pulse amplitude to be clearly distinguished from a typical  
 2 oscilloscope's noise floor. Considering the SiPM's capacitance of 60  
 3 pF, the estimated bandwidth of the amplifier is calculated as  
 4 approximately 120 MHz (corresponding to a rise time of 2.9 ns).  
 5 Thus, there is some distortion when a single pulse from the SiPM  
 6 passes through the amplifier due to the bandwidth being reduced  
 7 from 200 MHz to 120 MHz. However, the key factor that affects the  
 8 performance of ultralow photon flux links is the single pulse width,  
 9 which is dominated by the fall time (~40 ns). Thus, the predicted  
 10 increase in the rise time from 1.7 ns to 2.9 ns is acceptable.

11 4.3.2.3 Noise Analysis

12 A typical TIA noise model schematic is shown in Fig. 4.6 [170].  
 13 The total noise consists of amplifier voltage noise, amplifier current  
 14 noise, thermal noise  $i_{thermal}$  and shot noise  $i_{shot}$ .  $i_a$  is the input-referred  
 15 current noise and  $v_a$  is the input-referred voltage noise.



16

17 Figure 4.6 A typical TIA noise model schematic

18 The amplifier voltage noise and current noise can be combined  
 19 as an amplifier noise [171]:

1  $i_{\text{amplifier}} = \left( i_a + \pi C_i v_a f_{3dB} \right) \cdot \sqrt{f_{3dB}}$  The thermal noise mainly comes from the  
 (4.2)

2 trans-impedance load:

3  $i_{\text{thermal}} = \sqrt{\frac{4 k_B T_R f_{3dB}}{R_F}}$   
 (4.3)

4 Where  $k_B$  is the Boltzmann constant and  $T_R$  is the temperature of the  
 5 trans-impedance load. The shot noise can be obtained by:

6  $i_{\text{shot}} = \sqrt{2q I_{pho} f_{3dB}}$   
 (4.4)

7 Where  $q$  is the elementary charge and  $I_{pho}$  is the photocurrent into  
 8 the amplifier. Thus, the total noise is:

9  $i_{\text{noise}}^2 = i_{\text{amplifier}}^2 + i_{\text{thermal}}^2 + i_{\text{shot}}^2$   
 (4.5)

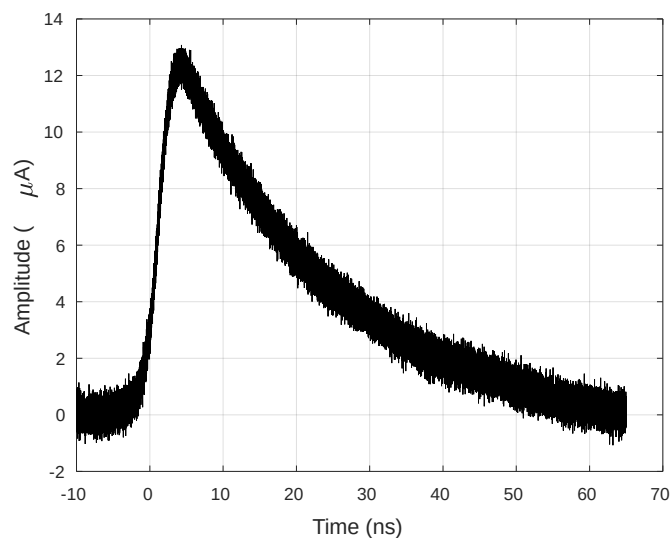
10 To verify the feasibility of using the OPA855-Q1 for the UV-C  
 11 SPD, the noise was estimated using the parameters in Table 4.1.

12 Table 4.1 Parameters and values used for simulating noise

Parameters	Values
Bandwidth	120 MHz
Input-referred current noise	2.5 pA/ $\sqrt{\text{Hz}}$
Input-referred voltage noise	0.98 nV/ $\sqrt{\text{Hz}}$
Temperature of the trans-impedance load	300 K
Trans-impedance load	1500 $\Omega$ (refer to section 4.3.2.2)

13 The input-referred current noise and input-referred voltage noise

1 were obtained from the datasheet of the TIA. The temperature of the  
2 trans-impedance load was assumed to be room temperature. The  
3 calculated current version of input amplifier noise, thermal noise  
4 and shot noise are 0.24  $\mu\text{A}$ , 36 nA and 24 nA, respectively. These  
5 noises were used to simulate a noisy received single pulse. The  
6 single pulse was also filtered using a first-order Butterworth filter  
7 with a cutoff frequency of 120 MHz, matching the amplifier  
8 bandwidth described in section 4.3.2.2. The corresponding  
9 simulation result is shown in Fig. 4.7. The results show the pulse  
10 shape is distinguishable after adding the simulated noise. The peak  
11 amplitude of the single pulse in Fig. 4.7 is slightly lower than that in  
12 Fig. 4.5. This is because the bandwidth is reduced from 200 MHz  
13 (see Fig. 4.5) to 120 MHz. The peak amplitude of a single pulse at  
14 the amplifier output is expected to be 18 mV with a 1500  $\Omega$  trans-  
15 impedance gain. However, considering the 50  $\Omega$  output resistance of  
16 the amplifier and the 50  $\Omega$  input impedance of the oscilloscope, the  
17 peak amplitude displayed on the oscilloscope is expected to be 9  
18 mV. Therefore, the amplifier design with OPA855-Q1 and a 1500  $\Omega$   
19 trans-impedance load should satisfy the requirements of the system.



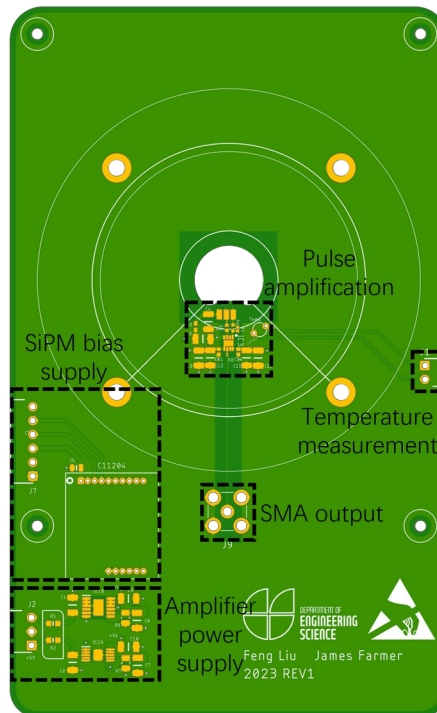
20

21 Figure 4.7. Simulated amplitude of a single pulse with noise versus

3  
4

1 time

2 4.3.2.4 PCB Production



3

4 Figure 4.8 Layout of PCB

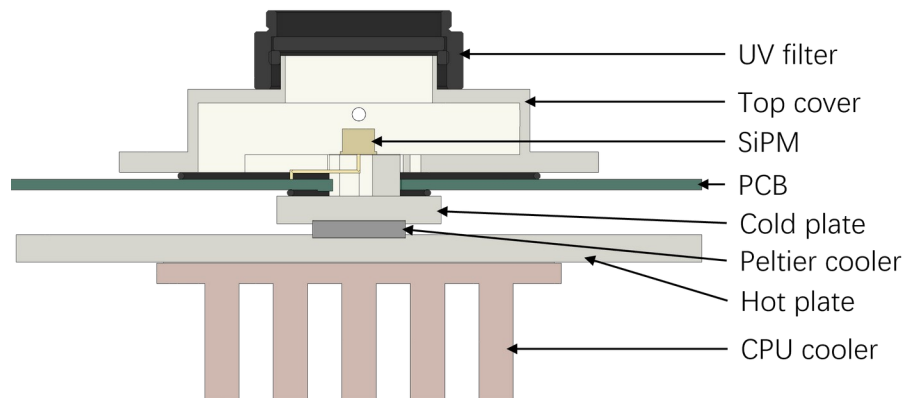
5 A PCB that contains the amplifier and SiPM bias was designed  
6 and produced. This is a 4-layer board, which allows multiple ground  
7 layers to prevent interference and ensure good low-noise  
8 performance. Fig. 4.8 shows the layout of PCB. The SiPM bias supply  
9 uses an integrated HAMAMATSU C11204-01 to provide a bias  
10 voltage to the SiPM. There are two regulators (LT3093 and LT3042)  
11 within the amplifier power supply, supplying +2.5 V and -2.5 V to  
12 the amplifier. A Texas OPA855-Q1 is used as a TIA to amplify pulses  
13 from the SiPM. The temperature of the SiPM is measured by a sensor  
14 and a signal is sent to the TEC controller. The output of the TIA is  
15 connected to the oscilloscope via a SMA connection.

16 **4.3.3 Liquid Cooling Design**

3

4

1 Chapter 3 describes a liquid cooling design with a separate SiPM  
2 and amplifier. The relatively long cable connecting the SiPM and the  
3 amplification circuit will introduce capacitance and potential  
4 interference. For these reasons, a modified liquid cooling design was  
5 developed, with the SiPM and amplifier mounted on a single PCB. A  
6 sectional view of the design is shown in Fig. 4.9.



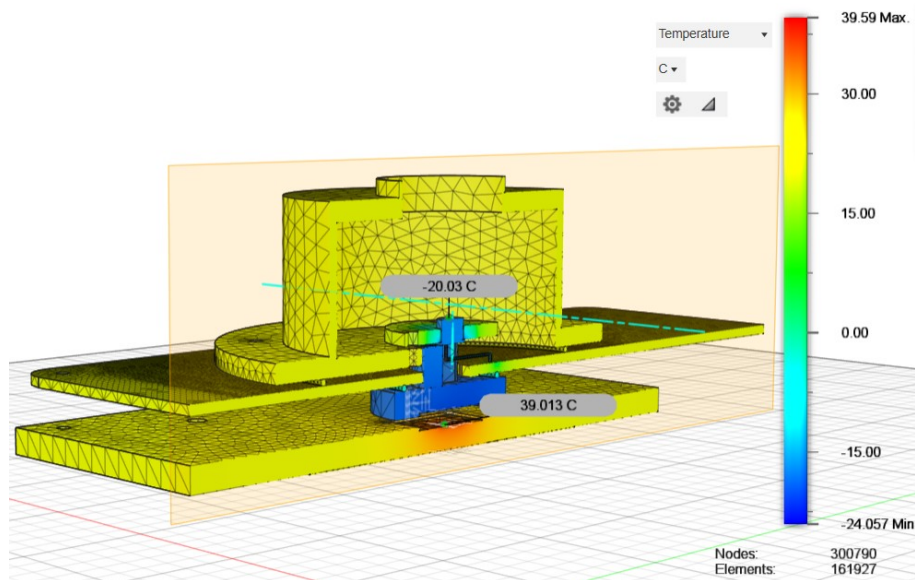
7

8 Figure 4.9 Sectional view of modified cooling design

9 The central component is a Peltier cooler (OTX20-68-F1A-1313-  
10 11-W2.25). On the cold side, a cold plate with a “pedestal” that  
11 passes through a hole in the PCB, establishes a thermal connection  
12 between the SiPM and the Peltier cooler. The cold plate is fabricated  
13 from ceramic aluminium oxide because of its high thermal  
14 conductivity ( $\sim 30$  W/mK) and effective electrical isolation. Because  
15 the temperature will be reduced to below  $0^\circ\text{C}$ , to prevent  
16 condensation, a plastic top cover and a UV filter (Thorlabs FGUV5M)  
17 are used to create a confined space around the SiPM. A vacuum  
18 pump (RS vacuum pump) is used to evacuate the air from the  
19 confined space. In addition, O-rings are used to ensure a good seal.

20 On the hot side, the heat flows to a liquid CPU cooler (MASTER  
21 LIQUID LITE 240) through the hot plate, which is also made of  
22 aluminium oxide. The purpose of the CPU cooler is to remove heat

1 from the hot side, rather than to directly cool the temperature of the  
2 SiPM. To achieve temperature control, an NTC thermistor (TDK  
3 B57540G1103F005) is placed on the cold plate close to the SiPM  
4 using thermal adhesive (Electrolube TBS20S). A TEC controller  
5 (Meerstetter Engineering TEC-1161-4A) obtains the temperature  
6 value from the thermistor through the PCB (see section 4.3.2.4) and  
7 dynamically drives the Peltier cooler to control the temperature of  
8 the SiPM.



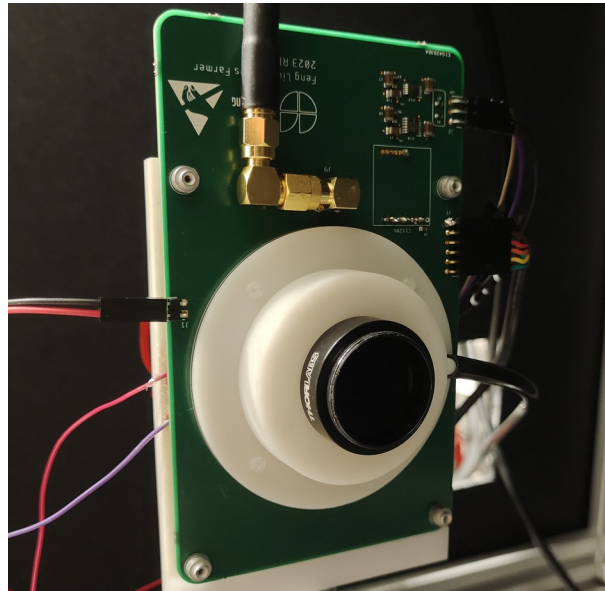
10 Figure 4.10 Sectional view of the liquid cooling structure with  
11 simulated temperature

12 A thermal simulation was carried out to estimate the  
13 performance of the liquid cooling design using Autodesk Fusion 360.  
14 The cold side (hot side) of the Peltier cooler was a cold (hot) heat  
15 source with a power of -1.6 W (10.95 W). This value was obtained  
16 from the datasheet of the Peltier cooler, assuming a driving current  
17 of 1.4 A. The convective heat transfer coefficient of the surface  
18 between the hot plate and the CPU cooler was  $1600 \text{ W/m}^2 \cdot \text{K}$   
19 (obtained from the datasheet of the CPU cooler). The convective  
20 heat transfer coefficient of other surfaces was  $20 \text{ W/m}^2 \cdot \text{K}$ , which is a

1 typical value of natural convection of air [172]. A sectional view of  
2 the liquid cooling structure with simulated temperature is shown in  
3 Fig. 4.10. The result shows the temperature of the SiPM is  
4 approximately  $-20\text{ }^{\circ}\text{C}$ , while the hot side of the Peltier cooler  
5 stabilizes at around  $40\text{ }^{\circ}\text{C}$ , thanks to the effective heat dissipation by  
6 the CPU cooler.

#### 7 4.3.4 UV-C SPD Characterization

8 An image of the UV-C SPD is shown in Fig. 4.11. The UV-C SPD  
9 was mounted on the optical table within a light-tight enclosure to  
10 ensure a dark operating environment.

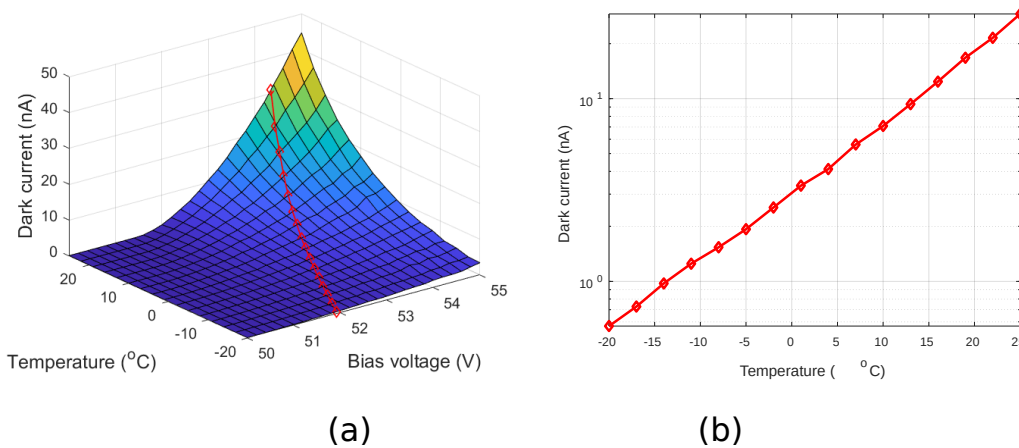


11

12 Figure 4.11 An image of the UV-C SPD

13 Dark current versus temperature was measured to test the UV-C  
14 SPD. The PCB and TEC controller were remotely controlled by a PC  
15 to change the bias voltage and temperature, respectively. The dark  
16 current was measured by a source meter (KEITHLEY 2636B). The  
17 bias voltage was varied (from  $54.34\text{ V}$  to  $51.91\text{ V}$ ) with temperature  
18 (from  $25\text{ }^{\circ}\text{C}$  to  $-20\text{ }^{\circ}\text{C}$ ) to ensure the overvoltage remained constant.  
19 This ensured constant PDE and gain.

1 Fig. 4.12 (a) shows the dark current of the SiPM changing with  
2 bias voltage and temperature. The red curve in Fig. 4.12 (b) shows  
3 the dark current versus temperature under the recommended bias  
4 voltage. The dark current reduces by approximately 11 dB with a 30  
5 °C temperature reduction, consistent with the result reported in  
6 [153] and section 3.4.4. The dark current at 25°C and -20°C are  
7 29.08 nA and 0.57 nA, respectively. Further reducing the  
8 temperature is challenging, limited by the capability of the Peltier  
9 cooler, CPU cooler and SiPM construction.



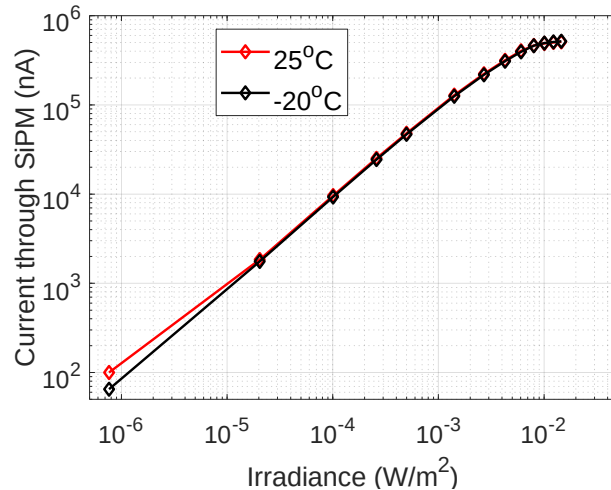
10

11

12 Figure 4.12 (a) Dark current versus bias voltage and temperature (b)  
13 Dark current versus temperature under recommended bias voltage

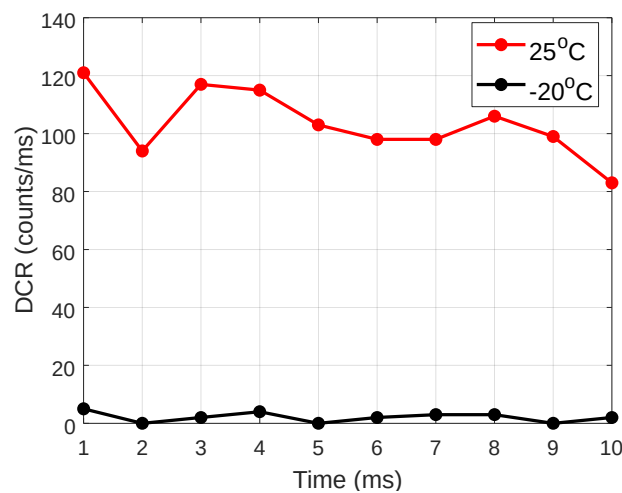
14 Then, the responsivity of the SiPM and DCR were measured to  
15 obtain the PDE and gain. A power meter (Newport 843-R with a UV-  
16 enhanced Si head 818 UV/DB) was used to measure the irradiance  
17 from the micro-LED. Subsequently, the UV-C SPD replaced the  
18 power meter at the same position. A source meter (KEITHLEY  
19 2636B) was used to measure the output current of the SiPM. The  
20 current from the SiPM, including both photocurrent and dark  
21 current, varying with irradiance at both 25 °C and -20 °C is shown in  
22 Fig. 4.13. It can be seen that the two curves overlap each other well,  
23 indicating that the PDE and gain remain constant at 25°C and -20°C.

1 The SiPM approaches saturation when the irradiance exceeds 10  
2 mW/m<sup>2</sup>, corresponding to a received power of 17 nW. The difference  
3 between the two curves at the beginning reflects different dark  
4 currents at these two temperatures.



5

6 Figure 4.13 Current from the SiPM versus irradiance at the aperture  
7 of the UV-C SPD



8

9 Figure 4.14 DCR versus time

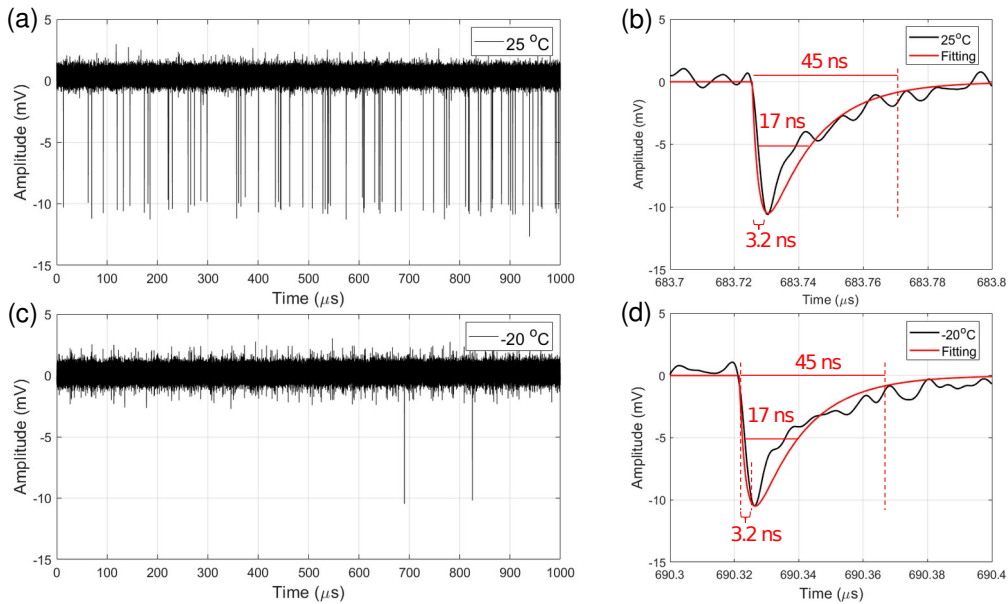
10 The output from the amplifier was captured by an oscilloscope  
11 (AGILENT MSO6104A) to obtain the DCR. The number of dark counts  
12 during a 10 ms period, summed every 1 ms, is shown in Fig. 4.14.  
13 The DCR at 25°C and -20°C are calculated as 103.4 kcps and 2.1

3

4

1 kcps, respectively. Considering the dark current values in Fig. 4.12,  
 2 the gain of the SiPM is calculated as  $1.76 \times 10^6$  and  $1.7 \times 10^6$  at 25°C and  
 3 -20°C. The PDE is calculated as 14.24% and 14.76% at 25°C and -  
 4 20°C.

5 The captured dark pulses at 25°C and -20°C are plotted in Fig.  
 6 4.15. Fig. 4.15 (a) and (c) represent dark pulses within a 1 ms  
 7 period, showing a significant reduction in the number of dark counts  
 8 with cooling. Fig. 4.15 (b) and (d) show two zoomed-in single pulses  
 9 at 25°C and -20°C, respectively.



10

11 Figure 4.15 Dark pulses during 1 ms period and zoomed-in single  
 12 pulses under 25°C and -20°C

13 Exponential fits to the pulse shape are also shown in Fig. 4.15.  
 14 These are defined by [160]:

$$A_f(t) = A_0 \cdot \left( e^{-\frac{t}{T_r}} - e^{-\frac{t}{T_f}} \right) \quad (4.6)$$

16 Where  $A_0$  is the amplitude of the single pulse.  $T_r$  and  $T_f$  are the  
 17 rise time and fall time of the single pulse, respectively. The fitting  
 18 curves show that both pulses have a peak amplitude of  
 19 approximately 10.5 mV, with a rise time (from 10% to 90% peak

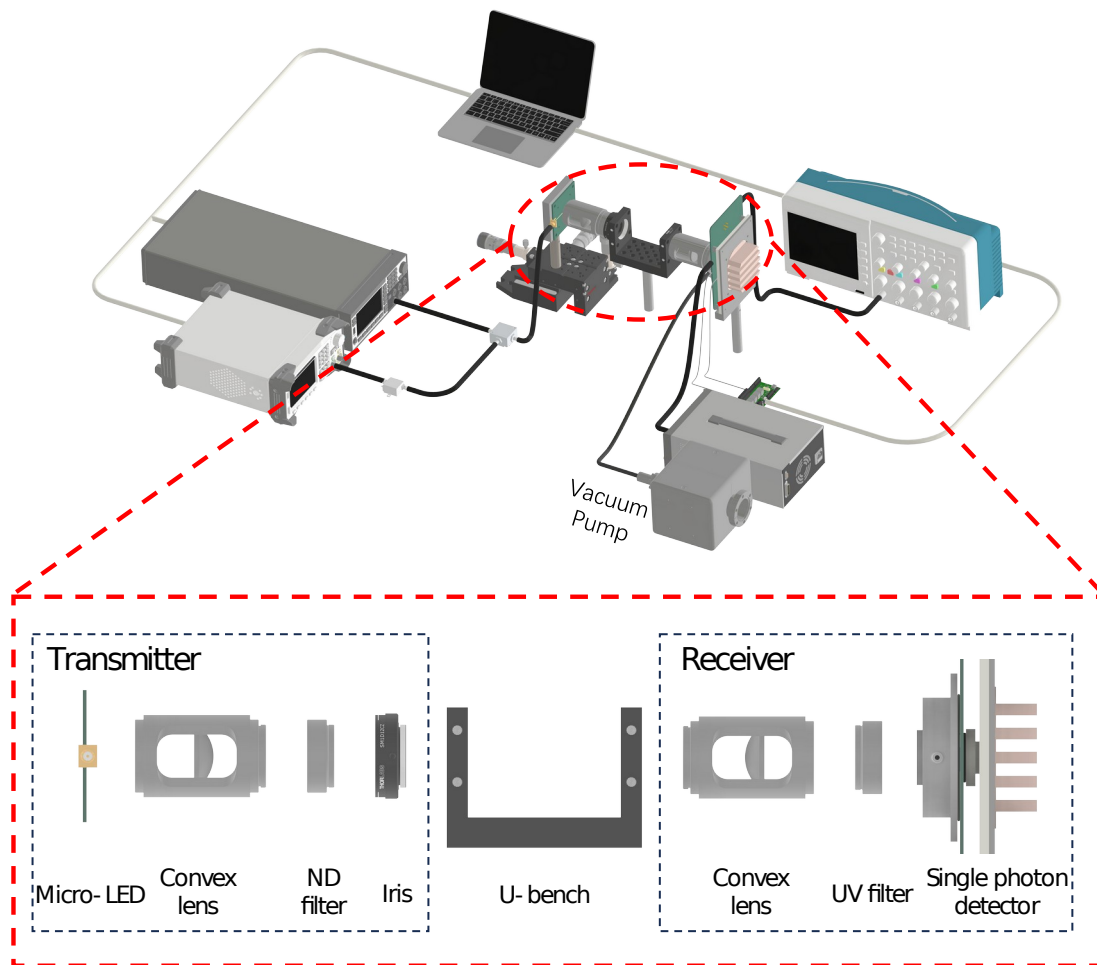
1 amplitude) of 3.2 ns. Thus, the practical 3-dB bandwidth of the UV-C  
2 SPD is estimated at 110 MHz, according to Eq. (4.1). The FWHM of  
3 the single pulse is 17 ns. The dead time of each pixel is  
4 approximately 45 ns, with the dominant contributor being the fall  
5 time, constrained by the SiPM's quenching time.

## 6 **4.4 ULTRALOW PHOTON FLUX OWC LINKS**

7 Ultralow photon flux links were demonstrated to test the  
8 sensitivity of the UV-C SPD and the feasibility of building a UV-C QKD  
9 link using the UV-C micro-LED and the UV-C SPD.

### 10 **4.4.1 Experimental Testbed**

11 The experimental setup for the ultralow photon flux links  
12 demonstration is shown in Fig. 4.16.



1

2 Figure 4.16 Experimental setup for ultralow photon flux links  
 3 demonstration

4 At the transmitter, a modulated signal was generated by an  
 5 arbitrary waveform generator (AWG, Tektronix AWG70002) and  
 6 amplified by an RF amplifier (iXblue DR-AN-20-MO) with 20 dB gain.  
 7 An RF amplifier is used, as the output voltage of the AWG is limited  
 8 to between 250 mV and 500 mV, which is insufficient to modulate  
 9 the micro-LED. Then, the amplified modulated signal and DC bias  
 10 from the source meter were combined by a Bias Tee. The combined  
 11 signal was connected to the micro-LED. A convex lens was placed in  
 12 a slotted tube (Thorlabs SM1L20C) and used to collimate the light  
 13 emitted from the micro-LED. A 3-axis stage (Thorlabs RBL13D/M)

3  
 4

1 was used to mount and adjust the micro-LED's position, ensuring  
2 precise collimation by the convex lens. An ND filter (Thorlabs  
3 NDUV30A) and an iris (Thorlabs SM1D12CZ) were employed to  
4 attenuate the optical power to a specific value. The diameter of the  
5 light beam was constrained to 6 mm by the iris. A U-bench (Thorlabs  
6 CBB1/M) connected the transmitter and the receiver for precise  
7 alignment.

8 At the receiver, a convex lens focused the collimated light onto  
9 the active area of the SiPM. The UV filter (Thorlabs FGUV5M) at the  
10 entrance to the SiPM permitted only UV light to enter. The received  
11 power can be calculated as:

$$12 \quad P_{RX} = P_{TX} \cdot T_{lens} \cdot T_{filter} \quad (4.7)$$

13 Where  $P_{TX}$  is the transmitted power.  $T_{lens}$  and  $T_{filter}$  are the  
14 transmission of the convex lens and UV filter, which are 92% and  
15 83%, respectively. These values were measured by placing the  
16 convex lens and UV filter (separately) between the UV-C micro-LED  
17 and a power meter and comparing transmitted power against the  
18 value with no component present.

19 The received pulses were displayed on the oscilloscope and  
20 captured for post-processing. The source meter, AWG, TEC  
21 controller, oscilloscope and PCB were all remotely controlled by a  
22 PC.

#### 23 **4.4.2 Cascaded Post-Processing Strategy**

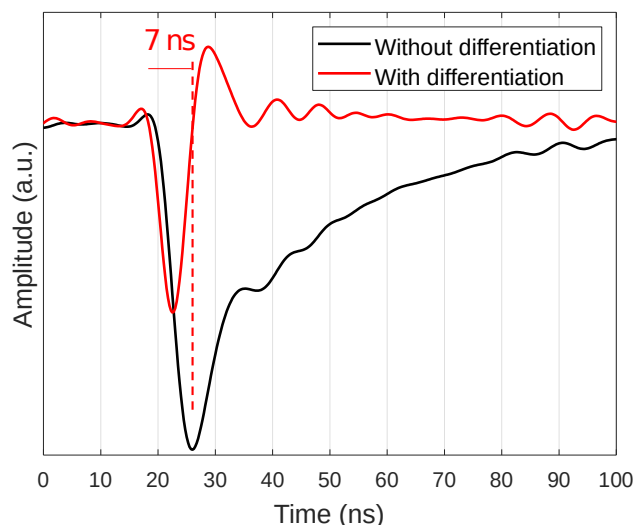
24 There are two post-processing strategies used:

- 25 ● Direct post-processing: The bit decision is achieved directly by  
26 comparing the minimum amplitude of the received signal within

1 a detection window with a decision threshold. (Note that the  
2 received signal is inverted, so the minimum signal represents a  
3 transmitted symbol)

4 ● Cascaded post-processing (CPP): The received signal is  
5 differentiated to mitigate the ISI due to the long fall time of  
6 pulses [173]. Then, the differentiated signal values are summed  
7 within a detection window. For OOK links, the bit decision is  
8 achieved by comparing the summed value with a threshold. For  
9 PPM links, the symbol decision is achieved by finding the  
10 minimum summed value among all time slots within a symbol.

11 The key to the CPP strategy is the differentiation. Previous  
12 research has proven that differentiation can efficiently mitigate the  
13 ISI caused by the long fall time of pulses [173, 174]. Fig. 4.17 shows  
14 a received pulse with and without differentiation.



15

16 Figure 4.17 Amplitude of a received pulse with and without  
17 differentiation

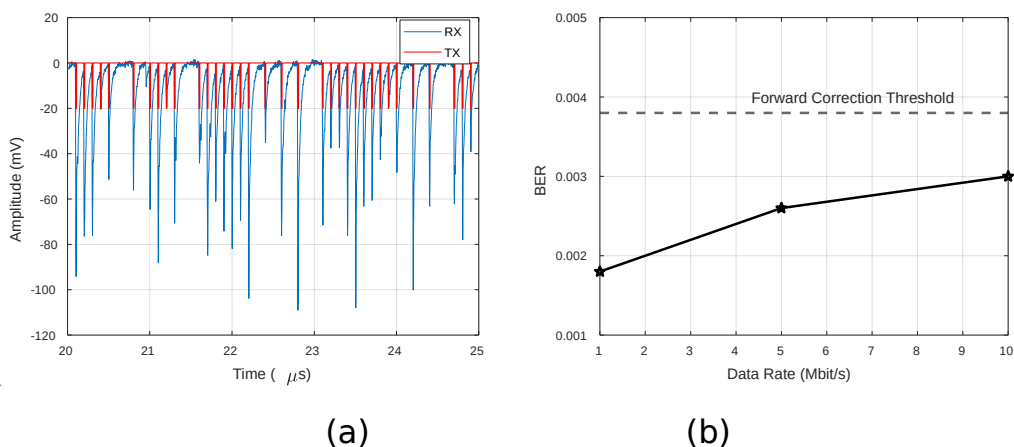
18 The time duration of a single pulse is reduced to approximately  
19 7 ns with differentiation, compared with 45 ns without  
20 differentiation. The effect of the positive peak following the negative

1 peak is negligible because only the points within the detection  
 2 windows are used for the symbol/bit decision. In our work, the  
 3 differentiation is achieved by taking the first-order derivative of the  
 4 received pulse. Higher differentiation orders have been used in  
 5 other work [174]. However, investigation showed that first-order  
 6 differentiation worked best in this case.

### 7 4.4.3 OOK Link Demonstration

#### 8 4.4.3.1 OOK links with direct post-processing

9 At the transmitter, a PRBS sequence with a length of  $2^{10} = 1024$   
 10 bits was generated by a PC and sent to the AWG, which then drove  
 11 the micro-LED to emit OOK signals. The data rate varied from 1  
 12 Mbit/s to 10 Mbit/s and the pulse width was set at 10 ns. At the  
 13 receiver, the OOK signal was detected by the UV-C SPD. The  
 14 detection window was set to 10 ns to match the transmitted pulse  
 15 width. The decision threshold was fixed at -9 mV (approximately  
 16 90% of the peak value of single pulses). The average received  
 17 photon number per bit was fixed at  $\sim 30$ , corresponding to  $\sim 0.22$  nW  
 18 received power at a data rate of 10 Mbit/s [53].



19

20

21

22

Figure 4.18 (a) Synchronized received signal and transmitted  
 sequence (b) BER versus data rate for OOK with direct post-

3

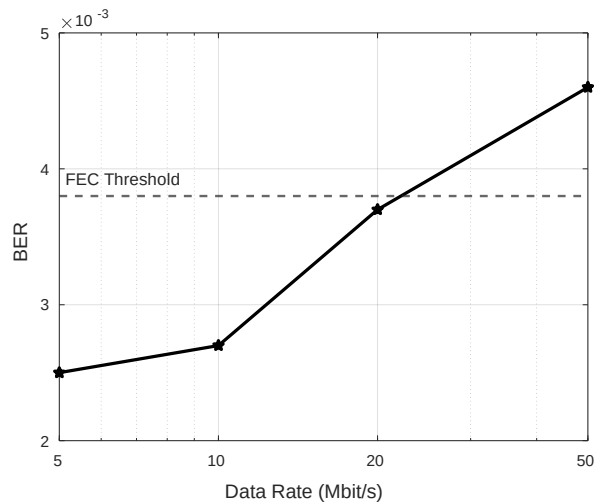
4

1 processing

2 Fig. 4.18 (a) shows a synchronized received signal and a  
3 transmitted sequence for a data rate of 10 Mbit/s. Note that the  
4 received data values are negative, as an inverting amplifier is used  
5 on the PCB board. The -20 mV and 0 mV levels in the transmitted  
6 sequence represent "1" and "0", respectively. The measured BER  
7 versus data rate is shown in Fig. 4.18 (b). When the data rate is 10  
8 Mbit/s, BER is measured as  $3 \times 10^{-3}$ , which is less than the FEC  
9 threshold of  $3.8 \times 10^{-3}$ . The BER increases with the data rate because  
10 the ISI caused by afterpulses increases [52].

#### 11 4.4.3.2 OOK links with CPP

12 At the transmitter, an OOK signal (generated by a PRBS  
13 sequence with a length of  $2^{10} = 1024$  bits) with a data rate ranging  
14 from 5 Mbit/s to 50 Mbit/s, and a pulse width set at 10 ns, was  
15 transmitted by the UV-C micro-LED. At the receiver, the signal was  
16 received and post-processed using the CPP. There are 8 values  
17 within a detection window, due to a 2 ns detection window and a 4  
18 GSa/s sampling rate. After differentiation, 8 values (corresponding to  
19 a detection window) were summed and compared with a decision  
20 threshold. The decision threshold was fixed at "-10", which is  
21 approximately 8 times 90 % of the peak amplitude of a single-  
22 photon pulse after the differentiation. If the summed value was less  
23 than the threshold, the bit was "1", otherwise "0".



1

2 Figure 4.19 BER versus data rate for OOK with CPP

3 The measured BER versus data rate is plotted in Fig. 4.19. The  
4 result shows the BER increases as the data rate rises. The maximum  
5 data rate below the FEC threshold reaches 20 Mbit/s with a BER of  
6  $3.7 \times 10^{-3}$ . Compared with the result in Fig. 4.18 (b), due to the CPP,  
7 the maximum data rate is improved from 10 Mbit/s to 20 Mbit/s  
8 because of lower ISI.

#### 9 4.4.4 PPM Link Demonstration

10 A PPM modulation scheme was used to further reduce the  
11 transmitted power. PPM encodes information by transmitting a pulse  
12 within one of the multiple time slots in each symbol period. For  
13 example, an 8-PPM scheme uses 8 time slots per symbol, thus 3 bits  
14 per symbol. Previous research has demonstrated that the PPM  
15 scheme is an effective approach to improving power efficiency, a  
16 critical factor in building deep-space and underwater optical  
17 communication systems [160-162, 175]. However, it requires  
18 precise timing synchronization at the receiver and generally has  
19 lower spectral efficiency compared to simpler modulation schemes  
20 such as OOK.

3  
4

1 At the transmitter, a PRBS sequence was modulated onto a PPM  
 2 sequence with a length of  $2^7 = 128$  symbols to drive the micro-LED  
 3 (The AWG has a waveform memory of 16 Gsamples). The sequence  
 4 length was limited by the oscilloscope's memory, which supports up  
 5 to 4000000 samples per data capture. The pulse width of a time slot  
 6 within each symbol was set as 2 ns, considering the measured  
 7 micro-LED's E-O bandwidth of around 500 MHz. The average  
 8 transmitted photon number per symbol was 60 and the symbol rate  
 9 varied from 100 kilobaud per second (kb/s) to 50 Mb/s. The  
 10 corresponding PPM index  $M$ , bit rate, photons per bit and  
 11 transmitted power are shown in Tables 4.2 and 4.3. Considering the  
 12 maximum transmitted power of 2.16 nW and 6 mm beam diameter,  
 13 the irradiance is calculated as  $76 \mu\text{W}/\text{m}^2$ , less than the safety exempt  
 14 level  $1 \text{mW}/\text{m}^2$ .

15 Table 4.2 Parameters and values of 128-PPM links at the  
 16 transmitter

<b>Symbol rate (Mb/s)</b>	<b>0.1</b>	<b>0.2</b>	<b>0.5</b>	<b>1</b>	<b>2</b>	<b>3</b>
Bit rate (Mbit/s)	0.7	1.4	3.5	7	14	21
Photons per bit	8.5	8.5	8.5	8.5	8.5	8.5
Transmitted power (pW)	4.32	8.64	21.6	43.2	86.4	129.6

17

18 Table 4.3 Parameters and values of 8-PPM links at the transmitter

<b>Symbol rate (Mb/s)</b>	<b>2</b>	<b>5</b>	<b>10</b>	<b>20</b>	<b>40</b>	<b>50</b>
Bit rate (Mbit/s)	6	15	30	60	120	150
Photons per bit	20	20	20	20	20	20
Transmitted power (pW)	86.4	216	432	864	1728	2160

19 At the receiver, the detection window was set as 2 ns to match  
 20 the pulse width at the transmitter. The received signal was

1 synchronized to the transmitted PPM sequence by correlation. The  
2 received signal was converted to a received PPM sequence based on  
3 either direct post-processing or the CPP. Then, the received PPM  
4 sequence was demodulated and compared with the PRBS sequence  
5 in the transmitter to obtain the BER. The symbol error rate (SER)  
6 was obtained by comparing the transmitted PPM sequence and the  
7 received PPM sequence.

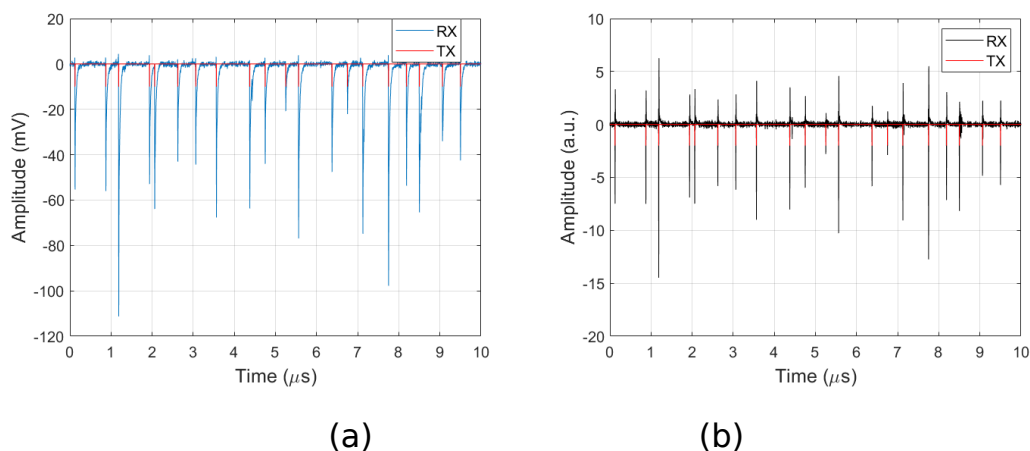
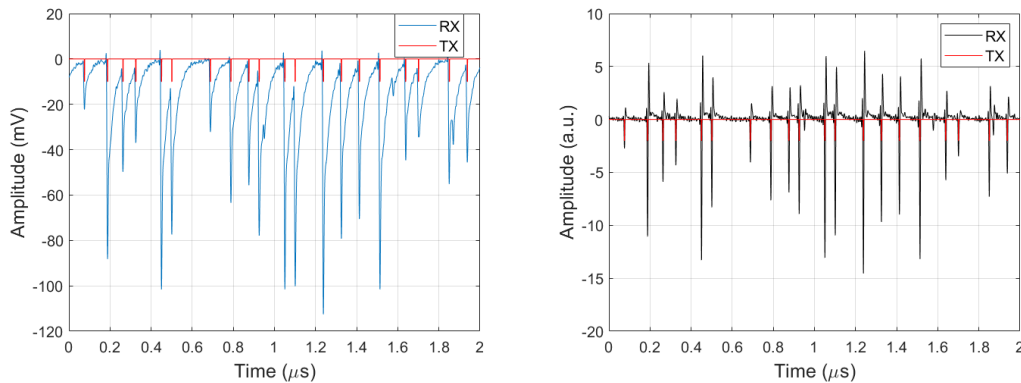
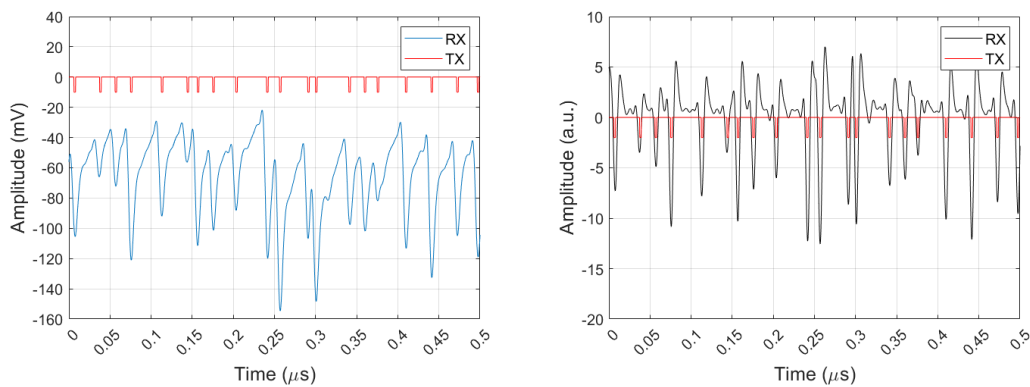


Figure 4.20 Synchronized received signal and transmitted PPM sequence for the 2 Mb/s 8-PPM link (a) received signal without CPP (b) received signal with CPP

The synchronized received signal (RX) and transmitted PPM sequences (TX) for 8-PPM links are illustrated in Fig. 4.20, 21 and 22, with symbol rates of 2 Mb/s, 10 Mb/s and 40 Mb/s, respectively. The -10 mV (or “-2”) and 0 mV (or “0”) in TX lines represent values of “1” and “0” in the transmitted PPM sequences, respectively. The amplitude of the received signal is variable because the number of detected photons in each symbol is changing, following a Poisson distribution. Comparing Fig. 4.20 (a), 4.21 (a) and 4.22 (a), the ISI of the received signal where there is no CPP is increasingly severe with the increasing symbol rate. However, the usage of CPP significantly mitigates the ISI, which can be seen in Fig. 4.22.

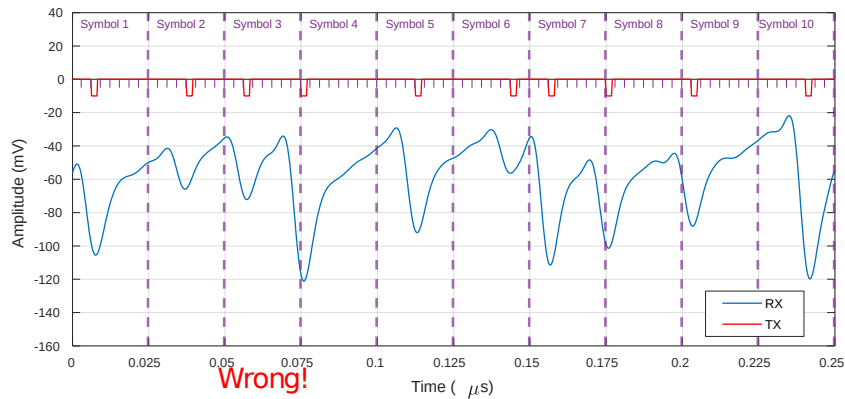


1  
 2 (a) (b)  
 3 Figure 4.21 Synchronized received signal and transmitted PPM  
 4 sequence for the 10 Mb/s 8-PPM link (a) received signal without CPP  
 5 (b) received signal with CPP



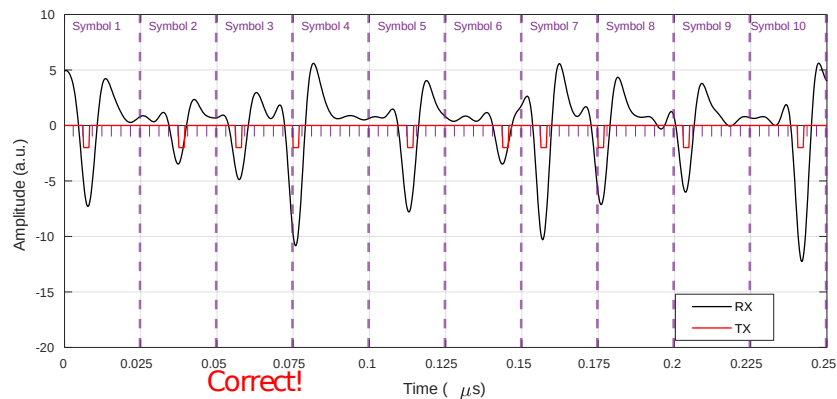
6  
 7 (a) (b)  
 8 Figure 4.22 Synchronized received signal and transmitted PPM  
 9 sequence for the 40 Mb/s 8-PPM link (a) received signal without CPP  
 10 (b) received signal with CPP

11 Fig. 23 shows both the transmitted sequence and received  
 12 signal (10 symbols) of the 40 Mb/s 8-PPM link in detail, further  
 13 highlighting the advantage of using CPP. The figure shows that the  
 14 detection of symbol 3 is incorrect without CPP, as the transmitted  
 15 time slot 3 is incorrectly decided as time slot 8. However, the usage  
 16 of CPP removed this incorrect detection.



1  
2

(a)



3  
4

(b)

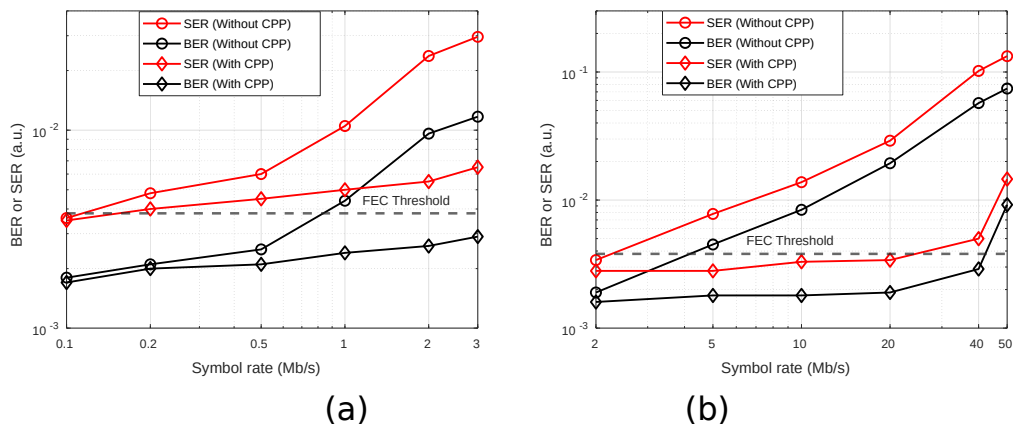
5 Figure 4.23 Ten synchronized symbols for the 40 Mb/s 8-PPM link (a)  
 6 without CPP (b) with CPP

7 The calculated SER and BER versus symbol rate are shown in  
 8 Fig. 4.24 (a) and Fig. 4.24 (b). Each point of the BER or SER curves  
 9 was calculated by transmitting  $100 \times 2^7 = 12800$  symbols. In this  
 10 context, each symbol in 128-PPM (8-PPM) represents 7 (3) bits,  
 11 meaning the bit rate is 7 (3) times the symbol rate. The SER  
 12 represents the probability that a transmitted symbol is incorrectly  
 13 received, while the bit error rate (BER) is the error probability per  
 14 bit. Since each symbol contains multiple bits, and a symbol error  
 15 does not imply that all bits are incorrect, the SER is typically higher  
 16 than the BER under the same conditions. The result shows that,  
 17 using direct post-processing, the BER increases from  $1.8 \times 10^{-3}$  to  
 18  $1.17 \times 10^{-2}$  and the SER increases from  $3.6 \times 10^{-3}$  to  $2.96 \times 10^{-2}$  as the

3  
4

1 symbol rate increases from 100 kb/s to 3 Mb/s for 128-PPM links.  
 2 Both SER and BER increase with the symbol rate because of the ISI  
 3 (due to the long fall time and afterpulses) [159].

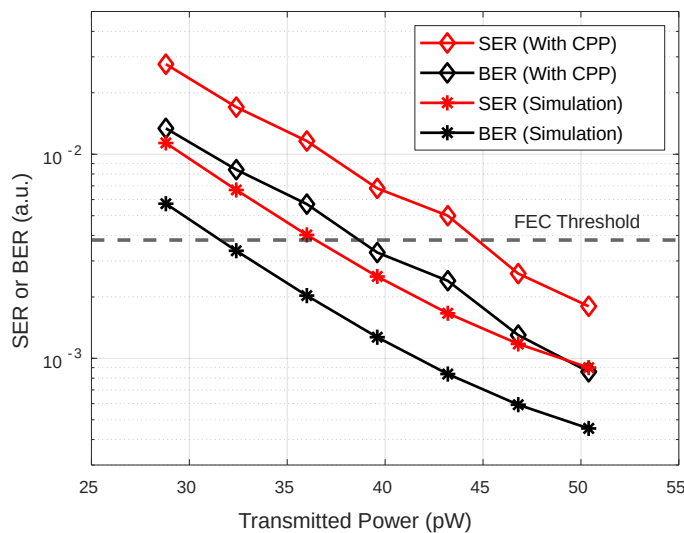
4 The CPP mitigates the ISI, so both SER and BER are reduced. For  
 5 the 3 Mb/s 128-PPM link, the BER is reduced to  $2.9 \times 10^{-3}$ , lower than  
 6 the FEC threshold. Due to the limitation of the transmitter  
 7 bandwidth, the symbol rate cannot be further increased. This is  
 8 because the minimum detection window will be larger than the time  
 9 slot required to support higher symbol rates. It can also be seen that  
 10 CPP contributes little improvement to performance at low symbol  
 11 rates. This is because the detection time slot is large enough to  
 12 accommodate the response of the UV-C SPD, so there is no ISI in the  
 13 received signals. For 8-PPM links, when the symbol rate rises from 2  
 14 Mb/s to 50 Mb/s, the BER increases from  $1.9 \times 10^{-3}$  to  $7.43 \times 10^{-2}$  and  
 15 the SER increases from  $3.4 \times 10^{-3}$  to  $1.33 \times 10^{-1}$ . With the CPP, the BER  
 16 of 40 Mb/s 8-PPM link decreases to  $2.9 \times 10^{-3}$ . In this case, a maximum  
 17 bit rate of 120 Mbit/s is achieved before ISI becomes too substantial.



18  
 19 (a)  
 20 Figure 4.24 (a) BER and SER versus symbol rate at 128-PPM links (b)  
 21 BER and SER versus symbol rate at 8-PPM links

22 To further explore the sensitivity limit of the UV-C SPD, the  
 23 average transmitted photon number per symbol was varied from 40

1 to 70 for 128-PPM links with a symbol rate of 1 Mb/s (corresponding  
 2 to transmitted power between 28.8 pW and 50.4 pW). Fig. 4.25  
 3 shows the BER and SER versus transmitted power for 1 Mb/s 128-  
 4 PPM links. Both BER and SER decrease with the increase in  
 5 transmitted power (the received power can be obtained according to  
 6 Eq. (4.7)). This is because the probability of symbols being received  
 7 with zero detected photons is reduced, due to the increasing mean  
 8 of the Poisson distribution. When the transmitted power is 38.85  
 9 pW, corresponding to 7.7 (54) transmitted photons per bit (symbol),  
 10 the BER is equal to the FEC threshold. Based on the Poisson  
 11 distribution, the standard quantum limit for PPM is 5 detected  
 12 photons per symbol for a BER of  $3.8 \times 10^{-3}$  [158]. In this work,  
 13 considering the losses due to convex lens transmission, UV filter  
 14 transmission and SiPM PDE, the 54 transmitted photons per symbol  
 15 correspond to 6.1 detected photons per symbol. To achieve the  
 16 quantum limit of 5 detected photons per symbol, the SiPM's  
 17 temperature or active area should be further reduced to minimize  
 18 the DCR, while advancements in manufacturing technology are  
 19 needed to mitigate the impact of afterpulses.



20

1 Figure 4.25 BER and SER versus transmitted power at 1Mb/s 128-  
2 PPM links with CPP

3 Additionally, simulation results are added to Fig. 4.25 to  
4 compare with the experimental result. The theoretical analysis of  
5 SER and BER refers to section 3.2.2. The simulation result performs  
6 better than the experimental result in terms of SER and BER. This is  
7 because the simulation does not consider ISI, due to the long fall  
8 time and afterpulses. Typically, an additional 7-10 pW of power is  
9 required to compensate for the effects of ISI at 1 Mb/s for 128-PPM  
10 links, in order to achieve similar BER and SER.

11 To further investigate the impact of long fall time and  
12 afterpulses, the simulated (see Fig. 4.25) and experimental results  
13 (see Fig. 4.24) for 128-PPM links at 1 Mb/s and 0.1 Mb/s are  
14 compared. Both the long fall time and afterpulses contribute to ISI at  
15 1 Mb/s, whereas only afterpulses contribute at 0.1 Mb/s (as the  
16 detection window is longer than the pulse duration). The simulated  
17 BER is calculated as  $8.4 \times 10^{-4}$  for both symbol rates. The experimental  
18 BER is  $2.4 \times 10^{-3}$  and  $1.7 \times 10^{-3}$  at 1 Mb/s and 0.1 Mb/s respectively. Thus,  
19 the difference in BER between experiment and simulation is reduced  
20 from  $1.56 \times 10^{-3}$  to  $8.6 \times 10^{-4}$ . This reduction occurs because the effect  
21 of long fall time is eliminated, leaving afterpulses as the source of  
22 the remaining difference. The change in SER follows a similar trend  
23 to that of the BER.

24 In addition to differentiation, the matched filter is another  
25 method to reduce ISI [160]. Based on using the matched filter  
26 or/and differentiation, there are four options: (1) differentiation, (2)  
27 matched filter, (3) matched filter followed by the differentiation and  
28 (4) differentiation followed by the matched filter.

1 The matched filter was implemented using a simple ‘integrate  
 2 and dump’ scheme. For each timeslot, the measured pulse shape  
 3 was multiplied by the simulated pulse shape, and the result was  
 4 integrated. Decisions were then based on whether the result was  
 5 above or below a threshold. Different simulated pulse shapes are  
 6 required, depending on whether the signal has been differentiated.  
 7 For options (2) and (3), data from Fig. 4.15 (d) is used, and data  
 8 from Fig. 4.17 is used for option (4).

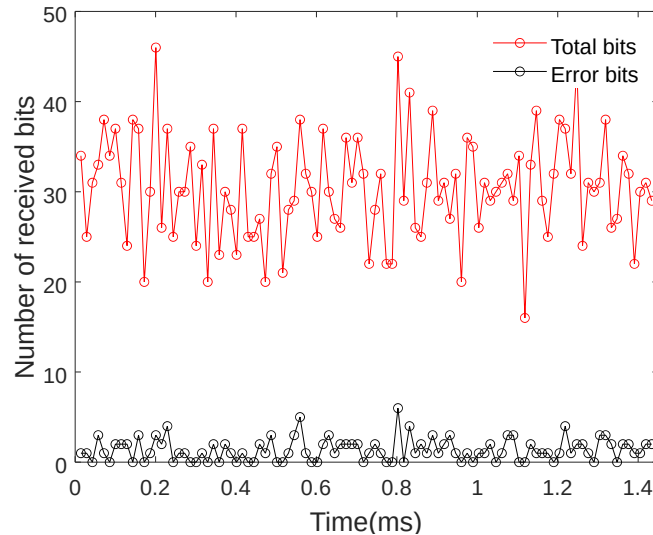
9 The SER and BER of 40 Mb/s 8-PPM links using these four  
 10 options are shown in Table 4.4. The result shows using only  
 11 differentiation achieves the best SER and BER performance. This is  
 12 because the shape of received pulses is changing due to the overlap  
 13 of multiple single pulses. Thus, an optimal matched filter does not  
 14 exist.

15 Table 4.4. SER and BER based on using a matched filter or/and  
 16 differentiation

Option	SER	BER
(1) Differentiation	0.0050	0.0029
(2) Matched filter	0.1130	0.0627
(3) Matched filter followed by the differentiation	0.2595	0.1316
(4) Differentiation followed by the matched filter	0.0298	0.0182

17 **4.4.5 Single-Photon Link Demonstration**

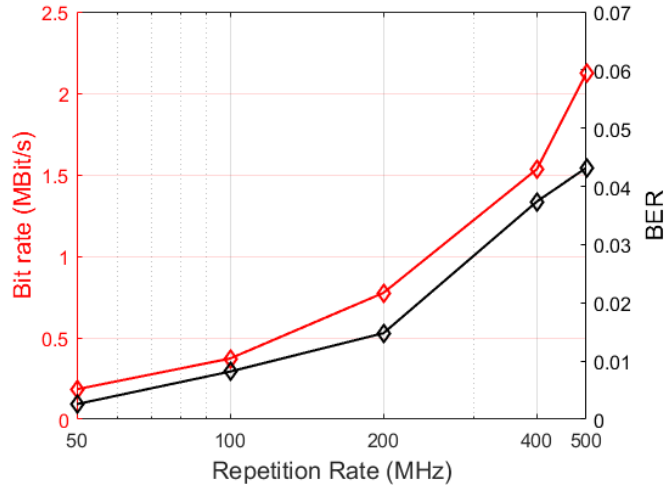
18 In addition to OOK and PPM links, a single-photon link was built  
 19 to verify the UV-C SPD’s single-photon counting ability towards  
 20 building a BB84 QKD system.



1  
 2 Figure 4.26 Number of received bits versus time (over 1.43 ms) for a  
 3 500 MHz repetition rate.

4 The demonstration of a single-photon link uses the same  
 5 approaches and methods for the OOK links with the CPP, as shown  
 6 in section 4.4.3.2. In this case, the transmitted optical power was  
 7 attenuated to 0.07 photon per bit “1”. Fig. 4.26 shows the number  
 8 of received bits per  $7 \times 2^{10}$  transmitted bits (corresponding to  $\sim 14.3$   
 9  $\mu\text{s}$ ) for a repetition rate of 500 MHz. The repetition rate was 500  
 10 MHz. The results show that the total numbers of received bits per  
 11  $7 \times 2^{10}$  transmitted bits fluctuate at around 30. This measured number  
 12 is similar to an expected number of  $7 \times 2^{10} \times 0.07 \times 0.92 \times 0.83 \times 0.1476 \approx 29$ ,  
 13 according to Eq. (4.7).

14 The BER and bit rate versus repetition rate are plotted in Fig.  
 15 4.27. BER increases from 0.26% to 4.32% with a repetition rate  
 16 ranging from 50 MHz to 500 MHz. A maximum bit rate of 2.12 Mbit/s  
 17 is achieved. Although the BER is insufficient for communications  
 18 purposes it is within the range that allows QKD operation, as will be  
 19 discussed in the next Chapter.



1

2 Figure 4.27 Bit rate and BER versus repetition rate for the single-  
3 photon link

#### 4 4.5 CONCLUSIONS

5 This Chapter presents work on establishing ultralow photon flux  
6 links based on a UV-C micro-LED and a UV-C SPD. The micro-LED has  
7 a bandwidth of over 400 MHz and a central wavelength of 276 nm.  
8 The UV-C SPD has a PDE of 14.76 % and a DCR of 2.1 kcps at a  
9 temperature of -20 °C. A minimum transmitted average photon  
10 number per bit of 7.7 was demonstrated in 1 Mb/s 128-PPM links  
11 and a maximum data rate of 120 Mbit/s was achieved in 40 Mb/s 8-  
12 PPM links. Moreover, a 500 MHz repetition rate single-photon link  
13 was demonstrated, resulting in a BER of 4.32% and a measured bit  
14 rate of 2.12 Mbit/s. This shows that UV-C SPD is sufficiently sensitive  
15 for establishing a Mbit/s-level UV-C QKD link using a UV-C micro-LED.

16 The next Chapter will report the modelling and demonstration of  
17 a UV-C wireless QKD link based on the single-photon link  
18 demonstrated in this Chapter.

---

# 1 **Chapter 5:**

## 2 **UV-C Short-Range Wireless B92 QKD**

### 3 **5.1 INTRODUCTION**

4 Quantum Key Distribution (QKD) is a cryptographic technique  
5 that ensures secure communication based on the principles of  
6 quantum mechanics [23]. Among the various QKD protocols shown  
7 in Chapter 2, the B92 protocol stands out as a simple and effective  
8 method for quantum encryption [90]. Compared with the BB84  
9 protocol, which utilizes two non-orthogonal bases (four states), the  
10 B92 protocol employs only two non-orthogonal quantum states for  
11 key generation. Thus, the B92 protocol is more hardware-efficient as  
12 it requires only half the number of light sources and detectors  
13 compared to the BB84 protocol. However, this comes at the cost of  
14 half the rate of key transmission.

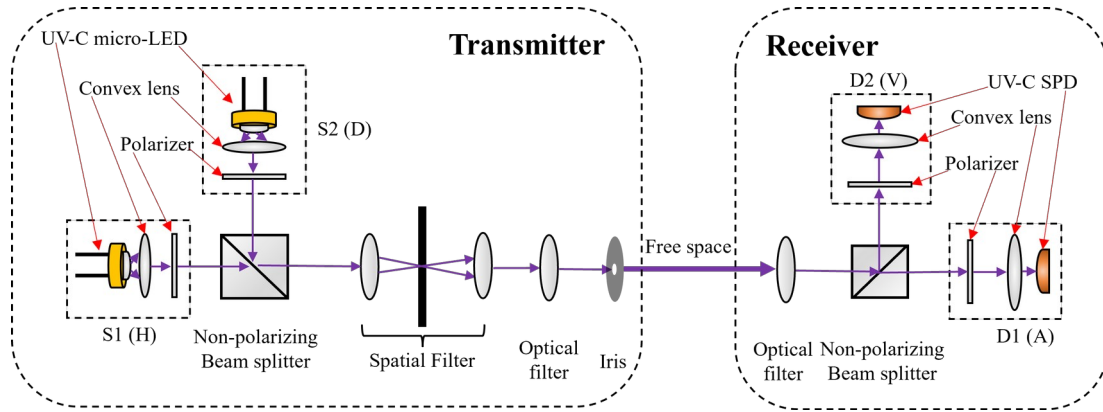
15 B92 QKD has been demonstrated in fibre-based systems, and  
16 theoretical analysis of wireless B92 QKD has been explored [176-  
17 180]. However, no experimental implementations of wireless B92  
18 QKD links have been reported to date. Chapter 4 presents single-  
19 photon links, using a UV-C micro-LED and a UV-C SPD. Based on this  
20 foundation, this Chapter will focus on implementing a wireless B92  
21 QKD link using these devices.

22 In this Chapter, a theoretical model of the UV-C wireless B92  
23 QKD system is first proposed. QBER & SKR estimation and  
24 simulation show the feasibility of a practical QKD system robust to  
25 sunlight interference. Then, a practical QKD transmitter and receiver

1 are presented and characterized. Horizontal (H) or diagonal (D)  
2 polarizations are used to encode bits in the transmitter, and anti-  
3 diagonal (A) or vertical (V) polarizations are used to decode bits in  
4 the receiver. The QKD links are demonstrated with an inferred SKR  
5 of over 1 Mbit/s. Finally, the ambient light rejection performance of  
6 our QKD system is analyzed, and the strategy for full robustness to  
7 sunlight is proposed.

## 8 **5.2 MODELLING OF UV-C WIRELESS B92 QKD**

9 A block diagram of a wireless B92 QKD system is shown in Fig.  
10 5.1. At the transmitter (Alice), two sources, S1 (H, horizontal  
11 polarization) and S2 (D, diagonal polarization), are used. Each  
12 source includes a UV-C micro-LED, a convex lens and a polarizer.  
13 The convex lens collimates light emitted from the UV-C micro-LED,  
14 followed by a linear polarizer (H or D) to determine the polarization  
15 of emitted photons. The collimated and polarized light from each  
16 source is combined by a non-polarizing beam splitter and directed to  
17 a spatial filter. The spatial filter contains two convex lenses and a  
18 pinhole. The filter ensures that the beams of light from S1 and S2  
19 are identical spatially. The filter operates as follows: the first convex  
20 lens focuses the light into the pinhole and the second convex lens  
21 recollimated the light. An optical filter (a narrow bandpass filter) is  
22 placed after the spatial filter to reshape the spectrum of S1 and S2  
23 to ensure they are identical spectrally. The diameter of the  
24 transmitter beam is defined by an iris placed at the transmitter  
25 output. Identical output intensity is achieved by adjusting the  
26 driving voltage of S1 and S2.



1

2

Figure 5.1 Structure of a wireless B92 QKD system

3

At the receiver, an optical filter (a narrow bandpass filter) passes  
 4 the wavelength of incident light and rejects the ambient light at  
 5 other wavelengths. Then, a non-polarizing beam splitter splits the  
 6 incoming light. The output beams from the beam splitter are  
 7 received by two UV-C SPDs. The polarizer (A, anti-diagonal or V,  
 8 vertical) determines the detection polarization, and a convex lens  
 9 focuses the light onto the active area of two UV-C SPDs.

10

The total link loss is made up of the geometric loss and optical  
 11 loss due to optics. The collimated light emitted from the transmitter  
 12 is modelled as a Gaussian beam. The corresponding power density  
 13 follows:

14

$$I(L) = I_0 \left( \frac{W_0}{W(L)} \right)^2 \exp\left( \frac{-2L^2}{W(L)^2} \right) \text{ where } L \text{ is the link length between the}$$

(5.1)

15

transmitter and receiver.  $W_0 = \frac{\lambda}{\pi \theta n}$  is the waist radius, where  $\theta$  is the  
 16 beam's half FOV,  $\lambda$  is the wavelength and  $n$  is the refractive index of

17

the atmosphere.  $W(L)$  is given by  $\sqrt{1 + \left( \frac{L}{L_0} \right)^2}$ , where  $L_0 = \frac{\pi W_0^2 n}{\lambda}$ .  $I_0$  is the

2

3

1 power density factor which is given by  $\frac{2P_0}{\pi W_0^2}$ , where  $P_0$  is the initial  
 2 beam power. Thus, the geometric transmission is calculated by  
 3 integration (assuming the transmitter and receiver are aligned):

$$4 \quad T_g(L) = \begin{cases} \int_0^{r_c} 2\pi r I(L) dr, L \cdot \tan(\theta) \geq r_c \\ \int_0^{L \cdot \tan(\theta)} 2\pi r I(L) dr, L \cdot \tan(\theta) \leq r_c \end{cases} \quad (5.2)$$

5 where  $r_c$  is the radius of the receiver's collection area.

6 The calculation of QBER is as follows. The photon number in  
 7 each qubit follows the Poisson distribution due to a single-photon  
 8 level qubit. The probability of getting no less than one detected  
 9 photon per qubit is [52]:

$$10 \quad P_s = 1 - e^{-PDE \cdot N_p \cdot T_g(L) \cdot T_o(L)} \quad (5.3)$$

11 where  $N_p$  is the transmitted average photon number per qubit.  $T_o(L)$   
 12 is the total transmission of all optics used in the receiver.

13 The noise consists of noise due to ambient light and dark  
 14 counts. The noise due to ambient light comes from sunlight  
 15 interference, which is defined by the ASTM 1.5 Spectra [51]. The  
 16 optical power from the ambient light is calculated as:

$$17 \quad U_a = I_a A_c \Delta\lambda \frac{\tan^2\left(\frac{FOV}{2}\right)}{2} \quad (5.4)$$

18 where  $I_a$  is the irradiance of sunlight and  $A_c$  is the receiver's

1 collection area, in this case the area of the convex lens. The  $FOV$  is  
 2 the receiver's field of view (FOV) and  $\Delta\lambda$  is the optical bandwidth of  
 3 the receiver's optical filter. Then, the probability of detecting pulses  
 4 created from ambient light per bit is:

$$5 \quad P_a = \frac{U_a \lambda}{hc} \cdot PDE \cdot t_d \quad (5.5)$$

6 where  $h$  is Planck's constant and  $t_d$  is the duration of the detection  
 7 window (similar to Eq. (3.1)). The probability of detecting dark  
 8 counts per bit is:

$$9 \quad P_d = 1 - e^{-DCR \cdot t_d} \quad (5.6)$$

10 Then, the QBER can be calculated as [21]:

$$11 \quad QBER = \frac{\frac{1}{2}(P_{\text{dark}} + P_a)}{P_s + \frac{1}{2}(P_{\text{dark}} + P_a) + n_i} \quad (5.7)$$

12 The coefficient 1/2 indicates that dark counts and pulses created  
 13 by ambient light have a 50% chance of occurring in the UV-C SPD  
 14 detection channel which will contribute to errors.  $n_i$  represents the  
 15 error due to system imperfection, such as polarization state  
 16 imperfection due to component (i.e. beamsplitter) performance,  
 17 optical misalignment and timing jitter [21]. A typical value of  $n_i$  is  
 18 1%.

19 The SKR is calculated based on the QBER. The transmitter sends  
 20  $N$  qubits to the receiver, with only  $N_\mu$  qubits are used for key  
 21 transmission. Within these  $N_\mu$  qubits,  $N_\mu^S$  is the number of qubits

1 using matched polarizations in the transmitter and receiver, such as  
 2 horizontal  $\rightarrow$  in the transmitter while anti-diagonal  $\leftarrow$  in the receiver  
 3 (see details in section 2.2.1.2). Within these  $N_\mu^s$  bits,  $K_\mu^s$  is the  
 4 number of qubits where the receiver correctly detects the received  
 5 photons. The normalized SKR is the number of generated secure  
 6 keys when one qubit is transmitted, which can be obtained from  
 7 [181, 182]:

$$8 \quad SKR \geq q \left[ -Q_\mu H(E_\mu) + Q_1 [1 - H(e_1)] \right] \quad (5.8)$$

9 where  $q$  is  $N_\mu^s/N$  and  $Q_\mu$  is  $K_\mu^s/N_\mu^s$ .  $H(x) = -x \log_2 x - (1-x) \log_2 (1-x)$  is the  
 10 Shannon binary entropy function and  $E_\mu$  is equal to  $QBER$ .  $Q_1$  and  $e_1$   
 11 are the gain and QBER of single-photon qubits respectively. When  
 12 we assume all loss and errors come from qubits with a single  
 13 photon,  $e_1$  is [182]:

$$14 \quad e_1 = \frac{Q_\mu E_\mu}{Q_1} \quad (5.9)$$

15 The value of  $Q_\mu$  and  $Q_1$  are obtained, assuming the transmitter  
 16 has a weak coherent source which generates Poisson distribution  
 17 pulses. The density function of the Poisson distribution  $\rho(i)$  is:

$$18 \quad \rho(i) = \frac{\mu^i}{i!} e^{-\mu} \quad (5.10)$$

19 where  $i$  represents  $i$ -photon qubits, where each qubit contains  $i$   
 20 photons, and  $\mu$  is the average number of photons per qubit. The link  
 21 transmission of  $i$ -photon qubits  $Y_i$  is:

$$1 \quad Y_i = 1 - (1 - PDE \cdot T_g(L) \cdot T_o(L))^i \quad (5.11)$$

2 Thus, the probability of detecting  $i$ -photon qubits is:

$$3 \quad Q_i = Y_i \frac{\mu^i}{i!} e^{-\mu} \quad (5.12)$$

4 Thus,  $Q_\mu$  can be obtained:

$$5 \quad Q_\mu = \sum_{i=1}^{\infty} Y_i \frac{\mu^i}{i!} e^{-\mu} \quad (5.13)$$

6 Therefore, considering Eq. (5.7), (5.9), (5.12), (5.13) and using  
7 the result in Eq. (5.8), the normalized SKR can be obtained.

8 A simulation was carried out, based on the above system model,  
9 to investigate the potential of establishing a practical UV-C wireless  
10 B92 QKD link. The parameters and corresponding values are listed  
11 in Table 5.1. At the transmitter, an average photon number per  
12 qubit of 0.1 and a transmitter FOV of  $1^\circ$  was used. At the receiver,  
13 the detection window is set as 2 ns. The error correction efficiency  
14 and system imperfection are typical values used in QKD  
15 demonstrations. Other values are all obtained based on commercial  
16 devices.

17 Table 5.1 Parameters and values for UV-C wireless B92 QKD  
18 simulation

Part	Parameter	Value
Transmitter	Average photon number per qubit	0.1
	Half transmitter FOV	$1^\circ$
Receiver	Detection window	2 ns

---

Optical bandwidth (set by optical filter)	10 nm
Diameter of convex lens	50.8 mm
Half receiver FOV	1°
PDE (HAMAMATSU S13360-1350 SPL)	14.76 %
Error correction efficiency	1/1.22 [64]
System imperfection	1 % [21]
(a) Transmission of optical filter	60%
(b) Transmission of BS	49.49%
(c) Reflectance of BS	49.46%
(d) Transmission of polarizer	86.54%
(e) Transmission of convex lens	92.46%

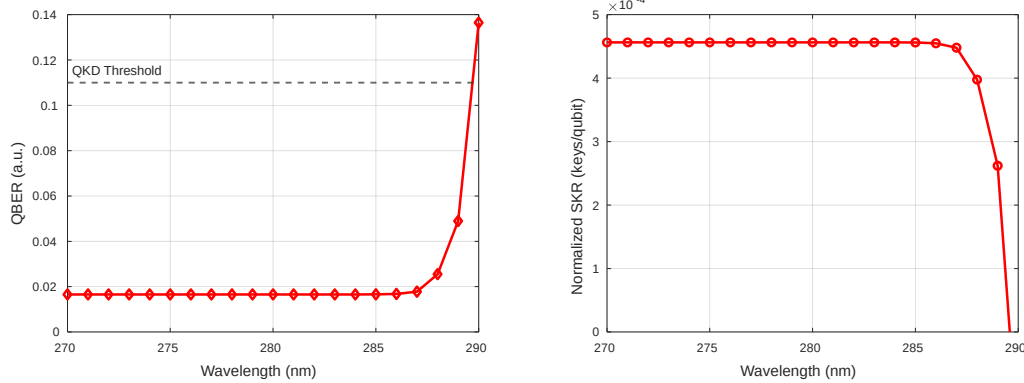
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1 Note: The optics used in (a) to (e) are: (a) Edmund 280 nm OD4 10  
 2 nm bandpass filter, (b) Thorlabs BSW20R, (c) Thorlabs BSW20R, (d)  
 3 Thorlabs GLB10-UV and (e) Thorlabs LA4052-UV. Parameter values  
 4 are from the respective data sheets.

5 The simulation also assumes: (1) no eavesdroppers interrupt the  
 6 QKD link (2) Decoy states are not applied, so all qubits are used for  
 7 generating keys (3) The transmitter and receiver are aligned and (4)  
 8 The PDE is constant with varying wavelengths.

9 The operating wavelength of a UV-C QKD system was first  
 10 investigated. The link length and DCR are set as 1 m and 5 kcps,  
 11 respectively. The wavelength used is varied from 270 nm to 290 nm,  
 12 in order to examine the influence of ambient light. Fig. 5.2 (a) shows  
 13 the QBER versus wavelength. It can be seen that the QBER is nearly  
 14 unchanged between 270 nm and 286 nm, as the DCR and system  
 15 imperfection dominate the total noise, and both of these are  
 16 constants. When wavelength increases to longer than 286 nm, the  
 17 ambient interference dominates the total noise. This is because the

1 power density of ambient sunlight increases extremely quickly  
 2 between 280 nm and 300 nm, as indicated by the ASTM 1.5 Spectra.  
 3 Thus, the QBER rises over the QKD threshold, and the QKD system  
 4 will not be feasible.



5

6

(a)

(b)

7 Figure 5.2 (a) QBER versus wavelength (b) SKR versus wavelength

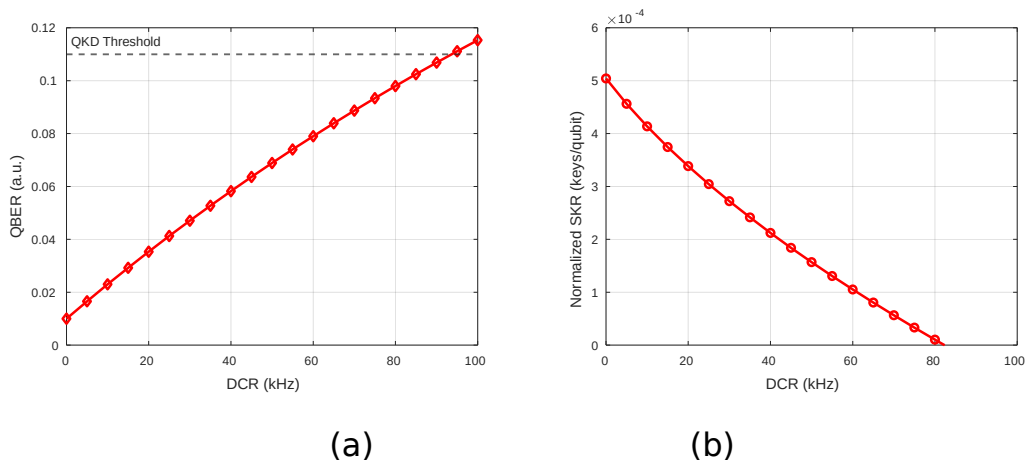
8 Fig. 5.2 (b) shows the plot of the normalized SKR vs wavelength,  
 9 where the normalized SKR is defined as the ratio of the number of  
 10 final secure keys to the total transmitted qubits. The result shows  
 11 that the operating wavelength must be less than 289 nm.

12 The next challenge is to find an acceptable value of DCR. A link  
 13 was simulated for a wavelength of 280 nm and a link length of 1 m  
 14 (other values in Table 5.1). The QBER and SKR versus DCR are  
 15 shown in Fig. 5.3 (a) and (b), respectively. The result shows the  
 16 QBER rises from 1 % to approximately 11.5 %, as the DCR increases  
 17 from 0 kcps to 100 kcps. Fig 5.3 shows the required DCR for a  
 18 theoretical QKD link should be below 80 kcps to obtain a non-zero  
 19 SKR. Previously reported practical wireless QKD links typically  
 20 maintain a QBER below 5% to ensure a reliable SKR [47, 48, 119].  
 21 Therefore, a DCR below 30 kcps is a more realistic target, with lower  
 22 values being preferable. Notably, in previous Chapters, we already

2

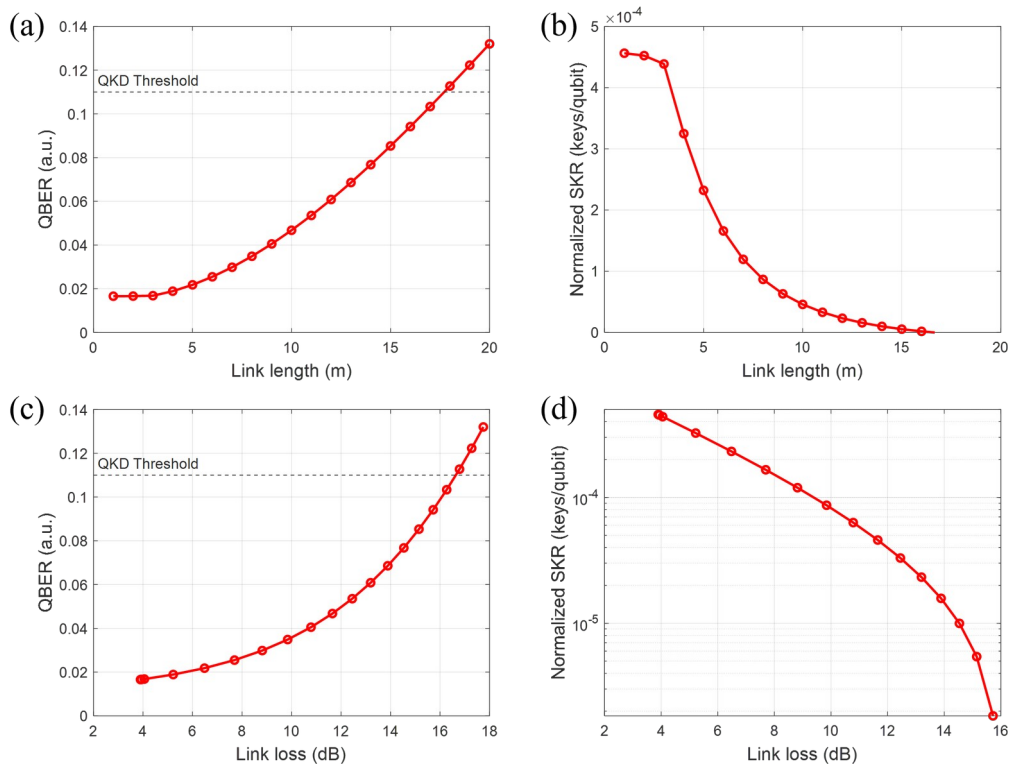
3

1 achieved a DCR below 5 kcps, which corresponds to a QBER of  
 2 approximately 1.7 % (as shown in Fig. 5.3 (a)).



5 Figure 5.3 (a) QBER versus DCR (b) SKR versus DCR

6 The last factor to consider is the link length/link loss. The  
 7 wavelength and DCR are set as 280 nm and 5 kcps respectively, and  
 8 the QBER vs link length is estimated and shown in Fig. 5.4 (a). The  
 9 result shows that QBER is almost unchanged when the link length is  
 10 less than 3m. The reason is that the geometric loss is unchanged, as  
 11 all the light from the transmitter is incident on the receiver. When  
 12 the link length is more than 3m, the beam spot at the receiver is  
 13 larger than the collection area. The increasing geometric loss then  
 14 increases rapidly, which results in a rapid rise in QBER. The  
 15 corresponding plot of normalized SKR vs link length is in Fig. 5.4 (b).  
 16 It can be seen that the maximum link length must be less than 16 m  
 17 for operation below the QKD threshold. Furthermore, the link length  
 18 is converted into link loss, considering both geometric loss and  
 19 optical loss. The corresponding QBER and SKR performance is shown  
 20 in Fig. 5.4 (c) and (d).



1

2 Figure 5.4 (a) QBER versus link length (b) SKR versus link length (c)  
3 QBER versus link loss (b) SKR versus link loss

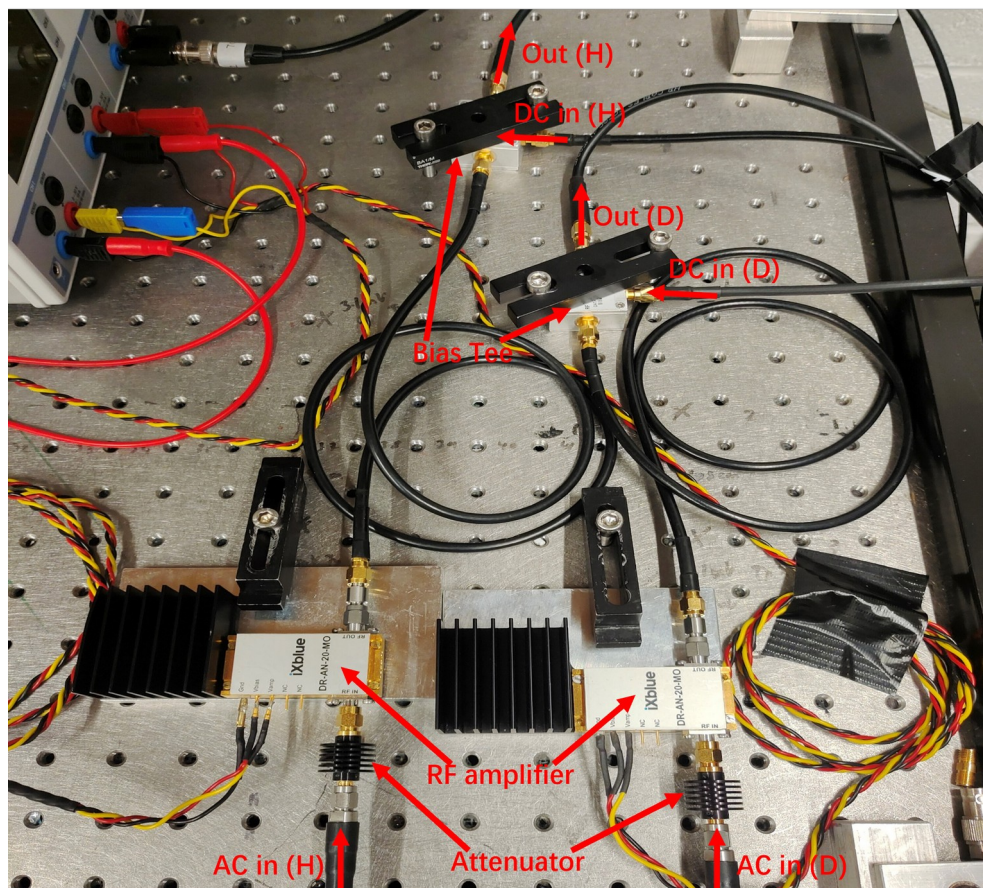
4 In conclusion, based on simulation results, the requirements for  
5 building a UV-C short-range wireless B92 QKD system are an  
6 operating wavelength of less than 289 nm and a DCR of below 30  
7 kcps. The central wavelength of the UV-C micro-LED and the DCR of  
8 the UV-C SPD presented in Chapter 4 are 276 nm and 2.1 kcps  
9 respectively, and these satisfy the requirements determined by  
10 simulation. In the next section, the elements of an experimental  
11 system are described, beginning with the transmitter.

## 12 5.3 TRANSMITTER SETUP AND CHARACTERIZATION

13 The experimental setup of the QKD transmitter consists of both  
14 an electrical and an optical subsystem. Fig. 5.5 (a) and (b) shows  
15 the photograph and block diagram of the electrical subsystem. This

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3

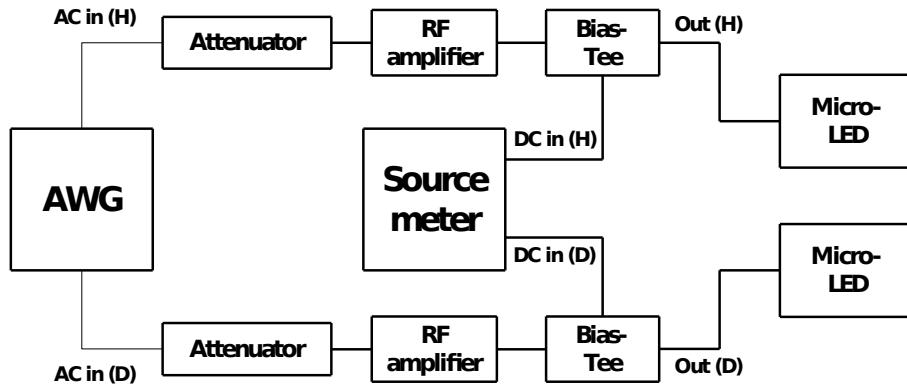
1 uses two channels (H and D) for driving two separate micro-LEDs,  
2 each of which generates light with a specific polarization, either  
3 horizontal (H) or diagonal (D). In each channel, the AC signal from  
4 an AWG (Tektronix AWG70002A) is amplified by an RF amplifier  
5 (iXblue DR-AN-20-MO). An attenuator is used to prevent high voltage  
6 damage to the micro-LED (The voltage supplied to the micro-LED  
7 must be less than 12 V according to Figure 4.2). The amplified AC  
8 signal is combined with a DC signal from a source meter (KEITHLEY  
9 2636B) by a Bias Tee (Mini-Circuits ZFBT-6GW), and the combined  
10 signal is used to drive the micro-LED.



11

12

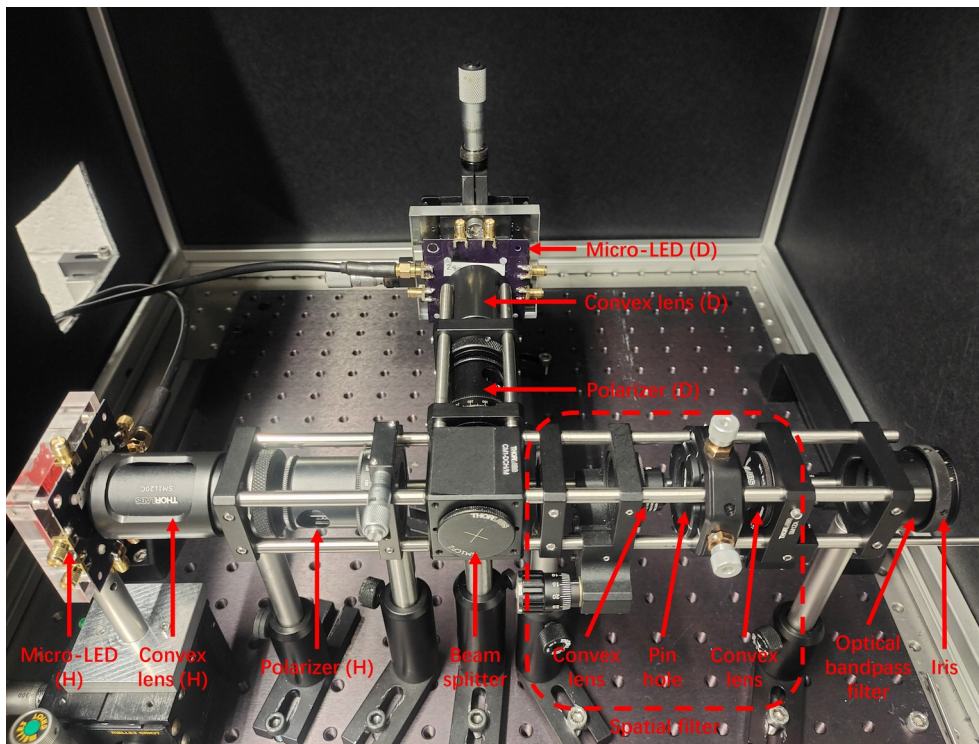
(a)



1  
2  
3  
4

(b)

Figure 5.5 QKD transmitter electrical subsystem (a) Photograph (b) Block diagram



5

6 Figure 5.6 Photograph of the QKD transmitter optical subsystem

7 Fig 5.6 shows a photograph of the QKD transmitter optical  
8 subsystem. Two micro-LEDs (H and D) are used as sources for the  
9 two QKD channels. In each channel, the light emitted from the  
10 micro-LED is collimated by a convex lens (Thorlabs LA4052-UV).  
11 Each micro-LED is mounted on a 3-axis stage (Thorlabs RBL13D/M)  
12 and aligned with a convex lens, which is placed in a slotted tube

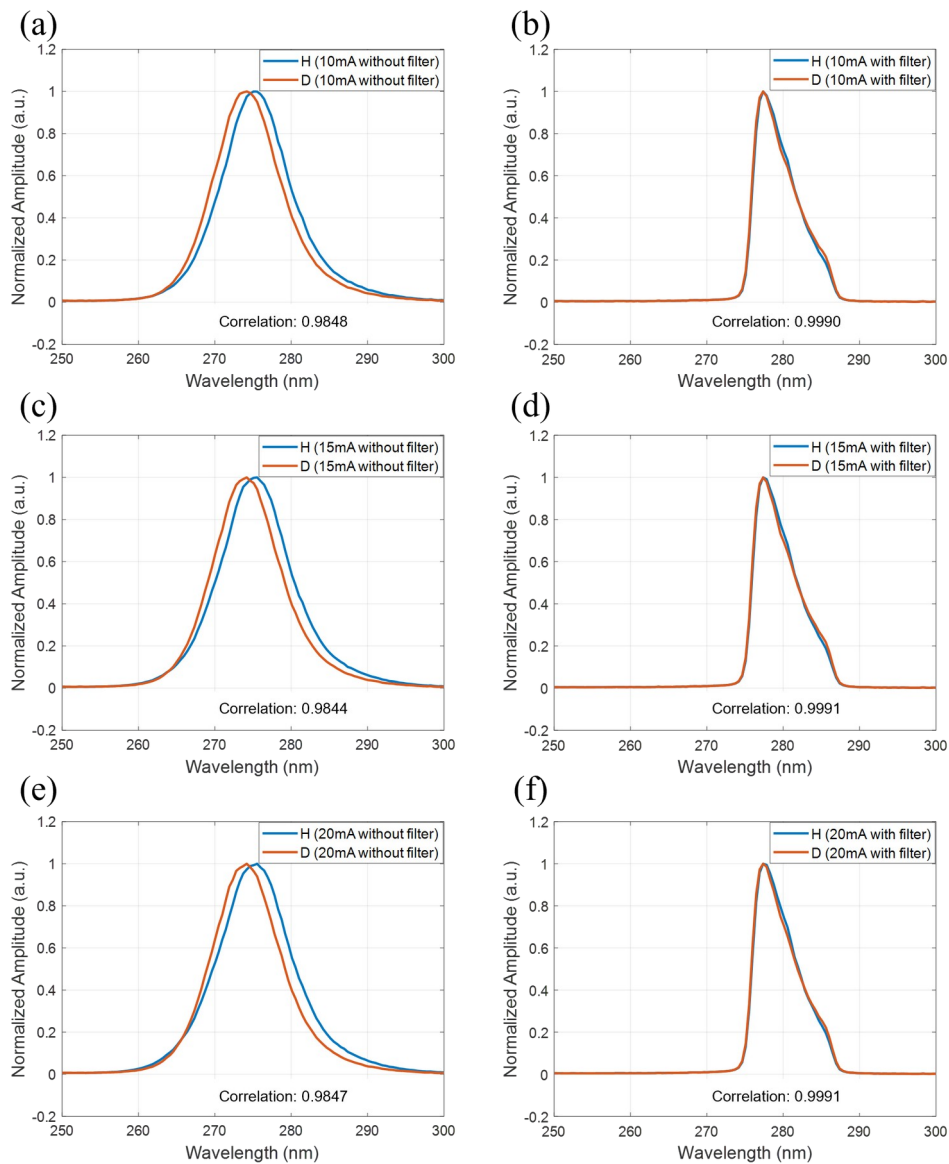
2  
3

1 (Thorlabs SM1L30C). The polarization is defined by a crystal linear  
2 polarizer (Thorlabs GLB10-UV) and the light from two channels is  
3 combined by a non-polarizing beam splitter (Thorlabs BSW20R). The  
4 combined light propagates through a spatial filter, which consists of  
5 two convex lenses (Thorlabs LA4647-UV-ML) and a pinhole (Thorlabs  
6 P100K). The positions of two convex lenses and a pinhole are  
7 adjusted by Z-axis Translation Mounts (Thorlabs SM1ZA) and XY  
8 Translation Mounts (CXY1A), respectively. Finally, an optical  
9 bandpass filter (Edmund 280nm hard coated OD4 10nm bandpass  
10 filter) reshapes the spectrum, and an iris constrains the beam  
11 aperture to 2 mm. An ND2 filter (Thorlabs NDUV20A) is placed at the  
12 end of the QKD transmitter to attenuate the transmitted light power  
13 to a single-photon level. All the optics are mounted on a movable  
14 base plate, which is fixed on an optical table. The whole QKD  
15 transmitter is enclosed, preventing the UV-C light from potentially  
16 causing hazards to operators.

17 The I-V-I curves and bandwidth of the micro-LED used have been  
18 measured as part of the work described in Chapter 4, and the data  
19 shown in Fig. 4.2 and 4.3 are used here. QKD requires sources that  
20 are spectrally indistinguishable, so a series of experiments were  
21 undertaken to assess this. First, the spectrum of sources without a  
22 narrow bandpass filter was measured by placing an optical spectrum  
23 analyzer (Ocean Insight OFX00951) at the output of the transmitter  
24 and measuring the spectrum. The driving current of two micro-LEDs  
25 was varied from 10 mA to 20 mA and the spectrum measured in  
26 both cases.

27 Then, the narrow bandpass filter was placed at the end of the  
28 transmitter and the spectrum of filtered sources was measured with  
29 the same process. Fig. 5.7 shows the spectrum of two sources with

1 and without the filter. The results show the spectrum of two sources  
 2 without the optical bandpass filter is different so that Eve can  
 3 distinguish the light from the two sources and eavesdrop on  
 4 information. However, the reshaped spectrum using the optical  
 5 bandpass filter is indistinguishable, which ensures the security of  
 6 QKD.



7

8 Figure 5.7 The spectrum of the S1 and S2 micro-LEDs (a) 10 mA  
 9 current without filter (b) 10 mA current with filter (c) 15 mA current  
 10 without filter (d) 15 mA current with filter (e) 20 mA current without

2  
 3

1 filter (f) 20 mA current with filter

2 Additionally, the correlation value between the two curves in  
3 each figure was calculated (see Fig. 5.7) to quantitatively assess the  
4 indistinguishability. The correlation is obtained by:

$$5 \quad corr = \frac{\sum (x - \bar{x})(y - \bar{y})}{\sqrt{\sum (x - \bar{x})^2} \sqrt{\sum (y - \bar{y})^2}} \quad (5.14)$$

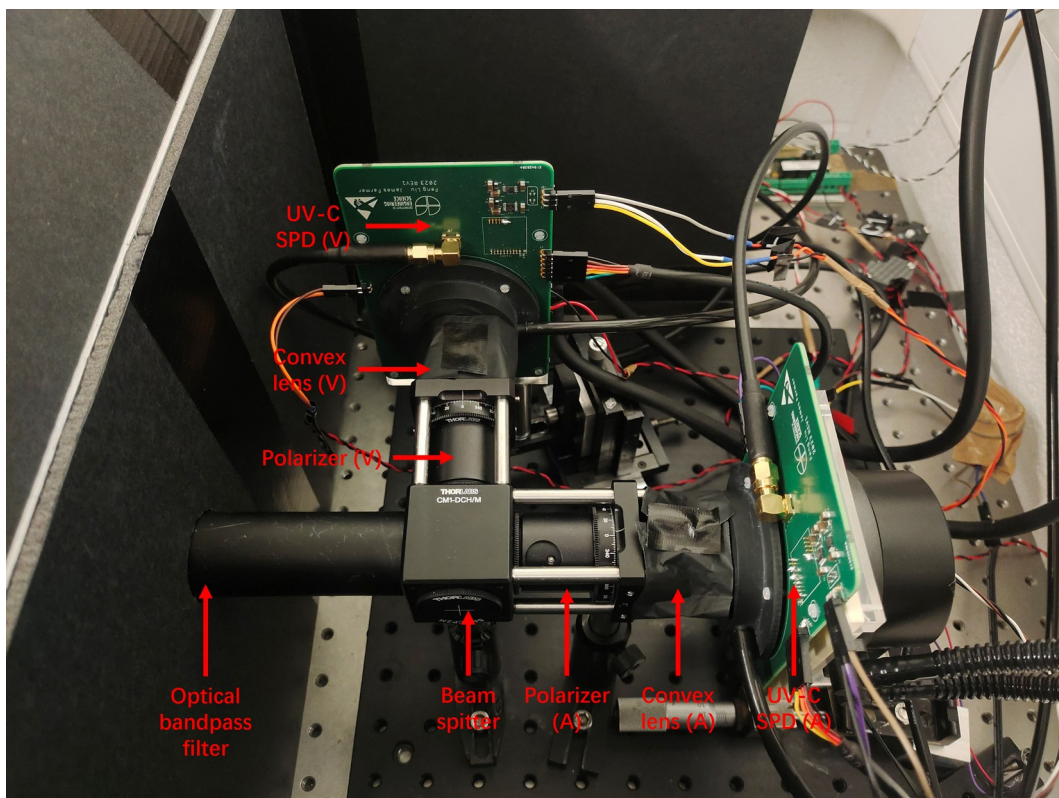
6 where  $x$  and  $y$  are values of two curves.  $\bar{x}$  and  $\bar{y}$  are the means of  
7 two curves. The result shows that the correlation value between the  
8 spectrum of the two sources is improved by using the filter.

## 9 **5.4 RECEIVER SETUP AND CHARACTERIZATION**

10 Fig. 5.8 shows the experimental setup of the QKD receiver. An  
11 optical bandpass filter (Edmund 280nm hard coated OD4 10nm  
12 bandpass filter) is placed at the input of the receiver to block  
13 ambient light at wavelengths different from the light from the  
14 transmitter. The incident light is split by a non-polarizing beam  
15 splitter (Thorlabs BSW20R) into two channels (A and V). In each  
16 channel, a crystal linear polarizer (Thorlabs GLB10-UV) is rotated to  
17 a particular polarization (diagonal or antidiagonal). A convex lens  
18 (Thorlabs LA4052-UV) collects light and focuses it on the active area  
19 of a UV-C SPD. Each UV-C SPD is mounted on a 3-axis stage, which  
20 adjusts the SPD's position to achieve maximum received light  
21 power. The design and operation details of the UV-C SPD can be  
22 found in Section 4.3. The signals generated by the two UV-C SPDs  
23 are sent to two channels of the oscilloscope (AGILENT MSO6104A).  
24 The corresponding data displayed on the two channels are collected  
25 by a PC for postprocessing. The QKD receiver is mounted on a

1 movable base plate to allow alignment and is enclosed in a light-  
2 tight box. The distance between the transmitter and receiver is 25  
3 cm.

4 Experiments were undertaken to characterize receiver  
5 performance. DCR is a key factor affecting the performance of a  
6 QKD system. The measurement of the two UV-C SPDs' DCRs is  
7 carried out using the same method presented in Section 4.3.3. The  
8 temperature is reduced from 25°C to -15°C with the bias voltage  
9 varied from 54.34 V to 52 V to maintain a constant overvoltage.  
10 Table 5.2 presents the DCR measurement. The result shows the DCR  
11 of two UV-C SPDs is reduced from 103.4 kcps and 110.5 kcps to 3  
12 kcps and 2.9 kcps, respectively.



13

14 Figure 5.8 The experimental setup of the QKD receiver

15 Table 5.2 DCR of SPDs under different conditions

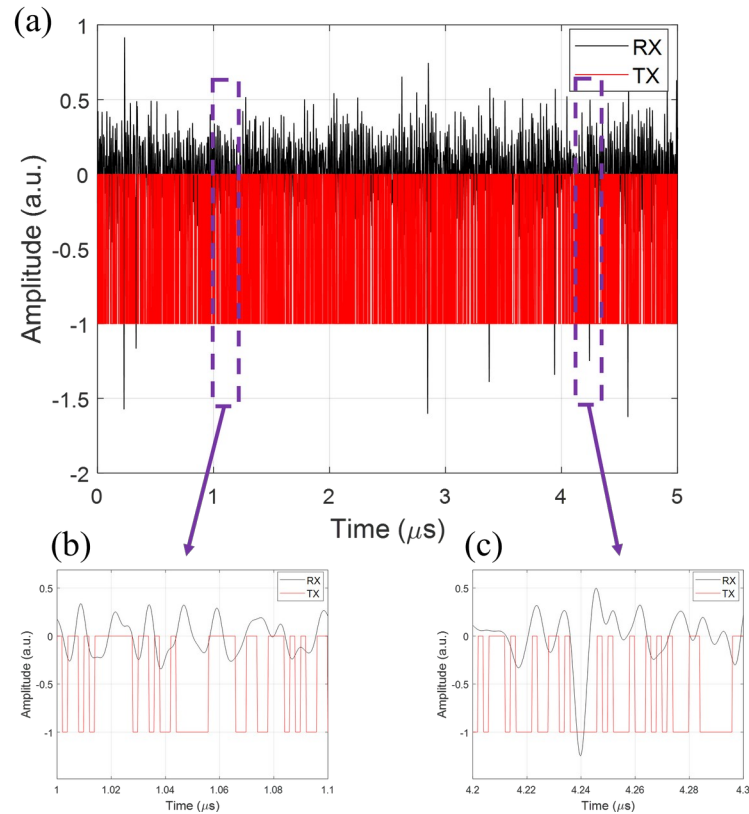
	DCR at 25°C	DCR at -15°C
SPD (A)	103.4 kcps	3 kcps
SPD (V)	110.5 kcps	2.9 kcps

## 1 5.5 UV-C WIRELESS B92 QKD DEMONSTRATION

2 At the transmitter, a PRBS sequence with a length of 4000 bits  
3 was generated by a PC and sent to the AWG. A bit “1” was used to  
4 generate a pulse to modulate the micro-LED corresponding to the  
5 (H) polarization and a bit “0” was used to generate a pulse to  
6 modulate the micro-LED corresponding to the (D) polarization. The  
7 pulse width was set as 2 ns to match the measured micro-LED’s E-O  
8 bandwidth of around 500 MHz. The average photon number per  
9 qubit was set as 0.1, and the transmitted repetition rate was varied  
10 from 100 MHz to 500 MHz. The corresponding transmitted power is  
11 shown in Table 5.3.

12 Table 5.3 Transmitted power versus repetition rate

Repetition rate	100 MHz	200 MHz	400 MHz	500 MHz
Transmitted power	7.1 pW	14.2 pW	28.4 pW	35.5 pW



1

2 Figure 5.9 Synchronized received pulses (channel A) and the  
3 corresponding transmitted qubits (channel H)

4 At the receiver, the received pulses from the two UV-C SPDs  
5 were captured by an oscilloscope for post-processing using the CPP  
6 technique (see section 4.4.2). For the data from each UV-C SPD, the  
7 detection window was fixed at 2 ns and synchronization between  
8 transmitted sequence and received pulses was achieved by  
9 correlation. Fig. 5.9 (a) shows the synchronized received pulses  
10 (from channel A) and the corresponding transmitted PRBS sequence  
11 (from channel H) over a period of 1 μs for a pulse repetition rate of  
12 500 MHz. The received pulses have been differentiated, following  
13 the differentiation process in the CPP procedure shown in Chapter 4.  
14 The values “-1” and “0” in the red line represent “1” and “0” in the  
15 transmitted sequences, respectively. The diagram shows very few  
16 qubits resulting in pulses detected by the receiver. This is due to the

2  
3

1 low probability of a photon being generated from each electrical  
 2 pulse (0.1), the link loss (see section 5.2) and PDE (14.76 %). The  
 3 left zoomed-in diagram (Fig. 5.9 (b)) shows no pulses generated in  
 4 the receiver and the right zoomed-in one (Fig. 5.9 (c)) shows one  
 5 pulse occurring at around 4.24  $\mu$ s. The average peak amplitude of  
 6 received pulses over this 5  $\mu$ s duration (7 pulses) is approximately “-  
 7 1.4”.

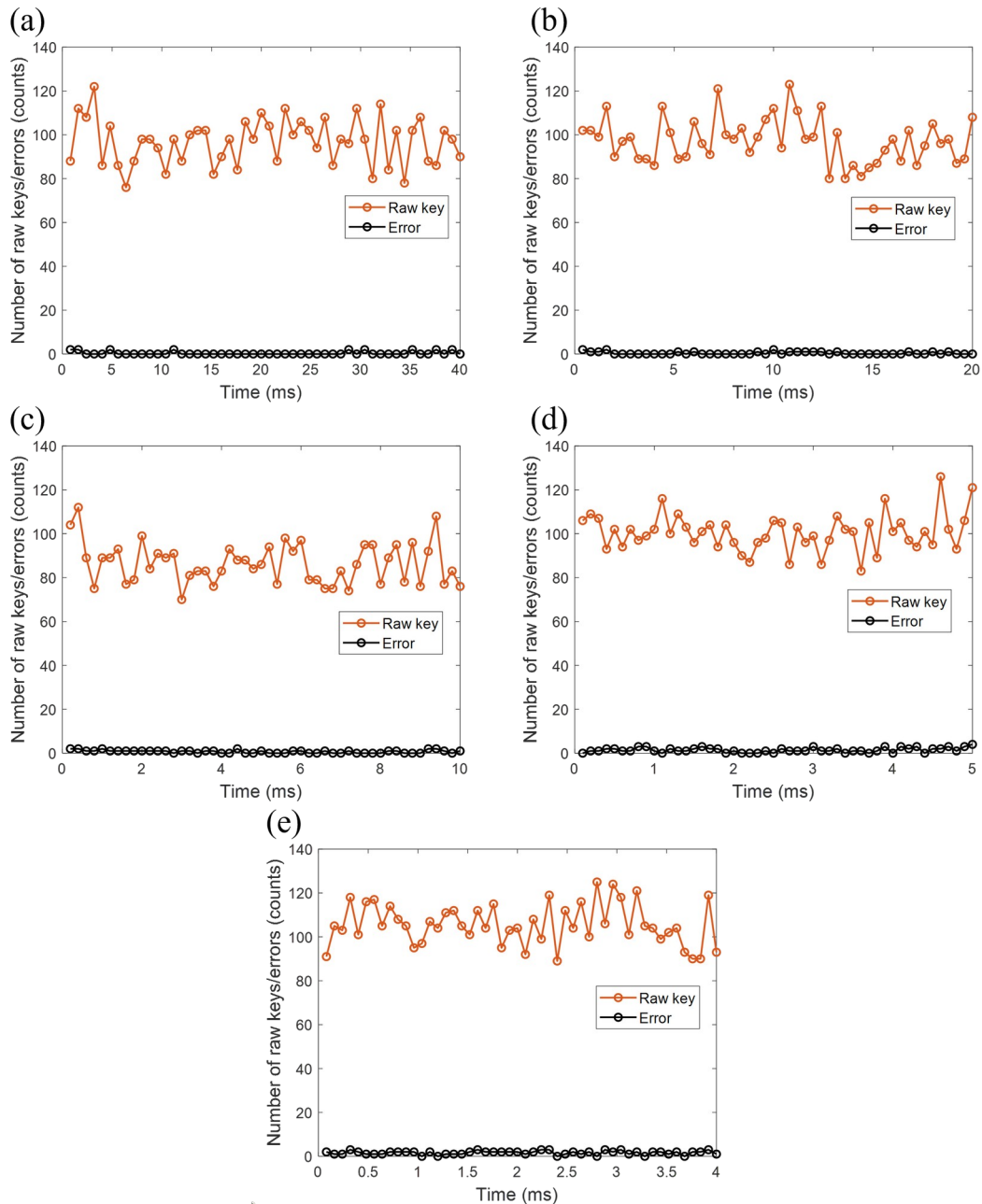
8 The peak amplitude of a single pulse is a little less than “-1”  
 9 (see Fig. 5.9) and each detection window (2 ns) consists of 8  
 10 sampling points. The decision threshold was set as -8. If the values  
 11 of each pulse summarized within a detection window are less than  
 12 the threshold, the bit is set to “1”, otherwise “0”. If the  
 13 corresponding two qubits obtained from two different UV-C SPDs are  
 14 both “1”, the received bit is “1” for UV-C SPD (A) and “0” for UV-C  
 15 SPD (V). Then, the received sequence is compared with the  
 16 transmitted sequence to obtain the raw key (transmitter and  
 17 receiver use the matched polarizations) and QBER. Finally, error  
 18 correction and privacy amplification were undertaken theoretically  
 19 to infer the SKR.

20 Table 5.4 Summary of raw key rate and QBER

Repetition rate	50 MHz	100 MHz	200 MHz	400 MHz	500 MHz
QBER (%)	0.38	0.37	0.77	1.56	1.76
Raw key rate	0.12135	0.24435	0.4357	1.0198	1.3405

21 For each repetition rate, 40000  $\times$  50 qubits were transmitted and  
 22 detected to obtain the raw key rate and QBER. Table 5.4 shows the  
 23 measured raw key rate and QBER. The results show that the QBER is  
 24 nearly unchanged with a repetition rate from 50 MHz to 100 MHz.  
 25 Then, the QBER increases because of the emergence and growth of

1 ISI. Fig. 5.10 shows the number of raw keys/errors per 40000  
 2 transmitted qubits versus time. The results show the number of raw  
 3 keys fluctuates at around 100 per 40000 qubits.



4

5 Figure 5.10 Number of raw keys/errors (per 40000 transmitted  
 6 qubits) versus time with a repetition rate of (a) 50 MHz (b) 100 MHz  
 7 (c) 200 MHz (d) 400 MHz (e) 500 MHz

8 Error correction and privacy amplification are carried out

2  
 3

1 according to [181] through the classical channel. For error  
 2 correction, redundant information is added to implement an efficient  
 3 correction protocol. Based on the Shannon theorem [183], the ratio  
 4 between the minimum number of redundant bits  $L_{EC}$  needed to  
 5 correct a raw key of length  $L_{raw}$  is:

$$6 \quad \frac{L_{EC}}{L_{raw}} = -E_{\mu} \log_2 E_{\mu} - (1 - E_{\mu}) \log_2 (1 - E_{\mu}) \quad (5.14)$$

7 where  $E_{\mu}$  is equal to QBER. However, Shannon's theorem is  
 8 conditional on the infinite length of raw keys. To get close to the  
 9 Shannon limit, a more efficient bidirectional code can be used [184],  
 10 which uses  $f(e)L_{EC}$  bits for error correction. Values of  $f(e)$  that  
 11 correspond to particular values of QBER are listed in Table 5.5.

12 Table 5.5 values of  $f(e)$  taken from [184]

$E_{\mu}$	$f(e)$
1%	1.16
5%	1.16
10%	1.22
15%	1.35

13 For privacy amplification, the corrected keys are shorted by  
 14 using hash functions to remove the effect of eavesdroppers [181,  
 15 185]. The proportion that the corrected keys need to be shorted  
 16  $L_{PA}/L_{raw}$  is given by:

$$1 \quad \frac{L_{PA}}{L_{raw}} = \begin{cases} \log_2(1 + 4E_\mu - 4E_\mu^2) & \text{for } E_\mu \leq 1/2 \\ 1 & \text{for } E_\mu > 1/2 \end{cases} \quad (5.15)$$

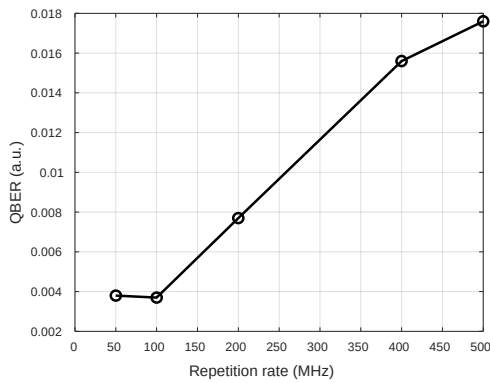
2 Thus, the SKR is calculated as:

$$3 \quad SKR = L_{raw} - L_{EC} - L_{PA} \quad (5.16)$$

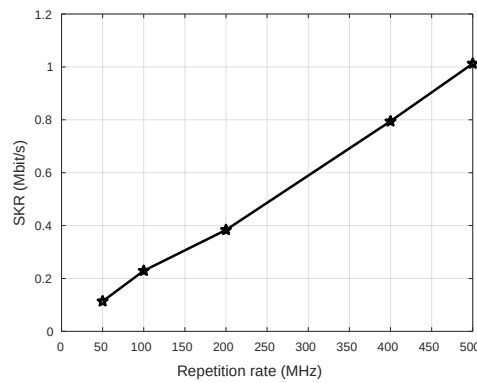
4 Table 5.6 presents the SKR versus repetition rate. The results  
 5 show that the SKR increases with the repetition rate, and the  
 6 maximum value achieved is 1.0127 Mbit/s, with a repetition rate of  
 7 500 MHz. The QBER and SKR versus repetition rate are shown in Fig.  
 8 5.11.

9 Table 5.6 SKR versus repetition rate

Repetition Rate	50 MHz	100 MHz	200 MHz	400 MHz	500 MHz
SKR (Mbit/s)	0.11373	0.22937	0.38374	0.79444	1.0127



10 (a)



11 (b)

12 Figure 5.11 UV-C B92 QKD performance under different repetition  
 13 rate (a) QBER versus repetition rate (b) SKR versus repetition rate

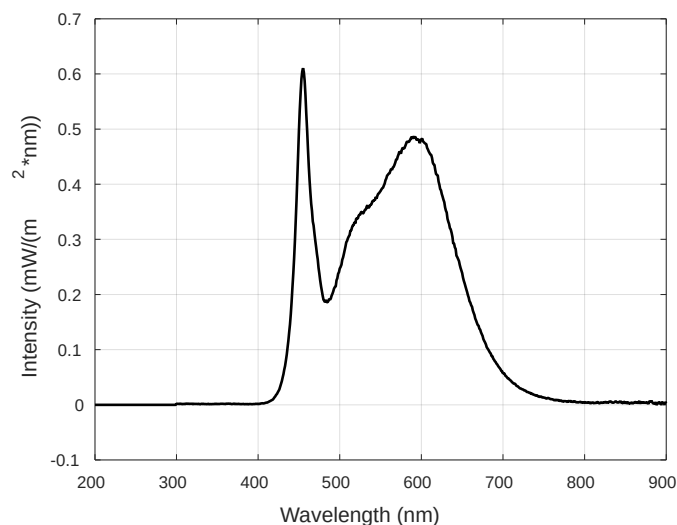
14 **5.6 RESISTANCE TO AMBIENT LIGHT**

15 These results show that the link functions well with no ambient

1 light incident on the detector. In this section, the influence of  
2 ambient light on the system performance will be assessed.

3 The major sources of ambient light are artificial lighting and  
4 sunlight. Most artificial lighting uses white LEDs (or is likely to use  
5 them in the future), so it is essential to evaluate the use of the QKD  
6 system under such illumination. In this work, the white LED-based  
7 ambient light is provided by the artificial lighting in the laboratory  
8 where experiments are conducted. The characterization of this  
9 LED's light is presented first below.

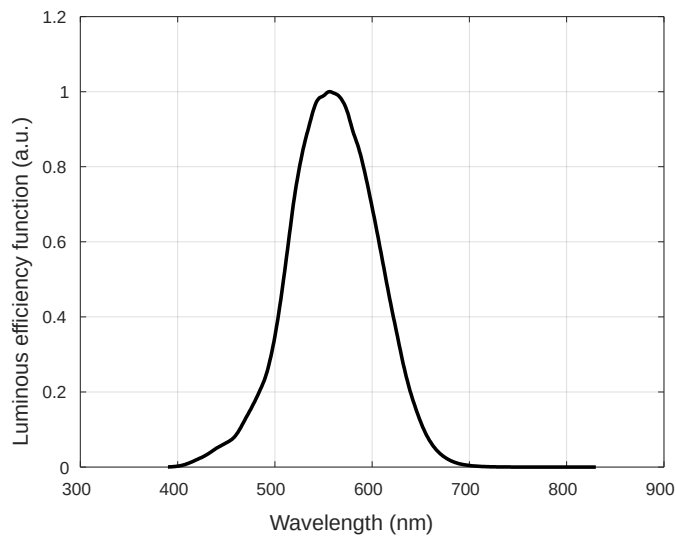
10 The spectrum of LED-based ambient light at the input of the  
11 QKD receiver was measured by the optical spectrum analyzer  
12 (Ocean Insight OFX00951). This is calibrated using a cosine  
13 corrector provided by the manufacturer to measure the irradiance.  
14 The analyzer input fibre was positioned horizontally to align with the  
15 orientation of the QKD receiver. The LED light was turned up to the  
16 maximum brightness, and the measured LED-based ambient light  
17 spectrum is shown in Fig. 5.12. The result shows the spectrum  
18 ranges from  $\sim 400$  nm to  $\sim 800$  nm, and the total horizontal  
19 irradiance is  $87.4 \text{ mW}/\text{m}^2$ .



20

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1 Figure 5.12 Optical spectrum of the ambient light from LED  
2 lighting



3

4 Figure 5.13 Luminous efficiency function [186]

5 Additionally, the illuminance is calculated based on the  
6 Commission Internationale de l'Eclairage (CIE) photopic luminosity  
7 function [187, 188]. The CIE 2008 curve (see Fig. 5.13) defines the  
8 luminous efficiency function, considering light ranging from 390 nm  
9 to 830 nm [186].

10 Based on the CIE 2008 curve, the illuminance can be obtained  
11 from irradiance:

$$12 \quad \text{Illuminance} = K \int_{390}^{830} V(\lambda) E(\lambda) d\lambda \quad (5.17)$$

13 where  $V(\lambda)$  is the light irradiance and  $E(\lambda)$  is the luminous efficiency  
14 function.  $K$  is  $683 \text{ lm/W}$ , representing the human eye's ability to  
15 perceive color. Thus, the horizontal illuminance of the LED at the  
16 input of the QKD receiver is calculated to be 29.8 lux at maximum  
17 brightness. In order to verify this, an illuminance meter (Light Meter

2

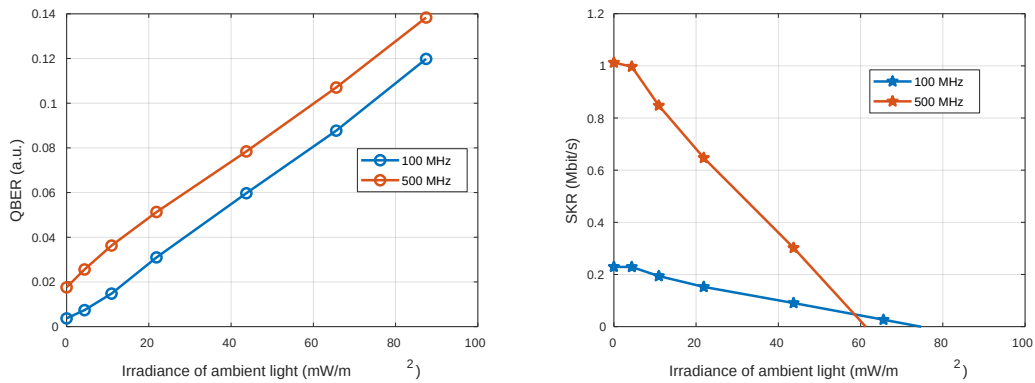
3

1 DT-1308) was used to measure the horizontal illuminance at the  
2 input of the QKD receiver. The measured illuminance was 22 lux,  
3 which is lower than the estimated 29.8 lux.

4 Once the irradiance and luminance from the LED were  
5 determined, the QKD link was operated whilst varying this ambient  
6 illumination. The horizontal irradiance of the LED-based ambient  
7 light at the input of the QKD receiver was varied from  $4.37 \text{ mW/m}^2$  to  
8  $87.4 \text{ mW/m}^2$ . Other parameters are the same as those used in section  
9 5.5. The measured QBER and SKR versus irradiance at a repetition  
10 rate of 100 MHz and 500 MHz are shown in Fig. 5.14. Note that the  
11 QBER and SKR results with no LED illumination are obtained from  
12 Fig. 5.11. The results show the QBER increases linearly with  
13 irradiance following Eq. (5.7). The maximum QBER reaches 11.98 %  
14 and 13.83 % for 100 MHz and 500 MHz repetition rates, respectively.  
15 These two values already exceed the QKD threshold of 11 %. The  
16 SKR decreases as a consequence of the higher QBER. The irradiance  
17 resulting in a QBER of 11 % is  $74.7 \text{ mW/m}^2$  and  $61.3 \text{ mW/m}^2$  for 100 MHz  
18 and 500 MHz repetition rates.

19 The corresponding illuminance of the LED-based ambient light is  
20 calculated according to Eq. (5.17). The QBER versus illuminance is  
21 plotted in Fig. 5.15. The illuminance resulting in a QBER of 11 % is  
22 25.5 lux and 20.9 lux for 100 MHz and 500 MHz repetition rates.

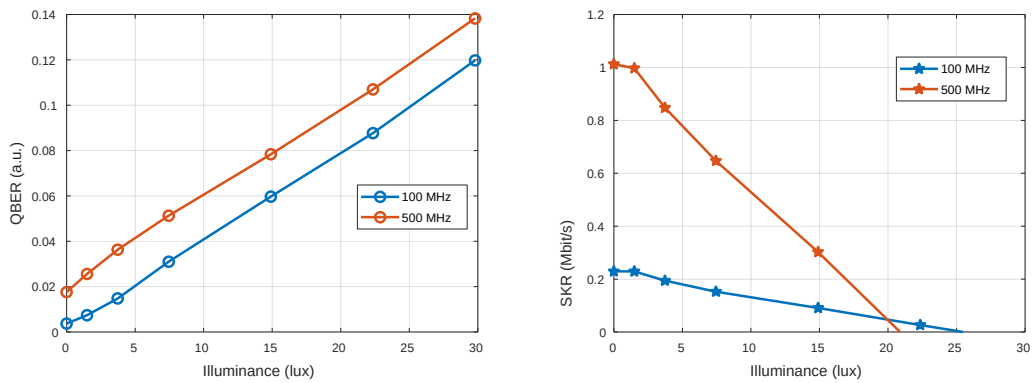
23 The LED light rejection of our QKD system is limited because the  
24 optical bandpass filter in the receiver can only block ambient light  
25 up to 600 nm, whereas the SiPM used in the UV-C SPD is sensitive  
26 up to 900 nm (see Fig. 4.4). Consequently, the ambient light in the  
27 600-900 nm region contributes significant interference, leading to a  
28 high QBER.



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3 Figure 5.14 UV-C B92 QKD performance under different irradiance of  
4 ambient light (a) QBER versus irradiance (b) SKR versus irradiance



5

6

7 Figure 5.15 UV-C B92 QKD performance under different illuminance  
8 of ambient light (a) QBER versus illuminance (b) SKR versus  
9 illuminance

10 **5.7 SUNLIGHT INTERFERENCE REJECTION DESIGN**

11 The LED light rejection performance of the QKD system was  
12 presented in Section 5.6. However, the spectrum of LED light (see  
13 Fig. 5.12) is different from the sunlight (see the ASTM 1.5 Spectra in  
14 Fig. 2.11). The illuminance of sunlight is 115650 lux, calculated  
15 based on the ASTM 1.5 Spectra and Eq. (5.17), while the LED in  
16 section 5.6 provides illuminance less than 100 lux. This means the

2

3

1 QKD link would be very likely to suffer much more severe  
2 interference from sunlight in outdoor applications.

3 A simulation was undertaken to estimate the ambient count rate  
4 caused by sunlight and LED light across the 200-900 nm spectrum.  
5 The ambient count rate is defined as the number of detection  
6 counts created by ambient light (sunlight or LED light) per second.  
7 The sunlight spectrum used was the ASTM 1.5 Spectra and the LED  
8 light spectrum was the same as that shown in Fig. 5.12, with an  
9 irradiance of  $61.3\text{mW}/\text{m}^2$  (corresponding to a QBER of 11 % at 500  
10 MHz repetition rate in section 5.6).

11 The model and parameters in Fig. 5.1 and Table 5.1 are applied  
12 here, with the following modifications: (1) the operating wavelength  
13 is set to 280 nm; (2) the DCR is 3 kcps; (3) the PDE versus  
14 wavelength follows Fig.4.4; (4) Total transmission (see section 5.2)  
15 versus wavelength is obtained based on device datasheets (Table  
16 5.7 shows device numbers).

17 Table 5.7 Optics and corresponding device numbers

Optical filter	Edmund 280 nm OD4 10 nm bandpass filter
Beam splitter	Thorlabs BSW20R
Polarizer	Thorlabs GLB10-UV
Convex lens	Thorlabs LA4052-UV
UV filter	Thorlabs FGUV5M

18 A plot of the simulated ambient count rate versus wavelength is  
19 shown in Fig. 5.16. The results show both sunlight and LED light  
20 create severe interference in the region from 650 nm to 900 nm, in  
21 the region where the SiPM is sensitive. The total ambient count

1 rates created by the LED light and sunlight are 68.368 kcps and  
2 6.529 Gcps, respectively. This means that the maximum tolerable  
3 ambient count rate is 68.368 kcps, while sunlight generates  
4 approximately 95,500 times this threshold ( $6.529 \text{ Gcps} / 68.368 \text{ kcps}$ ).

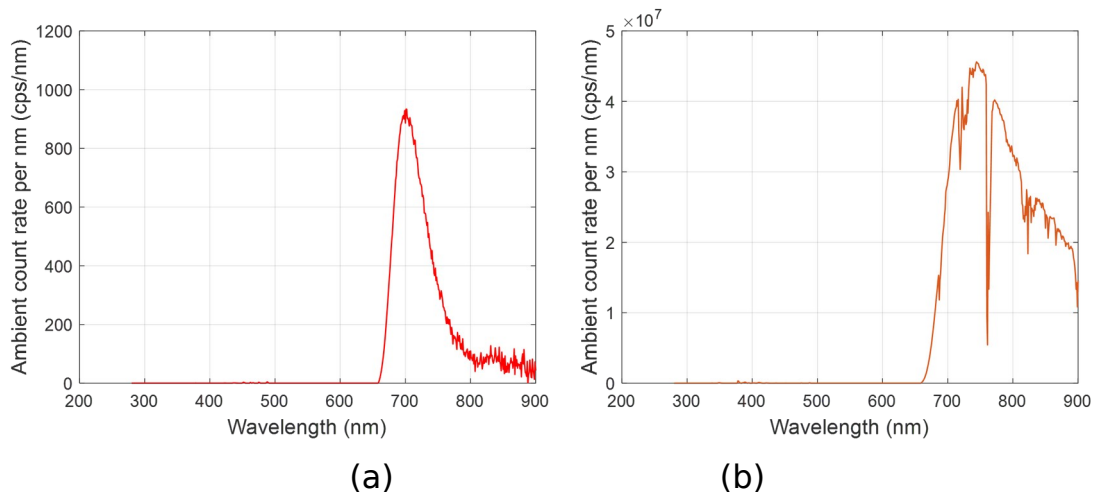
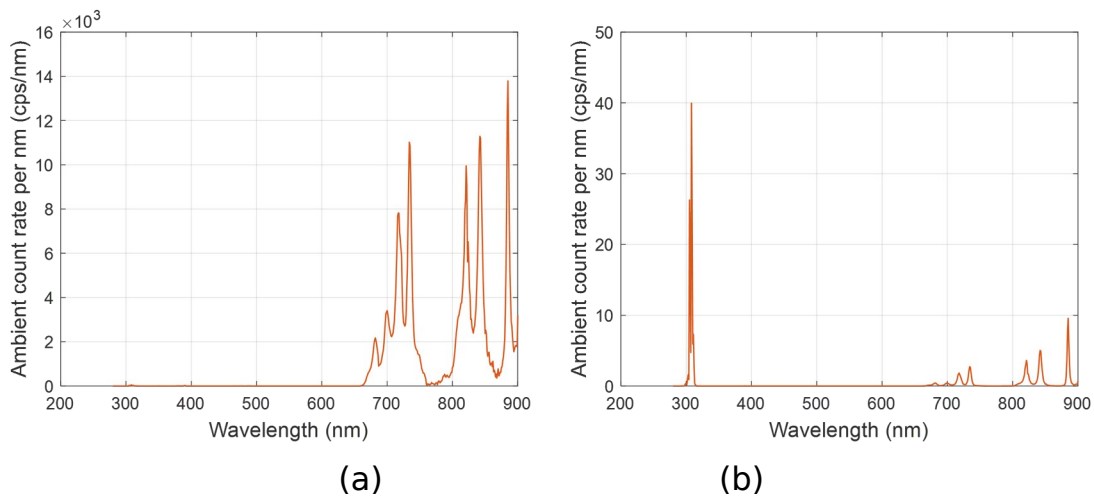


Figure 5.16 Simulated ambient count rate versus wavelength (a)  
LED light (b) Sunlight

9 Therefore, to realize indoor applications at realistic light levels  
10 and outdoor applications exposed to sunlight, the interference  
11 between 650 nm and 900 nm must be mitigated. Using additional  
12 optical filters at the entrance of the receiver to further attenuate the  
13 ambient light is a potential approach. A total attenuation of  $OD \geq 5$   
14 between 650 nm and 900 nm is needed. One potential solution is a  
15 short-pass filter, an example would be an ASAHI SPECTRA XUV0310.  
16 This filter can block light from 310 nm to 1000 nm with OD 4  
17 attenuation, while the transmission between 270 nm and 310 nm is  
18 over 80 %.

19 The transmission of this filter versus wavelength was introduced  
20 into the simulation model. The simulation was conducted to  
21 examine the feasibility of using this short-pass filter. Fig. 5.17 shows  
22 the ambient count rate from sunlight versus wavelength with one or

1 two short-pass filters. The result shows the total ambient count rate  
2 is 654.4 kcps with one short-pass filter. This is still higher than the  
3 maximum tolerable ambient count rate of 68.368 kcps. However,  
4 with two short-pass filters, the total ambient count rate is further  
5 reduced to 273 cps. Moreover, the corresponding QBER is calculated  
6 as 1.71 %. This shows the potential of our QKD system operating  
7 under sunlight conditions.



8  
9  
10  
11  
12

(a) (b)  
Figure 5.17 Simulated ambient count rate versus wavelength  
with (a) one short-pass filter (b) two short-pass filters (ASAHI  
SPECTRA XUV0310)

## 13 5.8 CONCLUSIONS

14 This Chapter presents a UV-C short-range wireless B92 QKD  
15 system. To the best of our knowledge, this is the first record of a  
16 QKD system working in the UV-C region. A theoretical model of a  
17 B92 QKD system is described, and simulation results show the  
18 potential of a practical short-range wireless QKD system with an  
19 operating wavelength of less than 289 nm. Then, a practical B92  
20 QKD transmitter and receiver are described and characterized. The  
21 UV-C wireless QKD links are demonstrated with a repetition rate  
22 from 50 MHz to 500 MHz. The QBER increases from 0.38 % to 1.76

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3

1 %, and the inferred SKR increases from 113.73 kbit/s to 1.0127 Mbit/  
2 s. Additionally, the LED ambient light was characterized, and our  
3 QKD system can work at a maximum LED light irradiance of  
4  $74.7\text{ mW/m}^2$  and  $61.3\text{ mW/m}^2$  for 100 MHz and 500 MHz repetition rates,  
5 respectively. These correspond to 25.5 lux and 20.9 lux respectively,  
6 which is lower than a normal indoor LED illuminance level of  $\sim 500$   
7 lux. Finally, a strategy of using short-pass filters to achieve sunlight  
8 rejection is proposed. An ambient count rate of 273 cps and a QBER  
9 of 1.71 % can be expected with two additional short-pass filters.  
10 Thus, our UV-C wireless QKD system has the potential to be robust  
11 to both sunlight and normal indoor LED light with the addition of  
12 short-pass filters.

13

# 1 **Chapter 6:**

## 2 **Conclusions and Future Work**

### 3 **6.1 CONCLUSIONS**

4 This thesis reports an in-depth investigation into UV-C  
5 communications. Given the challenges associated with OWC  
6 operating in the visible and near-infrared spectrum, this thesis  
7 focuses on exploiting the "solar blind" region within the UV-C  
8 spectrum. This region offers a unique opportunity due to its  
9 extremely low sunlight irradiance, which significantly reduces the  
10 impact of sunlight interference and enhances both the sensitivity  
11 and reliability of photon-counting communications. Three key  
12 contributions of this thesis are as follows.

- 13 ● The development of two photon-counting receivers based on  
14 SiPMs (see Chapters 3 and 4). Unlike previously reported UV  
15 photon-counting receivers [128-132], one advantage of these  
16 two receivers is that they are designed based on commercial  
17 devices. The receiver in Chapter 3 features a cooled SiPM along  
18 with a separate amplification circuit. This receiver shows a PDE  
19 of 15.38 %, a DCR of 35 kcps and a bandwidth of 5 MHz. The  
20 receiver in Chapter 4 (UV-C SPD) integrates the cooled SiPM with  
21 the PCB. Compared with the receiver in Chapter 3, the UV-C SPD  
22 exhibits improved performance, including a DCR of 2.1 kcps and  
23 a bandwidth of 110 MHz. Its excellent UV-C sensitivity allows it to  
24 play a crucial role in quantum communication or conventional  
25 communication scenarios with significant link losses.

- 1 ● The establishment of UV-C ultralow photon flux OWC links,  
2 including OOK links and PPM links (see Chapter 4). For OOK links,  
3 the maximum data rate achieved was 20 Mbit/s with a received  
4 photon number per bit of 30. In terms of PPM links, the  
5 maximum data rate reaches 120 Mbit/s, meeting eye safety  
6 standards. Additionally, a minimum transmitted photon number  
7 per bit of 7.7 was demonstrated in a 1 Mb/s 128-PPM link,  
8 marking the first reported Mbit/s-level UV-C wireless  
9 communication link operating with fewer than 10 transmitted  
10 photons per bit [189, 190]. The demonstration of these ultralow  
11 photon flux links explores the UV-C spectrum and provides a  
12 novel approach for exploring UV-C wireless communication  
13 scenarios facing significant link losses.
- 14 ● The demonstration of what is, to the best of our knowledge, the  
15 first wireless QKD link operating in the UV-C region (see Chapter  
16 5), which has the potential to be robust to ambient light. This  
17 provides a new solution for securing communications for mobile  
18 terminal devices. The QKD link infers a record SKR of over 1  
19 Mbit/s based on the measured raw keys. The QKD system can  
20 tolerate at most levels of  $74.7\text{ mW/m}^2$  (equal to 25.5 lux) white LED  
21 ambient light, and a design for rejecting sunlight and normal  
22 indoor LED light was also proposed and analysed.

## 23 **6.2 FUTURE WORK**

24 Ultralow photon flux OWC and QKD links operating in the UV-C  
25 region have been demonstrated in this thesis, using micro-LEDs and  
26 UV-C SPDs. However, further performance enhancements remain to  
27 be explored, particularly in achieving higher data rates and  
28 improved robustness to sunlight interference. Three potential future

1 improvements are outlined below.

### 2 **6.2.1 Higher-Speed Ultralow Photon Flux Links**

3 The highest data rate of demonstrated ultralow photon flux links  
4 in this thesis is 120 Mbit/s, as detailed in Section 4.4.4, specifically  
5 with the 40 Mb/s 8-PPM link. The primary factors obstructing further  
6 enhancement of data rates include the bandwidth of the transmitter  
7 (UV-C micro-LED) and receiver (UV-C SPD). Several potential  
8 solutions are as follows.

9 For the transmitter, reducing the size of the micro-LED can  
10 further increase the modulation bandwidth. UV-C micro-LEDs with  
11 approximately 1 GHz bandwidth have been reported with a diameter  
12 of less than 40  $\mu\text{m}$  [147].

13 For the receiver, the bandwidth of both SiPM and amplifier  
14 should be considered. Reducing the pixel size of the SiPM can  
15 improve its bandwidth. HAMAMATSU offers SiPMs with a pixel size of  
16 25  $\mu\text{m}$ , compared to the 50  $\mu\text{m}$  pixel size used in the UV-C SPD.  
17 Moreover, Onsemi provides SiPMs with pixel sizes of 20 or 10  $\mu\text{m}$   
18 [133, 191]. The fast output of these SiPMs has a rise time of  $\sim 0.3$   
19 ns, corresponding to a bandwidth of over 1 GHz. The bandwidth of  
20 an amplifier depends on its GBWP, the SiPM's capacitance and the  
21 trans-impedance load according to Eq. 3.7. The 8 GHz GBWP of  
22 Texas OPA855-Q1 is currently the highest available in the market.  
23 The trans-impedance load must be large enough to allow single-  
24 photon pulses to be observed by commercial oscilloscopes. Thus,  
25 reducing the active area of the SiPM to decrease its capacitance is a  
26 feasible solution. In addition, the cascade amplifier design is another  
27 potential solution because of its easily achieved high bandwidth.

### 28 **6.2.2 Sunlight Interference Rejection**

1 In addition to the strategy using additional filters in section 5.7,  
2 another approach is to use a holographic grating and a slit to select  
3 the UV-C light to enter the QKD receiver. Due to the diffraction of the  
4 holographic grating, the light of different wavelengths will be  
5 reflected in different directions. A slit is then used to select the  
6 desired wavelengths to pass through to the QKD receiver. For our  
7 UV-C wireless QKD system, light ranging from 270 nm to 290 nm  
8 can be selected, while wavelengths exceeding 290 nm will be  
9 blocked. Compared with the strategy using additional filters, this  
10 approach increases the complexity of the QKD receiver but can fully  
11 eliminate ambient light interference from wavelengths outside the  
12 working range.

### 13 **6.2.3 QKD Integrated with Tracking**

14 The UV-C wireless B92 QKD system presented in this thesis is  
15 static. To enable terminal mobile applications, a tracking system  
16 should be integrated with the existing static wireless QKD system.  
17 Many studies have been conducted to combine QKD systems with  
18 tracking systems [61, 192-195]. A typical tracking system's  
19 structure uses two mirrors, which can be automatically adjusted to  
20 steer the signal beam from the transmitter towards the receiver. A  
21 practical tracking system is more complex, but ways to enhance this  
22 static link with approaches for mobile data links would be of value.

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